An Instability in the Standard Model of Cosmology creates the **Anomalous Acceleration** without Dark Energy

Blake Temple, UC-Davis

Harvard University April 8, 2016

Collaborators: Joel Smoller and Zeke Vogler

The 1999 observations of redshift vs luminosity for type IA supernovae in nearby galaxies won the Nobel Prize because they discovered the

### **Anomalous Acceleration:**

The universe is expanding faster than the Standard Model of Cosmology (SM), based on Einstein's original theory of General Relativity, allows.

The only way to preserve the Cosmological Principlethat on the largest length scale the universe is described by a Friedmann Space-Time which holds no special placeis to add the Cosmological Constant to Einstein's equations as a source term. Its interpretation is

Dark Energy.

A best fit among Friedmann Space-Times with Dark Energy leads to the conclusion that the universe is a critical k=0 Friedmann Space-Time with Seventy Percent Dark Energy

 $\Omega_{\Lambda} \approx .7$ 

### 2007 PI talk in Relativity Session at AMS National Meeting in New Orleans:

We proposed the idea that a Simple Wave from the Radiation Epoch of the Big Bang might account for the Anomalous Acceleration of the Galaxies Without Dark Energy

(Our Motivation)

The Radiation Epoch:
After Inflation
until about

30,000 years after the Big Bang

is evolution by

Relativistic Compressible Euler Equations

The *p*-system with  $p = \frac{c^2}{3}\rho$ 

### PURE RADIATION

Stefan-Bolzman Law:  $ho = a \, T^4$ 

$$\rho = a T^4$$

(No Contact Discontinuities)

$$p = \frac{c^2}{3}\rho$$

The p-system with:

- Enormous sound speed  $\sigma \approx .57c$
- Enormous modulus of Genuine Nonlinearity

Every characteristic field contributes to Decay in the sense of Glimm and Lax

It is reasonable to expect fluctuations would decay to simple wave patterns by the End of Radiation

This was our Starting Assumption

### Stages of the Standard Model:

Uncoupling of Matter and Radiation

$$t \approx 3 \times 10^5$$

### Inflation

 $10^{-35}s$  $10^{-30}s$ 

### Pure Radiation

$$10^{-30} \text{ to } 3 \times 10^5 \text{ } yrs$$

$$p = \frac{c^2}{3}\rho$$

(Neglect Radiation Pressure)

$$p \approx 0$$

Time of CMB

379,000 yr

(Relativistic p-system)

Pursuing this Idea...

...we discovered that there is only one way

the Einstein equations can both perturb the Friedmann spacetimes and also reduce to ODE' when

$$p = \frac{c^2}{3}\rho$$

### ...we identified a

I-parameter family of Self-Similar Waves that perturb the Standard Model during the Radiation Epoch-And proposed that these might induce an Anomalous Acceleration at a later time.

We set out our ideas in PNAS in 2009 and Memoirs of the AMS in 2011

# Our interest is in the possible connection between these waves and the Anomalous Acceleration.

In Fact: This family of self-similar solutions was already from a different point of view...

### Cahill and Taub:

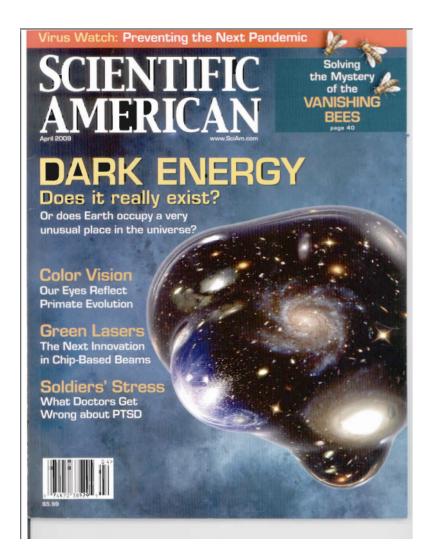
Commun Math Phys., 21, 1-40 (1971)

Extended by others, esp. Carr and Coley, Survey:

Physical Review D, 62, 044023-1-25 (1999)

### Around 2007:

Other research groups began exploring the possibility that the anomalous acceleration might be due to the earth lying near the center of a large region of



**Under-Density** 

We first saw publication in 2009

### The record is clear on one thing:

No one before us proposed this family of waves as a mechanism that could account for the **Anomalous Acceleration** without Dark Energy

We have now accomplished our goal of bringing the effects of these perturbations of the SM (waves) up to present time to compare with Dark Energy.

There are several surprises...
...in this talk I present what
we have found...

 We identify a instability in the SM based on a new (closed) asymptotic ansatz for local perturbations of the critical k=0 Friedmann Spacetime when p=0.

 The instability naturally creates a central region of accelerated uniform expansion on the scale of the supernova data within Einstein's original theory, without Dark Energy.

 The phase portrait of the instability is universal in the sense that it describes every smooth, spherically symmetric perturbation near the center, when p=0.

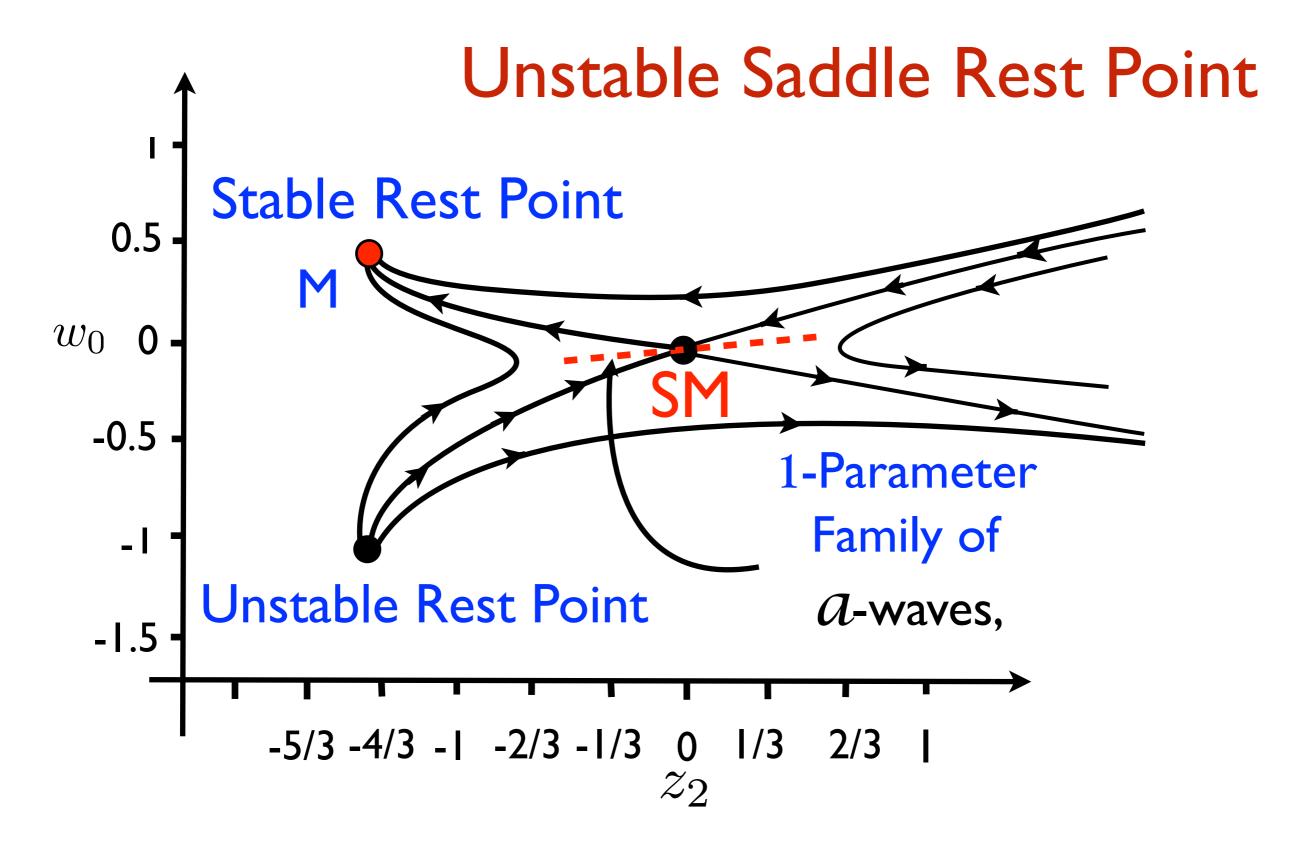
 The region of accelerated uniform expansion is one order of magnitude larger in extent than expected.

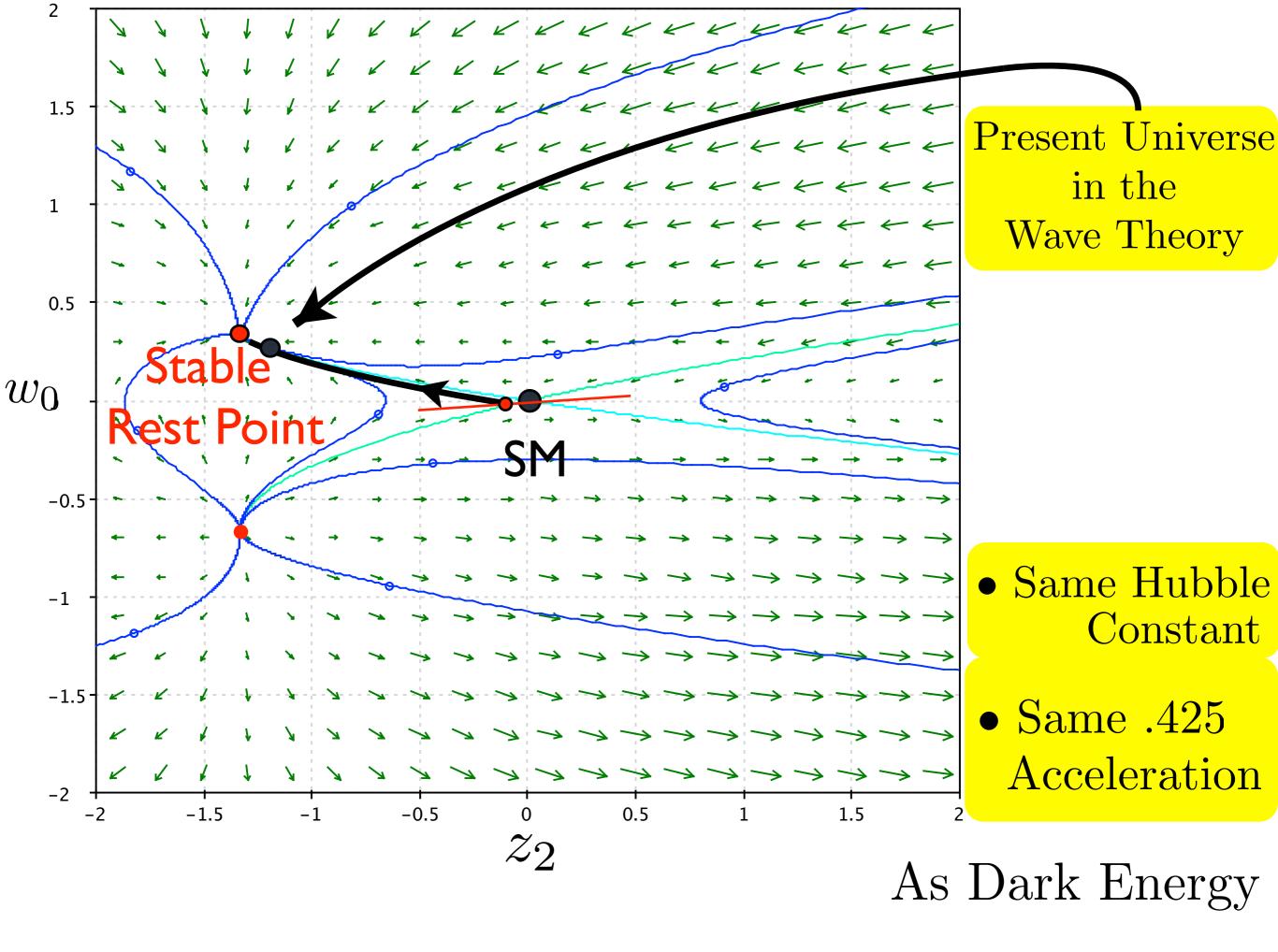
 The instability is triggered by our time asymptotic perturbations of SM from the

Radiation Epoch when: 
$$p = \frac{c^2}{3}\rho$$

- Surprisingly—The perturbations at the end of radiation do not directly cause the Anomalous Acceleration as we originally conjectured in PNAS.
- Rather—It is the non-trivial phase portrait of the instability they trigger when p=0 that creates the later accelerations.

