

## E. CREMMER

Laboratoire de Physique Théorique de l'Ecole Normale Supérieure

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<sup>+</sup>Laboratoire Propre du CNRS, associé à l'Ecole Normale Supérieure et à l'Université de Paris-Sud. Postal address : 24 rue Lhomond, 75231 PARIS cedex 05 (France)

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## Dedicated to Joël Scherk

#### 1 WHY 5 DIMENSIONS ?

In supersymmetry the consideration of theories in dimensions D > 4 has been very (ruitful. In particular the supersymmetric N=4 Yang-Mills theory has been derived from the N=1 supersymmetric Yang-Mills theory in 10 dimensions (Gliozzi, Scherk & Olive, 1977; Brink, Scherk & Schwarz, 1977). More recently, starting from the N=1 supergravity in 11 dimensions (Cremmer, Julia & Scherk, 1978) the N=8 supergravity in 4 dimensions has been derived with its unexpected symmetries E<sub>7</sub> global x SU(8) local (Cremmer & Julia, 1978 and 1979).

We would like today to concentrate on supergravities in 5 dimensions for essentially four reasons :

(i) For extended supergravities (especially N=8) the structure is simpler in 5 dimensions than in 4 dimensions because all the invariances are invariances of the Lagrangian instead of invariances of equations of motion. (This is related to the duality transformations on vector fields for the theories in 4 dimensions). This could therefore lead to a better understanding of these extended supergravities.

(ii) From the theories in 5 dimensions we can obtain spontaneously broken supersymmetric theories in 4 dimensions by a generalized dimensional reduction (Scherk & Schwarz, 1979). In particular, spontaneously broken N=8 supergravity with 4 mass parameters has been constructed in this way (Cremmer, Scherk & Schwarz, 1979).

(iii) The knowledge of the theory on-shell in 5 dimensions allows one to have an off-shell formulation in 4 dimensions modulo some differential constraints on the fields using the dimensional reduction by Legendre transformation (Schnius, Stelle & West, 1980). In particular, an off-shell formulation of extended N\*8 supersymmetry has been derived (Cremmer, Ferrara, Stelle & West, 1980).

(iv) From the "Lagrangian builder" point of view it shows how the conjecture of the bosonic symmetries allows one to construct

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the complete theory up to a few coefficients.

The plan of my talk will be the following :

(1) I shall give some notation and definitions of symplectic spinors in 5 dimensions.

(2) I shall give the particle contents of all supergravities in 5 dimensions.

(3) I shall briefly recall some facts about global and local symmetries in supergravity.

(4) The main part of the talk will be devoted to the description and construction of the N=8 supergravity in 5 dimensions.

(5) Finally, I shall give the consistent sets of truncation which lead to the N=6, 4 and 2 supergravities in 5 dimensions.

2 SYMPLECTIC SPINORS

The metric of the 5-dimensional spacetime is

$$\mathcal{D}_{rs} = (+, -, -, -, -)$$

The X matrices are defined by their anticommutation relations

 $\left\{Y_{r_{1}}Y_{s}\right\} = 2 p_{rs}$ 

Yo, Y<sub>4</sub>, Y<sub>2</sub>, Y<sub>3</sub> are the same as in 4 dimensions and are pure imaginary, Since  $(v_5)^2 = +1$  we must define

Y<sub>4</sub> = λ V5 which is real. This shows that there are no Majorana spinors in 5 dimensions. Ve and V5 are antisymmetric and VA, V2, V3 are symmetric. The five V matrices are related by

where Frature is the totally antisymmetric product of  $\forall r \forall s$   $\forall t \forall u \forall v$  and  $\exists rsture$  is the usual Levi-Civita symbol with 5 indices ( $\xi_{01234}$  \* +1).

In 5 dimensions the N extended supersymmetry algebra can only be defined for even N and it has a natural isomorphism which is the USp(N) symplectic symmetry (compact)

 $\{\overline{Q}^{a}_{a}, Q^{b}_{p}\} = \mathcal{R}^{ab}(\mathcal{Y}^{\mu})_{pa}P_{\mu}$ 

$$(1...N) = 0...4$$
;  $\alpha = 1...N$ 

The charges  $\Phi_{a}^{a}$  (and consequently the spinor fields) satisfy a generalized Msjorana condition

$$\begin{aligned}
\varphi_{\alpha}^{a} &= \zeta_{5} \overline{\varphi}_{\alpha}^{ta} \\
\text{where } \zeta_{5} \text{ satisfies } \qquad \zeta_{5} \forall_{\alpha} \zeta_{5}^{-1} = 1 \\
\overline{\varphi}^{a} &= (\varphi_{a}^{*})^{\frac{1}{2}} \forall_{0}
\end{aligned}$$

 $\mathfrak{L}^{ab}$  is the real symplectic metric and is used to raise or lower indices

$$d^{a} = -3^{ap} d_{p}$$

from which we deduce  $\overline{Q}_{\alpha} = \mathcal{R}_{\alpha} b \overline{Q}^{b} = - (Q^{\alpha})^{*t} \delta_{0}$ 

