

Emerging patterns of New Physics with and without Lepton Flavour Universal contributions

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Based on: [\[arXiv: 1903.09578\]](#) + [\[arXiv: 1809.08447\]](#)



bsll, September 2019, Lyon

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de Barcelona

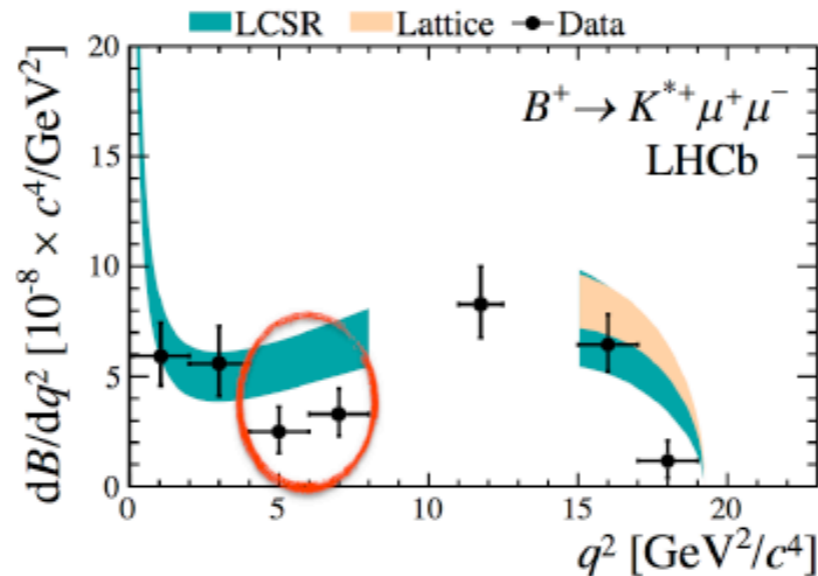
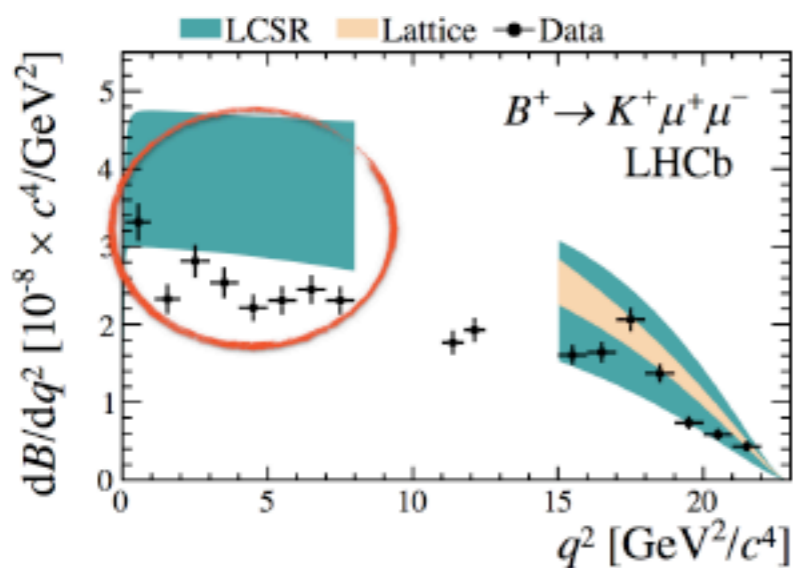
Outline

- ▶ Motivation
 - ▶ Pattern of New Physics from $b \rightarrow sll$ anomalies
- ▶ Anatomy of Fits
- ▶ State-of-the-art (2017 vs 2019)
- ▶ What's new?
- ▶ Discussion on models
- ▶ Conclusions and Outlook

Motivation

Pattern of $b \rightarrow sll$ anomalies

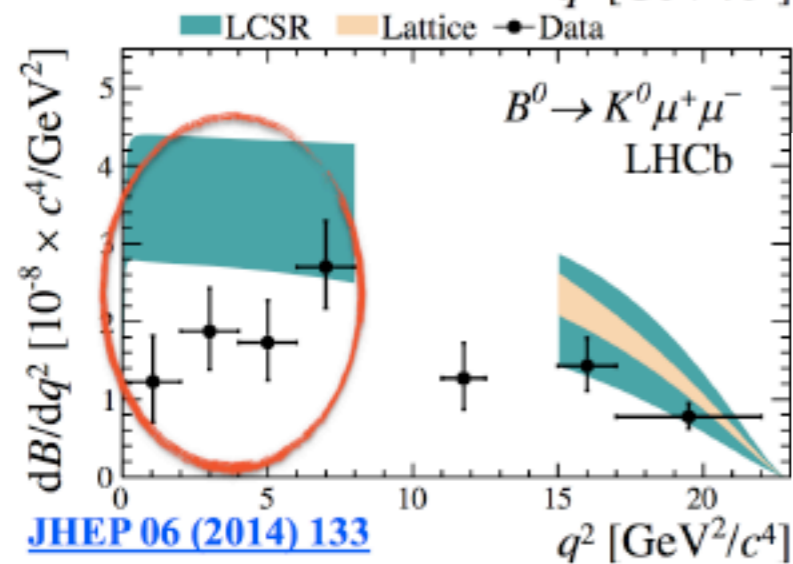
Coherent pattern of deviations \longrightarrow deficit in muons in different channels @ large and low recoil:



1st bin SM-like (\mathcal{C}_7 dominates & radiative corrections are constrained)



Issue: Excess in the electron channel in 1st bin explains deficit in R_{K^*}



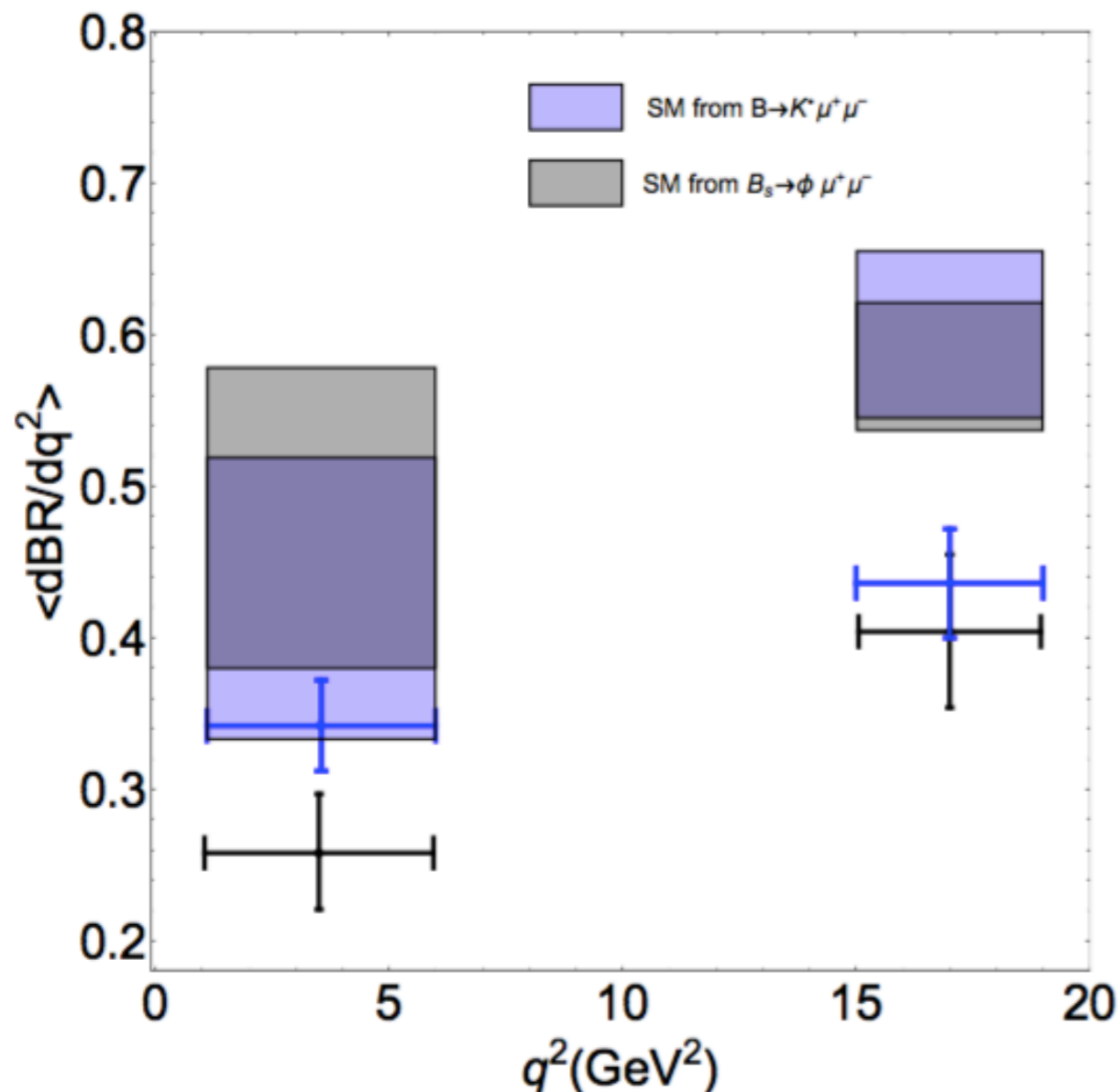
[JHEP 06 \(2014\) 133](#)

$b \rightarrow s\mu^+\mu^-$ ($\times 10^7$)	bin	SM	EXP	Pull
$\text{BR}(B^0 \rightarrow K^0\mu^+\mu^-)$	[15,19]	0.91 ± 0.12	0.67 ± 0.12	+1.4
$\text{BR}(B^0 \rightarrow K^{*0}\mu^+\mu^-)$	[16,19]	1.66 ± 0.15	1.23 ± 0.20	+1.7
$\text{BR}(B^+ \rightarrow K^{*+}\mu^+\mu^-)$	[15,19]	2.59 ± 0.25	1.60 ± 0.32	+2.5
$\text{BR}(B_s \rightarrow \phi\mu^+\mu^-)$	[15,18.8]	2.20 ± 0.17	1.62 ± 0.20	+2.2

Pattern of $b \rightarrow sll$ anomalies

[Algueró et al JHEP07 (2019) 096]

Tension observed in $\mathcal{B}(B_s \rightarrow \phi\mu\mu)$ between Large and Low Recoil



Low- q^2 : Pred. from sum rules using full FF with BSZ

High- q^2 : Pred. from lattice

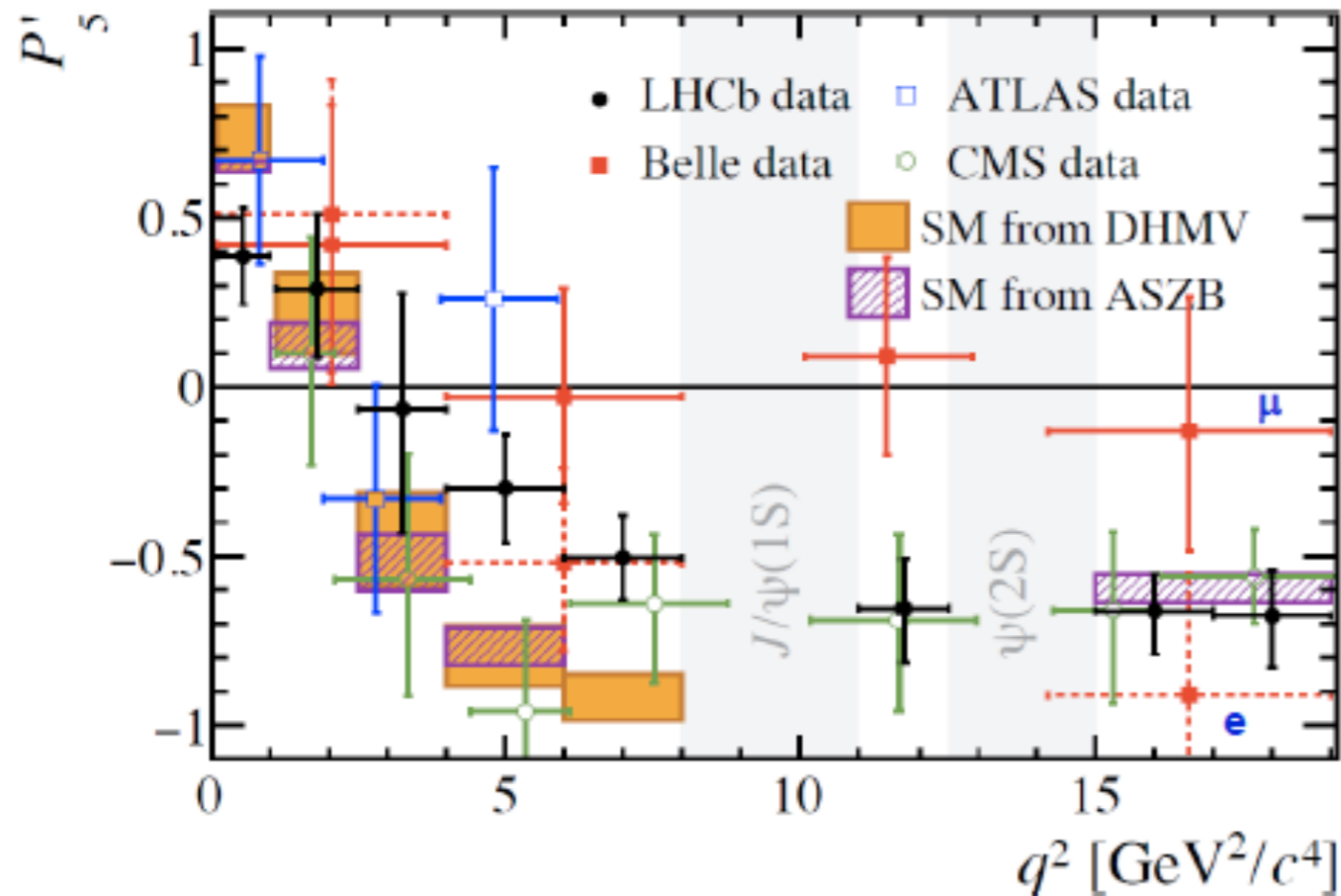
- No computations using B meson DA of KMPW available for $B_s \rightarrow \phi\mu\mu$
- At Low Recoil, $\mathcal{B}(B \rightarrow K^* \mu \mu)$ slightly larger than $\mathcal{B}(B_s \rightarrow \phi\mu\mu)$ (data follows same trend)
- **BUT:** At Large Recoil, $\mathcal{B}(B_s \rightarrow \phi\mu\mu)$ larger than $\mathcal{B}(B \rightarrow K^* \mu \mu)$ (data is reversed)

More data needed! to clarify if problem in Th. Prediction or statistical fluctuation

Pattern of $b \rightarrow sll$ anomalies

$$\frac{d^4\Gamma(\bar{B}_d \rightarrow K^*(\rightarrow K\pi)\mu^+\mu^-)}{dq^2 d\cos\theta_\ell d\cos\theta_K d\phi} = \frac{9}{32\pi} \sum_i J_i(q^2) f_i(\theta_\ell, \theta_K, \phi) \longrightarrow \text{Optimized angular observables}$$

[DMRV, JHEP1301(2013)048]



$$(q^2 = s_{\mu\mu})$$

$$P'_5 = \frac{J_5}{2\sqrt{-J_{2s}J_{2c}}} = P_5^\infty(1 + \mathcal{O}(\alpha_s\xi_\perp) + p.c.)$$

Optimized = Cancellation of SFF $\xi_{\parallel,\perp}$ @ LO

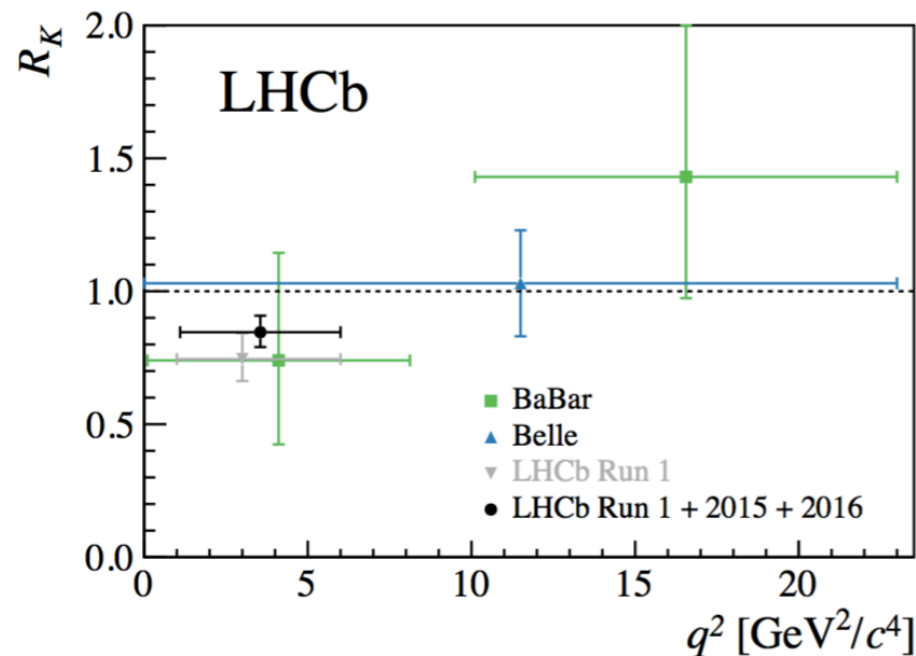
Sensitivity to hadronic uncertainties suppressed

- **LHCb** & **Belle** consistent
- **ATLAS** also shows deviation
- **CMS** compatible with SM but different procedure \longrightarrow Suggestion: extract correlations of F_L, P_1, P'_5 from same PDF. Use analytical integration of 3D PDFs

Pattern of $b \rightarrow sll$ anomalies

Anomalies also in **Lepton Flavour Universality Violating (LFUV)** observables

[LHCb: 1406.6482]

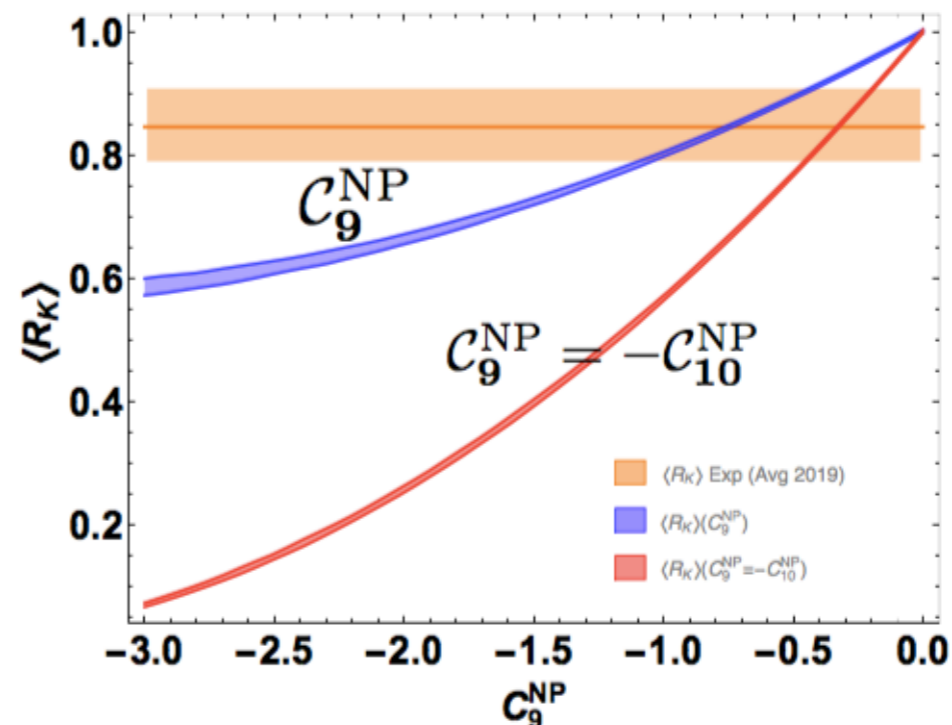


$$R_{K_{\text{LHCb}}}^{[1,6]} = 0.846^{+0.060+0.016}_{-0.054-0.014}$$

Still at 2.5σ from SM

- $R_K = 1$ in the SM
- NP couples $\neq \mu$ and e

Coherent deviation (deficit in muons)



Clean observable within NP

Pattern of $b \rightarrow sll$ anomalies

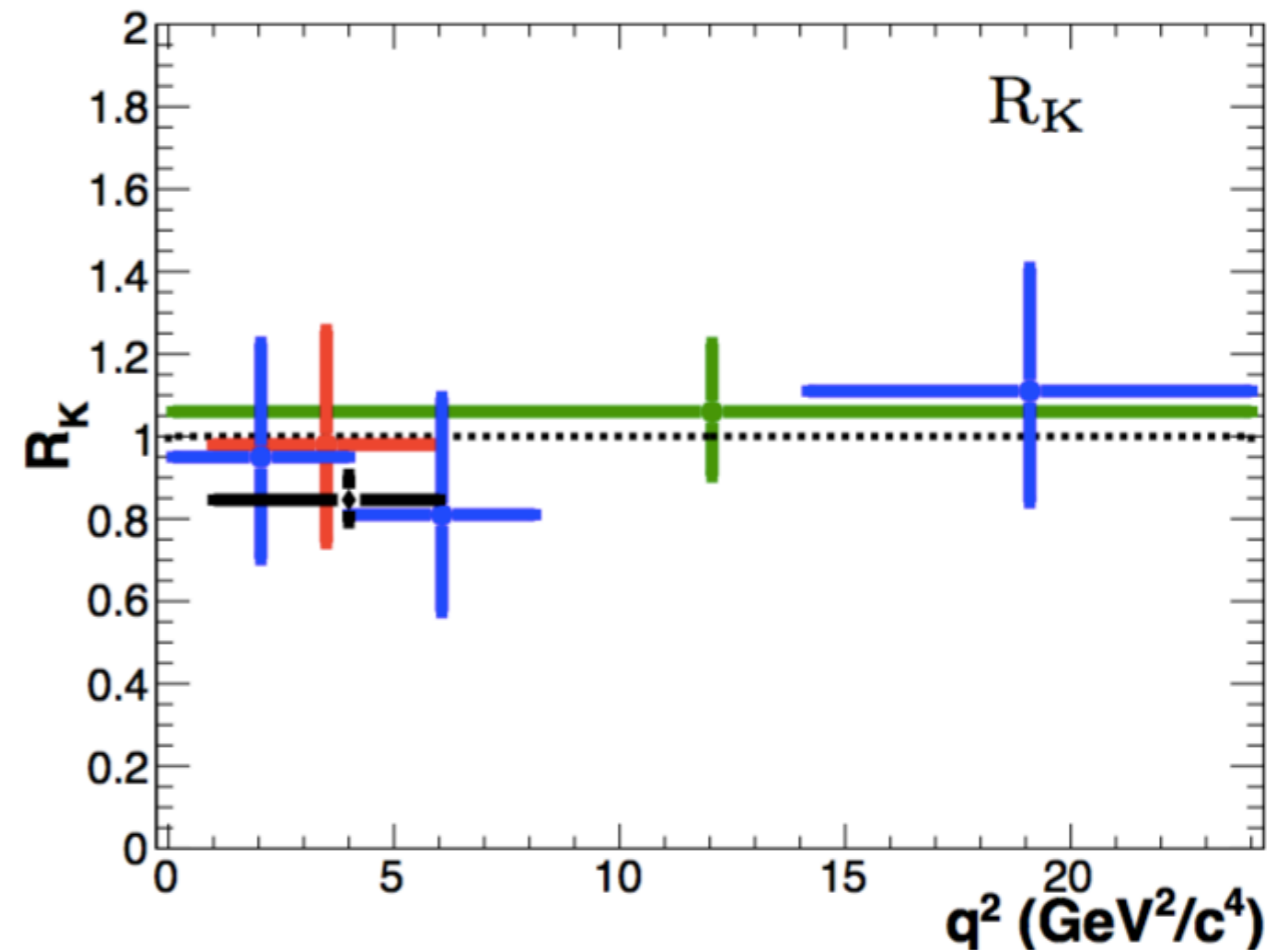
Belle released new measurement of R_K in the low and high kinematic range