 The phase portrait of the instability places the SM at a classic...





 The region of accelerated uniform expansion introduces precisely the same range of quadratic corrections to red-shift vs luminosity as does the cosmological constant in the theory of DE.

$$H_0 d_\ell = z + Q z^2 + O(z^3)$$

$$.25 \le Q \le .425 \le .5$$

 The results lead naturally to a testable alternative to Dark Energy within Einstein's original theory...

Without the Cosmological Constant.

 Our Proposal: The AA is due to a local under-dense perturbation of the SM on the scale of the supernova data, arising from time-asymptotic perturbations of SM from the Radiation Epoch that trigger an instability in the SM when the pressure drops to zero.

 A calculation shows the cubic correction is of the same order, but of a different sign, than the cubic correction in DE theory...

$$H_0 d_{\ell} = z + .425z^2 - .180z^3$$

Dark
Energy

$$H_0 d_\ell = z + .425z^2 + .359z^3$$

Wave Theory

 We address ONLY the anomalous acceleration...
 further assumptions regarding space-time far from the center would be required to connect the theory with other measurements...

## INTRODUCTION TO COSMOLOGY

### Edwin Hubble (1889-1953)

Hubble's Law (1929):

``The galaxies are receding from us at a velocity proportional to distance'



Universe is Expanding

Based on Redshift vs Luminosity

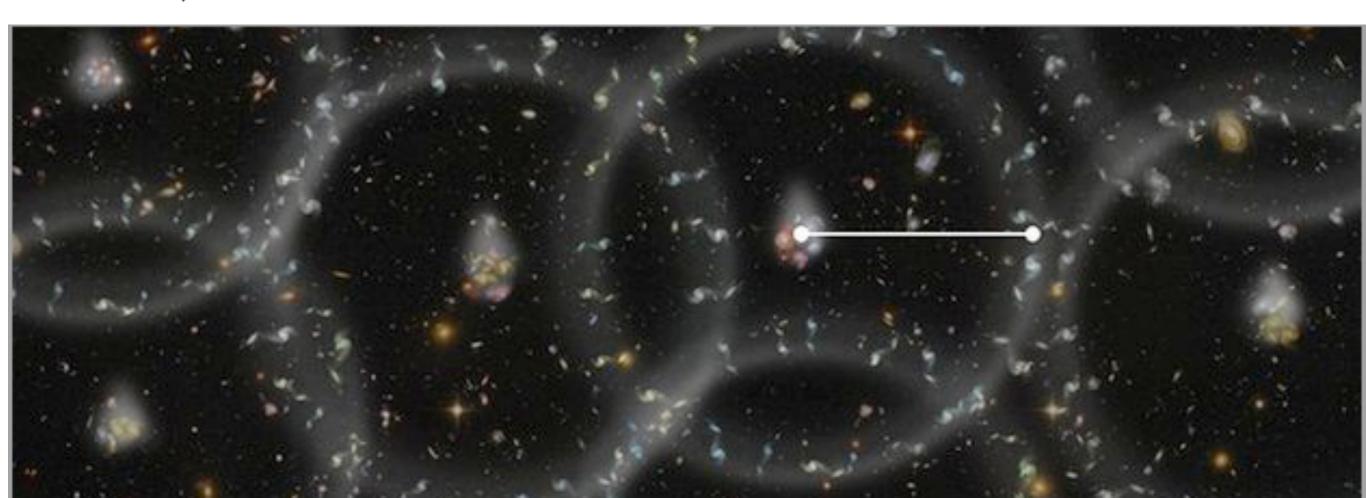
### Universe measured to 1% accuracy

By James Morgan Science reporter, BBC News, Washington DC

Astronomers have measured the distances between galaxies in the universe to an accuracy of just 1%.

This staggeringly precise survey - across six billion light-years - is key to mapping the cosmos and determining the nature of dark energy.

The new gold standard was set by **BOSS** (the Baryon Oscillation Spectroscopic Survey) using the Sloan Foundation Telescope in New Mexico, US.

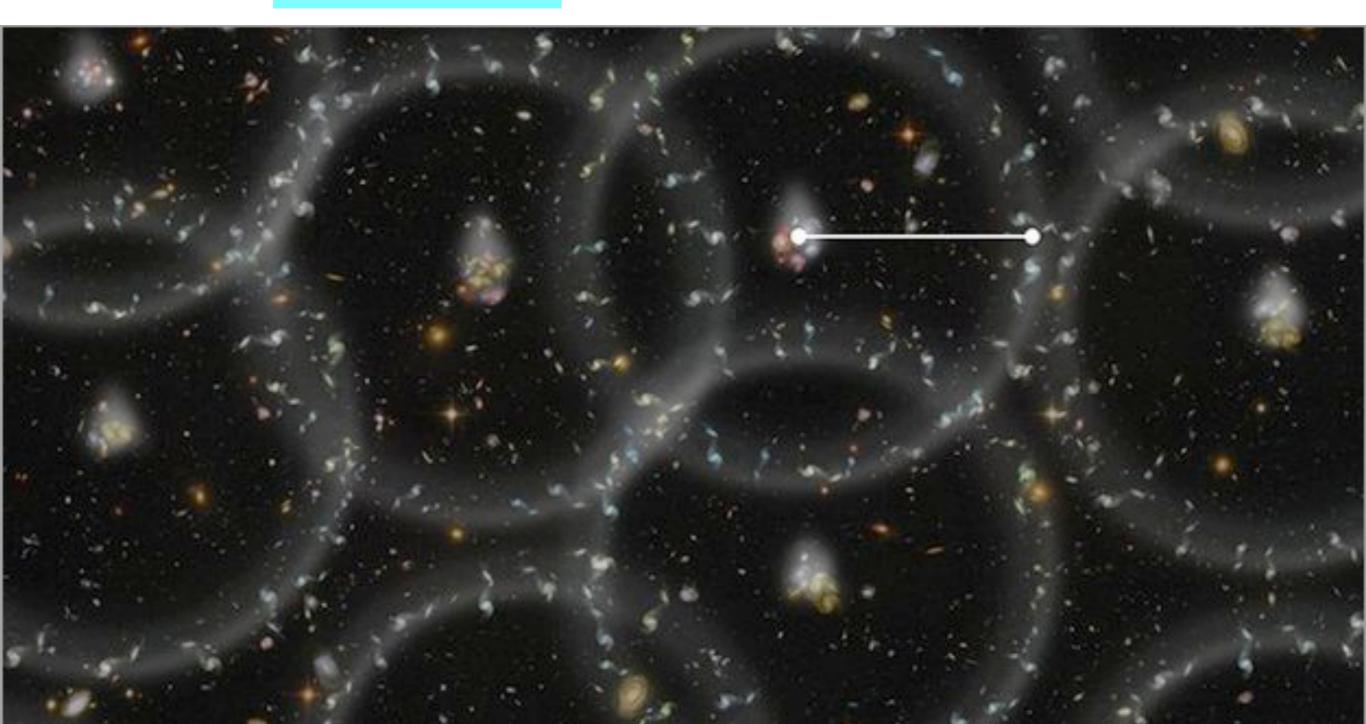


Frozen ripples
The BOSS team used baryon acoustic oscillations (BAOs) as a "standard ruler" to measure intergalactic distances.

BAOs are the "frozen" imprints of pressure waves that moved through the early universe - and help set the distribution of galaxies we see today.

"Nature has given us a beautiful ruler," said Ashley Ross, an astronomer from the University of Portsmouth.

"The ruler happens to be half a billion light years long, so we can use it to measure distances precisely, even from very far away."

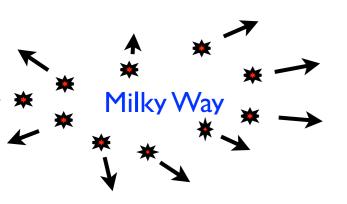


Conclude: The universe appears (and is assumed) uniform on a scale of about 1/20th

the distance across the visible universe

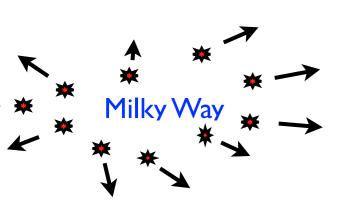
$$\xi = \frac{r}{ct} \approx \frac{.5 \text{ billion lightyear}}{13 \text{ billion lightyear}} \approx .04 \leq .05$$





500 million light-years  $\approx$  Uniform Density

Cosmic Length Scales



500 million light-years  $\approx$  Uniform Density

Cosmic Length Scales

• 50 million light-years  $\approx$  Separation between clusters of galaxies



500 million light-years  $\approx$  Uniform Density

Cosmic Length Scales

ullet 50 million light-years pprox Separation between clusters of galaxies



\*\*\*\*\* To million light-years  $\approx$  diameter of a cluster

500 million light-years  $\approx$  Uniform Density

Cosmic Length Scales

50 million light-years 

Separation between clusters of galaxies



ullet I million light-years pprox separation between galaxies in a cluster

500 million light-years  $\approx$  Uniform Density

Cosmic Length Scales

50 million light-years 

 Separation between clusters of galaxies



ullet I million light-years pprox separation between galaxies in a cluster

100 thousand light-years  $\approx$  distance across Milky Way

500 million light-years  $\approx$  Uniform Density

Cosmic Length Scales

■ 50 million light-years ≈ Separation between
 Local clusters of galaxies



ullet I million light-years pprox separation between galaxies in a cluster

100 thousand light-years  $\approx$  distance across Milky Way

ullet 28 thousand light-years pprox distance to galactic center

$$1 + z = \frac{R(t_0)}{R(t)}$$

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$$\xi \approx .1$$

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  $ct = (.1)(ct_0) \approx (.1)\frac{c}{H_0}$   $z \approx .07$ 



$$z \approx .07$$

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  $ct = (.1)(ct_0) \approx (.1)\frac{c}{H_0}$   $z \approx .07$ 



About a tenth of the distance to the Hubble Radius corresponds to about  $z \approx .07$ 

$$1 + z = \frac{R(t_0)}{R(t)}$$

$$p = 0$$

$$\frac{R(t_0)}{R(t)} = \left(\frac{t_0}{t}\right)^{2/3}$$

$$1 + z = \left(\frac{t_0}{t}\right)^{2/3}$$

$$ct = (.35)\frac{c}{H_0} \qquad \qquad z = 1$$

$$1 + z = \frac{R(t_0)}{R(t)}$$

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$$\frac{R(t_0)}{R(t)} = \left(\frac{t_0}{t}\right)^{2/3}$$

$$1 + z = \left(\frac{t_0}{t}\right)^{2/3}$$

$$ct = (.35)\frac{c}{H_0} \qquad \qquad z = 1$$

z=1 corresponds to about a "third of the way across the visible universe..."