We can choose  $C_5 = 3_0 3_5$ . In this case the symplectic spinors are defined by

$$\varphi^* = \gamma_5 (\varphi_a)^*$$

From these definitions we deduce the important property of bilinear expressions in Fermi fields :

$$\Psi^* \mathscr{Y}_{\mu_1} \dots \mathscr{Y}_{\mu_n} \mathscr{X}^\flat = \widetilde{\mathscr{X}}^\flat \mathscr{Y}_{\mu_1} \dots \mathscr{Y}_{\mu_n} \Psi^\flat \quad , \quad \mathsf{TM}$$

Finally let us give the Fierz transformation in 5 dimensions :

$$\overline{\epsilon}_{A} \overline{\epsilon}_{2} \overline{\epsilon}_{3} \overline{\epsilon}_{4} = -\frac{1}{4} \left\{ \overline{\epsilon}_{A} \overline{\epsilon}_{4} \overline{\epsilon}_{3} \overline{\epsilon}_{2} + \overline{\epsilon}_{A} \overline{\delta}_{F} \overline{\epsilon}_{4} \overline{\epsilon}_{3} \overline{\delta}^{F} \overline{\epsilon}_{2} - \frac{1}{2} \overline{\epsilon}_{A} \overline{\delta}_{F} \overline{\epsilon}_{4} \overline{\epsilon}_{3} \overline{\delta}^{F} \overline{\epsilon}_{2} \right\}$$

## 3 SUPERCRAVITIES IN 5 DIMENSIONS

The physical states of 2N extended supersymmetric mussless multiplets in 5 dimensions are classified by USp(2N) (compact) as the states of the massive multiplet with central charge in 4 dimensions. It is the 5th dimension which is related to the central charge in 4 dimensions.

In the simplest multiplets, the representations of USp(2N) which appear are the antisymmetric and traceless tensors  $R^{obc}$  with

with  $m \leq N$ . For  $N < m \leq 2N$  an antisymmetric traceless tensor is automatically zero because the Levi-Civita tensor with 2N indices can be written in terms of  $\Omega_{ab}$ 

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$$\mathcal{E}_{a_1 a_2 \dots a_{2N+n} a_{2N}} \sim \mathcal{L}_{[a_1 a_2} \mathcal{L}_{a_3 a_4 \dots a_{2N-n} a_{2N}]}$$
  
The fields also satisfy the same kind of generalized reality condi-  
tion as the spinor charges

$$A_{\mu}^{ab} = (A_{\mu ab})^{*}$$
$$\chi^{abc} = \chi_{5} (\chi_{abc})^{*}$$

The lowest spin supermultiplet for 2N supersymmetry has states from spin s up to s=N (SU(2) is the little group of Lorentz group in 5 dimensions) and has the following content

$$\phi^{a}$$
  $\phi^{ab}$   $\phi^{ab}$   $\phi^{ab}$ 

where all  $\phi$  are antisymmetric and traceless. Other multiplets can be obtained by combining this multiplet with states of angular momentum J and an arbitrary representation of USp(2N) (Ferrara & Zumino, 1979).

This allows a simple construction of the representations of extended supergravities in S dimensions. They are given, in the following table, as well as the lowest spin supermultiplet.

As in 4 dimensions the N=8 supergravity multiplet is also the lowest spin supermultiplet.

## 4 GLOBAL AND LOCAL SYMMETRIES IN SUPERGRAVITY

The dimensional reduction shows that for maximal extended super-

gravities in D dimensions obtained from N=1 supergravity in 11 dimensions (Cremmer & Julia, 1979; Cremmer, 1980), the theory is invariant under the product of a non-compact global group and a compact local group

SL(11 - D,R)global x SO(11 - D)local

SO(11 - D) local acts on Fermi fields and scalar fields. SL(11 - D,R)global acts on tensor fields and scalar fields. The scalar fields which come from the metric in 11 dimensions are described by the coset GL(11 - D,R)/SO(11 - D) (after a Weyl rescaling). We expect that all scalar fields can be described in this geometric way by a coset G/H (i.e. a matrix of G defined up to a local transformation of H) as the vielbein  $\mathfrak{G}_{L}^{-1}$  is described by the coset GL(D,R)/SO(D - 1, 1). If G is non compact, there is no problem of positivity if H is the maximal compact subgroup of G. The symmetries G and H can be conjectured by simple counting arguments if we temember that H is the maximal group linearly realised on all fields. (This H is the diagonal subgroup of  $G_{\text{global}} \times H_{\text{local}} \supset$ H global  $\times H_{\text{global}} \rightarrow$  H).

We shall give below the content and the symmetries of maximal supergravities in D=9...3 after duality transformations which convert a p tensor field into a (D - 2 - p) tensor field

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Let us note that in 3 dimensions there is no degree of freedom for the graviton and the gravitinos. The underlined tensor fields need duality transformations to form a representation of the global group. The global symmetry will not be a symmetry of the Legrangian but only of the equations of motion : the symmetry will exchange the Bianchi identity for the field strength of the tensor with its equation of motion.