[Belle: EPS-HEP Conference 2019]

$$R_{K_{\text{Belle}}}^{[1,6]} = 0.98_{-0.23}^{+0.27} \pm 0.06$$

$$R_{K_{\text{Belle}}}^{[q^2 > 14.18]} = 1.11_{-0.26}^{+0.29} \pm 0.07$$

Data from neutral and charged currents



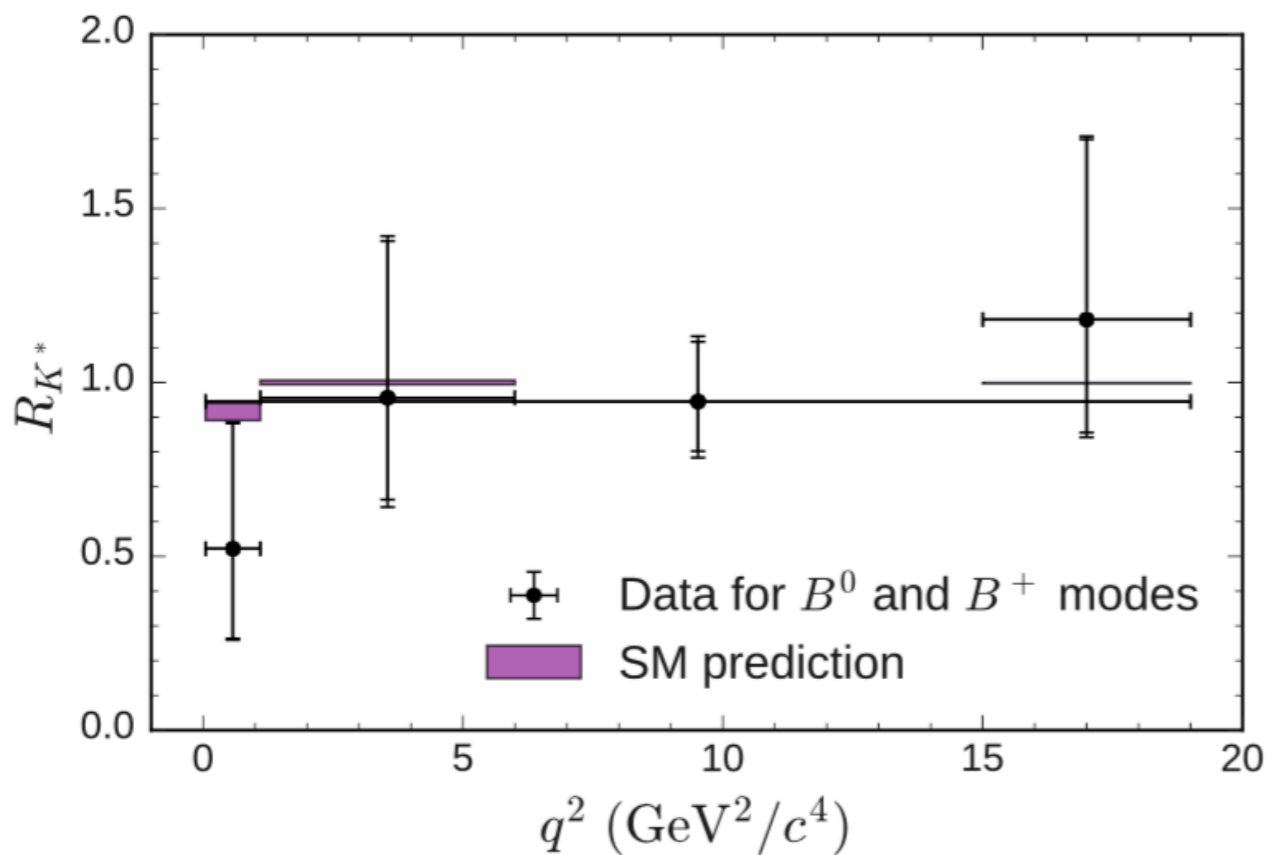
Long bin **[1,6]** compatible with SM but also with LCHb (large errors)

Bin	R_{K^+}	R_{K^0}	R_K
$0.1 < q^2 < 4$	$0.92_{-0.24}^{+0.27} \pm 0.05$	$1.5_{-1.0}^{+1.2} \pm 0.1$	$0.95_{-0.24}^{+0.27} \pm 0.06$
$4 < q^2 < 8.12$	$1.22_{-0.37}^{+0.42} \pm 0.07$	$0.50_{-0.30}^{+0.39} \pm 0.03$	$0.81_{-0.23}^{+0.28} \pm 0.05$
$1 < q^2 < 6$	$1.31_{-0.31}^{+0.34} \pm 0.07$	$0.53_{-0.33}^{+0.44} \pm 0.03$	$0.98_{-0.23}^{+0.27} \pm 0.06$
$q^2 > 14.18$	$1.08_{-0.27}^{+0.30} \pm 0.06$	$1.52_{-0.97}^{+1.23} \pm 0.10$	$1.11_{-0.26}^{+0.29} \pm 0.07$
whole q^2	$1.04_{-0.15}^{+0.16} \pm 0.06$	$1.25_{-0.44}^{+0.50} \pm 0.08$	$1.06_{-0.14}^{+0.15} \pm 0.07$

Pattern of $b \rightarrow sll$ anomalies

Another **LFUV** observable showing a sizeable deviation:

$$R_{K^{*0}} = \frac{\text{Br}(B^0 \rightarrow K^{*0} \mu^+ \mu^-)}{\text{Br}(B^0 \rightarrow K^{*0} J/\Psi(\rightarrow \mu^+ \mu^-))} \bigg/ \frac{\text{Br}(B^0 \rightarrow K^{*0} e^+ e^-)}{\text{Br}(B^0 \rightarrow K^{*0} J/\Psi(\rightarrow e^+ e^-))}$$



$$R_{K^{*0}} = \begin{cases} 0.66_{-0.07}^{+0.11} \pm 0.03 & q^2 \in [0.045, 1.1] \text{ GeV}^2 \\ 0.69_{-0.07}^{+0.11} \pm 0.05 & q^2 \in [1.1, 6] \text{ GeV}^2 \end{cases}$$

[LHCb: 1705.05802]

Belle combined data from neutral and charged channels:

$$R_{K^*}^{[0.045, 1.1]} = 0.52_{-0.26}^{+0.36} \pm 0.05$$

$$R_{K^*}^{[1.1, 6]} = 0.96_{-0.29}^{+0.45} \pm 0.11$$

$$R_{K^*}^{[15, 19]} = 1.18_{-0.32}^{+0.52} \pm 0.10$$

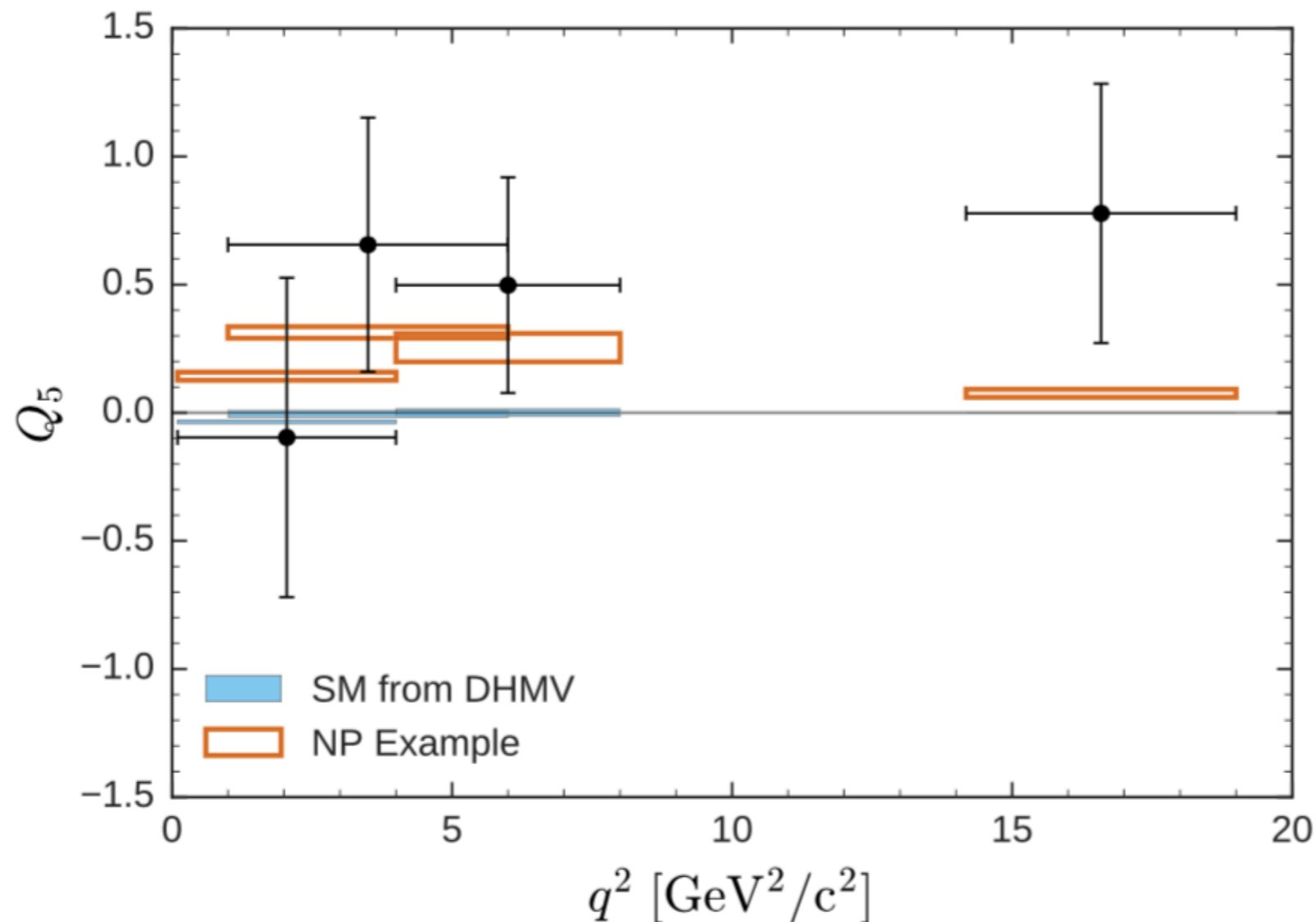
[Belle: 1904.02440]

- More complex structure, 6-8 amplitudes and 7 FF
- If NP or $q^2 < 1 \text{ GeV}^2$ **hadronic uncertainties return**

Pattern of $b \rightarrow sll$ anomalies

Belle, PRL118 (2017)

Another **LFUV** observable: $Q_5 = P'_{5\mu} - P'_{5e}$



Only measured by Belle:

$$Q_5^{[1,6]} = 0.656 \pm 0.485 \pm 0.103$$

More precision needed!
but coherent with present anomalies ($C_{9\mu}^{\text{NP}} < 0$)

It **must** be included in analysis



LFUV test

Anatomy of the Fits

Weak Effective Hamiltonian

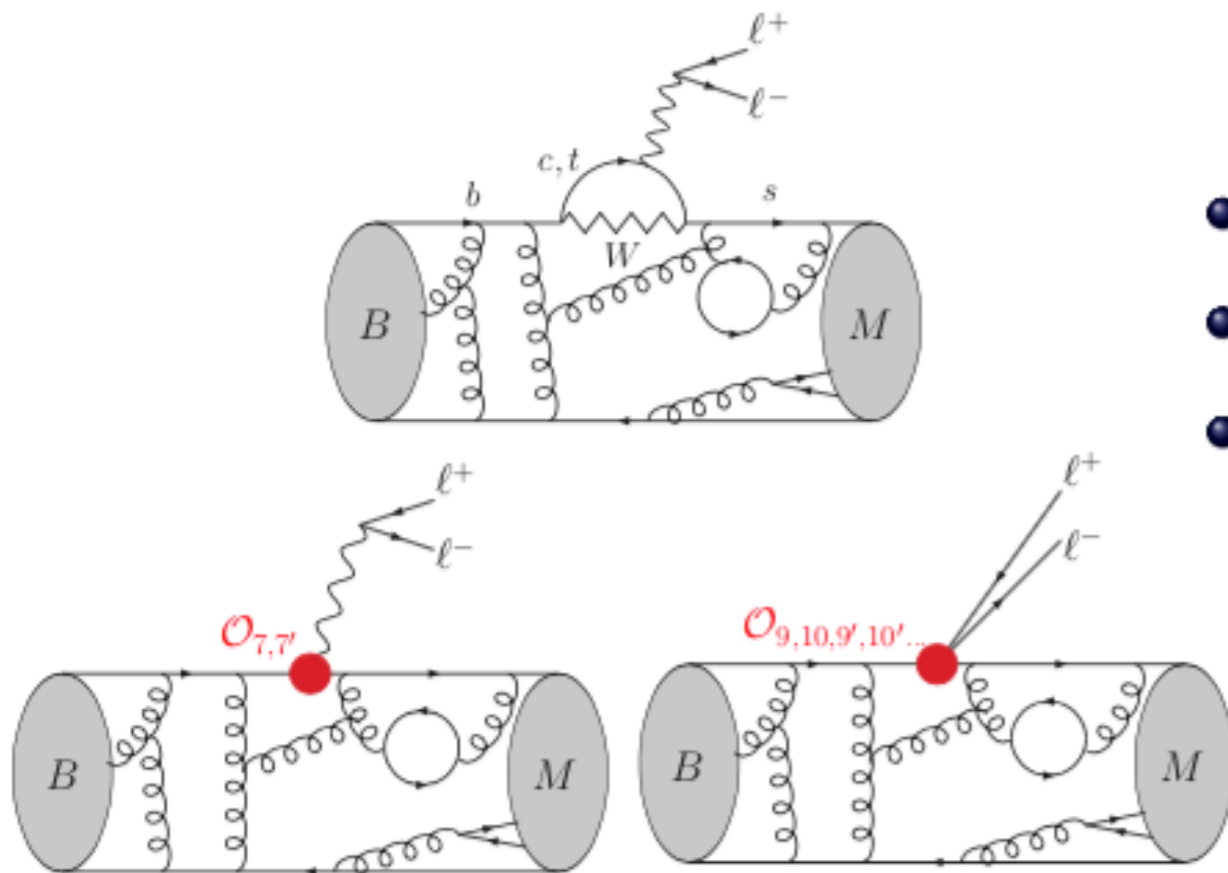
$$\mathcal{H}_{\Delta F=1}^{SM} \propto \sum V_{ts}^* V_{tb} C_i \mathcal{O}_i + \dots$$

separate short and long distances ($\mu = m_b = 4.8$)

- $\mathcal{O}_7 = \frac{e}{16\pi^2} m_b \bar{s} \sigma^{\mu\nu} (1 + \gamma_5) F_{\mu\nu} b$ [real or soft photon]
- $\mathcal{O}_9 = \frac{e^2}{16\pi^2} \bar{s} \gamma_\mu (1 - \gamma_5) b \bar{l} \gamma^\mu l$ [$b \rightarrow s \mu \mu$ via Z /hard $\gamma \dots$]
- $\mathcal{O}_{10} = \frac{e^2}{16\pi^2} \bar{s} \gamma_\mu (1 - \gamma_5) b \bar{l} \gamma^\mu \gamma_5 l$ [$b \rightarrow s \mu \mu$ via Z]

$$C_7^{SM} = -0.29, \quad C_9^{SM} = 4.1, \quad C_{10}^{SM} = -4.3$$

$A = C_i$ (short dist) \times Hadronic quantities (long dist)



Interesting directions:

$$C_9 = -C_{10} \Rightarrow L_q \otimes L_\ell$$

$$C_{9'} = -C_{10'} \Rightarrow R_q \otimes L_\ell$$

$$C_9 = -C_{9'} \Rightarrow A_q \otimes V_\ell$$

Global Fits

180 observables from (LHCb, Belle, ATLAS and CMS, no CP-violating obs)

- $B \rightarrow K^* \mu\mu$ ($P_{1,2}, P'_{4,5,6,8}, F_L$ in 5 large-recoil bins + 1 low-recoil bin)+available electronic obs.

...latest update $\text{Br}(B \rightarrow K^* \mu\mu)$ in small bins.

...LHCb results on R_{K^*}

- $B_s \rightarrow \phi \mu\mu$ ($P_1, P'_{4,6}, F_L$ in 3 large-recoil bins + 1 low-recoil bin)

+ 2 bins of R_K from Belle

- $B^+ \rightarrow K^+ \mu\mu, B^0 \rightarrow K^0 \ell\ell$ (BR) ($\ell = e, \mu$) (**new average $R_K = 0.846^{+0.060+0.016}_{-0.054-0.014}$**)

- $B \rightarrow X_s \gamma, B \rightarrow X_s \mu\mu, B_s \rightarrow \mu\mu$ (BR).

- Radiative decays: $B^0 \rightarrow K^{*0} \gamma$ (A_I and $S_{K^* \gamma}$), $B^+ \rightarrow K^{*+} \gamma, B_s \rightarrow \phi \gamma$

- ▶ **Belle measurements** for the isospin-averaged but lepton-flavour dependent ($Q_{4,5} = P_{4,5}^{\mu} - P_{4,5}^{e}$):
[3rd test of LFUV]

$$P_i^{\ell} = \sigma_+ P_i^{\ell}(B^+) + (1 - \sigma_+) P_i^{\ell}(\bar{B}^0) \quad \sigma_+ = 0.5 \pm 0.5$$

similar treatment of **new Belle isospin-averaged result on R_{K^*} (3-bins)**

- ▶ **ATLAS** measurement of whole basis of P_i and **CMS** measurements of P_1 and P'_5 .

- ▶ **ATLAS update** of $B_s \rightarrow \mu\mu$ (averaged with LHCb & CMS) and latest f_{B_s} lattice update.

State-of-the-art 2019

Fit Results

Frequentist approach: $C_i = C_i^{SM} + C_i^{NP}$, with C_i^{NP} assumed to be real (no CPV)

[JHEP01(2018)093]

1D Hyp.	All					LFUV				
	Best fit	1 σ	2 σ	Pull _{SM}	p-value	Best fit	1 σ	2 σ	Pull _{SM}	p-value
$C_{9\mu}^{NP}$	-1.11	[-1.28, -0.94]	[-1.45, -0.75]	5.8	68	-1.76	[-2.36, -1.23]	[-3.04, -0.76]	3.9	69
$C_{9\mu}^{NP} = -C_{10\mu}^{NP}$	-0.62	[-0.75, -0.49]	[-0.88, -0.37]	5.3	58	-0.66	[-0.84, -0.48]	[-1.04, -0.32]	4.1	78
$C_{9\mu}^{NP} = -C'_{9\mu}$	-1.01	[-1.18, -0.84]	[-1.34, -0.65]	5.4	61	-1.64	[-2.13, -1.05]	[-2.52, -0.49]	3.2	32
$C_{9\mu}^{NP} = -3C_{9e}^{NP}$	-1.07	[-1.24, -0.90]	[-1.40, -0.72]	5.8	70	-1.35	[-1.82, -0.95]	[-2.38, -0.59]	4.0	72

[1903.09578]

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$C_{9\mu}^{NP} = -C_{10\mu}^{NP}$	-0.46	[-0.56, -0.37] [-0.66, -0.28]	5.2	55.6 %	-0.40	[-0.53, -0.29] [-0.63, -0.18]	4.0	74.0 %
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$C_{9\mu}^{NP}$ preferred in All Fit

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1D Hyp.	All					LFUV				
	Best fit	1 σ	2 σ	Pull _{SM}	p-value	Best fit	1 σ	2 σ	Pull _{SM}	p-value
$C_{9\mu}^{NP}$	-1.11	[-1.28, -0.94]	[-1.45, -0.75]	5.8	68	-1.76	[-2.36, -1.23]	[-3.04, -0.76]	3.9	69
$C_{9\mu}^{NP} = -C_{10\mu}^{NP}$	-0.62	[-0.75, -0.49]	[-0.88, -0.37]	5.3	58	-0.66	[-0.84, -0.48]	[-1.04, -0.32]	4.1	78
$C_{9\mu}^{NP} = -C'_{9\mu}$	-1.01	[-1.18, -0.84]	[-1.34, -0.65]	5.4	61	-1.64	[-2.13, -1.05]	[-2.52, -0.49]	3.2	32
$C_{9\mu}^{NP} = -3C_{9e}^{NP}$	-1.07	[-1.24, -0.90]	[-1.40, -0.72]	5.8	70	-1.35	[-1.82, -0.95]	[-2.38, -0.59]	4.0	72

[1903.09578]

1D Hyp.	All				LFUV			
	Best fit	1 σ / 2 σ	Pull _{SM}	p-value	Best fit	1 σ / 2 σ	Pull _{SM}	p-value
$C_{9\mu}^{NP}$	-0.98	[-1.15, -0.81] [-1.31, -0.64]	5.6	65.4 %	-0.89	[-1.23, -0.59] [-1.60, -0.32]	3.3	52.2 %
$C_{9\mu}^{NP} = -C_{10\mu}^{NP}$	-0.46	[-0.56, -0.37] [-0.66, -0.28]	5.2	55.6 %	-0.40	[-0.53, -0.29] [-0.63, -0.18]	4.0	74.0 %
$C_{9\mu}^{NP} = -C'_{9\mu}$	-0.99	[-1.15, -0.82] [-1.31, -0.64]	5.5	52.9 %	-1.61	[-2.13, -0.96] [-2.54, -0.41]	3.0	42.5 %
$C_{9\mu}^{NP} = -3C_{9e}^{NP}$	-0.87	[-1.03, -0.71] [-1.19, -0.55]	5.5	61.9 %	-0.66	[-0.90, -0.44] [-1.17, -0.24]	3.3	52.2 %

$C_{9\mu}^{NP}$ preferred in All Fit

$C_{9\mu}^{NP} = -C_{10\mu}^{NP}$ preferred in LFUV Fit

Same hierarchies except

$$C_{9\mu}^{NP} = -C'_{9\mu}$$

$$R_K \sim 1$$

Pull_{SM}: how much the SM is disfavoured with respect to a New Physics hypothesis to explain data.