#### Standard Model of Cosmology

• 1922 (Alexander Friedmann):

Derived FRW solutions of the Einstein equations: 3-space of constant curvature expanding in time:

$$ds^{2} = -dt^{2} + R(t)^{2} \left\{ \frac{dr^{2}}{1 - kr^{2}} + r^{2} d\Omega^{2} \right\}$$

 The Big Bang theory based on the FRW metric was worked out by <u>George Lemaître</u> in the late 1920's leading to Hubble's comfirmation of redshift vs luminoscity consistent with an FRW spacetime

Hubble's Constant 
$$\equiv H \equiv \frac{\dot{R}}{R}$$

■ In 1935: Howard Robertson and Arthur Walker derived Friedmann spacetime from the

### Copernican Principle:

"Earth is not in a special place in the Universe"

R-W: Friedmann uniquely determined by condition

Homogeneous and Isotropic about every point



Any point can be taken as r=0



Each t=const surface is a 3-space of constant scalar curvature

### Standard Model of Cosmology

Observations of the micro-wave background IMPLY

k = 0

"Critical expansion to within about 2-percent"

## The Friedmann metric when k=0:

$$ds^{2} = -dt^{2} + R(t)^{2} \left\{ dr^{2} + r^{2} d\Omega^{2} \right\}$$

The universe is infinite flat space  $\mathbb{R}^3$  at each fixed time:

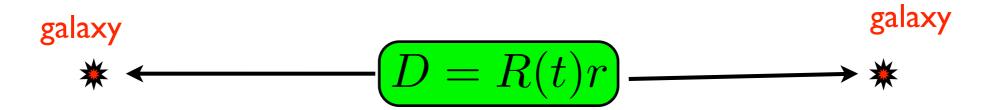
(Assumed to Apply on the Largest Length Scale)

### Standard Model of Cosmology

• FRW metric, k=0:

$$ds^{2} = -dt^{2} + R(t)^{2} \left\{ dr^{2} + r^{2} d\Omega^{2} \right\}$$

• D = Rr Measures distance between galaxies at each fixed t



Conclude:

$$\dot{D} = \dot{R}r = \frac{\dot{R}}{R}Rr = HD$$

$$\dot{D} = HD$$
 — Hubble's Law

Hubble's Constant 
$$\equiv H \equiv \frac{\dot{R}}{R}$$

Standard Model of Cosmology

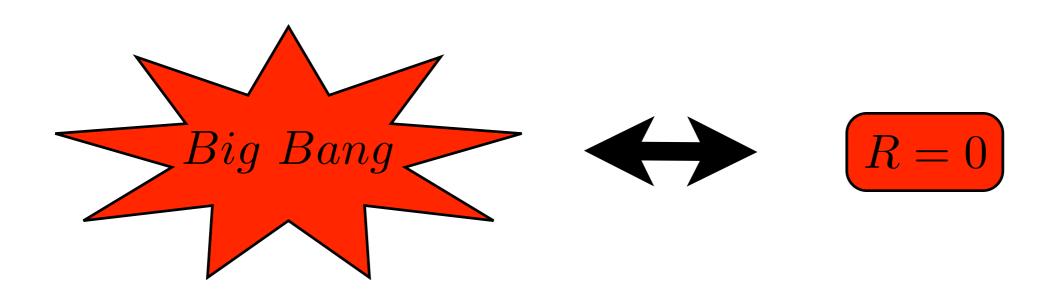
$$\left( ds^{2} = -dt^{2} + R(t)^{2} \left\{ dr^{2} + r^{2} d\Omega^{2} \right\} \right)$$

• Hubble's Law:

$$\dot{D} = HD$$

Conclude--

"The universe is expanding like a balloon"



### The Hubble "Constant" at present time

 The inverse Hubble Constant estimates the Age of the Universe

$$\frac{1}{H_0} \approx 10^{10} \text{ years} \approx \text{age of universe}$$

ullet is the distance of light travel since the Big Bang, a measure of the size of the visible universe

$$\frac{c}{H_0}$$
 = Hubble Length  $\approx 10^{10} \ lightyears$ 

### Measuring the Hubble Constant

 $m{D}$  Measures distance from Earth to distant galaxy at present time  $t_0$ 

$$H_0D = \dot{D} \qquad \text{Hubble's}$$
 Law 
$$& \\ & \\ D \approx d_\ell \equiv \text{luminosity distance} \\ \dot{D} \approx z \equiv \text{redshift factor} = \frac{\lambda_0 - \lambda_e}{\lambda_e}$$

$$H_0 d_{\ell} = z + \frac{1}{4}z^2 - \frac{1}{8}z^3 + O(z^4)$$

Friedmann k = 0

## Up until 1999, we could only measure the leading linear term:

$$H_0 d_\ell = z + \frac{1}{4} z^2 - \frac{1}{8} z^3 + O(z^4)$$
 Friedmann  $k = 0$ 

$$z << 1 \qquad H_0 \approx h_0 \, 100 \frac{km}{s \, mpc} \qquad h_0 \approx .68$$

 $mpc \approx 3.2$  million light years

"A galaxy at a distance of one mega-parsec is receding at about 68 kilometers per second..." The 1999 supernova data tested the dependence of the Hubble constant on time, and the results don't fit standard model...



"Anomalous Acceleration of Galaxies"



Introduction of "Cosmological Const" and "Dark Energy"

Dark energy is non-classical Negative pressure Anti-gravity effect

# The 1999 supernova data tested the dependence of the Hubble constant on time, and the results don't fit standard model...

$$H_0 d_{\ell} = z + \frac{1}{4} z^2 - \frac{1}{8} z^3 + O(z^4)$$

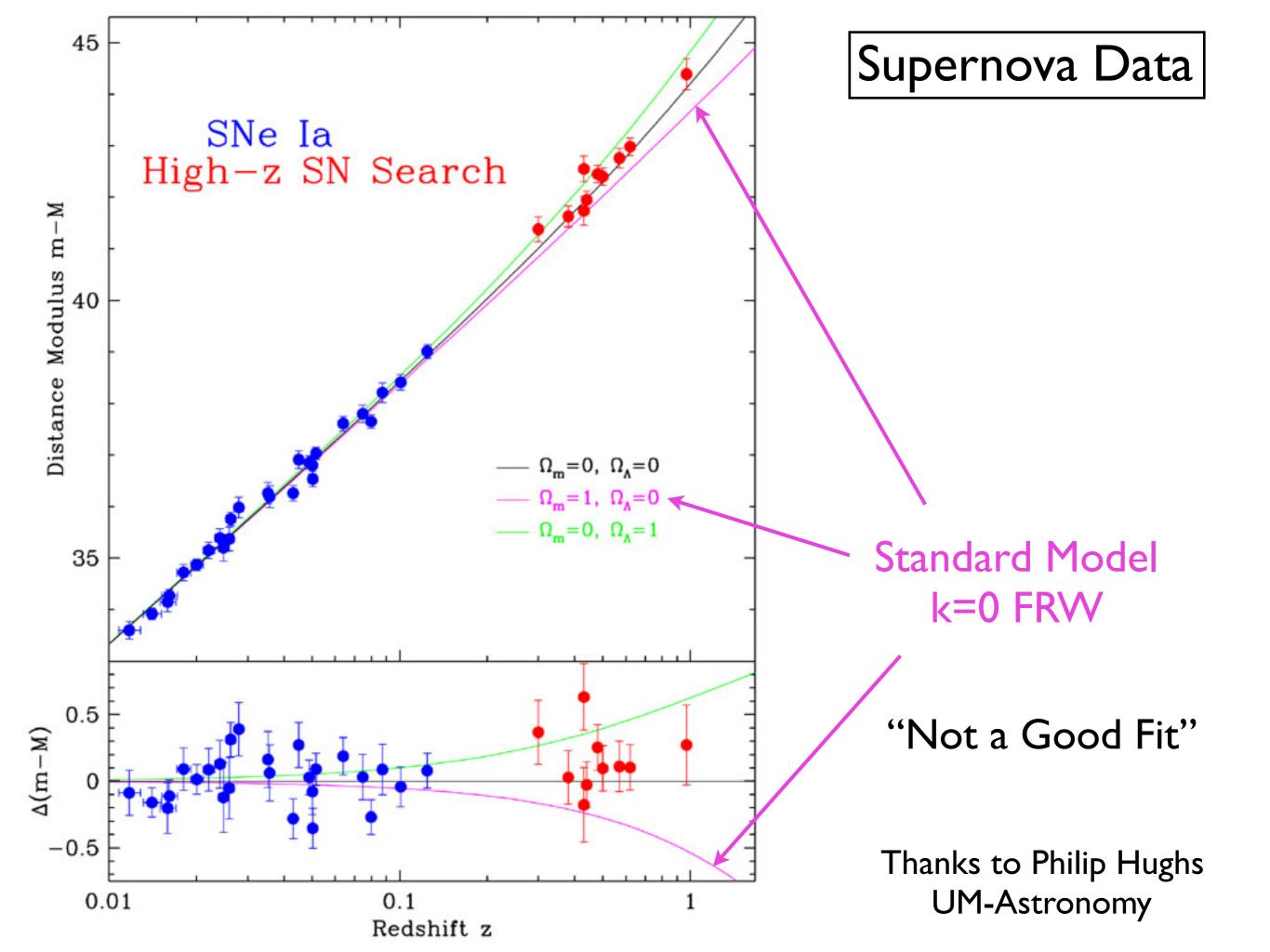
Friedmann k = 0

This is measured at about .425 not .25

The 1999 supernova data tested the dependence of the Hubble constant on time, and the results don't fit standard model...

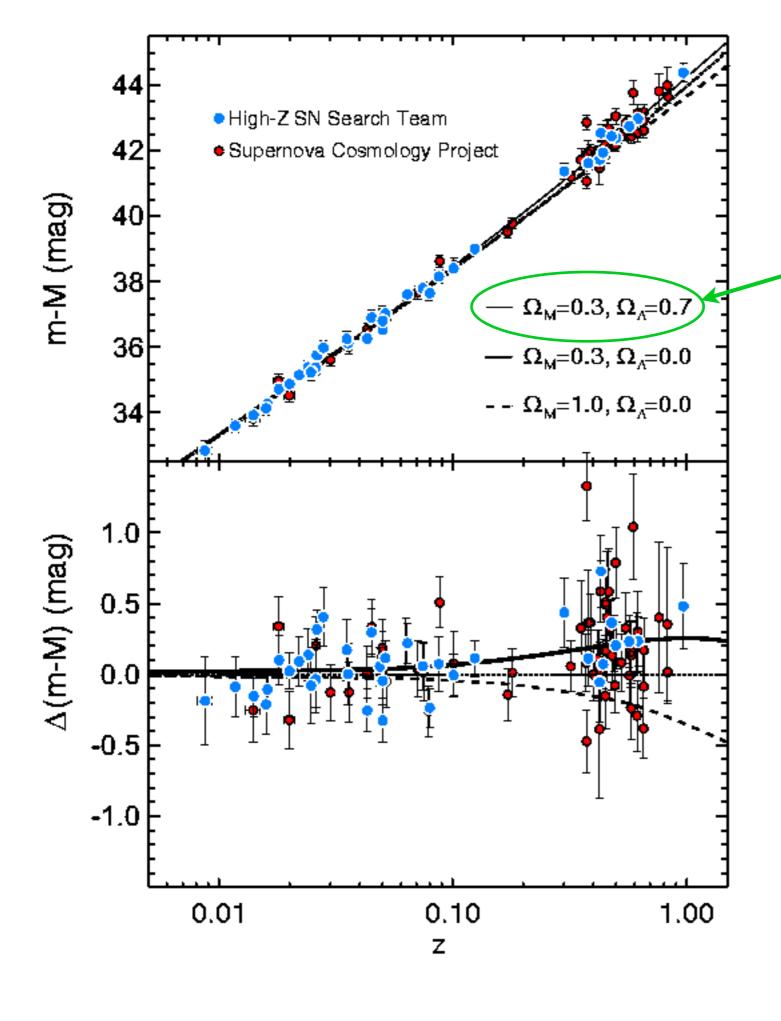
This is usually interpreted in terms of a Best Fit to Friedmann Universes with the Cosmological Constant

$$(k, \Omega_{\Lambda}) \rightarrow k = 0, \ \Omega_{\Lambda} \approx .7$$



That is: To preserve the Copernican Principle, that the Universe on the Largest Length Scale is evolving according to a Uniform Friedmann Spacetime with p=0, k=0A Cosmological Constant must be added To Einstein's Equations