It has been seen that in 4 dimensions, for all extended supergravities, the scalar fields are described by a coset, the local symmetry being U(N). In the same way, we can conjecture that all extended supergravities in 5 dimensions have a global symmetry G and a local symmetry USp(2N), the scalar fields being described by G/USp(2N). This gives the following table

- N=8 E<sub>6(+6)</sub>global 🕲 USp(8)local
- N=6 SU<sup>21</sup>(6)global Ø USp(6)local
- N=4 USp(4) x R global 🕙 USp(4)local
- N=2 USp(2)global & USp(2)local
- 5 N=8 SUPERGRAVITY IN 5 DIMENSIONS

As we have seen, the free particle spectrum is described by the fields  $A_{\mu\nu}$ ,  $A_{$ 

We have seen that we expect the theory to have a global symmetry  $E_{0}$  and a local symmetry USp(8). Let us first briefly describe  $E_{0}$ . It has 78 generators and the fundamental representation has dimension 27. We are interested in the non-compact form which has 42 non-compact generators and 36 compact ones which generate the maximal subgroup USp(8). The 27 representation acts in the vector space spanned by  $\mathbb{R}^{4,6}$  (  $d_{0}$  ( = 1...8) such that

$$Z^{(4)} = -Z^{(4)} = (Z_{40})^{2}$$

$$\Omega_{AB} Z^{AB} = 0$$
 ;  $Z_{AB} = \Omega_{AX} \Omega_{BB} Z^{XB}$ 

and the infinitesimal transformations of E are given by

where  $\Lambda^{a}\gamma$  is an antihermitian matrix such that  $\Lambda_{a}\chi$  is symmetric and  $\Sigma^{a}\beta^{a}\Sigma^{b}$  is totally antisymmetric, traceless and pseudo-

$$\Sigma^{\times \wedge YY} = (\Sigma_{\times \wedge YY})^*$$

There is no quadratic invariant for E :  $27 \times 27 \not> 1$  in particular  $2^{-n} Z_{\alpha | A}$  is not invariant for E. We can form an invariant from 27 x 27 where 27 is spanned by

which transforms under  $E_{L}$  by

.

$$\delta \tilde{z}_{AV} = \sqrt{4} \star \tilde{z}_{AV} + \sqrt{6} \star \tilde{z}_{Rs} - \Sigma_{RUSS} \tilde{z}^{AL}$$

Exp  $\Xi^{*/4}$  is an invariant under E. Both  $\Xi_{*}$ ,  $\Xi^{*/4}$  and  $\Xi_{*}$ ,  $\Xi^{*/4}$  are invariant under the subgroup  $\Xi^{*/4}$  and  $\Xi_{*}$ ,  $\Xi^{*/4}$  are invariant under the subgroup USp(8). There exists a trilinear invariant for E : 27 x 27 x 27 = 1+..

These properties of E6 are all we need to obtain the general structure of the theory.

The fields of the N=8 supergravity are :

- the graviton  $e_{\mu}^{(r)}$ , an element of GL(5,R)/SO(4, 1) the 8 gravitinos  $u_{\mu}^{(r)}$  which are in the representation 8 of
- USp(8) and singlets for E the 27 vector fields Arg 6 which are singlets for USp(8) and in
- the 27 representation of  $E_{c}$  the 48 spin 1/2 fields  $\chi^{442}$  which are in the representation 48
- of USp(8) and singlets for E. the 42 scalar fields will be described by an element de of the coset  $E_{6,+6}/USp(8)$  (78 - 36 = 42). It transforms as  $\overline{27}$  under E, and  $\overline{27}$  under USp(8). The indices  $\omega/3$  = 1...8 are the 'curved' indices of E, and  $\alpha, \beta = 1...8$  are the flat indices of USp(8) and  $\sqrt{3}$  = 1...8 are the flat indices of USp(8) are the flat indices of

The self-interaction of the scalar fields is described by a non linear 5 -model associated to the coset  $E_{g}/05p(8)$  and therefore by the Lagrangian

$$\mathcal{A}_{s} \sim \mathcal{D}_{\mu} \mathcal{V}_{\alpha\beta} \stackrel{ab}{\to} \mathcal{D}^{\mu} \tilde{\mathcal{V}}_{\alpha b} \stackrel{ab}{=} - \mathcal{T}_{\Gamma} \left( \tilde{\mathcal{V}} \quad \mathcal{D}_{\mu} \mathcal{V} \right)^{\mu}$$

where V is the inverse of V

$$N_{ab}^{ab} N_{ab}^{ab} = \frac{1}{2} (S_{a}^{c} S_{b}^{b} - S_{a}^{c} S_{b}^{c}) + \frac{1}{8} \Omega_{ab} \Omega^{cd}$$