→ A scenario with a large SM-pull ⇒ big improvement over SM and better description of data.

Fit Results

Frequentist approach: $C_i = C_i^{SM} + C_i^{NP}$, with C_i^{NP} assumed to be real (no CPV)

[JHEP01(2018)093]

1D Hyp.	All					LFUV				
	Best fit	1 σ	2 σ	Pull _{SM}	p-value	Best fit	1 σ	2 σ	Pull _{SM}	p-value
$C_{9\mu}^{NP}$	-1.11	[-1.28, -0.94]	[-1.45, -0.75]	5.8	68	-1.76	[-2.36, -1.23]	[-3.04, -0.76]	3.9	69
$C_{9\mu}^{NP} = -C_{10\mu}^{NP}$	-0.62	[-0.75, -0.49]	[-0.88, -0.37]	5.3	58	-0.66	[-0.84, -0.48]	[-1.04, -0.32]	4.1	78
$C_{9\mu}^{NP} = -C_{9\mu}'$	-1.01	[-1.18, -0.84]	[-1.34, -0.65]	5.4	61	-1.64	[-2.13, -1.05]	[-2.52, -0.49]	3.2	32
$C_{9\mu}^{NP} = -3C_{9e}^{NP}$	-1.07	[-1.24, -0.90]	[-1.40, -0.72]	5.8	70	-1.35	[-1.82, -0.95]	[-2.38, -0.59]	4.0	72

[1903.09578]

1D Hyp.	All				LFUV			
	Best fit	1 σ / 2 σ	Pull _{SM}	p-value	Best fit	1 σ / 2 σ	Pull _{SM}	p-value
$C_{9\mu}^{NP}$	-0.98	[-1.15, -0.81] [-1.31, -0.64]	5.6	65.4 %	-0.89	[-1.23, -0.59] [-1.60, -0.32]	3.3	52.2 %
$C_{9\mu}^{NP} = -C_{10\mu}^{NP}$	-0.46	[-0.56, -0.37] [-0.66, -0.28]	5.2	55.6 %	-0.40	[-0.53, -0.29] [-0.63, -0.18]	4.0	74.0 %
$C_{9\mu}^{NP} = -C_{9\mu}'$	-0.99	[-1.15, -0.82] [-1.31, -0.64]	5.5	59.9 %	-1.61	[-2.13, -0.96] [-2.54, -0.41]	3.0	42.5 %
$C_{9\mu}^{NP} = -3C_{9e}^{NP}$	-0.87	[-1.03, -0.71] [-1.19, -0.55]	5.5	61.9 %	-0.66	[-0.90, -0.44] [-1.17, -0.24]	3.3	52.2 %

$C_{9\mu}^{NP}$ preferred in All Fit

$C_{9\mu}^{NP} = -C_{10\mu}^{NP}$ preferred in LFUV Fit

Same hierarchies except

$$C_{9\mu}^{NP} = -C_{9\mu}'$$

$$R_K \sim 1$$

Pull_{SM}: how much the SM is disfavoured with respect to a New Physics hypothesis to explain data.

→ A scenario with a large SM-pull ⇒ big improvement over SM and better description of data.

Fit Results

Update in 2D Fits

[JHEP01(2018)093]

2D Hyp.	All			LFUV		
	Best fit	Pull _{SM}	p-value	Best fit	Pull _{SM}	p-value
$(C_{9\mu}^{\text{NP}}, C_{10\mu}^{\text{NP}})$	(-1.01,0.29)	5.7	72	(-1.30,0.36)	3.7	75
$(C_{9\mu}^{\text{NP}}, C_7')$	(-1.13,0.01)	5.5	69	(-1.85,-0.04)	3.6	66
$(C_{9\mu}^{\text{NP}}, C_{9'\mu})$	(-1.15,0.41)	5.6	71	(-1.99,0.93)	3.7	72
$(C_{9\mu}^{\text{NP}}, C_{10'\mu})$	(-1.22,-0.22)	5.7	72	(-2.22,-0.41)	3.9	85
$(C_{9\mu}^{\text{NP}}, C_{9e}^{\text{NP}})$	(-1.00,0.42)	5.5	68	(-1.36,0.46)	3.5	65
Hyp. 1	(-1.16,0.38)	5.7	73	(-1.68,0.60)	3.8	78
Hyp. 2	(-1.15, 0.01)	5.0	57	(-2.16,0.41)	3.0	37
Hyp. 3	(-0.67,-0.10)	5.0	57	(0.61,2.48)	3.7	73
Hyp. 4	(-0.70,0.28)	5.0	57	(-0.74,0.43)	3.7	72

[1903.09578]

2D Hyp.	All			LFUV		
	Best fit	Pull _{SM}	p-value	Best fit	Pull _{SM}	p-value
$(C_{9\mu}^{\text{NP}}, C_{10\mu}^{\text{NP}})$	(-0.91,0.18)	5.4	68.7 %	(-0.16,0.56)	3.4	76.9 %
$(C_{9\mu}^{\text{NP}}, C_{7'})$	(-1.00,0.02)	5.4	67.9 %	(-0.90,-0.04)	2.9	55.1 %
$(C_{9\mu}^{\text{NP}}, C_{9'\mu})$	(-1.10,0.55)	5.7	75.1 %	(-1.79,1.14)	3.4	76.1 %
$(C_{9\mu}^{\text{NP}}, C_{10'\mu})$	(-1.14,-0.35)	5.9	78.6 %	(-1.88,-0.62)	3.8	91.3 %
$(C_{9\mu}^{\text{NP}}, C_{9e}^{\text{NP}})$	(-1.05,-0.23)	5.3	66.2 %	(-0.73,0.16)	2.8	52.3 %
Hyp. 1	(-1.06,0.26)	5.7	75.7 %	(-1.62,0.29)	3.4	77.6 %
Hyp. 2	(-0.97,0.09)	5.3	65.2 %	(-1.95,0.25)	3.2	66.6 %
Hyp. 3	(-0.47,0.06)	4.8	55.7 %	(-0.39,-0.13)	3.4	76.2 %
Hyp. 4	(-0.49,0.12)	5.0	59.3 %	(-0.48,0.17)	3.6	84.3 %
Hyp. 5	(-1.14,0.24)	5.9	78.7 %	(-2.07,0.52)	3.9	92.5 %

$$\text{Hyp.1 : } (C_{9\mu}^{\text{NP}} = -C_{9'\mu}, C_{10\mu}^{\text{NP}} = C_{10'\mu})$$

$$\text{Hyp.2 : } (C_{9\mu}^{\text{NP}} = -C_{9'\mu}, C_{10\mu}^{\text{NP}} = -C_{10'\mu})$$

$$\text{Hyp.3 : } (C_{9\mu}^{\text{NP}} = -C_{10\mu}^{\text{NP}}, C_{9'\mu} = C_{10'\mu})$$

$$\text{Hyp.4 : } (C_{9\mu}^{\text{NP}} = -C_{10\mu}^{\text{NP}}, C_{9'\mu} = -C_{10'\mu})$$

$$\text{Hyp.5 : } (C_{9\mu}^{\text{NP}}, C_{9'\mu} = -C_{10'\mu})$$

- Small increase ($\sim 0.2\sigma$)
for some **RHC** scenarios
opposite to 1D cases

- R_K closer to SM prefers
 $C_{9'\mu} > 0$ and $C_{10'\mu} < 0$

- A $R_q \otimes L_\ell$ for $\mathcal{O}_{9',10'}$ clearly prefers
 V structure over L_ℓ for leptons

Fit Results

Update in 2D Fits

[JHEP01(2018)093]

2D Hyp.	All			LFUV		
	Best fit	Pull _{SM}	p-value	Best fit	Pull _{SM}	p-value
$(C_{9\mu}^{NP}, C_{10\mu}^{NP})$	(-1.01,0.29)	5.7	72	(-1.30,0.36)	3.7	75
$(C_{9\mu}^{NP}, C_7')$	(-1.13,0.01)	5.5	69	(-1.85,-0.04)	3.6	66
$(C_{9\mu}^{NP}, C_{9'\mu})$	(-1.15,0.41)	5.6	71	(-1.99,0.93)	3.7	72
$(C_{9\mu}^{NP}, C_{10'\mu})$	(-1.22,-0.22)	5.7	72	(-2.22,-0.41)	3.9	85
$(C_{9\mu}^{NP}, C_{9e}^{NP})$	(-1.00,0.42)	5.5	68	(-1.36,0.46)	3.5	65
Hyp. 1	(-1.16,0.38)	5.7	73	(-1.68,0.60)	3.8	78
Hyp. 2	(-1.15, 0.01)	5.0	57	(-2.16,0.41)	3.0	37
Hyp. 3	(-0.67,-0.10)	5.0	57	(0.61,2.48)	3.7	73
Hyp. 4	(-0.70,0.28)	5.0	57	(-0.74,0.43)	3.7	72

[1903.09578]

2D Hyp.	All			LFUV		
	Best fit	Pull _{SM}	p-value	Best fit	Pull _{SM}	p-value
$(C_{9\mu}^{NP}, C_{10\mu}^{NP})$	(-0.91,0.18)	5.4	68.7 %	(-0.16,0.56)	3.4	76.9 %
$(C_{9\mu}^{NP}, C_{7'})$	(-1.00,0.02)	5.4	67.9 %	(-0.90,-0.04)	2.9	55.1 %
$(C_{9\mu}^{NP}, C_{9'\mu})$	(-1.10,0.55)	5.7	75.1 %	(-1.79,1.14)	3.4	76.1 %
$(C_{9\mu}^{NP}, C_{10'\mu})$	(-1.14,-0.35)	5.9	78.6 %	(-1.88,-0.62)	3.8	91.3 %
$(C_{9\mu}^{NP}, C_{9e}^{NP})$	(-1.05,-0.23)	5.3	66.2 %	(-0.73,0.16)	2.8	52.3 %
Hyp. 1	(-1.06,0.26)	5.7	75.7 %	(-1.62,0.29)	3.4	77.6 %
Hyp. 2	(-0.97,0.09)	5.3	65.2 %	(-1.95,0.25)	3.2	66.6 %
Hyp. 3	(-0.47,0.06)	4.8	55.7 %	(-0.39,-0.13)	3.4	76.2 %
Hyp. 4	(-0.49,0.12)	5.0	59.3 %	(-0.48,0.17)	3.6	84.3 %
Hyp. 5	(-1.14,0.24)	5.9	78.7 %	(-2.07,0.52)	3.9	92.5 %

$$\text{Hyp.1 : } (C_{9\mu}^{NP} = -C_{9'\mu}, C_{10\mu}^{NP} = C_{10'\mu})$$

$$\text{Hyp.2 : } (C_{9\mu}^{NP} = -C_{9'\mu}, C_{10\mu}^{NP} = -C_{10'\mu})$$

$$\text{Hyp.3 : } (C_{9\mu}^{NP} = -C_{10\mu}^{NP}, C_{9'\mu} = C_{10'\mu})$$

$$\text{Hyp.4 : } (C_{9\mu}^{NP} = -C_{10\mu}^{NP}, C_{9'\mu} = -C_{10'\mu})$$

$$\text{Hyp.5 : } (C_{9\mu}^{NP}, C_{9'\mu} = -C_{10'\mu})$$

- Small increase ($\sim 0.2\sigma$)
for some **RHC** scenarios
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Fit Results

Update in 2D Fits

[JHEP01(2018)093]

2D Hyp.	All			LFUV		
	Best fit	Pull _{SM}	p-value	Best fit	Pull _{SM}	p-value
$(C_{9\mu}^{NP}, C_{10\mu}^{NP})$	(-1.01,0.29)	5.7	72	(-1.30,0.36)	3.7	75
$(C_{9\mu}^{NP}, C_7')$	(-1.13,0.01)	5.5	69	(-1.85,-0.04)	3.6	66
$(C_{9\mu}^{NP}, C_{9'\mu})$	(-1.15,0.41)	5.6	71	(-1.99,0.93)	3.7	72
$(C_{9\mu}^{NP}, C_{10'\mu})$	(-1.22,-0.22)	5.7	72	(-2.22,-0.41)	3.9	85
$(C_{9\mu}^{NP}, C_{9e}^{NP})$	(-1.00,0.42)	5.5	68	(-1.36,0.46)	3.5	65
Hyp. 1	(-1.16,0.38)	5.7	73	(-1.68,0.60)	3.8	78
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Hyp. 3	(-0.67,-0.10)	5.0	57	(0.61,2.48)	3.7	73
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[1903.09578]

2D Hyp.	All			LFUV		
	Best fit	Pull _{SM}	p-value	Best fit	Pull _{SM}	p-value
$(C_{9\mu}^{NP}, C_{10\mu}^{NP})$	(-0.91,0.18)	5.4	68.7 %	(-0.16,0.56)	3.4	76.9 %
$(C_{9\mu}^{NP}, C_{7'})$	(-1.00,0.02)	5.4	67.9 %	(-0.99,-0.04)	2.9	55.1 %
$(C_{9\mu}^{NP}, C_{9'\mu})$	(-1.10,0.55)	5.7	75.1 %	(-1.79,1.14)	3.4	76.1 %
$(C_{9\mu}^{NP}, C_{10'\mu})$	(-1.14,-0.35)	5.9	78.6 %	(-1.88,-0.62)	3.8	91.3 %
$(C_{9\mu}^{NP}, C_{9e}^{NP})$	(-1.05,-0.23)	5.3	66.2 %	(-0.73,0.16)	2.8	52.3 %
Hyp. 1	(-1.06,0.26)	5.7	75.7 %	(-1.62,0.29)	3.4	77.6 %
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Hyp. 3	(-0.47,0.06)	4.8	55.7 %	(-0.39,-0.13)	3.4	76.2 %
Hyp. 4	(-0.49,0.12)	5.0	59.3 %	(-0.48,0.17)	3.6	84.3 %
Hyp. 5	(-1.14,0.24)	5.9	78.7 %	(-2.07,0.52)	3.9	92.5 %

$$\text{Hyp.1 : } (C_{9\mu}^{NP} = -C_{9'\mu}, C_{10\mu}^{NP} = C_{10'\mu})$$

$$\text{Hyp.2 : } (C_{9\mu}^{NP} = -C_{9'\mu}, C_{10\mu}^{NP} = -C_{10'\mu})$$

$$\text{Hyp.3 : } (C_{9\mu}^{NP} = -C_{10\mu}^{NP}, C_{9'\mu} = C_{10'\mu})$$

$$\text{Hyp.4 : } (C_{9\mu}^{NP} = -C_{10\mu}^{NP}, C_{9'\mu} = -C_{10'\mu})$$

$$\text{Hyp.5 : } (C_{9\mu}^{NP}, C_{9'\mu} = -C_{10'\mu})$$

- Small increase ($\sim 0.2\sigma$) for some **RHC** scenarios opposite to 1D cases

- R_K closer to SM prefers $C_{9'\mu} > 0$ and $C_{10'\mu} < 0$

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Fit Results

Update in 2D Fits

[JHEP01(2018)093]

2D Hyp.	All			LFUV		
	Best fit	Pull _{SM}	p-value	Best fit	Pull _{SM}	p-value
$(C_{9\mu}^{NP}, C_{10\mu}^{NP})$	(-1.01,0.29)	5.7	72	(-1.30,0.36)	3.7	75
$(C_{9\mu}^{NP}, C_7')$	(-1.13,0.01)	5.5	69	(-1.85,-0.04)	3.6	66
$(C_{9\mu}^{NP}, C_{9'\mu})$	(-1.15,0.41)	5.6	71	(-1.99,0.93)	3.7	72
$(C_{9\mu}^{NP}, C_{10'\mu})$	(-1.22,-0.22)	5.7	72	(-2.22,-0.41)	3.9	85
$(C_{9\mu}^{NP}, C_{9e}^{NP})$	(-1.00,0.42)	5.5	68	(-1.36,0.46)	3.5	65
Hyp. 1	(-1.16,0.38)	5.7	73	(-1.68,0.60)	3.8	78
Hyp. 2	(-1.15, 0.01)	5.0	57	(-2.16,0.41)	3.0	37
Hyp. 3	(-0.67,-0.10)	5.0	57	(0.61,2.48)	3.7	73
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[1903.09578]

2D Hyp.	All			LFUV		
	Best fit	Pull _{SM}	p-value	Best fit	Pull _{SM}	p-value
$(C_{9\mu}^{NP}, C_{10\mu}^{NP})$	(-0.91,0.18)	5.4	68.7 %	(-0.16,0.56)	3.4	76.9 %
$(C_{9\mu}^{NP}, C_{7'})$	(-1.00,0.02)	5.4	67.9 %	(-0.99,-0.04)	2.9	55.1 %
$(C_{9\mu}^{NP}, C_{9'\mu})$	(-1.10,0.55)	5.7	75.1 %	(-1.79,1.14)	3.4	76.1 %
$(C_{9\mu}^{NP}, C_{10'\mu})$	(-1.14,-0.35)	5.9	78.6 %	(-1.88,-0.62)	3.8	91.3 %
$(C_{9\mu}^{NP}, C_{9e}^{NP})$	(-1.05,-0.23)	5.3	66.2 %	(-0.73,0.16)	2.8	52.3 %
Hyp. 1	(-1.06,0.26)	5.7	75.7 %	(-1.62,0.29)	3.4	77.6 %
Hyp. 2	(-0.97,0.09)	5.3	65.2 %	(-1.95,0.25)	3.2	66.6 %
Hyp. 3	(-0.47,0.06)	4.8	55.7 %	(-0.39,-0.13)	3.4	76.2 %
Hyp. 4	(-0.49,0.12)	5.0	59.3 %	(-0.48,0.17)	3.6	84.3 %
Hyp. 5	(-1.14,0.24)	5.9	78.7 %	(-2.07,0.52)	3.9	92.5 %

$$\text{Hyp.1 : } (C_{9\mu}^{NP} = -C_{9'\mu}, C_{10\mu}^{NP} = C_{10'\mu})$$

$$\text{Hyp.2 : } (C_{9\mu}^{NP} = -C_{9'\mu}, C_{10\mu}^{NP} = -C_{10'\mu})$$

$$\text{Hyp.3 : } (C_{9\mu}^{NP} = -C_{10\mu}^{NP}, C_{9'\mu} = C_{10'\mu})$$

$$\text{Hyp.4 : } (C_{9\mu}^{NP} = -C_{10\mu}^{NP}, C_{9'\mu} = -C_{10'\mu})$$

$$\text{Hyp.5 : } (C_{9\mu}^{NP}, C_{9'\mu} = -C_{10'\mu})$$

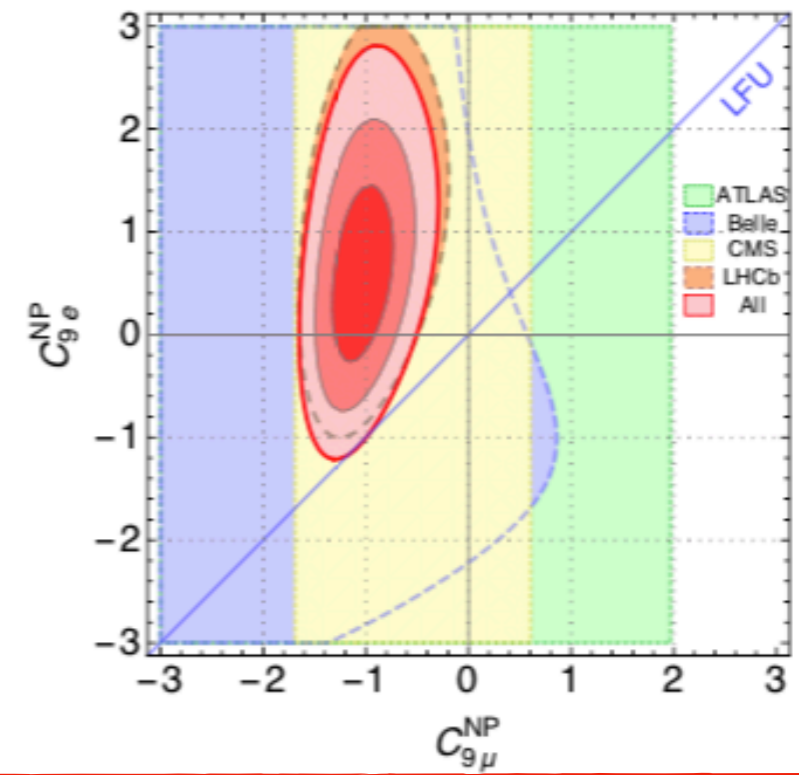
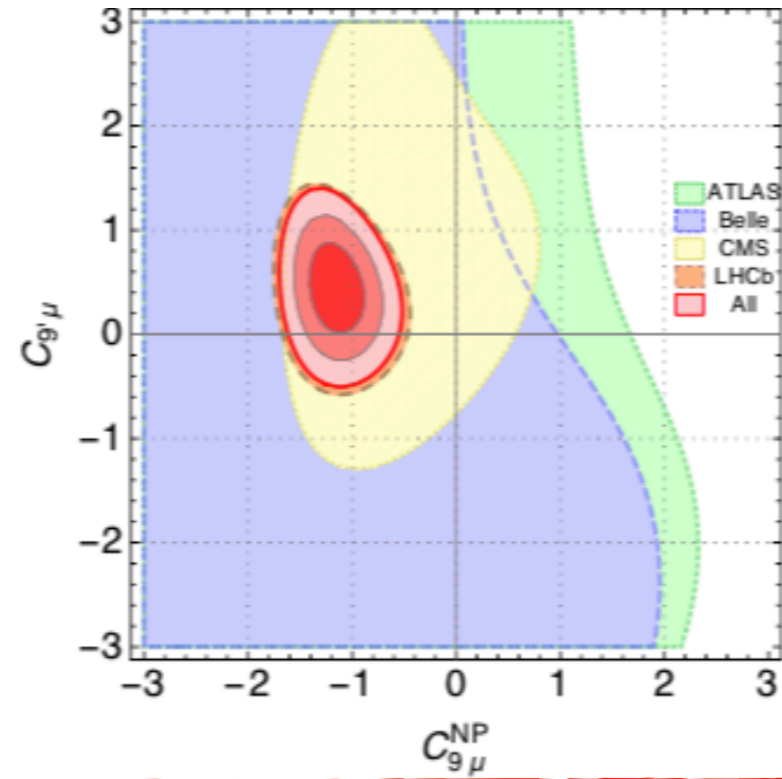
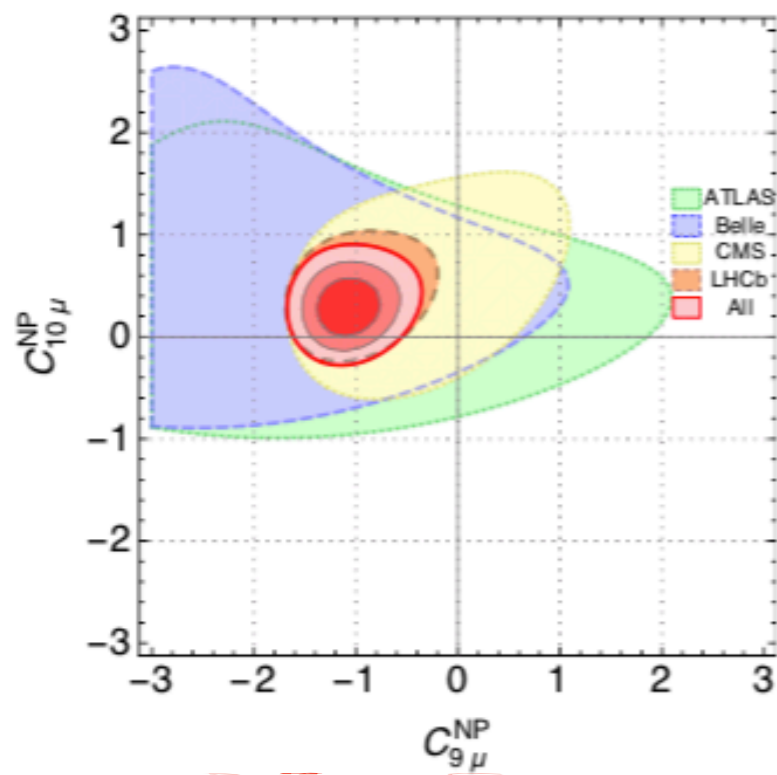
- Small increase ($\sim 0.2\sigma$) for some **RHC** scenarios opposite to 1D cases