The Physical Interpretation is Dark Energy



Thanks to Philip Hughs UM-Astronomy

### Einstein Equations for Friedmann:

Einstein Equations (1915):  $G_{ij} = \kappa T_{ij}$ 

 $G_{ij}$ =Einstein Curvature Tensor

$$T_{ij} = (\rho + p)u_iu_j + pg_{ij}$$
=Stress Energy Tensor (perfect fluid)

Einstein Equations for k=0 Friedmann metric:

$$H^2 = \frac{\kappa}{3}\rho$$

$$\dot{\rho} = -3(\rho + p)H$$



• Assume Einstein equations with a cosmological constant:

$$G_{ij} = 8\pi T_{ij} + \Lambda g_{ij}$$

• Assume 
$$k = 0$$
 FRW:  $ds^2 = -dt^2 + R(t)^2 \{dr^2 + r^2 d\Omega^2\}$ 

• Leads to:

$$H^2 = \frac{\kappa}{3}\rho + \frac{\kappa}{3}\Lambda$$

• Divide by  $H^2 = \frac{\kappa}{3} \rho_{crit}$ 

$$1 = \Omega_M + \Omega_{\Lambda}$$

• Best data fit leads to  $\Omega_{\Lambda} \approx .7$  and  $\Omega_{M} \approx .3$ 

Implies: The universe is 70 percent dark energy

More slowly...

$$H^2 = \frac{\kappa}{3}\rho + \frac{\kappa}{3}\Lambda$$

$$H^2 = \frac{\kappa}{3}\rho + \frac{\kappa}{3}\Lambda$$
 Constant in time

$$H^2 = \frac{\kappa}{3}\rho + \frac{\kappa}{3}\Lambda$$

Decreases to zero as  $t \to \infty$ 

$$\frac{H^2 = \frac{\kappa}{3}\rho + \frac{\kappa}{3}\Lambda}{H^2}$$

$$1 = \frac{\frac{\kappa}{3}\rho}{H^2} + \frac{\frac{\kappa}{3}\Lambda}{H^2}$$

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$$\Omega_{\Lambda}$$

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$$=\Omega_M+\Omega_\Lambda$$

Conclude...

$$1 = \frac{\frac{\kappa}{3}\rho}{H^2} + \frac{\frac{\kappa}{3}\Lambda}{H^2}$$

$$=\Omega_M+\Omega_\Lambda$$

$$\Omega_{\Lambda} \approx 0 \rightarrow 1$$
 as  $t \approx t_{rad} \rightarrow \infty$ 

$$t \approx t_{rad} \rightarrow \infty$$

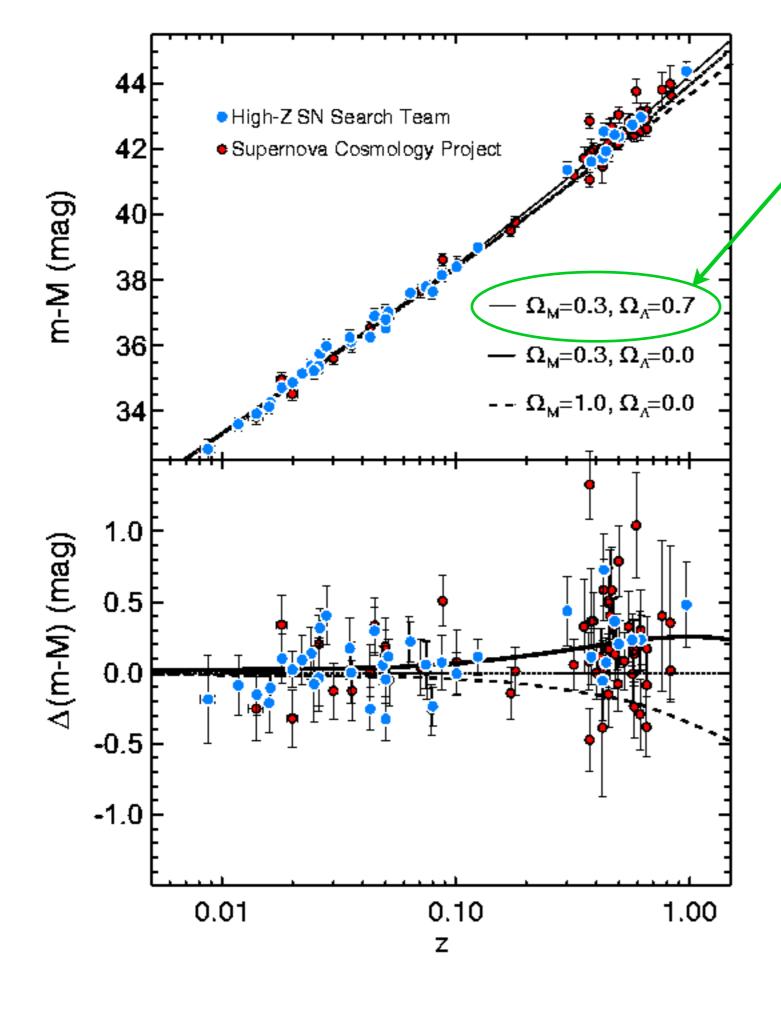
#### Incorporating Dark Energy into Friedmann

$$1 = \frac{\frac{\kappa}{3}\rho}{H^2} + \frac{\frac{\kappa}{3}\Lambda}{H^2}$$

$$=\Omega_M+\Omega_\Lambda$$

Best Fit...  $\Omega_{\Lambda} \approx .7$ 

$$\Omega_{\Lambda} \approx .7$$



# \*\* Best Fit: 70% Dark Energy 30% Classical Energy

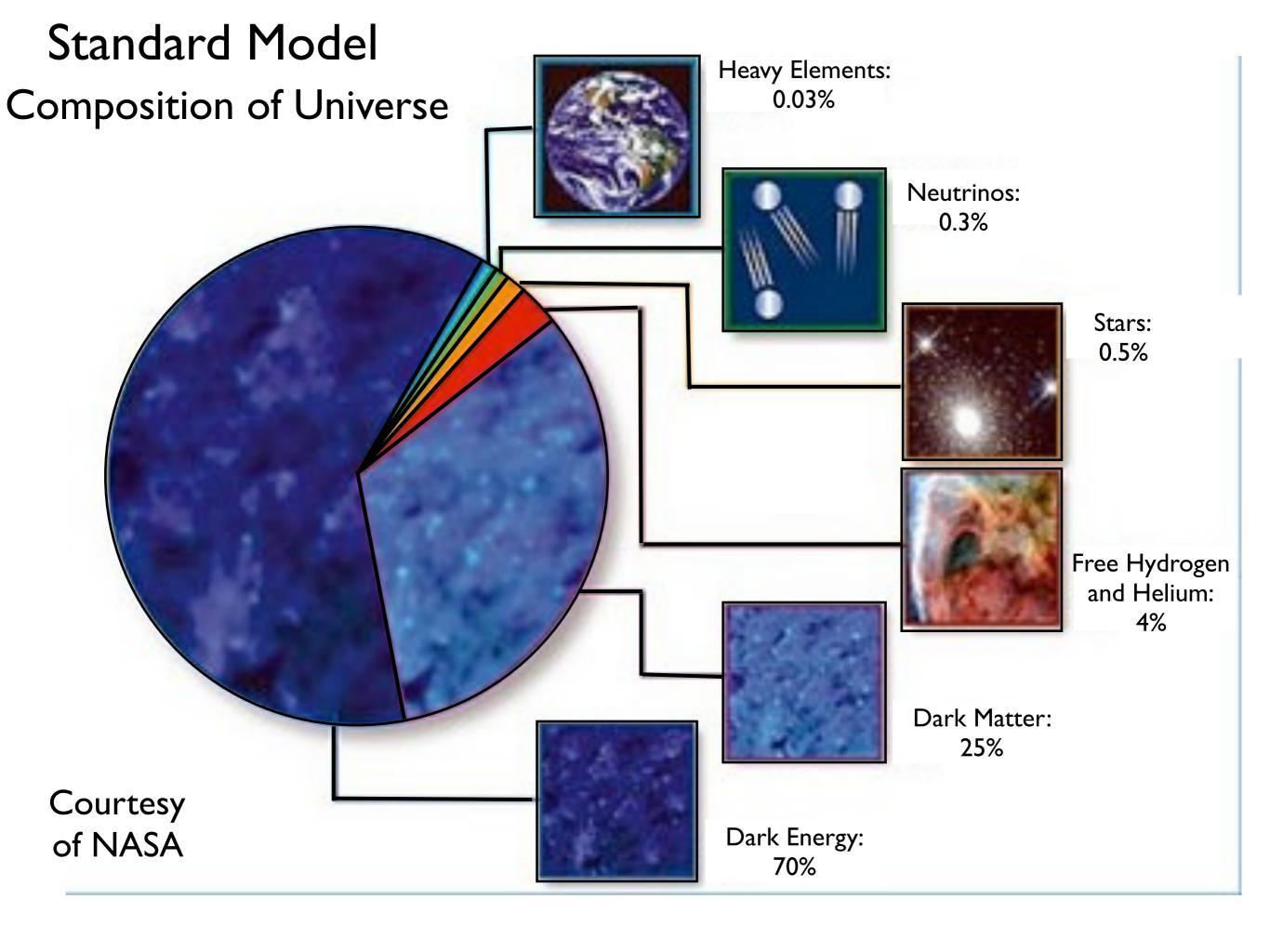
- m M = "Distance Modulus"M=absolute Magnitudem=apparent magnitude
- d=distance in parsecs:

$$m - M = 5 \log(d) - 5$$

z=redshift factor

$$1+z = \frac{\lambda_{emit}}{\lambda_{obs}}$$

•  $\Omega_m + \Omega_{\Lambda} = 1$  for a flat (k = 0) universe.



# The Question we Explore:

"Could the Anomalous Acceleration of the galaxies be due to the fact that we are looking outward into an expansion wave that formed during the Radiation Epoch of the Big Bang?"

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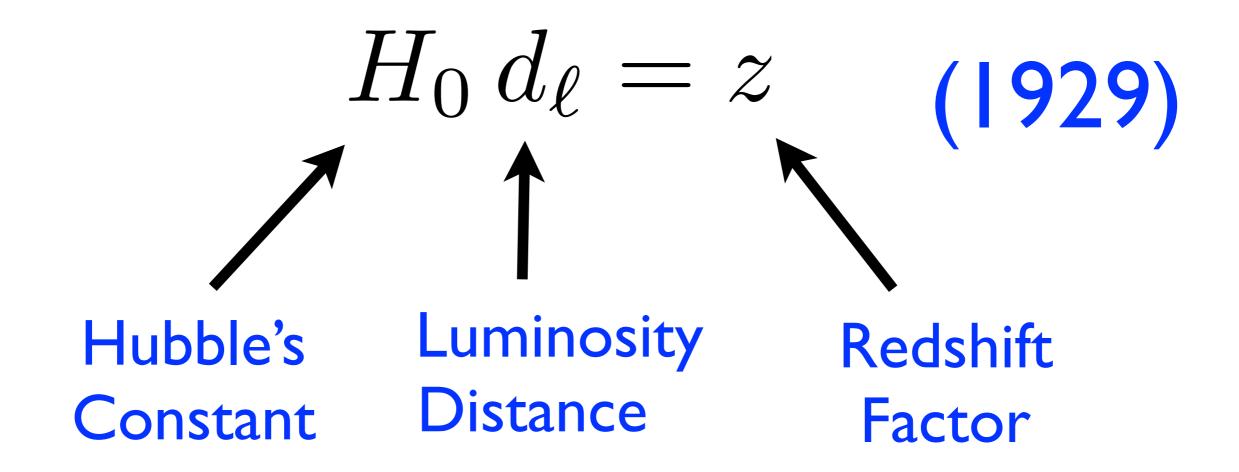
"Could the Anomalous Acceleration of the galaxies be due to the fact that we are looking outward into an expansion wave that formed during the Radiation Epoch of the Big Bang?"