D<sub>1</sub> is the covariant derivative with respect to USp(8) using the associated connexion  $\Sigma_{\mu,\alpha}$ . Since there is no kinetic term for  $\Sigma_{\mu,\alpha}$  we can solve its equations of motion. Since V is an element of E, we have the following decomposition for  $\nu^{-1}\partial_{\mu}V$ which is In the Lie algebra of E,

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where  $Q_{m,c}$  belongs to the Lie algebra of USp(8) and  $P_{m,m}$  about is in the orthogonal part to USp(8) (with respect to the Killing metric). For Some we get

The Lagrangian then becomes

 $P_{max}$  and  $P_{mated}$  are obviously invariant under  $E_6$ . If we restrict ourselves to the scalar fields, as in the case of general relative ity, we can describe them by a metric  $\{\sigma, \gamma\}$  instead of the 27-bein  $\forall \sigma_{1}$  (to be compared to  $\{\sigma, \gamma\}$  and  $e_{1}$ ). The metric is invariant under the local group USp(8) and covariant under  $E_{6}$ . It is given by

and is characterized by the property :

The Lagrangian is then written as

This metric must also be used to describe the interaction of the vector fields since there is no quadratic invariant for  $E_{1}$ . The generalized "kinetic" term for the vectors is then given by

As in 11 dimensions there also exists a trilinear gauge invariant coupling (up to a total derivative) of the vectors which is required by supersymmetry. Since there is a trilinear E, invariant J, we do not need the scalar metric (nor the metric tensor gange)

The couplings to the fermions can no longer be described by the metric, but require the 27-bein  $\mathcal{V}$ . The "kinetic" terms for the fermionic fields  $\mathcal{V}_{a}^{a}$  and  $\mathcal{X}^{abc}$  will be covariant with respect to the local Lorentz group SO(4, 1) with the consexion  $\omega_{\mu,rs}$  and the local group USp(8) with the connexion  $\varphi_{\mu\nu}^{b}$ 

$$D_{\mu} \Psi_{\rho}^{a} = (\partial_{\mu} \delta_{b}^{a} - Q_{\mu}^{a} b + \frac{1}{4} \omega_{\mu rs} \gamma^{rs} \delta_{b}^{a}) \Psi_{\rho}^{b}$$
$$D_{\mu} \gamma^{obc} = (\partial_{\mu} \delta_{d}^{b} - 3Q_{\mu}^{b} d + \frac{1}{4} \omega_{\mu rs} \gamma^{rs} \delta_{b}^{a}) \chi^{bdd}$$

As usual there exists a Noether-type coupling required by supersymmetry

The coupling of fermions to  $F_{\rm ev}$  must occur only through the E invariant (and scalar under the general coordinate transformation)

Let us note that  $x'_{y^2}$  can be written as

$$f_{v^2} \sim (F_{rs}^{ab})$$

but Fre in no longer a curl.

The supersymmetry transformation laws  $\delta \phi$  are conjectured to be covariant with respect to USp(8) and E. Therefore  $\delta c$  and  $\delta \phi$ are now defined up to numerical coefficients, quartic fermionic terms for  $\delta c$  and trilinear fermionic terms for  $\delta f_{\mu\nu\sigma}$  and  $\delta f_{\mu\nu\sigma}$ . In particular all the non-polynomial structure in the scalar fields is fixed. Supersymmetry is used to get rid of the remaining arbitrariness.

(i) Numerical coefficients (and Lorentz structure) in  $\delta\Psi$  and  $\delta\chi$  are determined by checking the supersymmetry invariance of  $\mathcal{Z}$  in the terms of the type  $\tilde{\mathcal{E}}\Psi$ ,  $\tilde{\mathcal{E}}\chi$ 

(ii) Quartic terms in  $\mathbb{Z}$  and trilinear terms in  $\mathbb{SY}$  and  $\mathbb{SX}$  are determined in two independent ways :

- we require supercovariant equations of motion for fermionic fields
- we require the closure of the supersymmetry algebra on the bosonic fields

 $[\delta_{\epsilon_1}, \delta_{\epsilon_1}] = \delta_{\epsilon_1} + \delta_{\epsilon_1} + \delta_{\epsilon_2} + \delta_{\epsilon_1} + \delta_{$ 

where  $\delta_G$  is the general coordinate transformation,  $\delta_{e'}$  a new supersymmetry transformation,  $\delta_{L}$  a local Lorentz transformation,  $\delta_{u-1,0}$  a local USn(8) transformation and  $\delta_{MAL}$  an Abelian value

 $S_{USp(R)}$  a local USp( $\vartheta$ ) transformation and  $S_{U(A)}$  an Abelian gauge transformation on the vector fields. At this stage only the  $\chi^4$  terms in  $\varkappa$  are still undetermined. They are determined by checking  $\delta \varkappa$  in the terms of the type  $\in \chi^4$  or by looking at the closure on fermionic fields which requires the fermionic equations of motion.