- R_K closer to SM prefers $C_{9'\mu} > 0$ and $C_{10'\mu} < 0$

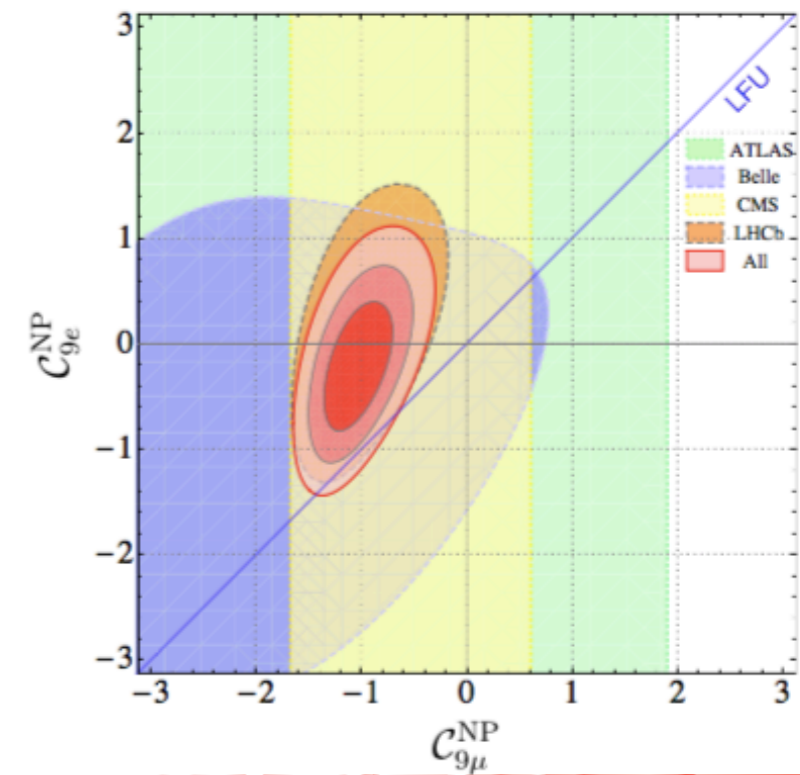
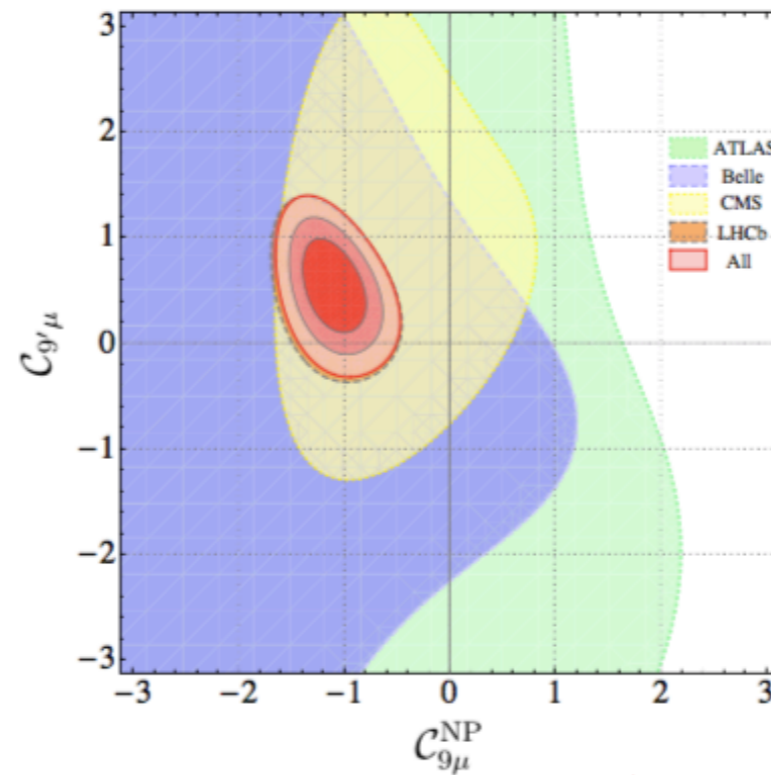
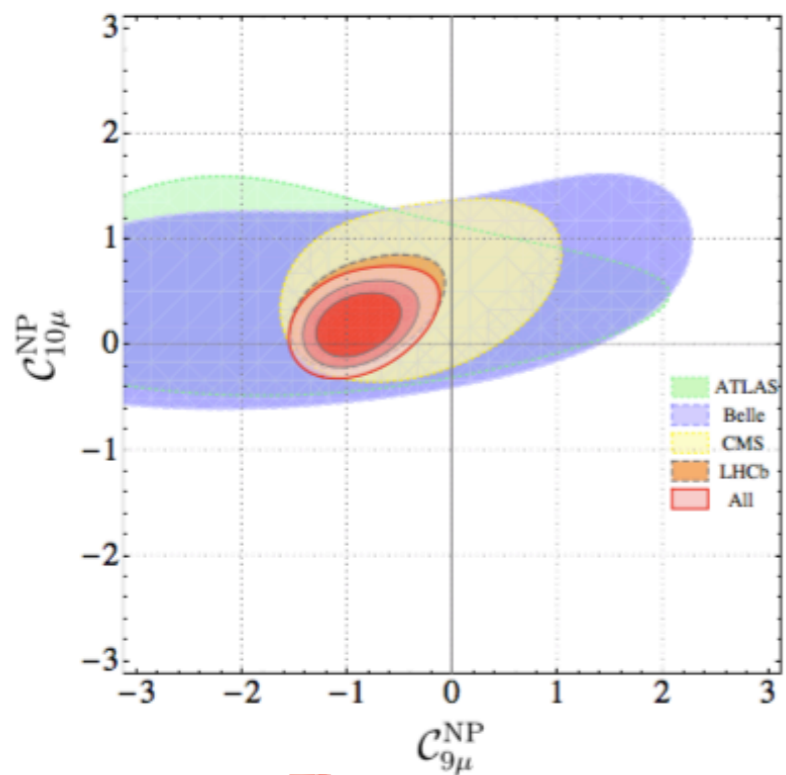
- A $R_q \otimes L_\ell$ for $\mathcal{O}_{9',10'}$ clearly prefers V structure over L_ℓ for leptons

- 2D Fits: 1σ CL for fit **All** and 3σ CL for individual experiments

[JHEP01(2018)093]



[1903.09578]



Fit Results

Concerning the 6D Fit:

[JHEP01(2018)093]

	C_7^{NP}	$C_{9\mu}^{\text{NP}}$	$C_{10\mu}^{\text{NP}}$	$C_{7'}$	$C_{9'\mu}$	$C_{10'\mu}$
Best fit	+0.03	-1.12	+0.31	+0.03	+0.38	+0.02
1 σ	[-0.01, +0.05]	[-1.34, -0.88]	[+0.10, +0.57]	[+0.00, +0.06]	[-0.17, +1.04]	[-0.28, +0.36]
2 σ	[-0.03, +0.07]	[-1.54, -0.63]	[-0.08, +0.84]	[-0.02, +0.08]	[-0.59, +1.58]	[-0.54, +0.68]

[1903.09578]

	C_7^{NP}	$C_{9\mu}^{\text{NP}}$	$C_{10\mu}^{\text{NP}}$	$C_{7'}$	$C_{9'\mu}$	$C_{10'\mu}$
Best fit	+0.01	-1.10	+0.15	+0.02	+0.36	-0.16
1 σ	[-0.01, +0.05]	[-1.28, -0.90]	[-0.00, +0.36]	[-0.00, +0.05]	[-0.14, +0.87]	[-0.39, +0.13]
2 σ	[-0.03, +0.06]	[-1.44, -0.68]	[-0.12, +0.56]	[-0.02, +0.06]	[-0.49, +1.23]	[-0.58, +0.33]

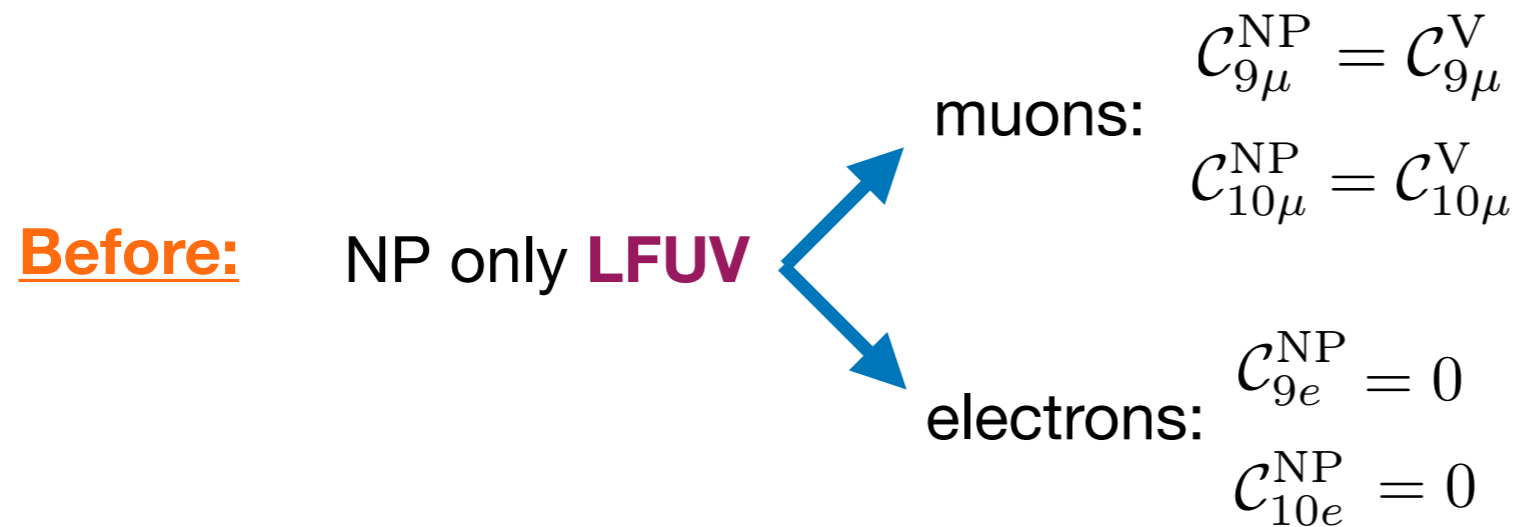
No significant changes except:

- 5.0 σ (2017) \longrightarrow **5.1 σ** (2019)
- Opposite sign for b.f.p. of $C_{10'}$ **but** $C_{10} - C_{10'}$ stays the same

What's new?

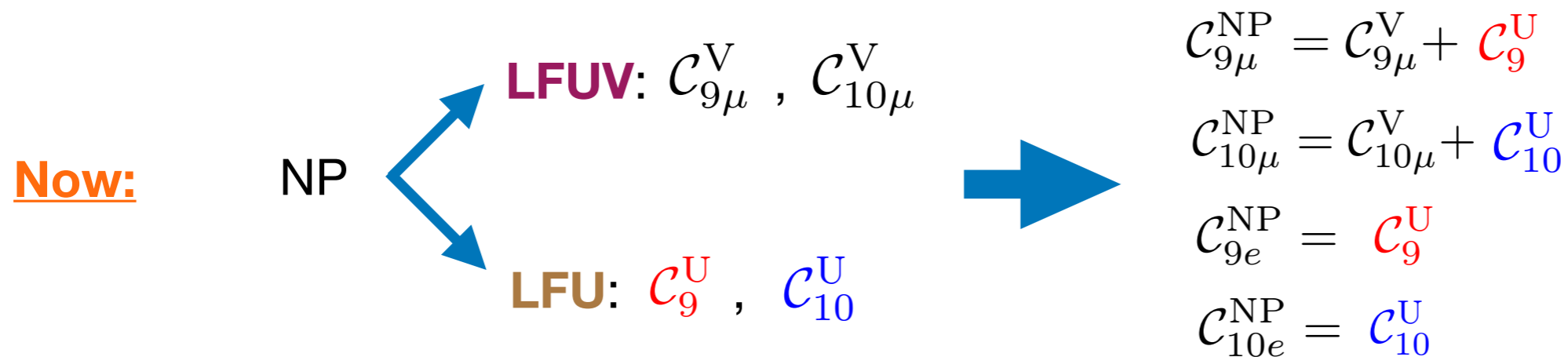
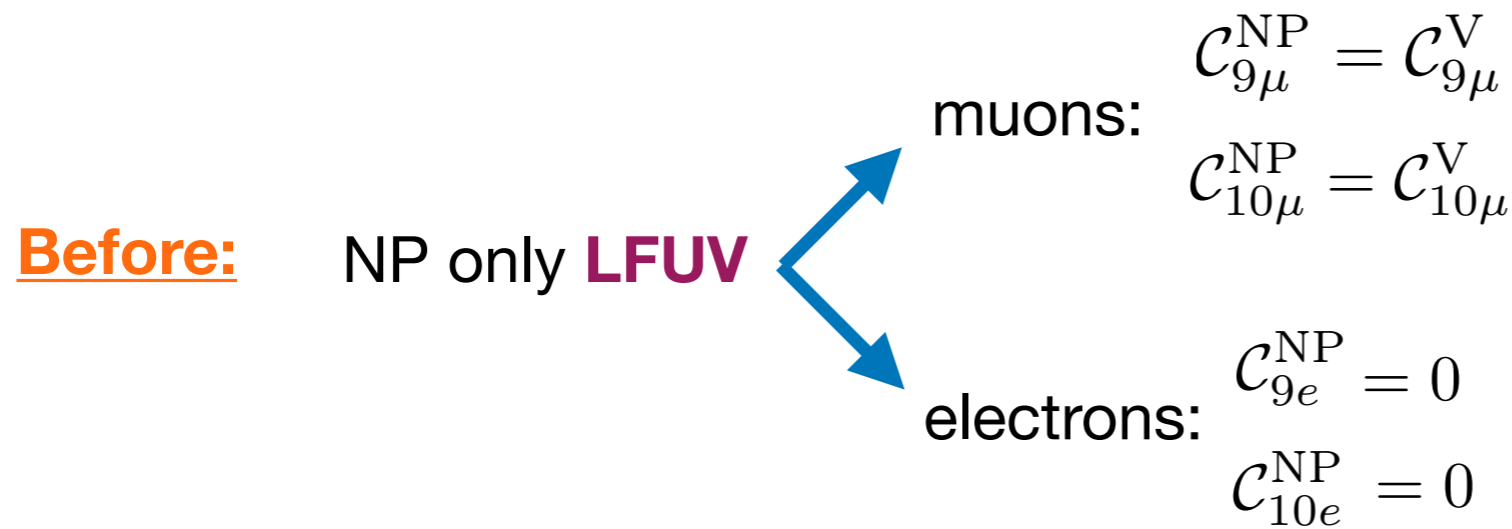
Are we overlooking LFU NP?

[Algueró, Capdevila, SDG, Masjuan, Matias, PRD'19]