\* The Einstein equations have been confirmed without the cosmological constant in every setting except cosmology...

Note: A general expansion wave has a center of expansion...

# Summary of our results for the Wave Theory

# Hubbles Law:



Measured value:

$$H_0 = h_0 \, \frac{100km}{s \, mpc}$$

$$h_0 \approx .68$$

# The 1999 Supernova data was refined enough to measure the quadratic correction to Hubble's Relation:

$$H_0 d_\ell = z + \underbrace{??}_{Z^2}$$

# Einstein's Equations: $G = \kappa T + \Lambda q$

$$G = \kappa T + \Lambda g$$

$$\Omega_M + \Omega_{\Lambda} = 1$$

Cosmological
Constant 1999

$$H_0 d_\ell = z + .25z^2 + O(z^3) \quad \frac{\text{Friedmann}}{\Omega_\Lambda = 0}$$

**Anomalous** 

Acceleration

$$H_0 d_{\ell} = z + .425z^2 + O(z^3)$$

$$\Omega_{\Lambda}=.7$$

WE PROVE: The Friedmann Universe is UNSTABLE

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A small wave perturbation at the end of radiation will expand to create a large region of accelerated uniform expansion at the Center of the Wave

This induces exactly the same range of quadratic corrections to redshift vs luminosity as does

Dark Energy

# The self-similar perturbations we identified at the end of the radiation epoch TRIGGER this instability when p=0

The self-similar perturbations we identified at the end of the radiation epoch TRIGGER this instability when p=0

# This induces exactly the same range of Q as does Dark Energy:

$$H_0 d_\ell = z + Q z^2 + O(z^3)$$

### Dark Energy

$$H_0 d_{\ell} = z + .25 \left(1 + \Omega_{\Lambda}\right) z^2 - .125 \left(1 + \frac{2}{3}\Omega_{\Lambda} - \Omega_{\Lambda}^2\right) z^3 + O(z^4)$$

$$.25 \le Q \le .5$$

as

$$\Omega_M + \Omega_{\Lambda} = 1$$

$$0 \leq \Omega_{\Lambda} \leq 1$$

• In the case  $\Omega_M = .3$ ,  $\Omega_{\Lambda} = .7$  this gives

$$H_0 d_\ell = z + .425 z^2 - .1804 z^3 + O(z^4)$$

$$H_0 d_\ell = z + Q(z_2, w_0)z^2 + C(z_2, w_0, w_2)z^3 + O(z^4)$$

#### Orbit evolves to a NEW STABLE REST POINT

• A Wave with Underdensity:

$$\frac{\rho_{SM} - \rho_{ssw}}{\rho_{SM}} = 7.45 \times 10^{-6}$$

$$H_0 d_\ell = z + .425 z^2 + .359 z^3 + O(z^4)$$

Conclusion: The cubic correction is of the same order, but of a different sign, from Dark Energy... A Testable Prediction!

$$H_0 d_\ell = z + .425z^2 - .180z^3$$

Dark Energy

$$H_0 d_\ell = z + .425z^2 + .359z^3$$

Wave Theory

#### Self-Similar Solutions

The Friedmann spacetimes admit self-similar expressions when  $p = \sigma^2 \rho$ 

$$ds^{2} = -B(\xi)dt^{2} + \frac{1}{A(\xi)}dr^{2} + r^{2}d\Omega^{2}$$

$$\xi = \frac{r}{ct}$$
 "Fractional Distance to the Hubble Radius"

$$\rho r^2 = z(\xi)$$
 "Dimensionless Density"

$$\frac{v}{\xi} = w(\xi)$$
 "Dimensionless Velocity"

#### The p=0 Friedmann Universe in Self-Similar Coordinates:

$$ds^{2} = -B_{F}(\xi)d\bar{t}^{2} + \frac{1}{A_{F}(\xi)}d\bar{r}^{2} + \bar{r}^{2}d\Omega^{2}$$

$$p = 0$$

$$A_F(\xi) = 1 - \frac{4}{9}\xi^2 - \frac{8}{27}\xi^4 + O(\xi^6)$$
$$D_F(\xi) \equiv \sqrt{A_F B_F} = 1 - \frac{1}{9}\xi^2 + O(\xi^4).$$

$$z_F(\xi) = \frac{4}{3}\xi^2 + \frac{40}{27}\xi^4 + O(\xi^6)$$
$$w_F \equiv \frac{v}{\xi} = \frac{2}{3} + \frac{2}{9}\xi^2 + O(\xi^4)$$

The  $p = \frac{1}{2}\rho$  Friedmann Universe in Self-Similar Coordinates:

$$p = \frac{c^2}{3}\rho$$

 $p = \frac{c^2}{3}\rho$  Pure Radiation  $\bar{\xi} \neq \xi$ 

$$ar{\xi} 
eq \xi$$

$$z_{1/3} = \frac{3}{4}\bar{\xi}^2 + \frac{9}{16}\bar{\xi}^4 + O(\bar{\xi}^6)$$

$$v_{1/3} = \frac{1}{2}\bar{\xi} + \frac{1}{8}\bar{\xi}^3 + O(\bar{\xi}^5)$$

$$A_{1/3} = 1 - \frac{1}{4}\bar{\xi}^2 - \frac{1}{8}\bar{\xi}^4 + O(\bar{\xi}^6)$$

$$D_{1/3} = 1 + O(\bar{\xi}^4)$$

The  $p = \frac{c^2}{3}\rho$  Friedmann Universe extends to I-parameter family of Self-Similar spacetimes that perturb the Standard Model during the Radiation Epoch:

The p=0 Friedmann Universe DOES NOT admit Self-Similar perturbations!

(Something has to give when p drops to zero!)

(The topic of our PNAS and MEMOIR)

# A 1-parameter family of solutions depending on

the Acceleration Parameter  $0 < a < \infty$ 

$$0 < a < \infty$$

$$p = \frac{1}{3}\rho$$

$$z_{1/3}^{a} = \frac{3a^{2}}{4}\bar{\xi}^{2} + \left[\frac{9a^{2}}{16} + 3a^{2}\left(V_{0} + A_{0}\right)\left(1 - a^{2}\right)\right]\bar{\xi}^{4} + O(\bar{\xi}^{6})$$

$$v_{1/3}^a = \frac{1}{2}\bar{\xi} + \left[\frac{1}{8} + V_0(1 - a^2)\right]\bar{\xi}^3 + O(\bar{\xi}^5)$$

$$A_{1/3}^a = 1 - \frac{a^2}{4}\bar{\xi}^2 - \left[\frac{a^2}{8} + a^2A_0(1 - a^2)\right]\bar{\xi}^4 + O(\bar{\xi}^6)$$

$$D_{1/3}^a = 1 + O(\bar{\xi}^4)$$

$$V_0 = \frac{2}{3}A_0 = \frac{1}{20}$$

The ANSATZ that triggers the instability when p=0:

#### The ANSATZ that triggers the instability:

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + \left\{\frac{2}{9} + w_2(t)\right\}\xi^2 + O(\xi^4),$$

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$$\xi = \frac{r}{ct}$$

"Fractional Distance to the Hubble Radius"

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

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$$\xi = rac{r}{ct}$$
 "Fractional Distance to the Hubble Radius"

$$z(t,\xi) = \rho r^2$$
 "Dimensionless Density"

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + \left\{\frac{2}{9} + w_2(t)\right\}\xi^2 + O(\xi^4),$$

$$\xi = \frac{r}{ct}$$
 "Fractional Distance to the Hubble Radius"

$$z(t,\xi) = \rho r^2$$
 "Dimensionless Density"

$$w(t,\xi) = \frac{v}{\xi}$$
 "Dimensionless Velocity"

### In a non-uniform spacetime:

 $\xi$  = `Fractional distance to the Hubble Radius' measures (approximately) how far out you would think you were if you believed you were at the center of a Friedmann spacetime...

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + \left\{\frac{2}{9} + w_2(t)\right\}\xi^2 + O(\xi^4),$$

# Only EVEN powers of $\xi$ ...

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

# Uniform Density out to errors $\xi^4$

$$z(t,\xi) = \rho r^2$$

$$\rho(t) \sim \frac{\left(\frac{4}{3} + z_2(t)\right)}{t^2} = \frac{f(t)}{t^2}$$

**THEOREM:** The p = 0 waves take the asymptotic form

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + \left\{\frac{2}{9} + w_2(t)\right\}\xi^2 + O(\xi^4),$$

where  $z_2(t), z_4(t), w_0(t), w_2(t)$  evolve according to the equations

$$-t\dot{z}_{2} = 3w_{0}\left(\frac{4}{3} + z_{2}\right),$$

$$-t\dot{z}_{4} = -5\left\{\frac{2}{27}z_{2} + \frac{4}{3}w_{2} - \frac{1}{18}z_{2}^{2} + z_{2}w_{2}\right\}$$

$$-5w_{0}\left\{\frac{4}{3} - \frac{2}{9}z_{2} + z_{4} - \frac{1}{12}z_{2}^{2}\right\},$$

$$-t\dot{w}_{0} = \frac{1}{6}z_{2} + \frac{1}{3}w_{0} + w_{0}^{2},$$

$$-t\dot{w}_{2} = \frac{1}{10}z_{4} + \frac{4}{9}w_{0} - \frac{1}{3}w_{2} + \frac{1}{24}z_{2}^{2} - \frac{1}{3}z_{2}w_{0}$$

$$-\frac{1}{3}w_{0}^{2} + 4w_{0}w_{2} - \frac{1}{4}w_{0}^{2}z_{2}.$$

$$H_0 d_\ell = z + Q(z_2, w_0)z^2 + C(z_2, w_0, w_2)z^3 + O(z^4)$$

$$25 \le Q \le .5$$

$$z_2' = -3w_0 \left(\frac{4}{3} + z_2\right)$$

$$w_0' = -\left(\frac{1}{6}z_2 + \frac{1}{3}w_0 + w_0^2\right)$$

$$H_0 d_{\ell} = z + Q(z_2, w_0) z^2 + C(z_2, w_0, w_2) z^3 + O(z^4)$$

$$.25 \le Q \le .5$$

as

$$z_2' = -3w_0 \left(\frac{4}{3} + z_2\right)$$

$$w_0' = -\left(\frac{1}{6}z_2 + \frac{1}{3}w_0 + w_0^2\right)$$

$$Q(z_2, w_0) = \frac{1}{4} + \frac{24w_0 + 45w_0^2 + 3z_2}{4(2+3w_0)^2}$$

$$H_0 d_\ell = z + Q(z_2, w_0)z^2 + C(z_2, w_0, w_2)z^3 + O(z^4)$$

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$$\frac{1}{4} = Q(0,0) \le Q \le Q(M) = \frac{1}{2}$$

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$$\frac{1}{4} = Q(0,0) \le Q \le Q(M) = \frac{1}{2} \quad \text{(Along orbit } SM \to M\text{)}$$