The Lagrangian is then written, (we have put K=1)

$$e^{-1} d = -\frac{1}{4} R(\omega) - \frac{i}{2} \overline{\Psi}_{\omega}^{*} y^{\mu\nu} p_{\nu} \Psi_{\mu} - \frac{1}{3} g^{\mu\nu} g^{\nu\sigma} g_{\nu\rho, \tau\delta} F^{\nu\rho}_{\nu\nu} \overline{F}^{\tau}_{\sigma}$$

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$$+\lambda^{a}\overline{\lambda}^{abc}\gamma^{a}D_{\mu}\gamma_{abc} + \frac{1}{24\mu}g^{a\nu}F_{\mu}abcd P_{\nu}abcd$$

$$-\frac{e^{-1}}{12} \mathcal{E}^{\mu\nu}\ell^{\nu\lambda}(F_{\mu\nu})^{d}\beta (\overline{\xi}_{er})^{\mu}\gamma (\overline{f}_{\lambda})^{\gamma}\alpha + \frac{i}{3\ell_{2}}P_{e}abcd \overline{\Psi}^{a}\gamma^{\mu}\gamma^{\mu}\gamma^{bd}$$

$$+\frac{i}{4} \Psi^{ab}_{\beta}F^{d}_{\mu\nu} \left\{\overline{\Psi}^{e}_{\alpha}\gamma^{\mu\nu}\gamma^{\mu}\gamma^{\mu}\gamma^{b} + \frac{1}{42}\overline{\Psi}^{e}_{e}\gamma^{\mu\nu}\gamma^{\mu}\gamma^{a}bc + \frac{1}{2}\overline{\lambda}_{acd}\gamma^{\mu\nu}\gamma^{bd}\right\}$$

$$+e^{i}\chi^{\mu}$$

 $a_4$  represents the quartic fermionic terms. Except for the  $\Psi^4$  terms it is not enough to replace  $\omega_{\rm mes}$ ,  $P_{\rm obsc}$  and  $F_{\rm b}$  by  $\omega_{\rm me}$ ,  $P_{\rm bec}$  and  $F_{\rm b}$  by  $\omega_{\rm me}$ ,  $P_{\rm bec}$  and  $F_{\rm b}$  by  $\omega_{\rm me}$ ,  $P_{\rm bec}$  and  $F_{\rm b}$  by  $\omega_{\rm me}$ ,  $P_{\rm bec}$  and  $F_{\rm b}$  by  $\omega_{\rm me}$ ,  $P_{\rm bec}$  and  $F_{\rm b}$  by  $\omega_{\rm me}$ ,  $P_{\rm bec}$  and  $F_{\rm b}$  by  $\omega_{\rm me}$ . The supercovariant extension to reason ball the quartic terms. For completeness we give  $a_4$  below : + 1 - " abe 8" 814 4 - Fre Her + 3 [ 3 and 8 4 abe 4 ab + 32 [ 7 obr (818-39er ) X d + Fee Yod - 1 x " + 2 x " + 2 x + + 1 [ x bed & x be x d to x, x eta - y wy x x x e x d to x, x eta - y wy x x x eta + I T abe Yes Xabe Rely Vis Xety ]

The supersymmetry transformation laws are given by

$$\begin{split} \delta e_{\mu} &= -i \, \xi^{a} \, \xi^{r} \, \Psi_{\mu a} \\ \tilde{\Psi}^{a\beta} c_{d} \, \delta \tilde{\Psi}_{a,ab} &= -2i \overline{2} \left( \overline{e}_{a} \, \chi_{bcd} \right) + \frac{3}{4} \, \mathcal{D}_{ab} \, \overline{e}_{e} \, \chi^{e}_{cd} \right) \\ \delta h_{\mu}^{a\beta} &= 2i \, \tilde{\Psi}^{a\beta}_{\ ab} \left( \xi^{e} \, \Psi_{\mu}^{b} + \frac{1}{2f_{2}} \, \overline{e}_{e} \, \chi_{\mu}^{abc} \right) \\ \delta \Psi_{\mu a} &= \left( D_{\mu} (\hat{\omega}) \delta_{a}^{b} + \psi_{\mu a}^{b} \right) 6_{b} - \frac{1}{6} \, \overline{f}_{rscb} \left( \delta^{rs} \xi + 2 Y \, \underline{e}_{a}^{s} \right) e^{b} \\ &+ \frac{i \sqrt{2}}{4} \left( 3 \, \overline{e}^{b} \, \Psi_{\mu}^{c} \, \chi_{abc} - \overline{e}^{b} \, \delta^{r} \, \Psi_{\mu}^{c} \, \chi_{r} \, \chi_{abc} \right) \\ &- \frac{i}{i2} \left( \chi_{\mu} + 2g_{\mu} \right) 6_{d} \, \overline{\chi}_{abc} \, \delta^{r} \, \chi^{bcd} \\ &- \frac{i}{i2} \left( \chi_{\mu} + 2g_{\mu} \right) 6_{d} \, \overline{\chi}_{abc} \, \delta^{r} \, \chi^{bcd} \\ &\delta \chi_{abc} &= \sqrt{2} \, \overline{F}_{\mu abcd} \, e^{d} - \frac{3}{2f_{2}} \, \delta^{rs} \left( \overline{f}_{rscb} \, 6_{c} \right) \\ &+ \frac{3i \sqrt{2}}{8} \left[ 3 \, 6_{3} \, \overline{\chi}_{a}^{b} \, \overline{\xi}^{r} \, \chi_{bd}^{c} \, - 6_{3} \, \overline{\chi}_{a}^{bfd} \, \chi_{fda} \, \overline{\chi}_{bd} \right] \\ &- \chi_{r} \, e_{3} \, \overline{\chi}_{a}^{b} \, \chi^{rr} \, \chi_{bd}^{c} \, f + \frac{3}{2} \, \epsilon_{b} \, \overline{\chi}_{rda}^{c} \, \chi^{b} \\ &+ \frac{3i \sqrt{2}}{8} \left[ 3 \, 6_{3} \, \overline{\chi}_{a}^{c} \, \chi^{rr} \, \chi_{bd}^{c} \, f + \frac{3}{2} \, \epsilon_{b} \, \overline{\chi}_{rda}^{c} \, \chi^{b} \right] \\ &- \chi_{r} \, e_{3} \, \overline{\chi}_{a}^{b} \, \chi^{rr} \, \chi_{bd}^{c} \, f + \frac{3}{2} \, \epsilon_{b} \, \overline{\chi}_{rda}^{c} \, \chi^{b} \, \xi^{c} \, \chi^{c} \, \chi^{$$