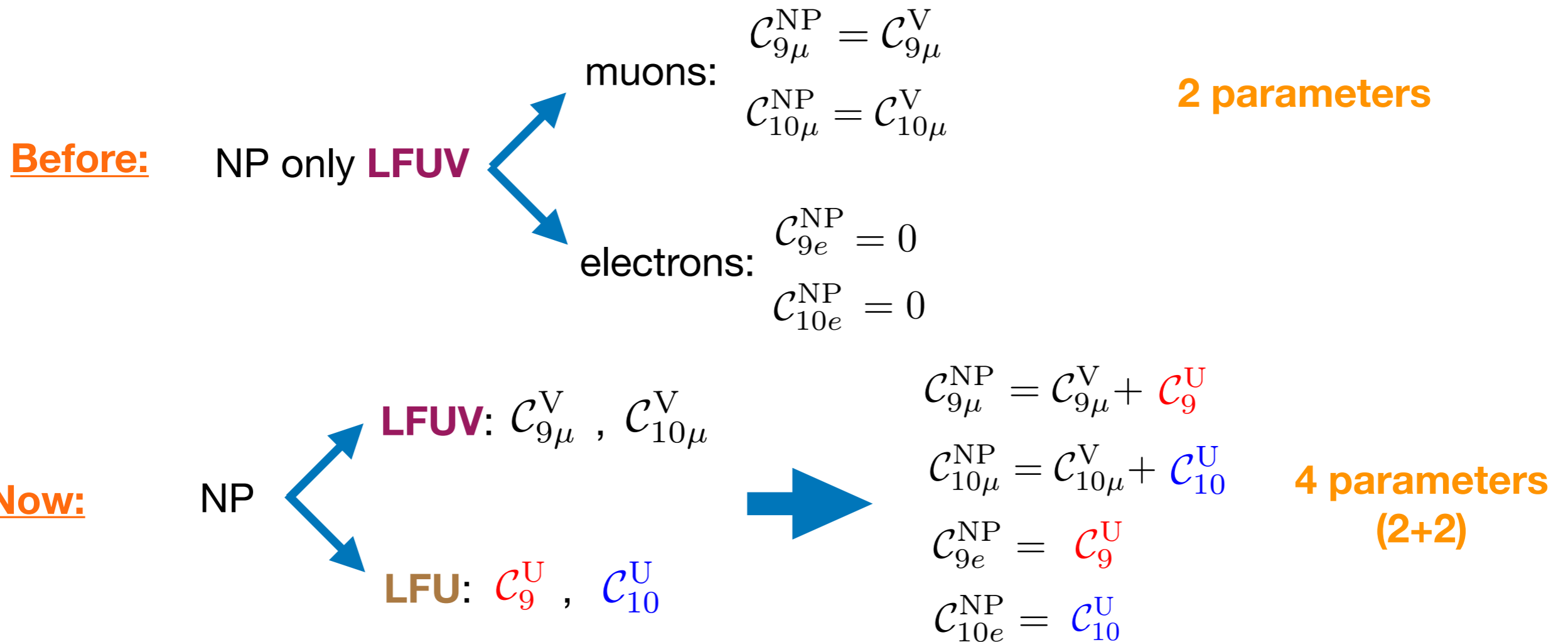
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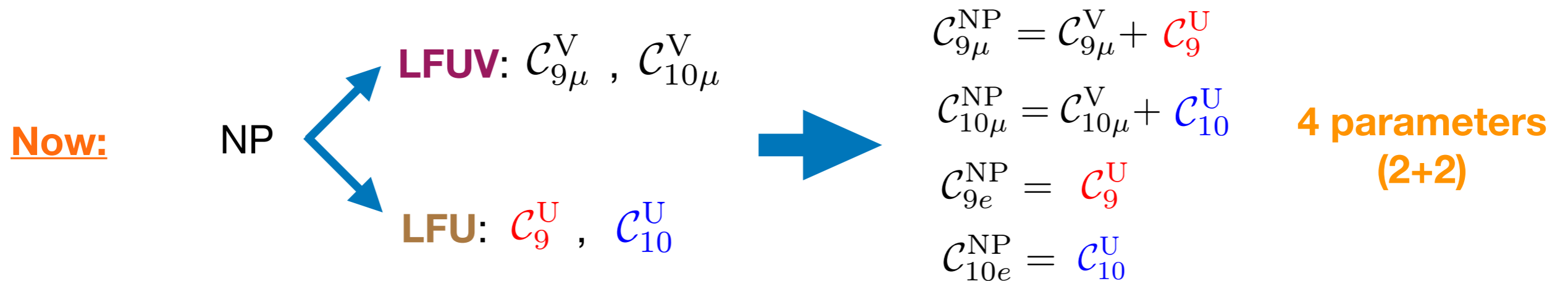
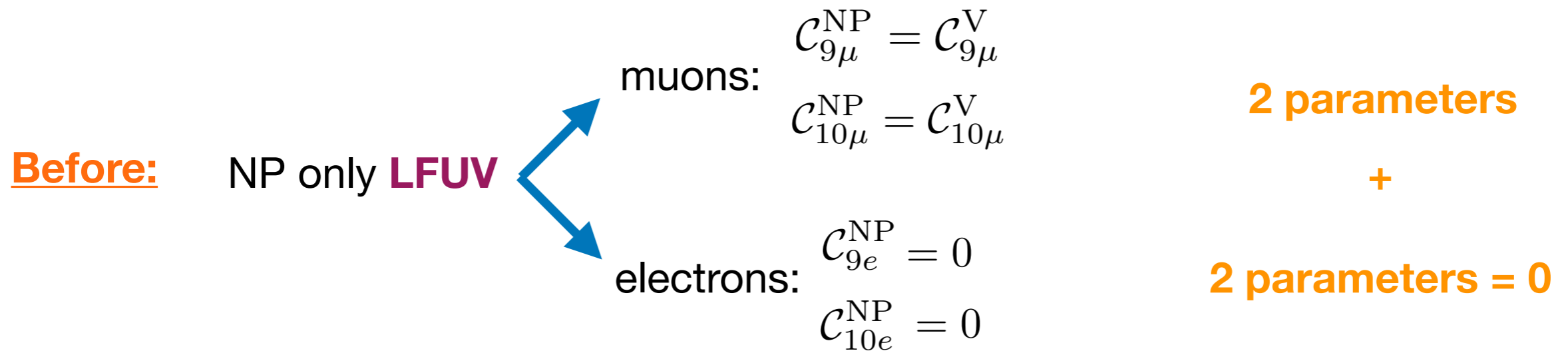
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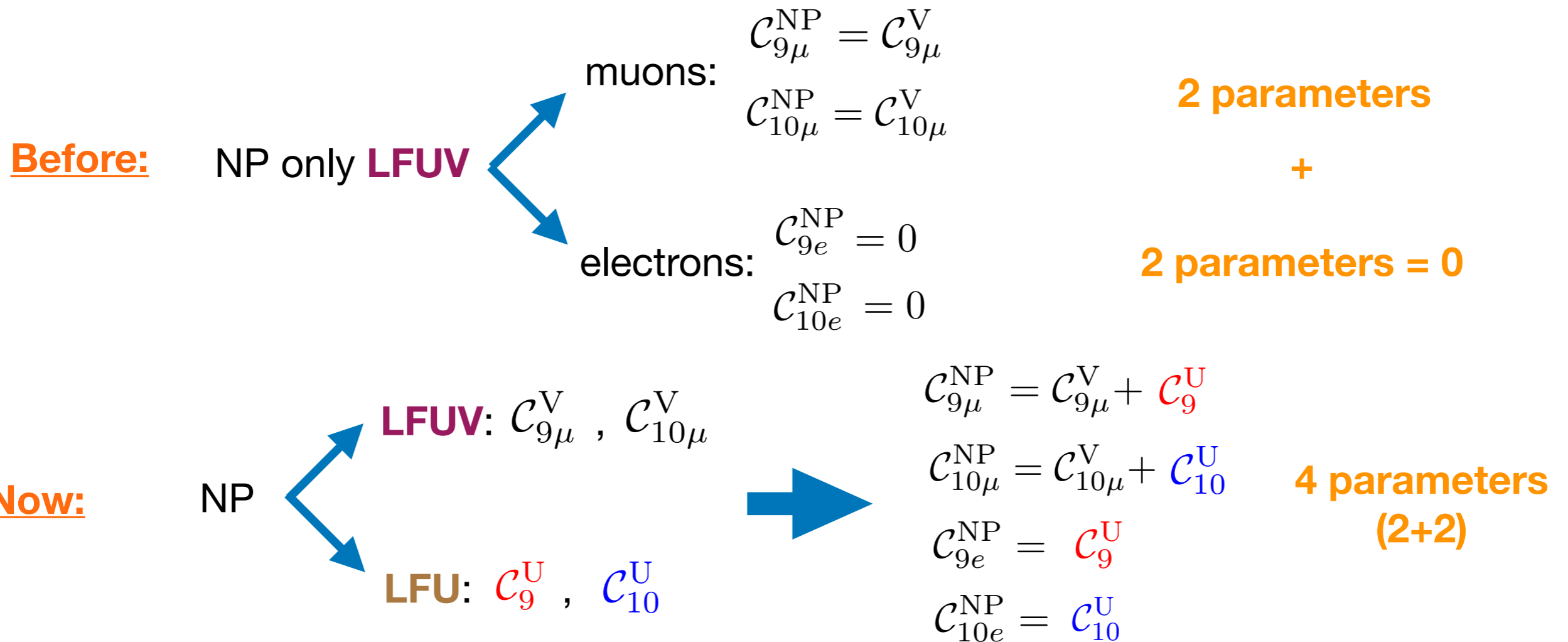
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Are we overlooking LFU NP?

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- Change of paradigm! Now same **LFU** piece in electrons, muons and tau different from **LFUV**
- Assumption: hadronic uncertainties well controlled $\rightarrow C_9^{\text{U}}$ is NP
- New mechanism to fulfill $B(B_s \rightarrow \mu^+ \mu^-) \propto |C_{10}^{\text{V}} + C_{10}^{\text{U}}|^2$

Are we overlooking LFU NP?

Update of scenarios with LFU

[Algueró et al, 1809.08447]

	Best-fit point	1σ	Pull _{SM}	p-value	
Sc. 5	$C_{9\mu}^V$	-0.16	[-0.94, +0.46]	5.8	78%
	$C_{10\mu}^V$	+1.00	[+0.18, +1.59]		
	$C_9^U = C_{10}^U$	-0.87	[-1.43, -0.14]		
Sc. 6	$C_{9\mu}^V = -C_{10\mu}^V$	-0.64	[-0.77, -0.51]	6.0	79%
	$C_9^U = C_{10}^U$	-0.44	[-0.58, -0.29]		
Sc. 7	$C_{9\mu}^V$	-1.57	[-2.14, -1.06]	5.7	72%
	C_9^U	+0.56	[+0.01, +1.15]		
Sc. 8	$C_{9\mu}^V = -C_{10\mu}^V$	-0.42	[-0.57, -0.27]	5.8	74%
	C_9^U	-0.67	[-0.90, -0.42]		

- Significances remain basically unchanged
- Sc. 6 (LFUV V-A, LFU V+A) still describes data well
- Sc. 7 **changed** → preference for LFUV- C_9 if only V contributions allowed
- Sc. 8 LFUV left-handed lepton structure provides better description with LFU in C_9

[Algueró et al, 1903.09578]

	Best-fit point	1σ	2σ	Pull _{SM}	p-value
Sc. 5	$C_{9\mu}^V$	-0.36	[-0.86, +0.10]	5.2	71.2%
	$C_{10\mu}^V$	+0.67	[+0.24, +1.03]		
	$C_9^U = C_{10}^U$	-0.59	[-0.90, -0.12]		
Sc. 6	$C_{9\mu}^V = -C_{10\mu}^V$	-0.50	[-0.61, -0.38]	5.5	71.0%
	$C_9^U = C_{10}^U$	-0.38	[-0.52, -0.22]		
Sc. 7	$C_{9\mu}^V$	-0.78	[-1.11, -0.47]	5.3	66.2%
	C_9^U	-0.20	[-0.57, +0.18]		
Sc. 8	$C_{9\mu}^V = -C_{10\mu}^V$	-0.30	[-0.42, -0.20]	5.7	75.2%
	C_9^U	-0.74	[-0.96, -0.51]		

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[Algueró et al, 1809.08447]

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LFU-NP dependent on structure of LFUV-NP

Are we overlooking LFU NP?

New scenarios with **LFU** in RHC

Scenario 9	$C_{9\mu}^V = -C_{10\mu}^V$	-0.57	[-0.73, -0.41]	[-0.87, -0.28]	5.0	60.2 %
	C_{10}^U	-0.34	[-0.60, -0.07]	[-0.84, +0.18]		
Scenario 10	$C_{9\mu}^V$	-0.95	[-1.13, -0.76]	[-1.30, -0.57]	5.5	69.5 %
	C_{10}^U	+0.27	[0.08, 0.47]	[-0.09, 0.66]		
Scenario 11	$C_{9\mu}^V$	-1.03	[-1.22, -0.84]	[-1.38, -0.65]	5.6	73.6 %
	$C_{10'}^U$	-0.29	[-0.47, -0.12]	[-0.63, 0.05]		
Scenario 12	$C_{9'\mu}^V$	-0.03	[-0.22, 0.15]	[-0.40, 0.32]	1.6	15.7 %
	C_{10}^U	+0.41	[0.21, 0.63]	[0.02, 0.83]		
Scenario 13	$C_{9\mu}^V$	-1.11	[-1.28, -0.91]	[-1.41, -0.71]	5.4	78.7 %
	$C_{9'\mu}^V$	+0.53	[0.24, 0.83]	[-0.10, 1.11]		
	C_{10}^U	+0.24	[0.01, 0.48]	[-0.21, 0.69]		
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$C_{10^{(\prime)}}^U$ contributions naturally from modified Z couplings



- Sc. 9 via Two-Higgs Doublet
- Sc. 11-13 via vector-like quarks @ tree level

- Extension of **LFU** contributions to RHC (motivated by 2D results) \Rightarrow natural connection to models
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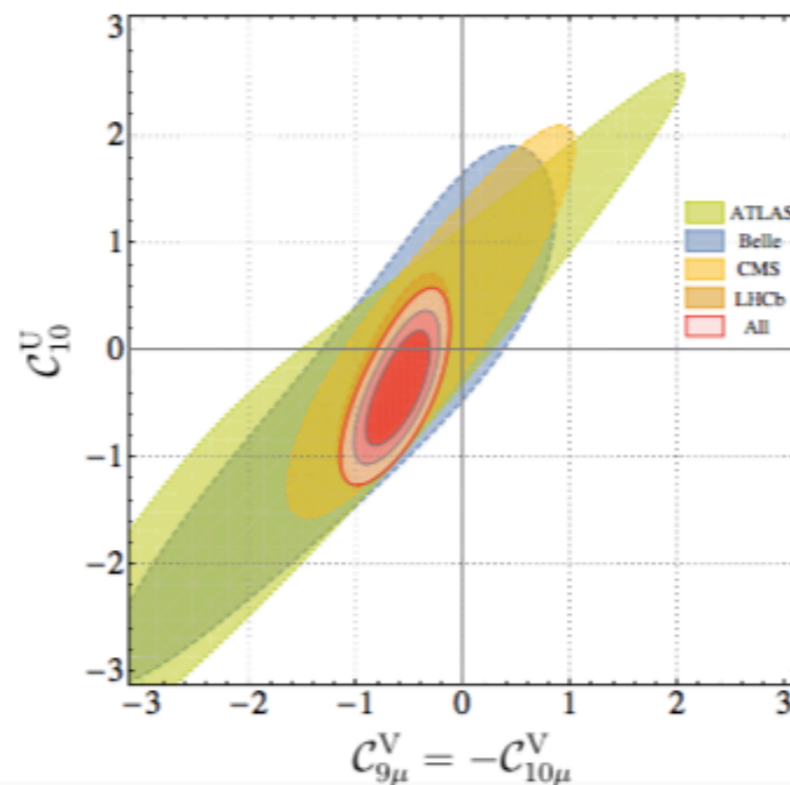
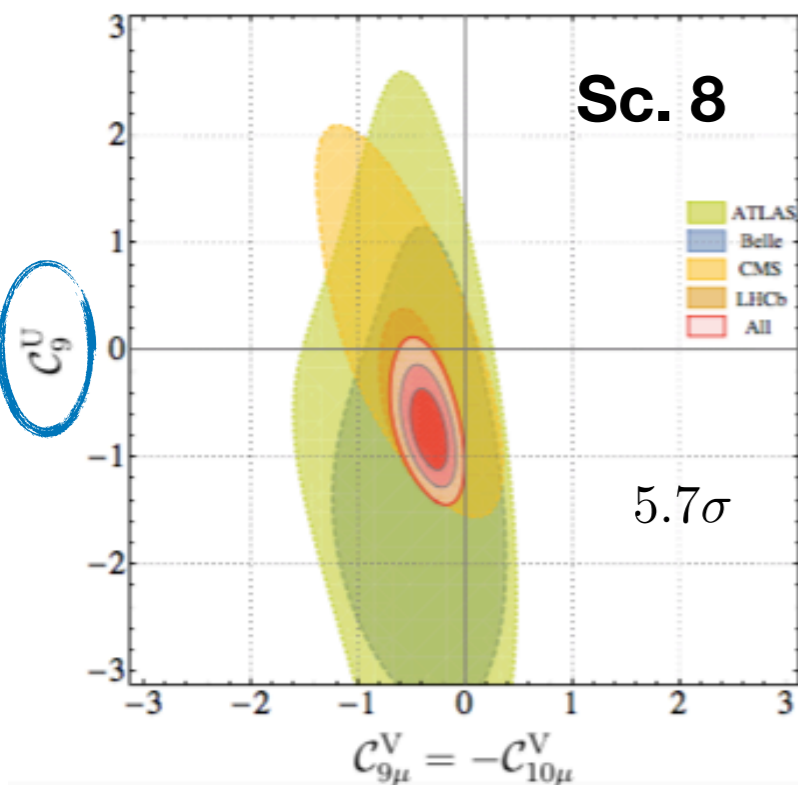
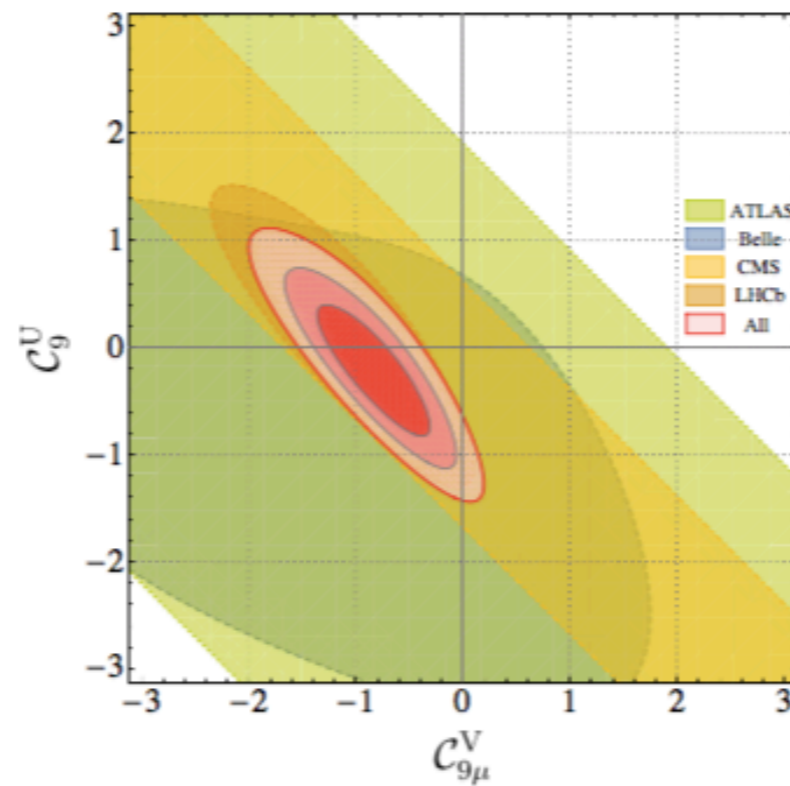
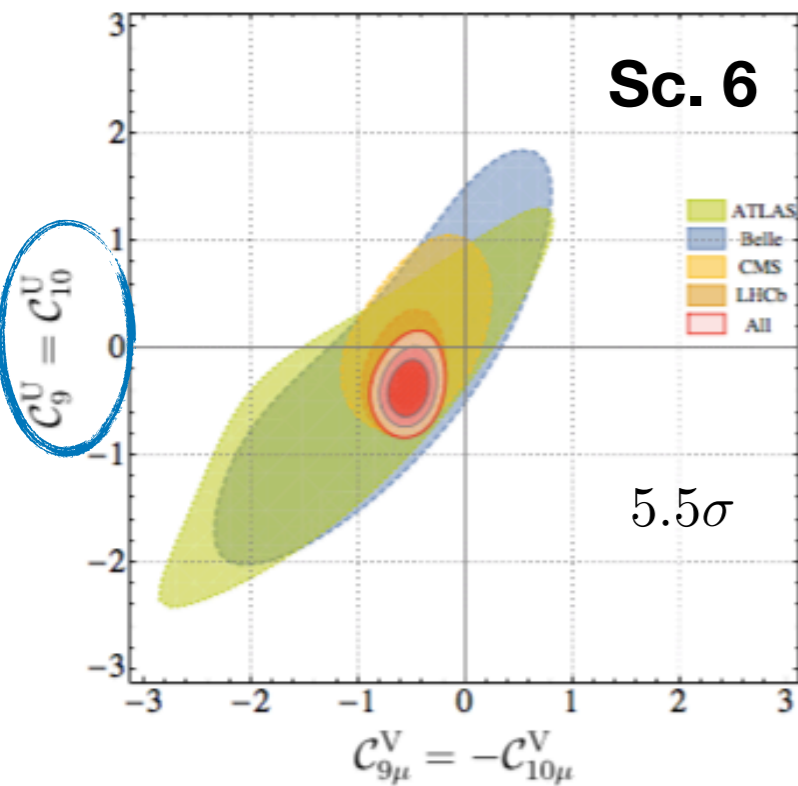
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Are we overlooking LFU NP?



Assuming loop-level scale of NP and no MFV

$$\Lambda_i^L \sim \frac{v}{s_w g} \frac{1}{\sqrt{2|V_{tb}V_{ts}^*|}} \frac{1}{|C_i^{NP}|^{1/2}}$$

Mild preference

Scenario 6:

$$\begin{aligned} C_{9\mu}^V &= -C_{10\mu}^V \\ C_9^U &= C_{10}^U \end{aligned}$$

LFUV-NP $L_q \otimes L_\ell$

$$\Lambda_i^{\text{LFUV}} \sim 4.0 \text{ TeV}$$

LFU-NP $L_q \otimes R_\ell$

$$\Lambda_i^{\text{LFU}} \sim 4.6 \text{ TeV}$$

Scenario 8:

$$\begin{aligned} C_{9\mu}^V &= -C_{10\mu}^V \\ C_9^U & \end{aligned}$$

LFUV-NP $L_q \otimes L_\ell$

$$\Lambda_i^{\text{LFUV}} \sim 5.1 \text{ TeV}$$

LFU-NP $L_q \otimes V_\ell$

$$\Lambda_i^{\text{LFU}} \sim 3.3 \text{ TeV}$$

[Matias, 19th Lomonosov Conf.]

Discussion on models

Discussion on models

Framework: SMEFT $\rightarrow \Lambda_{\text{NP}} \gg m_{t,W,Z}$ [Grzadkowski, Iskrzynski, Misiak, Rosiek; Alonso, Grinstein, Camalich]

Model independent connection between $b \rightarrow sll$ and $b \rightarrow c\tau\nu$ anomalies through **Sc. 8** (3.1 σ)

$$\begin{aligned} c_{9\mu}^V &= -c_{10\mu}^V \\ c_9^U & \end{aligned}$$



- SMEFT with $c^{(1)} = c^{(3)}$ in terms of $SU(2)_L$ inv. operators

$$\begin{cases} \mathcal{O}_{ijkl}^{(1)} = [\bar{Q}_i \gamma_\mu Q_j][\bar{L}_k \gamma^\mu L_l] \\ \mathcal{O}_{ijkl}^{(3)} = [\bar{Q}_i \gamma_\mu \sigma^I Q_j][\bar{L}_k \gamma^\mu \sigma^I L_l] \end{cases}$$

[Capdevila, Crivellin, SDG, Hofer, JM, PRL'18]

3rd lepton generation Op. $\rightarrow R_{D^{(*)}}$

2nd lepton generation Op. \rightarrow LFUV effect in $b \rightarrow s\mu^+\mu^-$

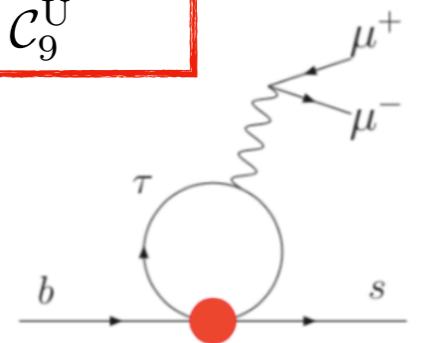
- Constraint from $b \rightarrow c\tau\nu$ + invariance \rightarrow enhance $b \rightarrow s\tau^+\tau^-$ **But** also mixes into \mathcal{O}_9



Rad Corr w/ insertion of $\mathcal{O}_{2333} \rightarrow c_9^U$

Assuming **generic flavour structure:**

$$c_9^U \approx 7.5 \left(1 - \sqrt{\frac{R_{D^{(*)}}}{R_{D^{(*)}\text{SM}}}} \right) \left(1 + \frac{\log(\Lambda^2/(1\text{TeV}^2))}{10.5} \right)$$



Discussion on models

$$\Lambda_{\text{NP}} = 2 \text{ TeV}$$

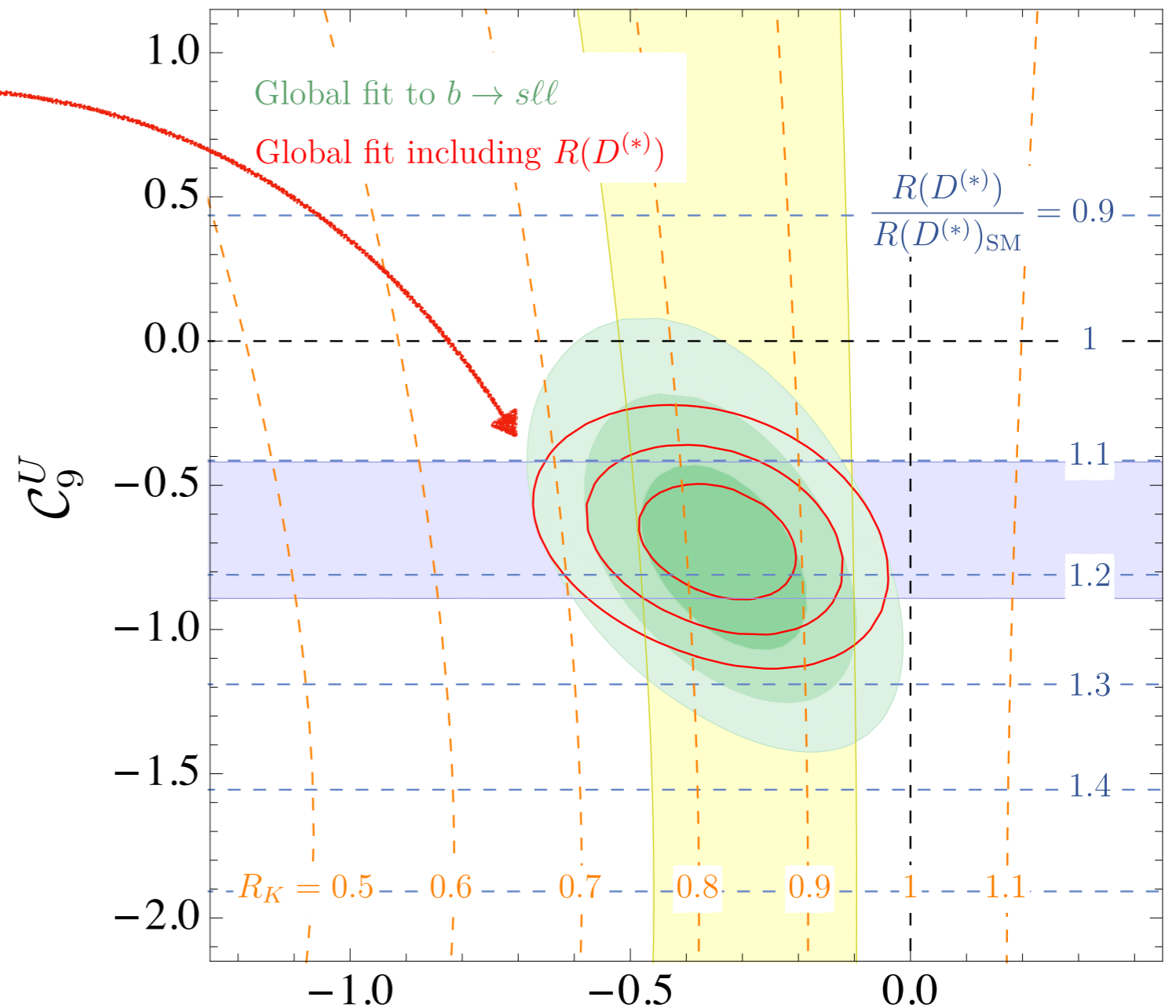
1-3 σ CL Regions including $R_{D^{(*)}}$ from Belle



Now Sc. 8 has **7 σ** due to inclusion of $R_{D^{(*)}}$

See explicit UV completions by A. Crivellin et al PRL 2019

Blue (yellow) band consistent with $R_{D^{(*)}}(R_K)$ @ 2 σ



Good agreement with results from Aebischer et al, arXiv:1903.10434

A possible candidate?