### Our Wave Theory

$$H_0 d_\ell = z + Q(z_2, w_0)z^2 + C(z_2, w_0, w_2)z^3 + O(z^4)$$

$$25 \le Q \le .5$$

$$z_2' = -3w_0 \left(\frac{4}{3} + z_2\right)$$

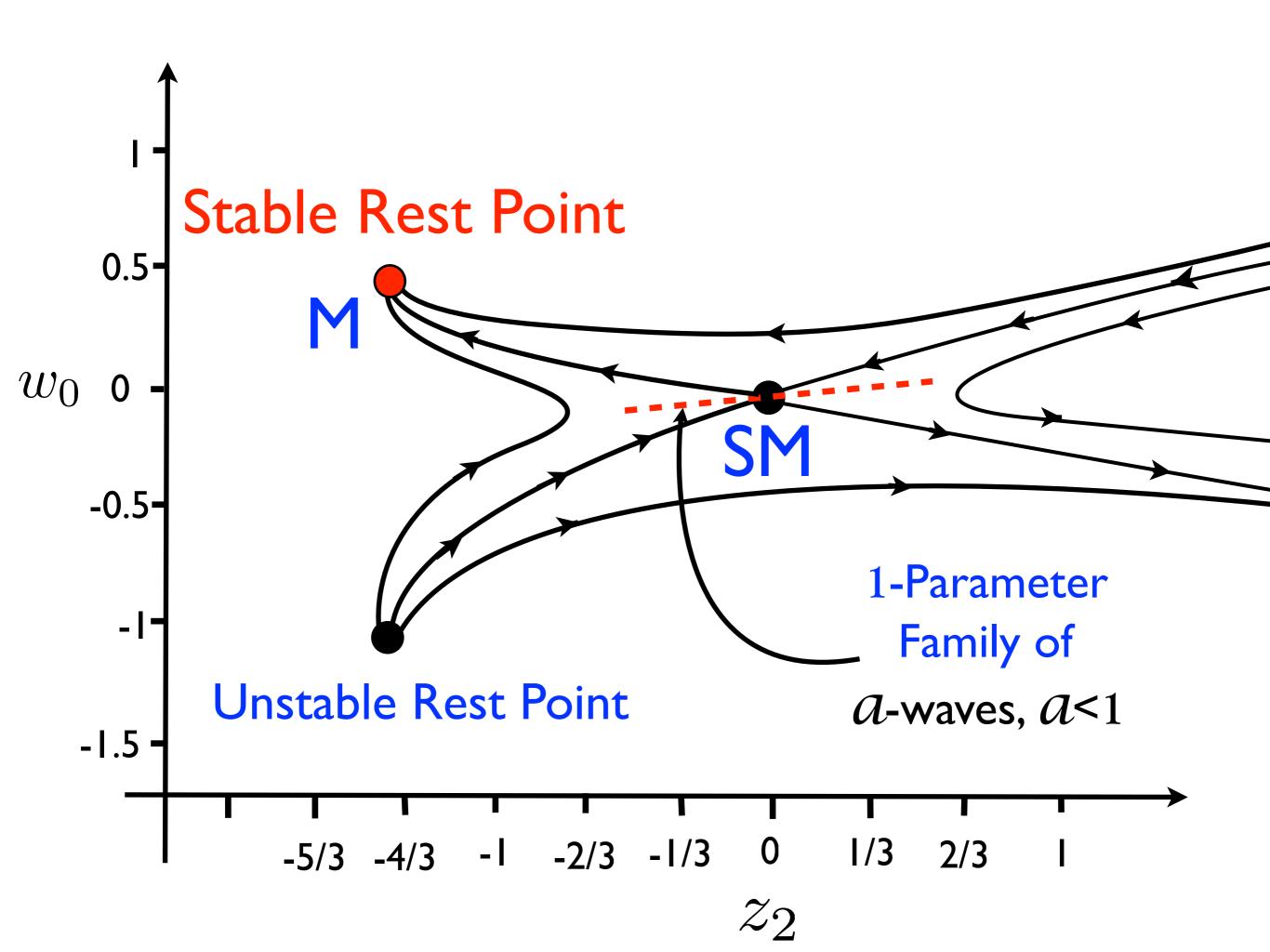
$$w_0' = -\left(\frac{1}{6}z_2 + \frac{1}{3}w_0 + w_0^2\right)$$

### Orbit evolves to a NEW STABLE REST POINT

• A Wave with Under-density:

$$\frac{\rho_{SM} - \rho_{ssw}}{\rho_{SM}} = 7.45 \times 10^{-6}$$

$$H_0 d_{\ell} = z + .425z^2 + .359z^3 + O(z^4)$$



Strategy: Use our equations to evolve the initial data for a-waves at the end of radiation to determine  $(a, T_*)$  that gives the correct anomalous acceleration.

I.e.,  $(a,T_*)$  that give the observed quadratic correction to redshift vs luminosity at present time

In the Standard Model p=0 at about

$$t_* \approx$$
 10,000-30,000

$$T_* \approx 9000^0 K$$

(Depending on theories of Dark Matter)

- Our simulation turns out to be entirely insensitive to the initial  $t_{st}$ ,  $T_{st}$
- I.e., we need only compute the value of the acceleration parameter that gives the correct anomalous acceleration.

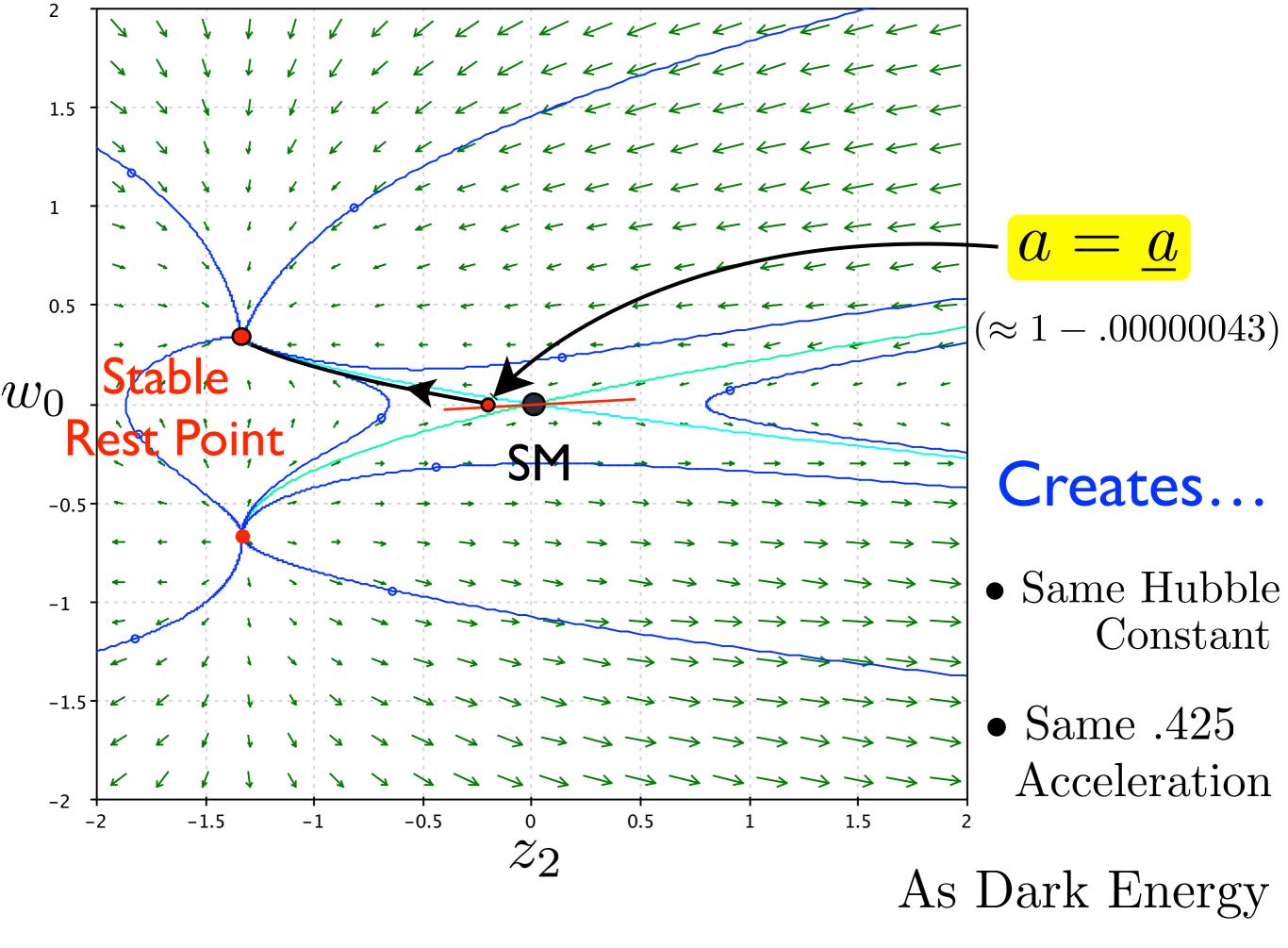
THE ANSWER: The value of the acceleration for the wave perturbation of SM that produces a quadradic correction of .425 at the present value of  $H_0$  is:

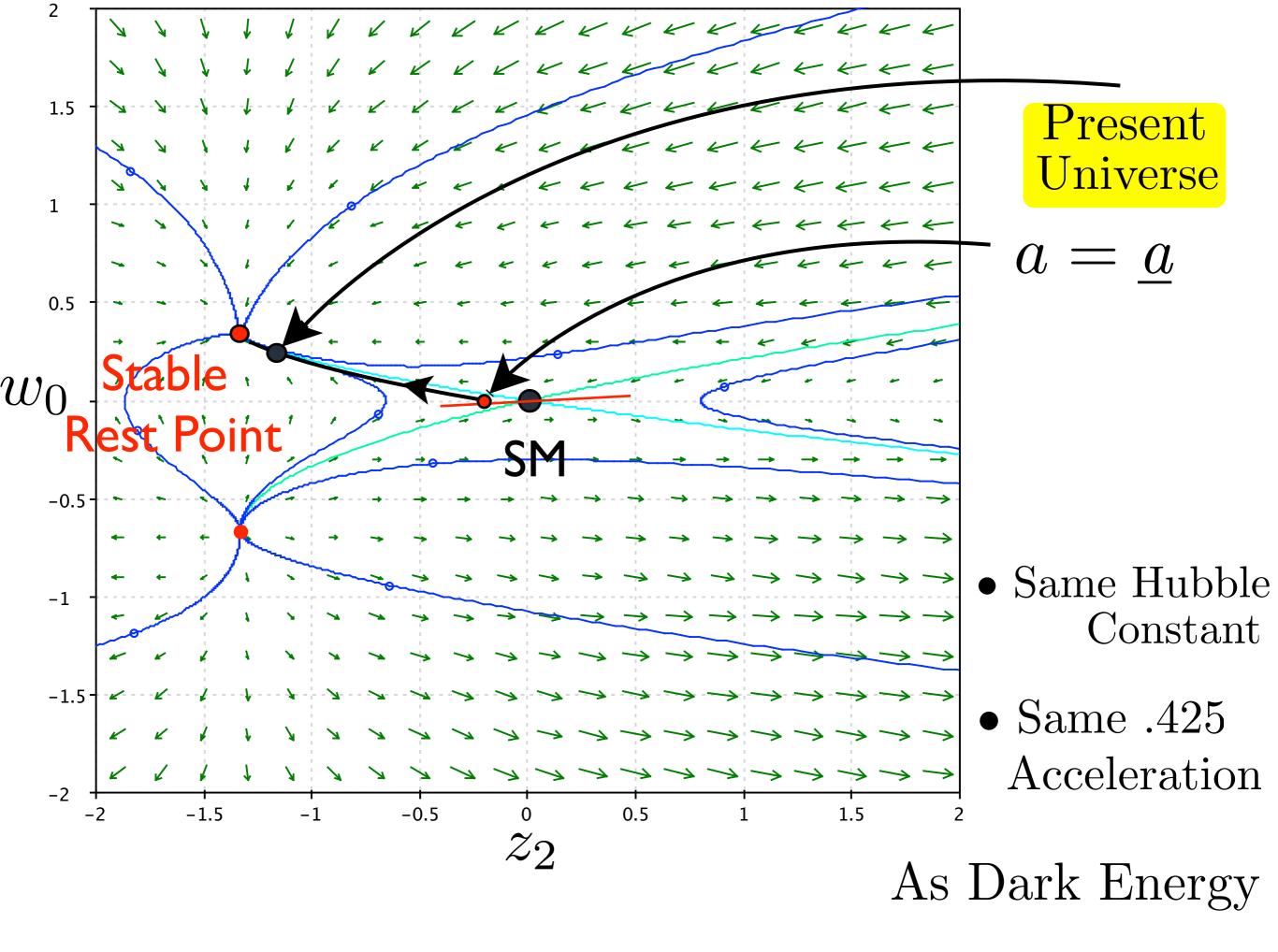
$$\underline{a} = 0.99999957 = 1 - \left(4.3 \times 10^{-7}\right)$$