with :

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$$\hat{\omega}_{mrs} = \omega_{mrs}^{0}(e) + \frac{i}{2} \left[ \bar{\Psi}_{\mu}^{a} Y_{s} \Psi_{ra} - \bar{\Psi}_{r}^{a} Y_{\mu} \Psi_{sa} + \bar{\Psi}_{s}^{a} Y_{r} \Psi_{\mu} \right] - \frac{i}{24} \bar{\chi}_{\mu rs}^{a} X_{ebr} \bar{\chi}_{ebr}^{a} \bar{\chi}_{ebr$$

The fermionic equations of motion are :

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$$\stackrel{i}{=} \chi^{n} \stackrel{j}{\supset} \chi_{abc} + \stackrel{i}{+} \chi^{mv} (\chi^{e}_{[ab} \stackrel{f}{=}_{uv} \stackrel{j}{=} d - True abc)$$
  
 $- \frac{1}{20} [\chi^{m} \chi_{e[ab} \stackrel{f}{\chi} \stackrel{j}{=} \chi^{efg} - \chi^{efg} - \chi^{efg} \stackrel{f}{=} \chi^{eg} \chi_{efb} \stackrel{f}{\chi} \stackrel{j}{=} \chi^{efg} \qquad + \frac{1}{24} \chi^{ee} \chi_{abc} \stackrel{f}{\chi} \stackrel{j}{=} \chi^{ee} \chi^{efg} - True abc] = 0$ 

 $\hat{D}_{\mu}$ ,  $\Psi_{\gamma}$  and  $\hat{D}_{\mu}$ ,  $Y_{\sigma bc}$  are supercovariant extensions of  $\hat{D}_{\mu}$ ,  $\Psi_{\gamma c}$ and  $\hat{D}_{\mu}$ ,  $Y_{\sigma bc}$  defined such that their variation by supersymmetry has no derivative of  $\in$ .

The algebra of supersymmetry is given by

$$[\delta_{\epsilon^{2}} \delta_{\epsilon^{2}}] = \delta^{\epsilon}(\lambda) + \delta^{2}(\epsilon_{\lambda}) + \delta^{2}(\Sigma_{\lambda}) + \delta^{n}(\lambda) +$$