A very promising candidate:

Vector leptoquark $SU(2)$ singlet:

$$U_1(3, 1, 2/3)$$

Coupled mainly to 3rd generation

See talk from J. Fuentes

Why:

- 1) Explains charged and neutral anomalies @ same time
- 2) Allows $C_{9\mu}^V = -C_{10\mu}^V$ and generates C_9^U
- 3) Avoids contributions to $b \rightarrow s\nu\bar{\nu}$
- 4) Safe concerning LHC searches

Conclusions and Outlook

- ▶ Coherent pattern of deviations in $b \rightarrow s\ell\ell$ and also **LFUV** (among other flavour anomalies)
- ▶ **Similar big picture (and pattern of NP)** after LHCb and Belle new measurements:
 - ▶ 1D hypotheses show same hierarchy (Global fit better explained by $C_{9\mu}^{\text{NP}}$ while LFUV observables prefer $C_{9\mu}^{\text{NP}} = -C_{10\mu}^{\text{NP}}$)
 - ▶ New emerging scenarios including RHC seen in 2D Fits
- ▶ New scope for NP \longrightarrow 2 sectors: **LFU+LFUV** \longrightarrow same level of significance
- ▶ **LFU-NP** structure dependent on **LFUV-NP** structure ($C_{9\mu}^{\text{V}} = -C_{10\mu}^{\text{V}}$ provides good description only in presence of C_9^{U}) \longrightarrow Need strategy to disentangle (Q_5)
- ▶ Promising candidate \longrightarrow Vector LeptoQuark **U(3,1,2/3)**

Thanks!

Back-up

Are we overlooking LFU NP?

- Change of paradigm! Now same **LFU** piece in electrons, muons and tau different from **LFUV**
- Assumption: hadronic uncertainties well controlled $\longrightarrow C_9^U$ is NP
- New mechanism to fulfill $B(B_s \rightarrow \mu^+ \mu^-) \propto |C_{10}^V + C_{10}^U|^2$

LFUV observables explained mainly by $C_{i\mu}^V$

LFD observables (like P_5') explained by $C_{i\mu}^V + C_i^U$

Our strategy: 4D \longrightarrow 3D \longrightarrow 2D

[Algueró et al, 1809.08447]

$$4D \rightarrow \{C_{9\mu}^V, C_{10\mu}^V, C_9^U, C_{10}^U\}$$

$$3D \rightarrow \{C_{9\mu}^V = -C_{10\mu}^V, C_9^U, C_{10}^U\}$$

$$3D \rightarrow \{C_{9\mu}^V, C_{10\mu}^V, C_9^U = C_{10}^U\}$$

$$2D \rightarrow \{C_{9\mu}^V = -C_{10\mu}^V, C_9^U = C_{10}^U\}$$

$$2D \rightarrow \{C_{9\mu}^V, C_9^U\}$$

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Correlation Matrices Fits

$$\text{Corr}(\mathcal{C}_{9\mu}^{\text{NP}}, \mathcal{C}_{10\mu}^{\text{NP}}) = \begin{pmatrix} 1.00 & 0.30 \\ 0.30 & 1.00 \end{pmatrix}$$

$$\text{Corr}(\mathcal{C}_{9\mu}^{\text{NP}} = -\mathcal{C}_{9'\mu}, \mathcal{C}_{10\mu}^{\text{NP}} = \mathcal{C}_{10'\mu}) = \begin{pmatrix} 1.00 & -0.17 \\ -0.17 & 1.00 \end{pmatrix}$$

$$\text{Corr}(\mathcal{C}_{9\mu}^{\text{NP}}, \mathcal{C}_{9'\mu}) = \begin{pmatrix} 1.00 & -0.39 \\ -0.39 & 1.00 \end{pmatrix}$$

$$\text{Corr}(\mathcal{C}_{9\mu}^{\text{NP}}, \mathcal{C}_{9'\mu} = -\mathcal{C}_{10'\mu}) = \begin{pmatrix} 1.00 & -0.34 \\ -0.34 & 1.00 \end{pmatrix}$$

$$\text{Corr}(\mathcal{C}_{9\mu}^{\text{NP}}, \mathcal{C}_{10'\mu}) = \begin{pmatrix} 1.00 & 0.33 \\ 0.33 & 1.00 \end{pmatrix}$$

$$\text{Corr}(\mathcal{C}_{9\mu}^{\text{NP}}, \mathcal{C}_{9e}^{\text{NP}}) = \begin{pmatrix} 1.00 & 0.51 \\ 0.51 & 1.00 \end{pmatrix}$$

$$\text{Corr}_{6\text{D}} = \begin{pmatrix} 1.00 & -0.34 & -0.07 & 0.06 & 0.02 & -0.03 \\ -0.34 & 1.00 & 0.24 & -0.06 & 0.04 & 0.24 \\ -0.07 & 0.24 & 1.00 & -0.13 & 0.61 & 0.59 \\ 0.06 & -0.06 & -0.13 & 1.00 & -0.13 & -0.08 \\ 0.02 & 0.04 & 0.61 & -0.13 & 1.00 & 0.85 \\ -0.03 & 0.24 & 0.59 & -0.08 & 0.85 & 1.00 \end{pmatrix}$$

Correlation Matrices Fits

$$\text{Corr}(C_{9\mu}^V, C_9^U = C_{10}^U, C_{10\mu}^V) = \begin{pmatrix} 1.00 & -0.93 & 0.91 \\ -0.93 & 1.00 & -0.94 \\ 0.91 & -0.94 & 1.00 \end{pmatrix}$$

$$\text{Corr}(C_{9\mu}^V = -C_{10\mu}^V, C_{10}^U) = \begin{pmatrix} 1.00 & 0.69 \\ 0.69 & 1.00 \end{pmatrix}$$

$$\text{Corr}(C_{9\mu}^V = -C_{10\mu}^V, C_9^U = C_{10}^U) = \begin{pmatrix} 1.00 & 0.17 \\ 0.17 & 1.00 \end{pmatrix}$$

$$\text{Corr}(C_{9\mu}^V, C_{10}^U) = \begin{pmatrix} 1.00 & 0.05 \\ 0.05 & 1.00 \end{pmatrix}$$

$$\text{Corr}(C_{9\mu}^V, C_9^U) = \begin{pmatrix} 1.00 & -0.85 \\ -0.85 & 1.00 \end{pmatrix}$$

$$\text{Corr}(C_{9\mu}^V, C_{10'}^U) = \begin{pmatrix} 1.00 & 0.20 \\ 0.20 & 1.00 \end{pmatrix}$$

Possible UV completions

- $SU(4) \times SU(3)' \times SU(2)_L \times U(1)_Y$ + Vector-like fermions
L. Di Luzio, A. Greljo, M. Nardecchia, arXiv:1708.08450
- $SU(4) \times U(2)_L \times SU(2)_R$ + vector-like fermions
L. Calibbi, AC, T. Li, arXiv:1709.00692
- $SU(4) \times SU(4) \times SU(4)$
M. Bordone, C. Cornella, J. Fuentes-Martin, G. Isidori, arXiv:1712.01368
- $SU(4) \times U(2)_L \times SU(2)_R$ including scalar LQs and light right-handed neutrinos
J. Heeck, D. Teresi, arXiv:1808.07492
- $SU(8)$ might even explain ε'/ε
S. Matsuzaki, K. Nishiwaki and K. Yamamoto, arXiv:1806.02312
- $SU(4) \times U(2) \times SU(2)_R$ in RS background
M. Blanke, AC, arXiv:1801.07256

Towards a decision tree:

R_K and Q_5

Disentangling Scenarios

Significances very similar \longrightarrow What is the best scenario?

Is there any observable that could disentangle among scenarios/hypotheses?

$$Q_{4,5} = P'_{4,5}{}^{\mu} - P'_{4,5}{}^e$$

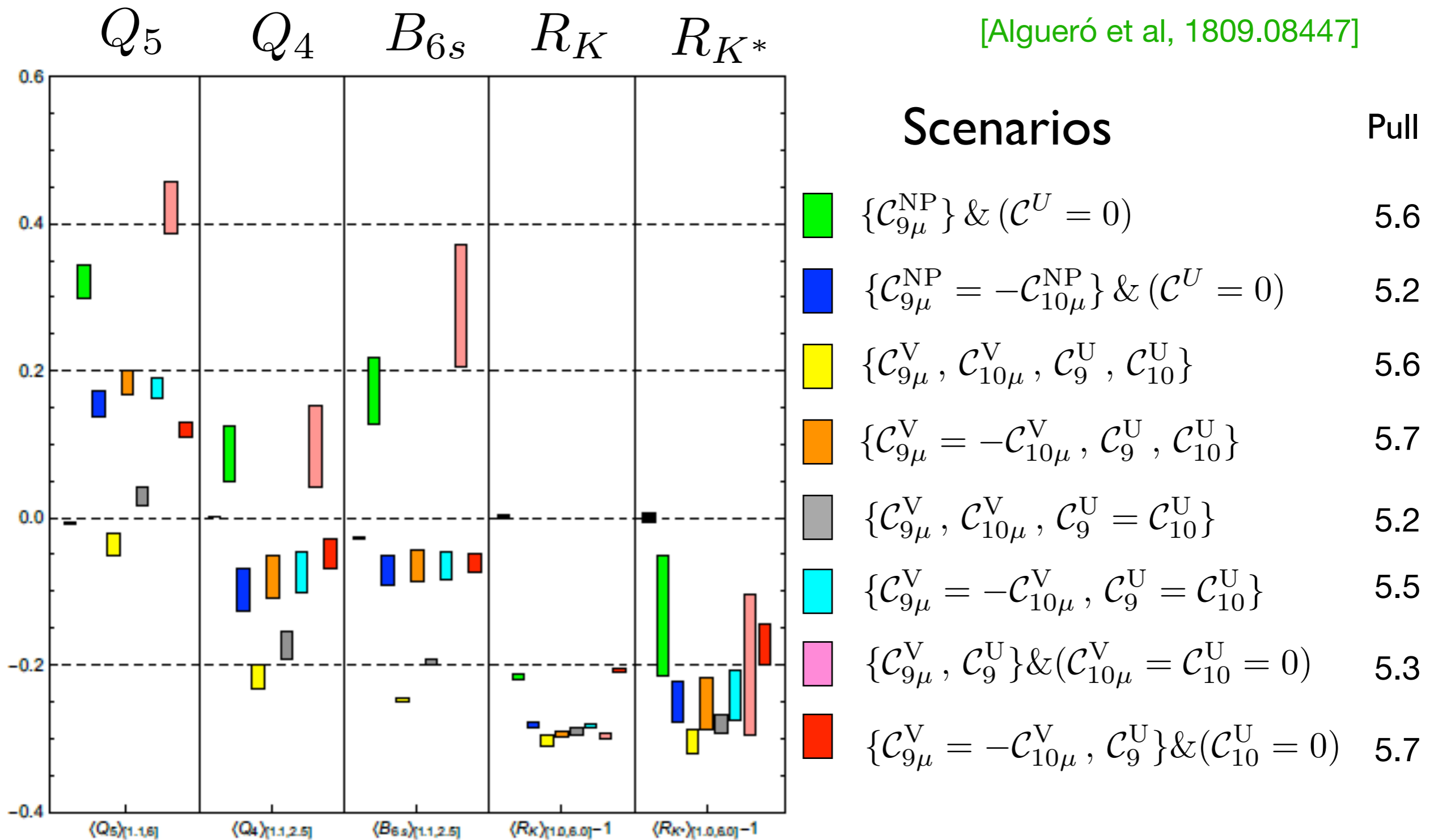
Several **LFUV** observables:

$$B_{6s} = \frac{J_{6s}^{\mu}}{J_{6s}^e} - 1$$

$$R_{K^{(*)}} = \frac{BR(B \rightarrow K^{(*)}\mu^+\mu^-)}{BR(B \rightarrow K^{(*)}e^+e^-)}$$

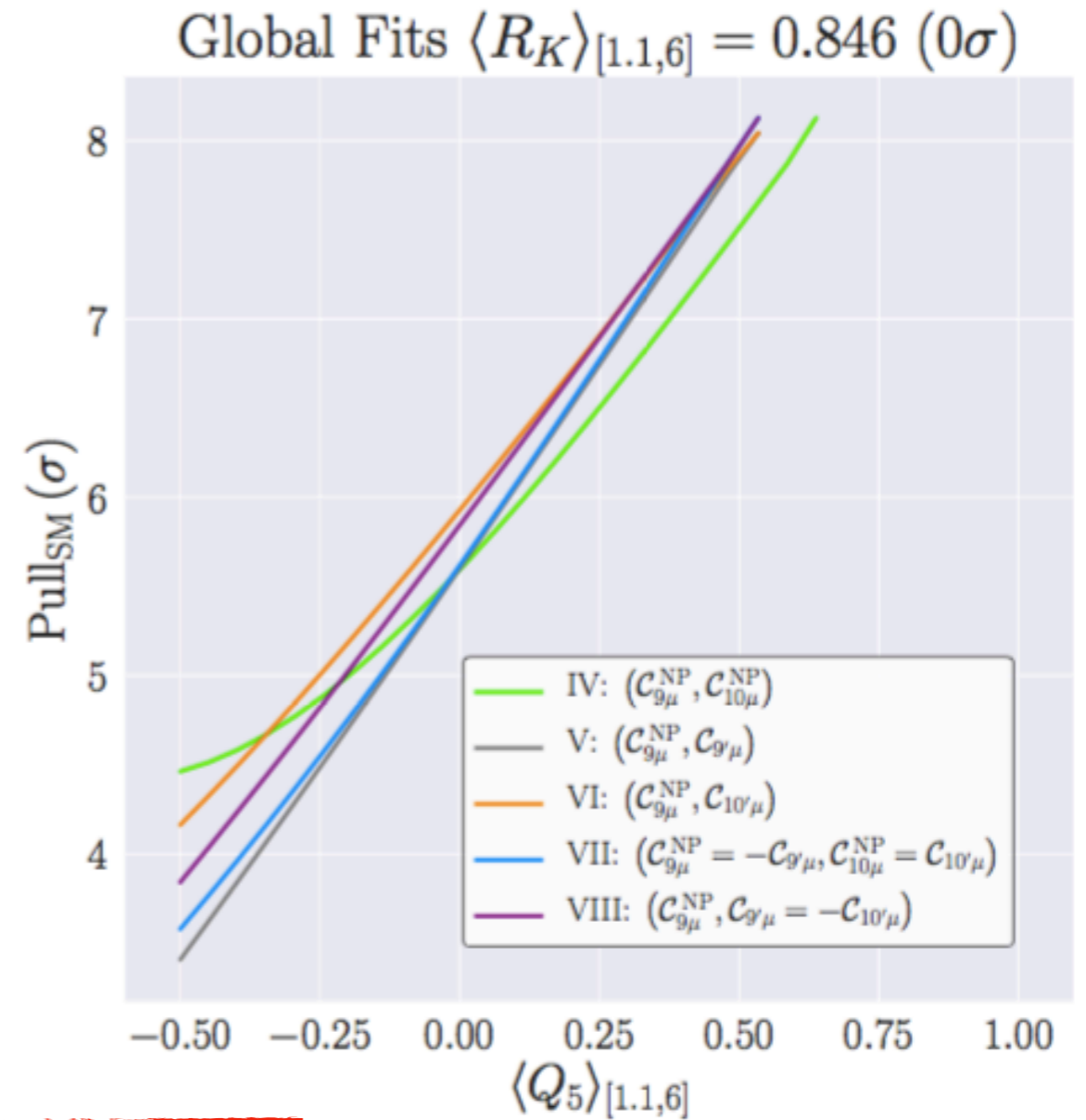
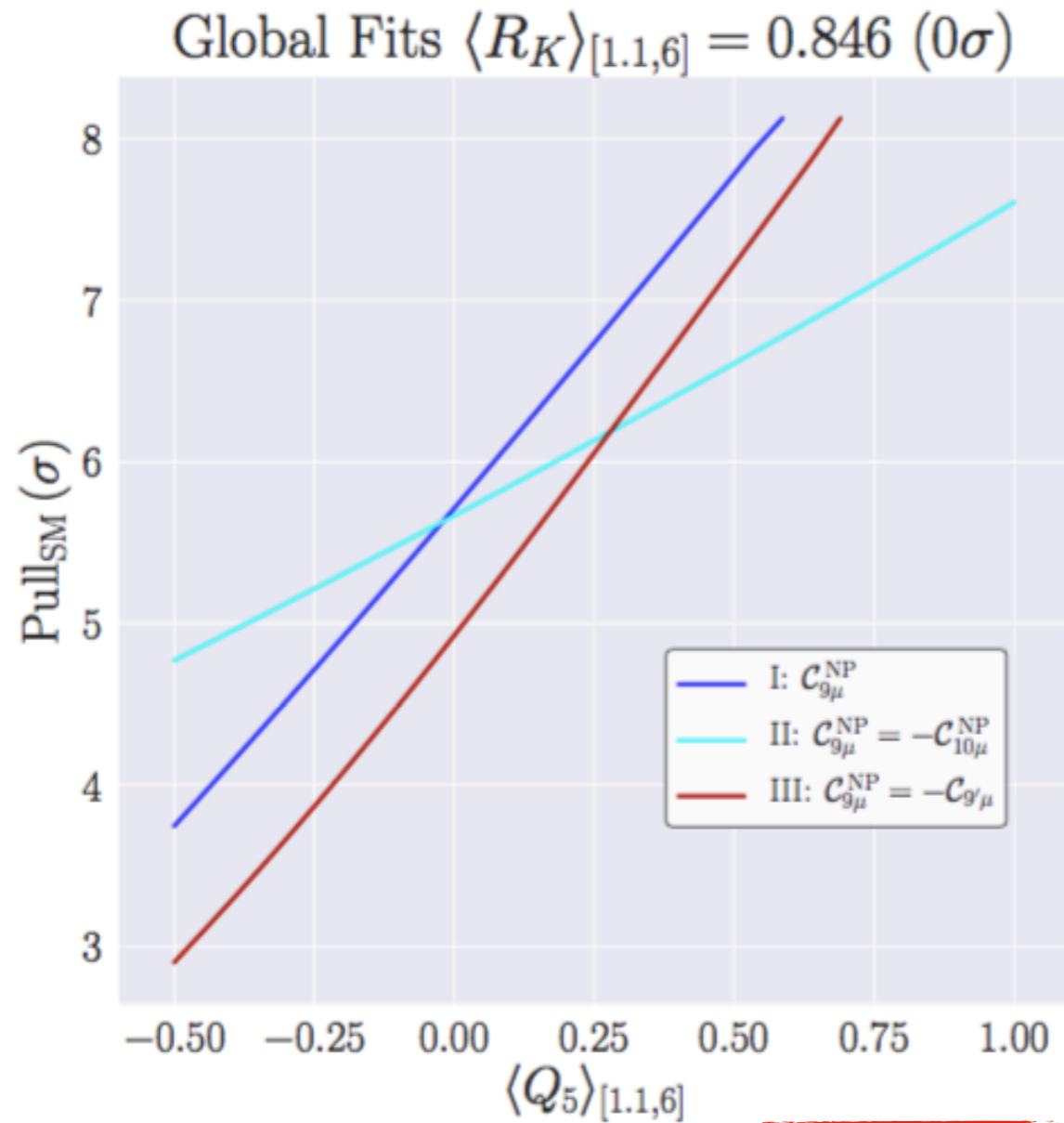
Disentangling Scenarios

[Algueró et al, 1809.08447]



Disentangling Scenarios

R_K cannot disentangle among scenarios **BUT** together with Q_5 : [Algueró et al JHEP07 (2019) 096]