$$H_0 d_\ell = z + .425z^2 + .359z^3$$

This corresponds to a relative under-density of

$$\frac{\rho_{SM} - \rho_{ssw}}{\rho_{SM}} = 7.45 \times 10^{-6}$$





The relative under-density at the end of radiation:

$$\frac{\rho_{SM} - \rho_{ssw}}{\rho_{SM}} = 7.45 \times 10^{-6}$$

The relative under-density at present time:

$$\frac{\rho_{ssw}(t_0)}{\rho_{SM}(t_0)} = .1438 \approx \frac{1}{7}.$$

### Conclude:

An under-density of one part in  $10^6$ at the end of radiation produces a seven-fold under-density at present time...

### CONCLUDE:

# The Standard Model is Unstable to Perturbation by this I-parameter family of Waves

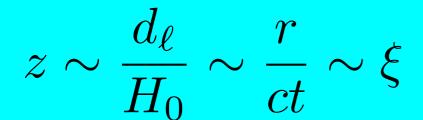
### Comparison with Dark Energy:

$$H_0 d_{\ell} = z + .425z^2 - .180z^3$$

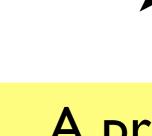
Dark Energy

$$H_0 d_\ell = z + .425z^2 + .359z^3$$

Wave Theory



Measures Fractional Distance to Hubble Radius z << 1



A prediction:
The wave contributes
MORE to the Anomalous
Acceleration
far from the center

### Neglecting $O(\xi^4)$ errors: The spacetime near the center evolves toward the Stable Rest Point

- The metric tends to Flat Minkowski Spacetime which is not co-moving with the fluid
- BUT: The evolution creates a uniformly expanding density near the center, which, neglecting relativistic corrections,

is CENTER-INDEPENDENT

(like Friedmann Spacetimes)

#### CONCLUDE:

### The wave creates a

### UNIFORMLY EXPANDING SPACETIME

with an

**ANOMALOUS ACCELERATION** 

in a

LARGE, FLAT, CENTER-INDEPENDENT

region near the center of the wave

### Neglecting errors $O(\xi^4)$ :

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + \left\{\frac{2}{9} + w_2(t)\right\}\xi^2 + O(\xi^4)$$

$$z \sim density \qquad w \sim velocity$$
 
$$\xi = \frac{r}{t} \sim fractional\ distance\ to\ Hubble\ Length$$

THEOREM: Neglecting  $O(\xi^4)$  errors, as the orbit tends to the Stable Rest Point:

• The Density drops FASTER than SM:

$$\rho_{WAVE}(t) = \frac{k_0}{t^3(1+\bar{w})} \quad \rho_{SM}(t) = \frac{4}{3t^2}$$

where  $\bar{w}(t)$  and  $k_0(t)$  change exponentially slowly.

The metric tends to FLAT MINKOWSKI:

$$ds^2 = -dt^2 + dr^2 + r^2 d\Omega^2$$

 $\underline{a} = 0.99999956 \approx 1 - 4.3 \times 10^{-7}$  such that:

 $\underline{a} = 0.99999956 \approx 1 - 4.3 \times 10^{-7}$  such that:

• The p=0 evolution starting from this initial data evolves to  $H=H_0, Q=.425$  at  $t=t_0$ , in agreement with Dark Energy at  $t=t_{DE}$ .

$$\underline{a} = 0.99999956 \approx 1 - 4.3 \times 10^{-7}$$
 such that:

- The p = 0 evolution starting from this initial data evolves to  $H = H_0$ , Q = .425 at  $t = t_0$ , in agreement with Dark Energy at  $t = t_{DE}$ .
- The cubic correction is C = 0.359 at  $t = t_0$ , while Dark Energy is C = -0.180 at  $t = t_{DE}$ .

$$\underline{a} = 0.99999956 \approx 1 - 4.3 \times 10^{-7}$$
 such that:

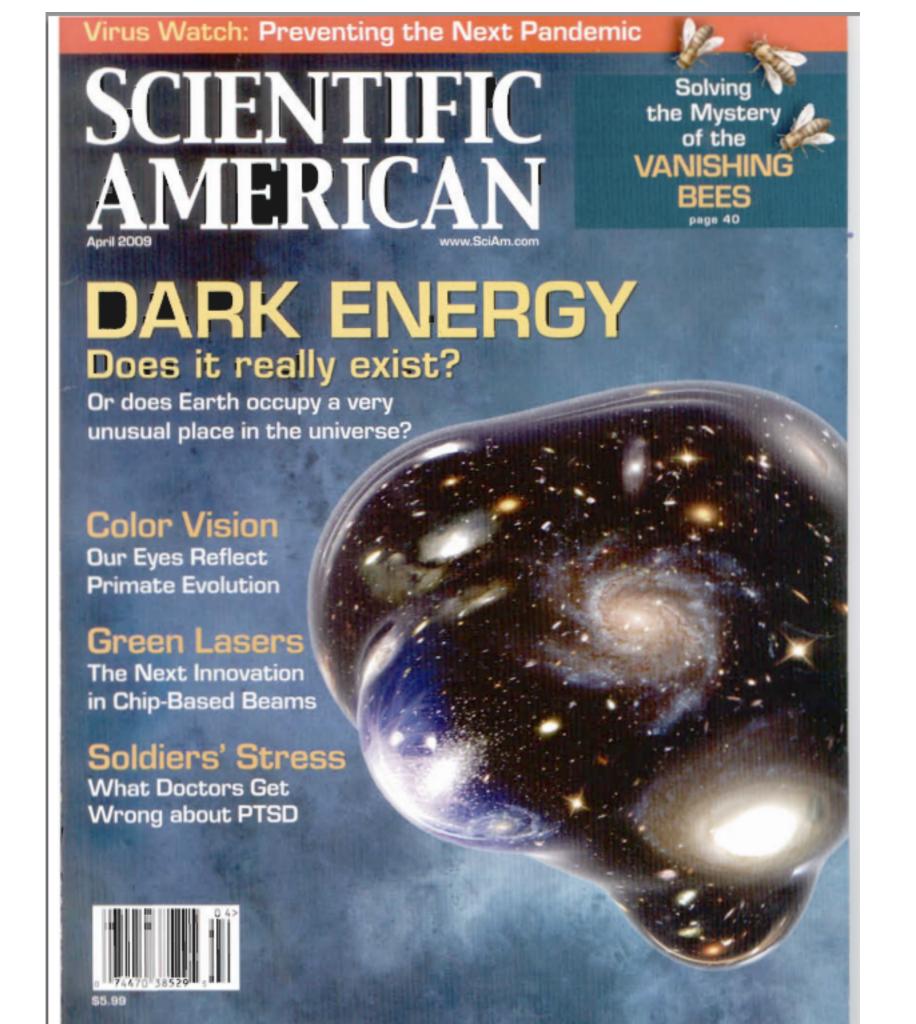
- The p = 0 evolution starting from this initial data evolves to  $H H_0$ , Q = .425 at  $t = t_0$ , in agreement with Dark Energy at  $t = t_{DE}$ .
- The cubic correction is C = 0.359 at  $t = t_0$ , while Dark Energy is C = -0.180 at  $t = t_{DE}$ .
- The ages of the universe are related by:

$$t_0 \approx (.95)t_{DE} \approx 1.38 \times t_{SM} = 1.38 \times (9.8 \times 10^9 yr)$$

### Around 2007:

Other research groups began exploring the possibility that the anomalous acceleration might be due to the earth lying near the center of a large region of Under-Density

We first saw publication in 2009



# This proposal is still taken seriously in Astrophysics

### Prokopek...2013 (Astrophysicist, Utrecht University)

Some of the more important discrepancies are as follows:

- the ΛCDM model predicts more galactic satellites (dwarf galaxies) than what has been observed [11] (this can be in part cured by a large merger rate, see however Ref. [12]);
- the Gaussian model for the origin of Universe's structure has difficulties in explaining the
  controversial large scale (dark) flow of galaxies [13] (even though the Planck satellite has not
  seen evidence of such flows in its data), and outliers such as the large relative speed in the
  Bullet Cluster collision [14];
- our Universe is supplied with a large number of voids, whose sizes and distribution may not
  be consistent with the ΛCDM model; moreover the voids should be filled with dwarfs and
  low surface brightness galaxies [15], which is not what has been observed [16];
- there are hints [17] that the structure growth rate is somewhat slower from that predicted by the  $\Lambda$ CDM model (alternatively we live in a universe with the equation of state parameter for dark energy  $w_{\text{de}} < -1$ );
- the disagreement between the Hubble Key Project and supernovae measurements of the Hubble constant [18, 19] and that obtained from the Planck data could be an indication that we live in an underdense region, whose size and magnitude would be difficult to reconcile with the standard ΛCDM with Gaussian initial perturbations (see however [20]).

# Details of our Analysis

### Main Steps:

- (I) Derivation of the p=0 Einstein equations in a new coordinate system aligned with the structure of the waves.
  - (2) A new ansatz for the Corrections to SM such that the asymptotic equations close.
  - (3) Putting the Initial Data from the Radiation Epoch into the gauge of our asymptotics.
- (4) The Redshift vs Luminosity determined by the Corrections.