with

$$\begin{split} \hat{J}_{\mu} &= -i \, \hat{\epsilon}_{n}^{c} \, \hat{V}_{\mu} \, \epsilon_{2c} \\ \hat{\epsilon}^{\prime a} &= -\hat{s}^{t} \, \hat{W}_{t}^{a} + \frac{i \, G}{4} \left( 3 \, \hat{\epsilon}_{nb} \, \epsilon_{2c} \, \bar{\chi}^{abc} - \tilde{\epsilon}_{nb} \, \hat{v}^{c} \, \epsilon_{2c} \, \bar{\chi}^{abc} \, \delta_{p} \right) \\ \bar{z}_{rs} &= 5^{t} \, \hat{\omega}_{b,rs} + \frac{i}{3} \, \bar{F}_{uv,ab} \, \bar{\epsilon}_{1}^{a} \left( \hat{v}_{rs}^{\mu\nu} + i \, \bar{\delta}_{u}^{\nu} \, \delta_{s}^{\nu} \right) \hat{\epsilon}_{2}^{b} \\ &- \frac{1}{6} \, \hat{\epsilon}_{1a} \, \hat{v}_{rst} \, \epsilon_{2b} \, \bar{\chi}^{a}_{cd} \, \hat{v}^{b} \, \chi^{bcd} \\ &+ \frac{1}{6} \, \bar{\epsilon}_{ia} \left( \hat{v}_{rstu} + \hat{v}_{rt} \, \bar{v}_{su} \right) \hat{\epsilon}_{2b} \, \bar{\chi}^{a}_{cd} \, \hat{v}^{tu} \, \chi^{bcd} \\ \bar{\Lambda}_{a}^{b} &= 5^{t} \, \bar{\varphi}_{ba}^{b} - \frac{8}{3} \left[ \left( \bar{\epsilon}_{2}^{b} \, \chi^{bdel} - \frac{3}{4} \, \frac{\Omega^{bc}}{2} \bar{\epsilon}_{2p} \, \chi^{del} \, \hat{\ell} \right) \\ & \quad \kappa (\bar{\epsilon}_{nja} \, \bar{\chi}_{cdej} + \frac{3}{4} \, \Omega_{pa}^{c} \, \bar{\epsilon}_{nt} \, \bar{\chi}_{dg}^{c} \right) - \hat{\epsilon}_{iee} \hat{\epsilon}_{2} \right] \\ \mathcal{U}^{d\beta} &= -3^{c} \, \bar{A}_{p}^{\alpha\beta} \, + 2i \, \tilde{v}^{\alpha\beta}_{ab} \, \bar{\epsilon}_{i}^{\alpha} \, \bar{\epsilon}_{2}^{b} \end{split}$$

We have seen that the conjectured  $E_6$  global O USp(8) local was the clues to construct the N=8 supergravity in 5 dimensions. There still remain some complications in the quartic fermionic terms. This could be a sign that there is still some structure to be discovered. We can hope that it would be easier to discover it in 5 than in 4

dimensions. Another problem could also be more easily solved in 5 dimensions : the construction of the multiplet containing the connexion of USp(8)  $\varphi_{a}^{b}$ . This is of crucial importance in 4 dimensions where we conjectured that the local SU(8) could become dynamical at the quantum level (Cremmer & Julia, 1979, 1980). Some conjectures have been made on this multiplet which could lead to a grand unified model based on SU(5) with 3 families (Ellis, Gaillard, Maiani & Zumino, 1980 ; Zumino, 1980).

6 SUPERGRAVITIES N=6, 4, 2

In order to derive supergravities N=6, 4, 2 from N=8 by consistent truncations, it is useful to choose a particular representation for Shap namely



We shall describe the consistent truncations for N=6, 4, 2 and give the complete results only for N=2.

## 6.1 N=6 Supergravity

The invariance of the theory is SO<sup>R</sup>(6)global x USp(6). We note by  $\Delta = 1...6$  the indices of USp(6) and by d = 1...6 the indices of SU<sup>5</sup>(6), we keep the fields  $\sqrt{78}$ ,  $\sqrt{36}$ ,  $\sqrt{2}$ ,  $\sqrt{2}$ ,  $\sqrt{2}$ ,  $\sqrt{2}$ ,  $\sqrt{3}$ , the traceless condition for N=8 into account

this implies that for  $\mathcal{V}^{-1}\partial_{\mu}\mathcal{V}$  (N=8) only the following components Q and P remain

with the trace conditions

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$$\mathcal{R}^{ab}\mathcal{R}^{cd}\mathcal{P}_{\mu abcd} = 0 \quad \mathcal{R}^{ab}\mathcal{P}_{\mu ab7g} = 0$$

$$\mathcal{R}^{ab}\mathcal{P}_{\mu abcd} + 2\mathcal{R}^{78}\mathcal{P}_{\mu 73cd} = 0$$

and  $P_{u}$  about totally antisymmetric : this is equivalent to an anti-symmetric  $P_{u}$  ab traceless which can be identified to  $P_{u}$  by a cor-responding to the 14 scalar fields ( $P_{u}$  about  $\sim E_{b}$  bound  $P_{u}$ ). The theory is invariant under the N=6 supersymmetry obtained by restricting  $\boldsymbol{\epsilon}$  to  $\boldsymbol{\epsilon}^{\alpha}$  ( $\boldsymbol{\epsilon}^{\gamma} = \boldsymbol{\epsilon}^{\otimes \pm} 0$ ). The content of the theory is as expected 1 graviton  $\boldsymbol{\epsilon}_{\mu}$ , 6 gravitinos  $\boldsymbol{\psi}_{\mu}$ , 14+1 vector fields  $\boldsymbol{A}_{\mu}$ , 14+6 spinor fields and 14 scalar fields.