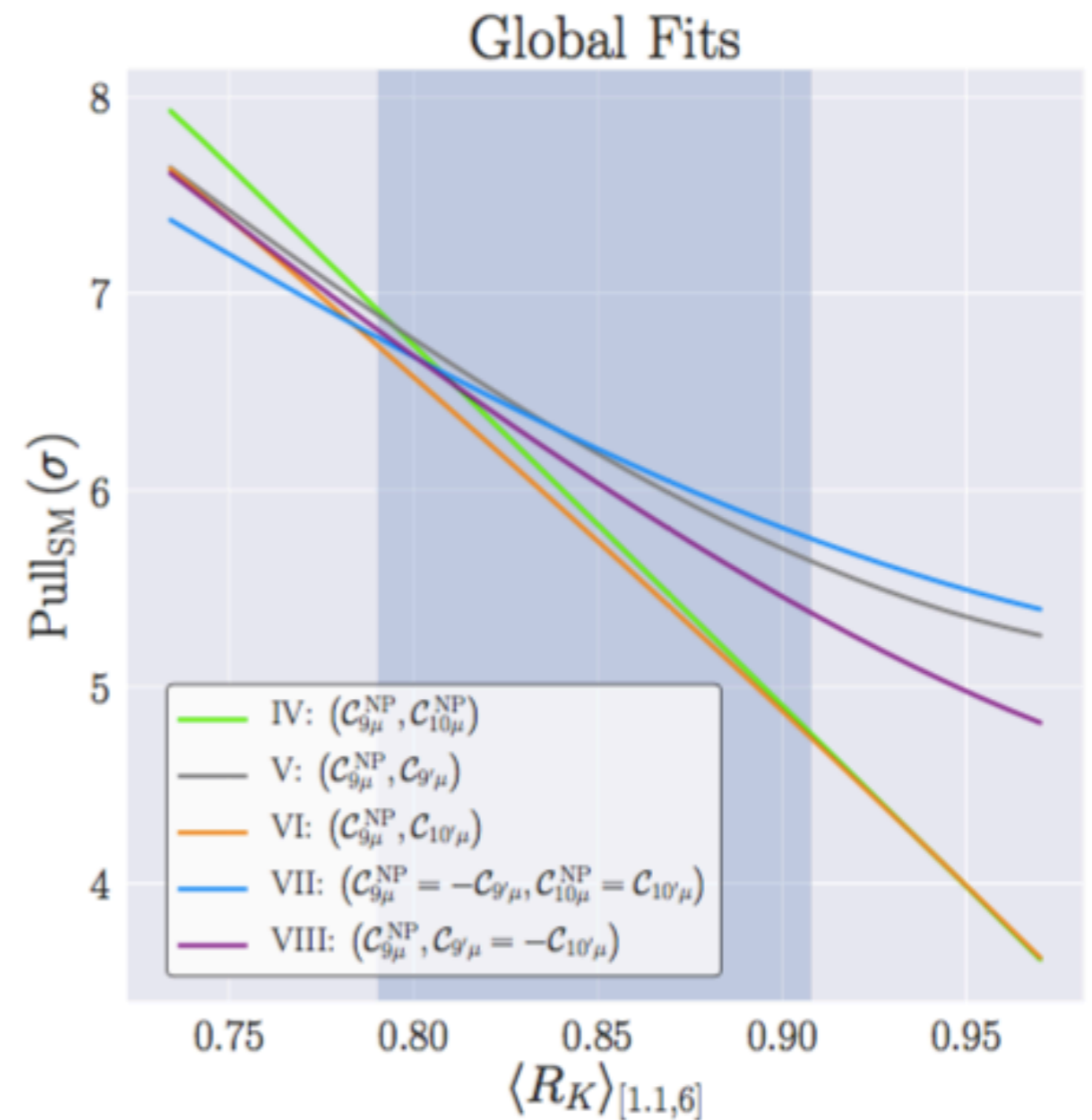
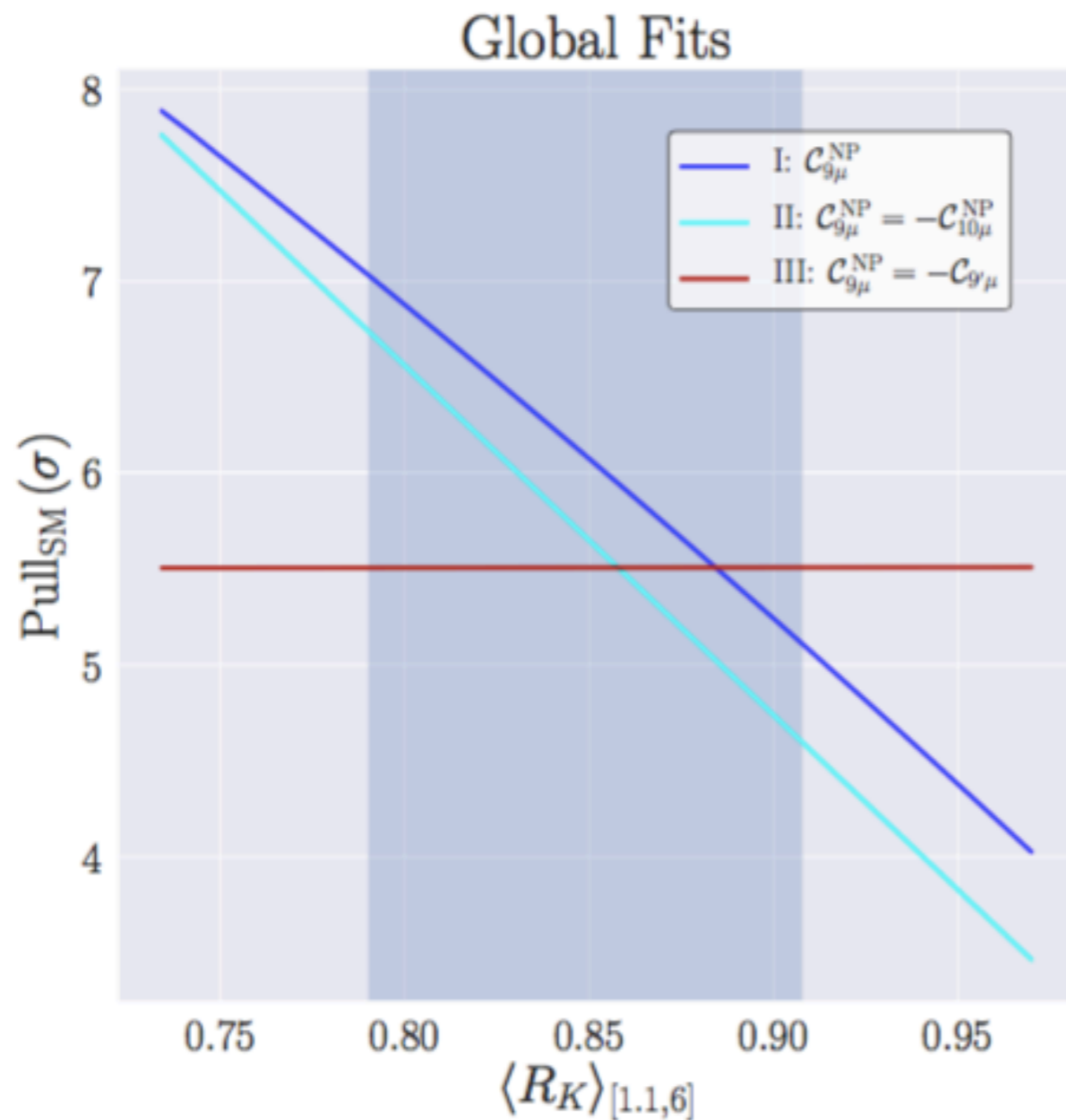


$$Q_5 > 0.2 \Rightarrow \mathcal{C}_{9\mu}$$

$$Q_5 < 0 \Rightarrow \mathcal{C}_{9\mu} = -\mathcal{C}_{10\mu}$$

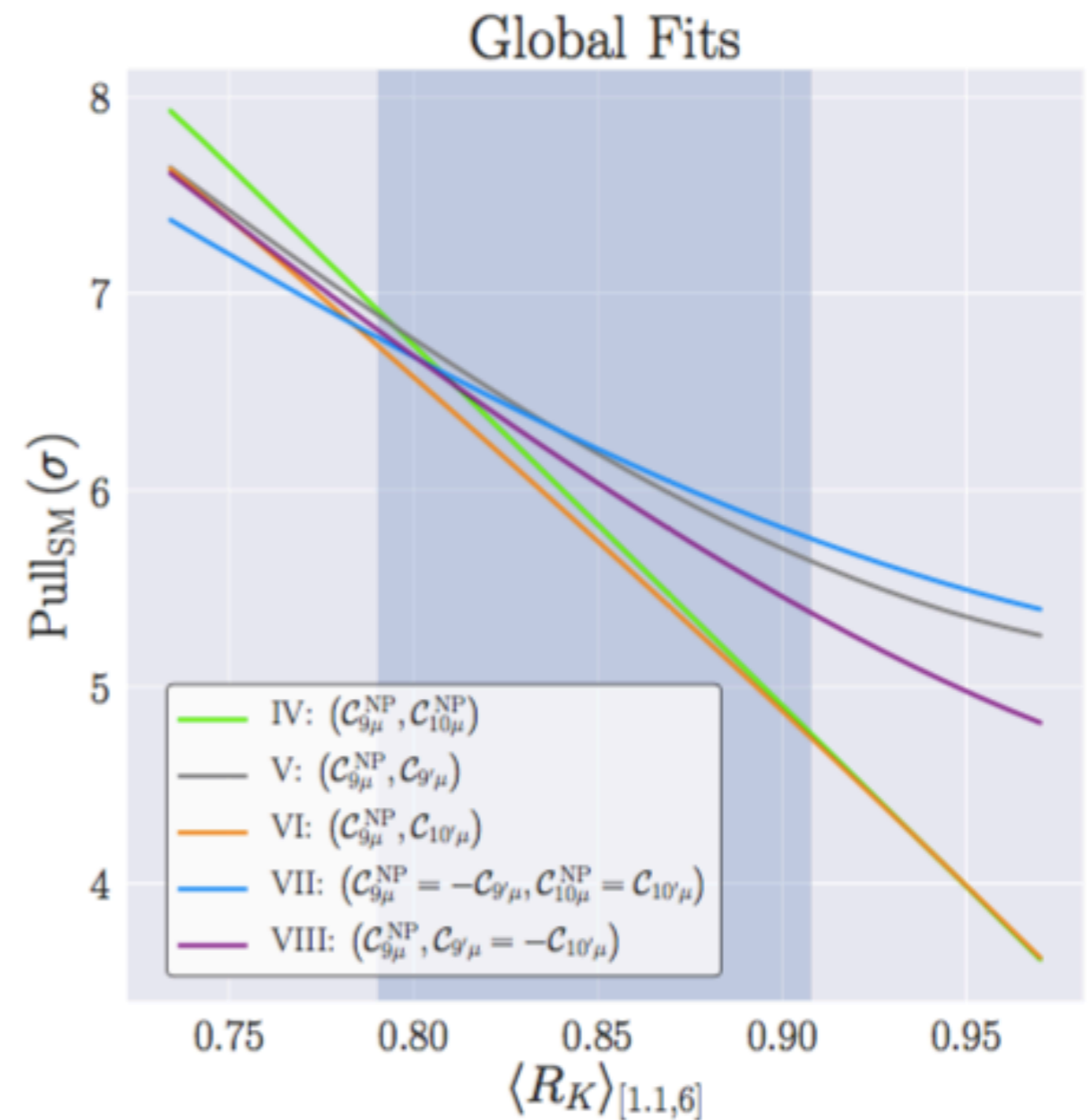
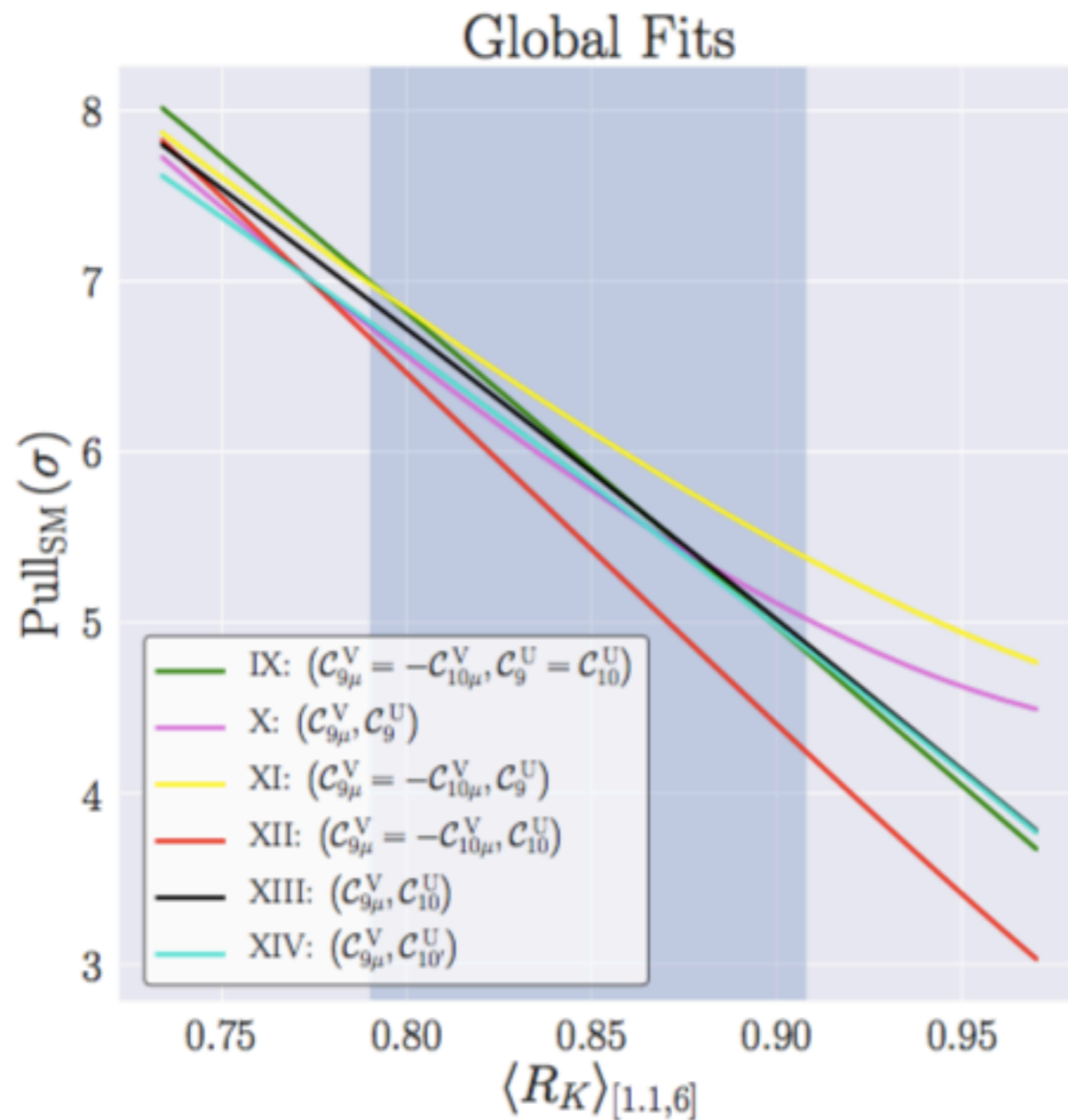
Disentangling Scenarios

[Alguero et al, 1902.04900]



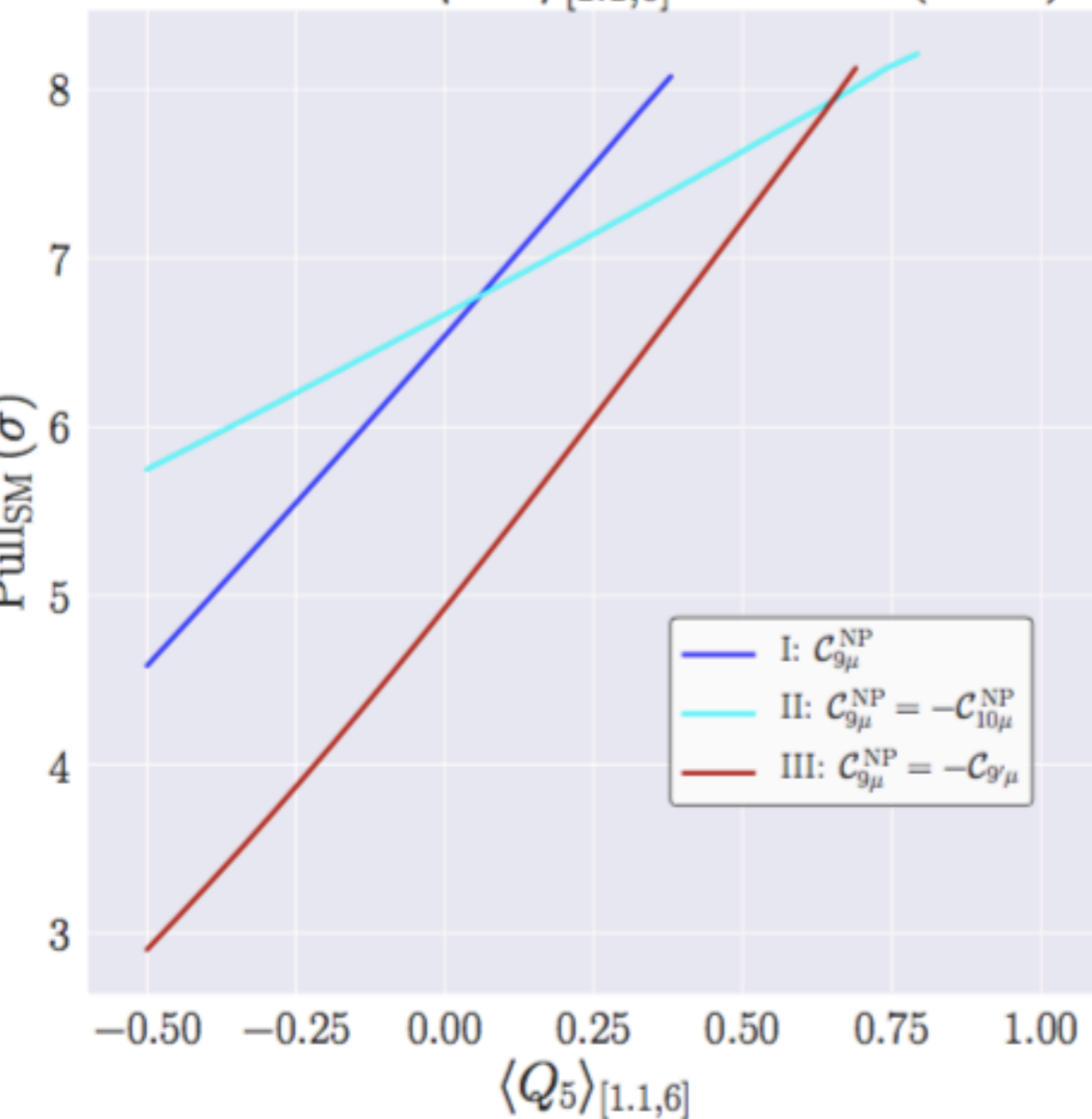
Disentangling Scenarios

[Alguero et al, 1902.04900]

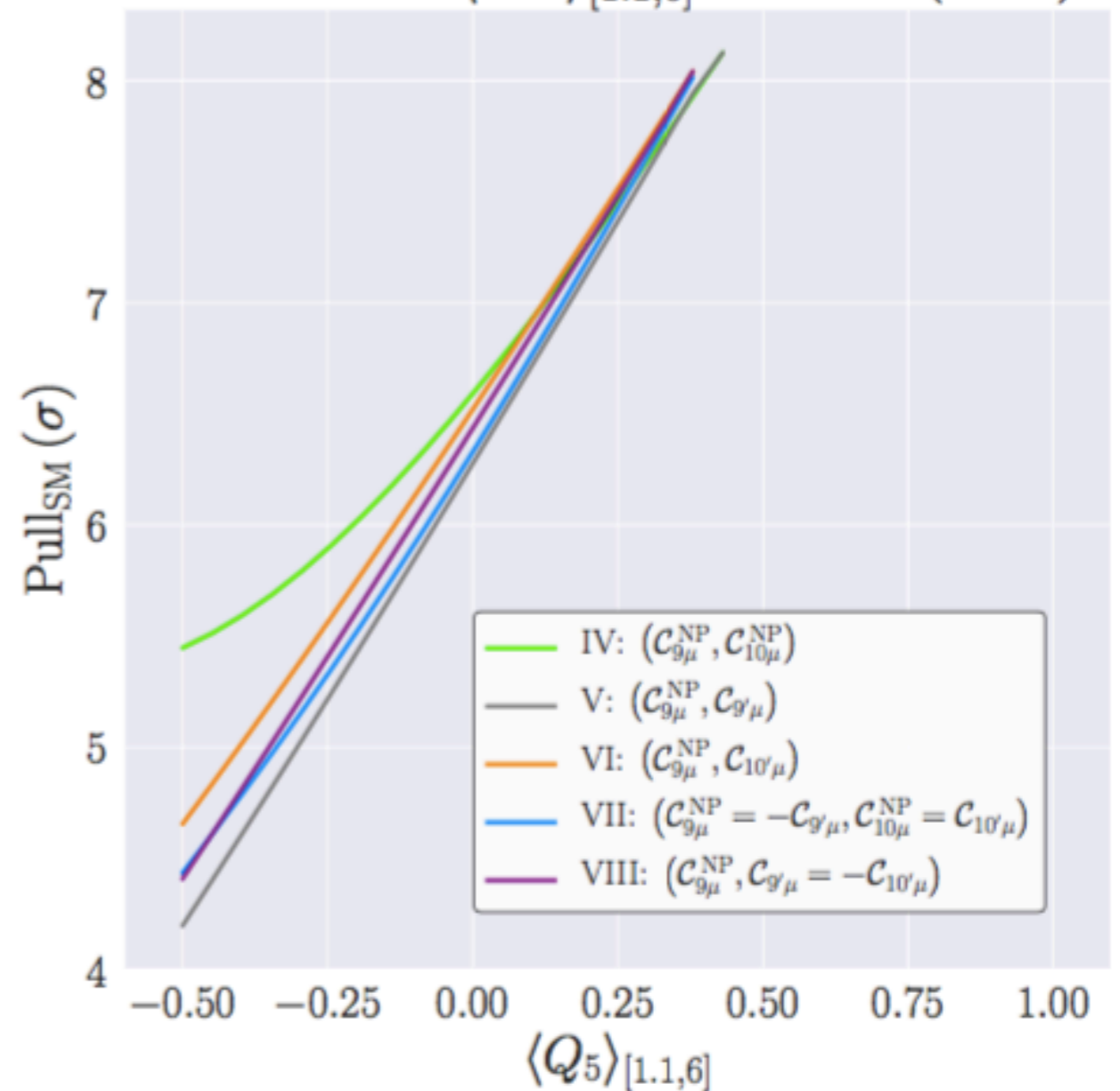


Disentangling Scenarios

Global Fits $\langle R_K \rangle_{[1.1,6]} = 0.789 (-1\sigma)$

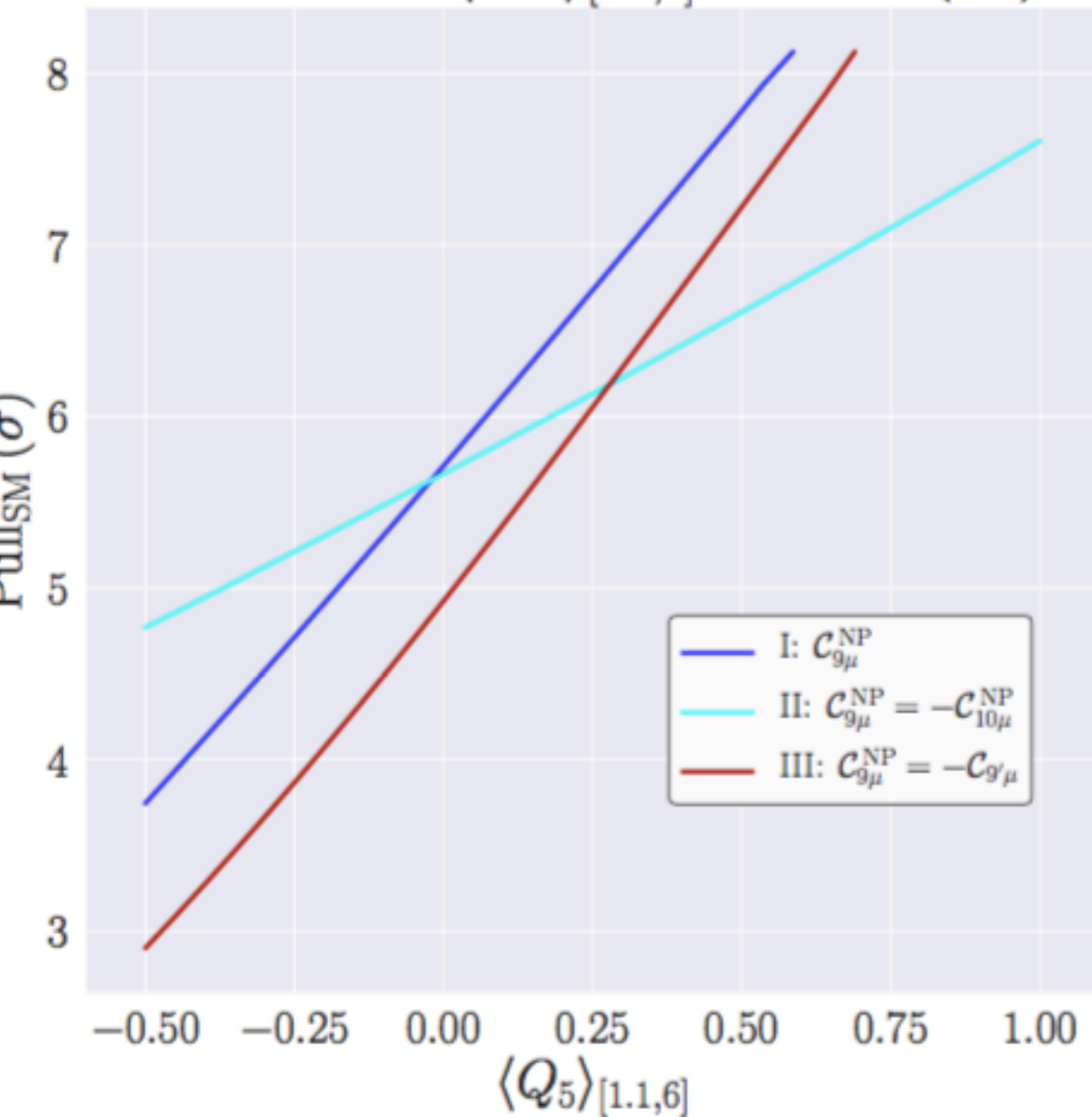


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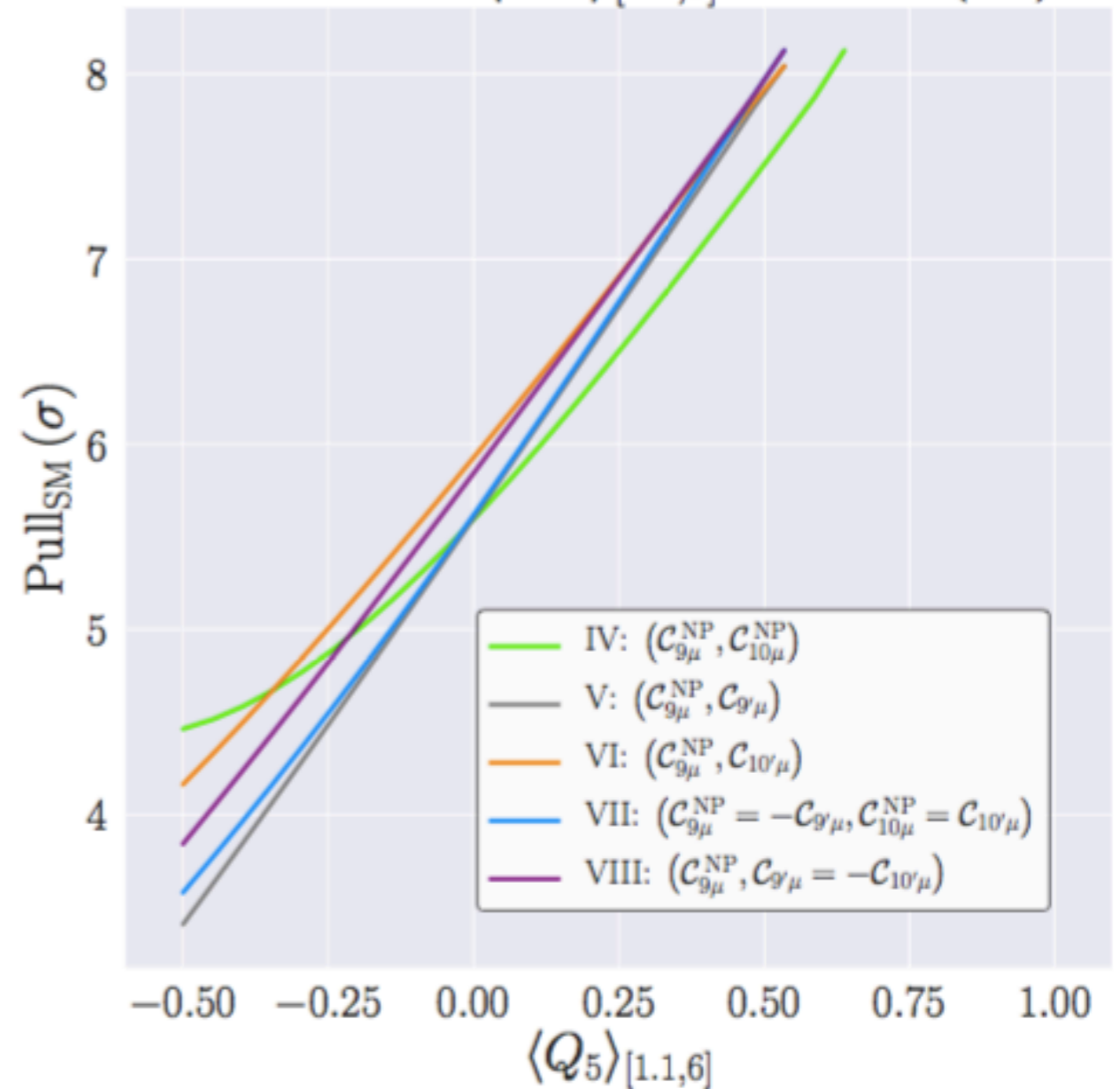


Disentangling Scenarios

Global Fits $\langle R_K \rangle_{[1.1,6]} = 0.846 (0\sigma)$

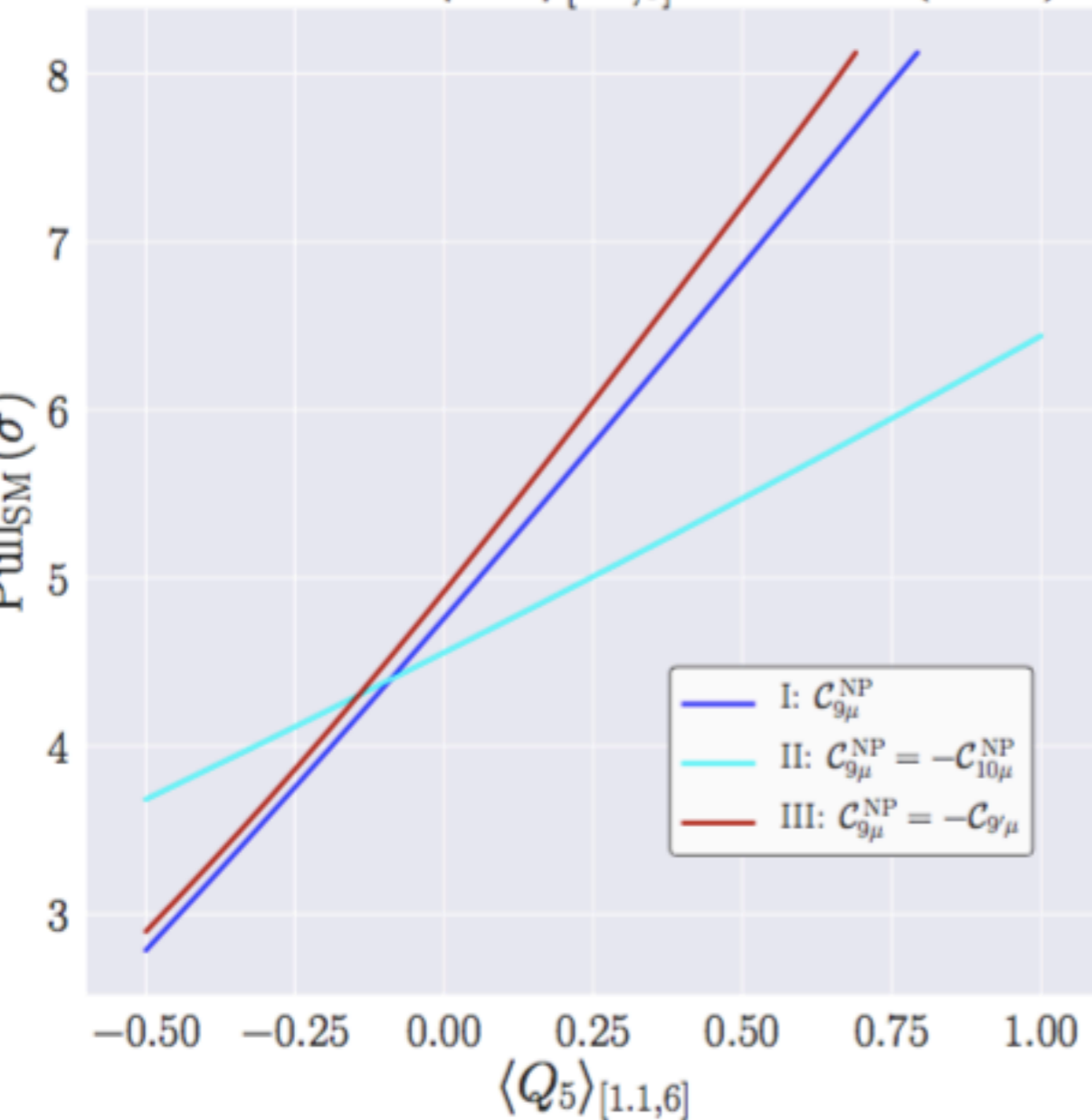


Global Fits $\langle R_K \rangle_{[1.1,6]} = 0.846 (0\sigma)$

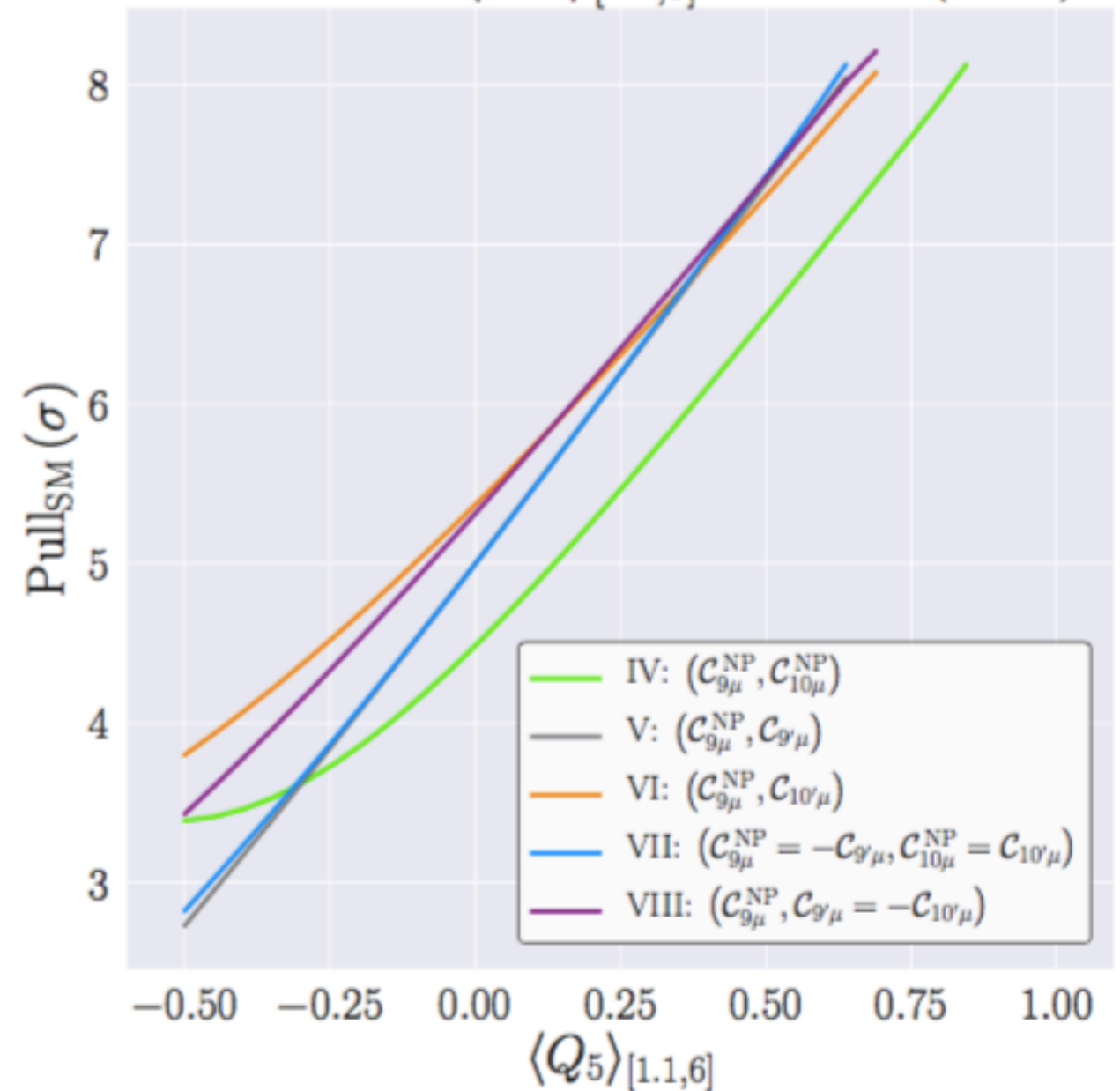


Disentangling Scenarios

Global Fits $\langle R_K \rangle_{[1.1,6]} = 0.908 (+1\sigma)$



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Inner tensions of the global fit

Inner tensions

- We can see how inner tensions behave under NP Hypotheses

- We define pull: $\text{pull}_i^{\text{obs}} = \sqrt{\chi_{\min}^2 - \chi_{\min \text{ w/o obs } i}^2}$

$$\chi^2 = \sum_{ij} (\mathcal{O}^{\text{th}} - \mathcal{O}^{\text{exp}})_i V_{ij}^{-1} (\mathcal{O}^{\text{th}} - \mathcal{O}^{\text{exp}})_j,$$

$$\chi_{\text{w/o obs } i}^2 = \sum_{i \neq j} (\mathcal{O}^{\text{th}} - \mathcal{O}^{\text{exp}})_i V_{ij}^{-1} (\mathcal{O}^{\text{th}} - \mathcal{O}^{\text{exp}})_j,$$

[Hyp. I] $\{\mathcal{C}_{9\mu}^{\text{V}}\} \rightarrow \{\mathcal{C}_{9\mu}^{\text{NP}}\}$

[Hyp. II] $\{\mathcal{C}_{9\mu}^{\text{V}} = -\mathcal{C}_{10\mu}^{\text{V}}\} \rightarrow \{\mathcal{C}_{9\mu}^{\text{NP}} = -\mathcal{C}_{10\mu}^{\text{NP}}\}$

[Hyp. III] $\{\mathcal{C}_{9\mu}^{\text{V}} = -\mathcal{C}_{9'\mu}\} \rightarrow \{\mathcal{C}_{9\mu}^{\text{NP}} = -\mathcal{C}_{9'\mu}^{\text{NP}}\}$

[Hyp. IV] $\{\mathcal{C}_{9\mu}^{\text{V}}, \mathcal{C}_{10\mu}^{\text{V}}\} \rightarrow \{\mathcal{C}_{9\mu}^{\text{NP}}, \mathcal{C}_{10\mu}^{\text{NP}}\}$

[Hyp. V] $\{\mathcal{C}_{9\mu}^{\text{V}} = -\mathcal{C}_{10\mu}^{\text{V}}, \mathcal{C}_9^{\text{U}} = \mathcal{C}_{10}^{\text{U}}\} \rightarrow \{\mathcal{C}_{9\mu}^{\text{NP}} = -\mathcal{C}_{10\mu}^{\text{NP}} + 2\mathcal{C}_{9e}^{\text{NP}}, \mathcal{C}_{9e}^{\text{NP}} = \mathcal{C}_{10e}^{\text{NP}}\}$

[Hyp. VI] $\{\mathcal{C}_{9\mu}^{\text{V}}, \mathcal{C}_9^{\text{U}}\} \rightarrow \{\mathcal{C}_{9\mu}^{\text{NP}}, \mathcal{C}_{9e}^{\text{NP}}\}$

[Hyp. VII] $\{\mathcal{C}_{9\mu}^{\text{V}} = -\mathcal{C}_{10\mu}^{\text{V}}, \mathcal{C}_9^{\text{U}}\} \rightarrow \{\mathcal{C}_{9\mu}^{\text{NP}} = -\mathcal{C}_{10\mu}^{\text{NP}} + \mathcal{C}_{9e}^{\text{NP}}, \mathcal{C}_{10\mu}^{\text{NP}}, \mathcal{C}_{9e}^{\text{NP}}\}$

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Reduction 👍

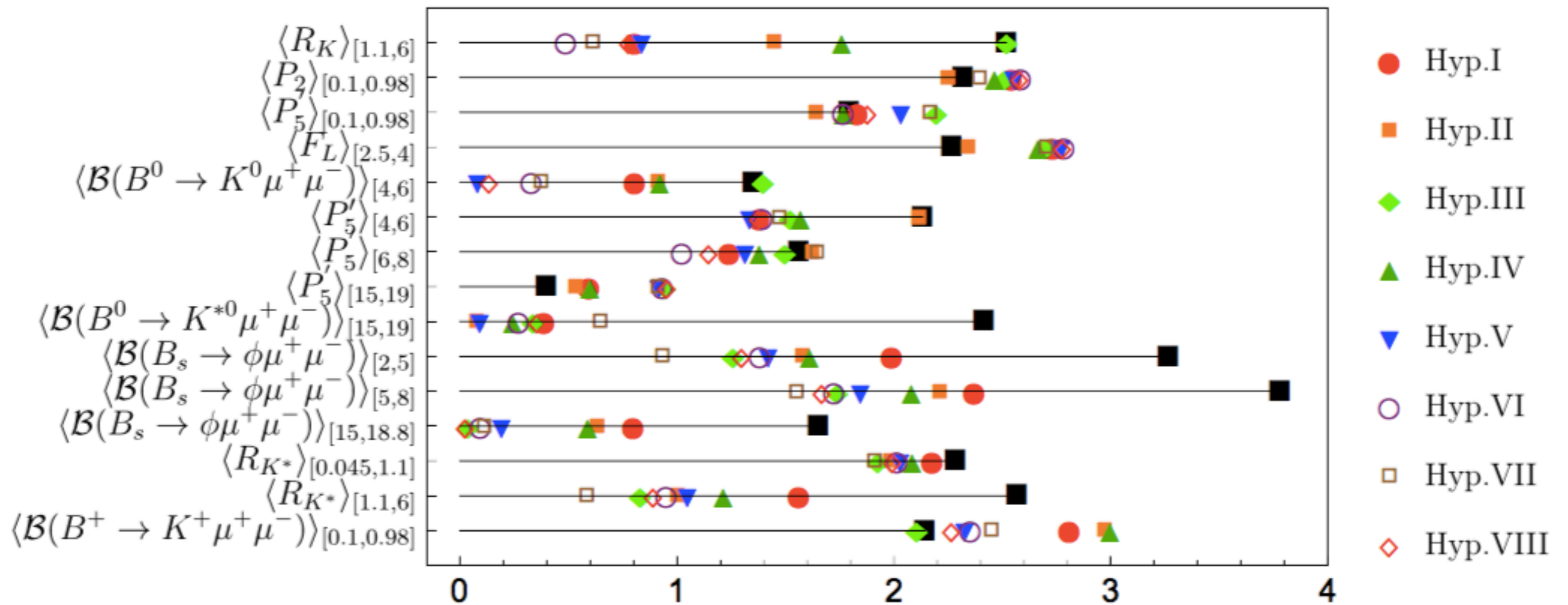
No reduction or worse 👎

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Inner tensions

- List of observables with larger $\text{pull}_i^{\text{obs}}$ in SM (black square)

[Alguero et al, 1902.04900]



Inner tensions

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