## 6.2 N=4 Supergravity

The symmetries are  $(USp(4) \times R)$  global  $\bigotimes USp(4)$  local. We keep  $(\alpha = 1...4)$ ;  $A_{\perp}^{A^{A}}$  ( $\alpha = 1...4$ )  $A_{\perp}^{S^{A}}$  with the condition

$$\Omega_{d,0} P_{\mu}^{d,0} + 4 \Omega_{56} A_{\mu}^{56} = 0$$

as well as  $A_{\mu\nu}^{78}$  defined for N=6 in terms of  $A_{\mu\nu}^{4/3}$  and  $A_{\mu\nu}^{56}$ . In the same way we keep Number and NSGC (Nagy being function of the previous ones) with the condition

$$\mathcal{Q}^{ab} \gamma_{abc} + 4 \mathcal{Q}^{sc} \gamma_{scc} = 0$$

This corresponds to 4 spin 1/2 since Yake ~ Eabed X<sup>d</sup> from which we deduce Xssc ~ Xc. On  $V^{e_0}$  by we make the same truncation as on  $A_{\mu}$ <sup>eff</sup>. This implies, for  $B_{\mu}$  and  $B_{\mu}$  ssed, the relation

This corresponds to I scalar field : Probel & Eabed \$

The remaining  $\varphi_{\mu a}^{\ b}$  has the form  $\psi^* \partial_{\mu} \psi$  where  $\psi$  is an element of USp(4) and therefore being a pure gauge it can be reabsorbed by redefining the fermionic field with the USp(4) transformation The theory is invariant under N=4 supersymmetry with parameter E\* The content is 1 graviton, 4 gravitinos, 5+1 vector fields, 4 spin 1/2 fields and 1 scalar field.

6.3 N=2 Supergravity

From N=4 we keep 
$$\Psi_{\mu\nu}$$
 (a =1, 2) Anabovith the relation  
 $\Omega^{ab} A_{\mu\nu} + 6 \Omega^{34} A_{\mu\nu} = 0$ 

There is no Xaba left and we can replace  $\mathcal{V}^{*}_{\alpha \gamma t}$  by "1". The truncation can also be directly made from N=8, keeping  $\mathcal{V}_{\mu \alpha \gamma}, \mathcal{A}_{\mu \gamma 2}$ ,  $\mathcal{A}_{\mu \gamma 5}$ ,  $\mathcal{A}_{\mu \gamma 5}$ , and  $\mathcal{A}_{\mu \gamma 2}$  with the relations

After renormalization of the vector fields we get the Lagrangian for  $N^{a_2}$  supergravity with field  $\mathbf{e}_{\mu}$ ,  $\Psi_{\mu}^{a_1}$  (a. =1, 2) and  $A_{\mu}$ .

$$e^{-1}d = -\frac{1}{4}R(\omega) - \frac{1}{2}\overline{\psi}^{a} \delta^{\mu\nu\ell} D_{\nu}(\omega;\tilde{\omega}) \Psi_{\ell a} - \frac{1}{4}F_{\mu\nu}F_{\ell\nu}g^{\mu\ell}g^{\nu\sigma} + \frac{e^{-1}}{6\sqrt{2}}E^{\mu\nu\ell}F_{\mu\nu}F_{\ell\nu}A_{\lambda} - \frac{\sqrt{2}}{46}(F_{\mu\nu} + \tilde{F}_{\mu\nu})\overline{\psi}^{\ell\nu} \epsilon_{\mu} \delta_{\mu}\Psi_{\ell}^{c}$$

with

and  $\omega$  is given by the 1st order formalism if  $\widehat{\omega}$  is defined as

After solving the equations of motion for  $\omega$  we get, as usual

$$\widetilde{\omega}_{nrs} = \omega_{nrs}^{\circ} (e) + \frac{i}{2} \left( \overline{\Psi}^{a} \delta_{s} \Psi_{ra} - \overline{\Psi}^{a} \delta_{\mu} \Psi_{sa} + \overline{\Psi}^{a}_{s} \delta_{r} \Psi_{\mu a} \right)$$

is invariant under the following N=2 supersymmetry transformation

$$\delta e_{\mu} = -i \vec{e}^{\alpha} \delta' \Psi_{\mu \alpha}$$
  

$$\delta \Psi_{\mu \alpha} = \left[ D_{\mu} (\hat{\omega}) + \frac{1}{4^{3}3} \hat{F}_{e_{\sigma}} (\chi^{e_{\sigma}} \delta_{\mu} + 2 \delta'^{e_{\sigma}} \delta_{\mu}^{\sigma}) \right] \epsilon_{\alpha}$$
  

$$\delta A_{\mu \alpha} = -\frac{\sqrt{3}}{4} \vec{e}^{\alpha} \Psi_{\mu \alpha}$$

All quartic terms are contained in the replacement of  $\omega$  and F by  $4\pm\omega$  and EtF in the bilinear fermionic terms. We note that N=2 supergravity in 5 dimensions has exactly the same structure as the N=1 supergravity in 11 dimensions where everything comes from. This should be compared with the partial purely geometric results obtained by D'Auria & Fré, 1980; D'Auria, Fré & Regge, 1980 for this N=2 supergravity in 5 dimensions.

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