# I. A New Formulationof the p=0Einstein Equations

The Einstein equations for spherically symmetric spacetimes take their Simplest Form Standard Schwarzschild Coordinates (SSC)

l.e.

$$ds^{2} = -D(t, \bar{r})d\bar{t}^{2} + E(\bar{t}, \bar{r})d\bar{t}d\bar{r} + F(\bar{t}, \bar{r})d\bar{r}^{2} + G(\bar{t}, \bar{r})d\Omega^{2}$$

$$ds^{2} = -D(t, \bar{r})d\bar{t}^{2} + E(\bar{t}, \bar{r})d\bar{t}d\bar{r} + F(\bar{t}, \bar{r})d\bar{r}^{2} + G(\bar{t}, \bar{r})d\Omega^{2}$$

### Transforms to SSC form:

$$ds^{2} = -D(t, \bar{r})d\bar{t}^{2} + E(\bar{t}, \bar{r})d\bar{t}d\bar{r} + F(\bar{t}, \bar{r})d\bar{r}^{2} + G(\bar{t}, \bar{r})d\Omega^{2}$$

### Transforms to SSC form:

$$(\bar{t},\bar{r}) \rightarrow (t,r)$$



$$ds^{2} = -D(t, \bar{r})d\bar{t}^{2} + E(\bar{t}, \bar{r})d\bar{t}d\bar{r} + F(\bar{t}, \bar{r})d\bar{r}^{2} + G(\bar{t}, \bar{r})d\Omega^{2}$$

### Transforms to SSC form:

$$(\bar{t},\bar{r}) \rightarrow (t,r)$$



$$ds^{2} = -B(t,r)dt^{2} + \frac{1}{A(t,r)}dr^{2} + r^{2}d\Omega^{2}$$

SSC

# The Equations In SSC

#### Standard Schwarzschild Coordinates

Four

$$\left\{ -r\frac{A_r}{A} + \frac{1-A}{A} \right\} = \frac{\kappa B}{A} r^2 T^{00} \tag{1}$$

$$\frac{A_t}{A} = \frac{\kappa B}{A} r T^{01} \tag{2}$$

$$\left\{r\frac{B_r}{B} - \frac{1-A}{A}\right\} = \frac{\kappa}{A^2}r^2T^{11} \tag{3}$$

$$-\left\{ \left(\frac{1}{A}\right)_{tt} - B_{rr} + \Phi \right\} = 2\frac{\kappa B}{A}r^2T^{22},\tag{4}$$

where

$$\Phi = \frac{B_t A_t}{2A^2 B} - \frac{1}{2A} \left(\frac{A_t}{A}\right)^2 - \frac{B_r}{r} - \frac{BA_r}{rA} + \frac{B}{2} \left(\frac{B_r}{B}\right)^2 - \frac{B}{2} \frac{B_r}{B} \frac{A_r}{A}.$$

$$(1)+(2)+(3)+(4)$$
 (weakly)  $(1)+(3)+div T=0$ 

### Theorem: (Te-Gr) The equations close in a "locally inertial" formulation of (1), (2) & Div T=0:

$$\begin{aligned}
\left\{T_{M}^{00}\right\}_{,0} + \left\{\sqrt{AB}T_{M}^{01}\right\}_{,1} &= -\frac{2}{r}\sqrt{AB}T_{M}^{01}, \\
\left\{T_{M}^{01}\right\}_{,0} + \left\{\sqrt{AB}T_{M}^{11}\right\}_{,1} &= -\frac{1}{2}\sqrt{AB}\left\{\frac{4}{r}T_{M}^{11} + \frac{(1-A)}{Ar}(T_{M}^{00} - T_{M}^{11})\right\} \\
&+ \frac{2\kappa r}{A}(T_{M}^{00}T_{M}^{11} - (T_{M}^{01})^{2}) - 4rT^{22}\right\}, \\
rA_{r} &= (1-A) - \kappa r^{2}T_{M}^{00}, \\
rB_{r} &= \frac{B(1-A)}{A} + \frac{B}{A}\kappa r^{2}T_{M}^{11}.
\end{aligned} \tag{3}$$

$$T_{M}^{00} = \frac{\rho c^{2} + p}{1 - \left(\frac{v}{c}\right)^{2}} \qquad T_{M}^{01} = \frac{\rho c^{2} + p}{1 - \left(\frac{v}{c}\right)^{2}} \frac{v}{c}$$

$$T_{M}^{11} = \frac{p + \left(\frac{v}{c}\right)^{2}}{1 - \left(\frac{v}{c}\right)^{2}} \rho c^{2} \qquad T^{22} = \frac{p}{r^{2}} \qquad v = \frac{1}{\sqrt{AB}} \frac{u^{1}}{u^{0}}$$

### Setting p=0:

$$T_{M}^{00} = \frac{\rho c^{2}}{1 - \left(\frac{v}{c}\right)^{2}} , \qquad T_{M}^{01} = \frac{\rho c^{2}}{1 - \left(\frac{v}{c}\right)^{2}} \frac{v}{c}$$

$$T_{M}^{11} = \frac{\rho c^{2}}{1 - \left(\frac{v}{c}\right)^{2}} \left(\frac{v}{c}\right)^{2}, \qquad T^{22} = 0$$

# Everything can be written in terms of $T_M^{00}$ and $\left(\frac{v}{c}\right)$ :

$$T_M^{01} = T_M^{00} \left(\frac{v}{c}\right), \qquad T_M^{22} = T_M^{00} \left(\frac{v}{c}\right)^2$$

## Substituting into the Equations gives:

$$\begin{split} \left(T_{M}^{00}\right)_{t} + \left\{\sqrt{AB}\left(\frac{v}{c}\right)\left(T_{M}^{00}\right)\right\}_{r} &= -\frac{2\sqrt{AB}}{r}\left(\frac{v}{c}\right)\left(T_{M}^{00}\right) \\ \left(\left(\frac{v}{c}\right)T_{M}^{00}\right)_{t} + \left\{\sqrt{AB}\left(\frac{v}{c}\right)^{2}T_{M}^{00}\right\}_{r} &= \\ &-\frac{\sqrt{AB}}{2r}\left\{4\left(\frac{v}{c}\right)^{2} + \frac{1-A}{A}\left(1-\left(\frac{v}{c}\right)^{2}\right)\right\}T_{M}^{00} \\ \frac{A'}{A} &= \frac{1}{r}\left(\frac{1}{A}-1\right) - \frac{\kappa r}{A}T_{M}^{00} \\ \frac{B'}{B} &= \frac{1}{r}\left(\frac{1}{A}-1\right) + \frac{\kappa r}{A}T_{M}^{00}\left(\frac{v}{c}\right)^{2} \end{split}$$

## Substituting into the Equations gives:

$$\begin{aligned}
\left(T_{M}^{00}\right)_{t} + \left\{\sqrt{AB} \left(\frac{\mathbf{v}}{c}\right) \left(T_{M}^{00}\right)\right\}_{r} &= -\frac{2\sqrt{AB}}{r} \left(\frac{\mathbf{v}}{c}\right) \left(T_{M}^{00}\right) \\
\left(\left(\frac{\mathbf{v}}{c}\right) T_{M}^{00}\right)_{t} + \left\{\sqrt{AB} \left(\frac{\mathbf{v}}{c}\right)^{2} T_{M}^{00}\right\}_{r} &= \\
&- \frac{\sqrt{AB}}{2r} \left\{4 \left(\frac{\mathbf{v}}{c}\right)^{2} + \frac{1-A}{A} \left(1 - \left(\frac{\mathbf{v}}{c}\right)^{2}\right)\right\} T_{M}^{00} \\
\frac{A'}{A} &= \frac{1}{r} \left(\frac{1}{A} - 1\right) - \frac{\kappa r}{A} T_{M}^{00} \\
\frac{B'}{B} &= \frac{1}{r} \left(\frac{1}{A} - 1\right) + \frac{\kappa r}{A} T_{M}^{00} \left(\frac{\mathbf{v}}{c}\right)^{2}
\end{aligned}$$

## Everything in terms of $T_M^{00}$ and $\binom{v}{c}$

## Substituting into the Equations gives:

$$(T_M^{00})_t + \left\{\sqrt{AB} \left(\frac{v}{c}\right) \left(T_M^{00}\right)\right\}_r = -\frac{2\sqrt{AB}}{r} \left(\frac{v}{c}\right) \left(T_M^{00}\right)$$

$$\left(\left(\frac{v}{c}\right) T_M^{00}\right)_t + \left\{\sqrt{AB} \left(\frac{v}{c}\right)^2 T_M^{00}\right\}_r =$$

$$-\frac{\sqrt{AB}}{2r} \left\{4 \left(\frac{v}{c}\right)^2 + \frac{1-A}{A} \left(1 - \left(\frac{v}{c}\right)^2\right)\right\} T_M^{00}$$

$$\frac{A'}{A} = \frac{1}{r} \left(\frac{1}{A} - 1\right) - \frac{\kappa r}{A} T_M^{00}$$

$$(1)$$

$$\frac{B'}{B} = \frac{1}{r} \left( \frac{1}{A} - 1 \right) + \frac{\kappa r}{A} T_M^{00} \left( \frac{v}{c} \right)^2$$

Note: Equations are Singular at r = 0

The 1/r singularity reflects the fact that waves coming into r=0 can amplify and blowup.

Since we are only interested in solutions representing outgoing, expanding waves, we look for natural changes of variables that regularize the equations at r = 0.

First: set c=1, collect v/r, and assume v/r smooth at r=0:

$$(T_M^{00})_t + r \left\{ \sqrt{AB} \left( \frac{v}{r} \right) T_M^{00} \right\}_r = 3\sqrt{AB} \left( \frac{v}{r} \right) T_M^{00}$$

$$\left(\frac{v}{r}\right)_t + r\sqrt{AB}\left(\frac{v}{r}\right)\left(\frac{v}{r}\right)_r = -\sqrt{AB}\left\{\left(\frac{v}{r}\right)^2 + \frac{1-A}{2Ar^2}\left(1 - r^2\left(\frac{v}{r}\right)^2\right)\right\}$$

$$\frac{A'}{A} = \frac{1}{r} \left( \frac{1}{A} - 1 \right) - \frac{\kappa r}{A} T_M^{00}$$

$$\frac{B'}{B} = \frac{1}{r} \left( \frac{1}{A} - 1 \right) + \frac{\kappa r}{A} T_M^{00} \left( \frac{v}{c} \right)^2$$

## Next: use (I) to eliminate $T_M^{00}$ from (2)

$$(T_M^{00})_t + \left\{\sqrt{AB} \left(\frac{v}{c}\right) \left(T_M^{00}\right)\right\}_r = -\frac{2\sqrt{AB}}{r} \left(\frac{v}{c}\right) \left(T_M^{00}\right) \qquad (\mathbf{I})$$

$$(\left(\frac{v}{c}\right) T_M^{00}\right)_t + \left\{\sqrt{AB} \left(\frac{v}{c}\right)^2 T_M^{00}\right\}_r =$$

$$-\frac{\sqrt{AB}}{2r} \left\{4 \left(\frac{v}{c}\right)^2 + \frac{1-A}{A} \left(1 - \left(\frac{v}{c}\right)^2\right)\right\} T_M^{00} \qquad (\mathbf{2})$$

$$\frac{A'}{A} = \frac{1}{r} \left(\frac{1}{A} - 1\right) - \frac{\kappa r}{A} T_M^{00}$$

$$\frac{B'}{B} = \frac{1}{r} \left( \frac{1}{A} - 1 \right) + \frac{\kappa r}{A} T_M^{00} \left( \frac{v}{c} \right)^2$$

l.e.

$$\frac{\left(\left(\frac{v}{c}\right)T_{M}^{00}\right)_{t} + \left\{\sqrt{AB}\left(\frac{v}{c}\right)^{2}T_{M}^{00}\right\}_{r}}{-\frac{\sqrt{AB}}{2r}\left\{4\left(\frac{v}{c}\right)^{2} + \frac{1-A}{A}\left(1 - \left(\frac{v}{c}\right)^{2}\right)\right\}T_{M}^{00}}$$
(2)

$$LHS = r\left(\frac{v}{r}\right) \left[ \left(T_{M}^{00}\right)_{t} + \left\{\sqrt{AB}r\left(\frac{v}{r}\right)T_{M}^{00}\right\}_{r} \right] + rT_{M}^{00}\left(\frac{v}{r}\right)_{t} + rT_{M}^{00}\sqrt{AB}\left(\frac{v}{r}\right)\left(r\left(\frac{v}{r}\right)\right)_{r}$$

## Substitute (1) into (2):

#### Obtain:

$$-2\sqrt{AB} \left(\frac{v}{r}\right)^{2} r T_{M}^{00} + r T_{M}^{00} \left(\frac{v}{r}\right)_{t}$$

$$+r T_{M}^{00} \sqrt{AB} \left(\frac{v}{r}\right) \left(r \left(\frac{v}{r}\right)\right)_{r}$$

$$(2)$$

$$= -\frac{\sqrt{AB}}{2r} \left\{ 4 \left( \frac{v}{c} \right)^2 + \frac{1-A}{A} \left( 1 - \left( \frac{v}{c} \right)^2 \right) \right\} T_M^{00}$$

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$$+r T_{M}^{00} \sqrt{AB} \left(\frac{v}{r}\right) \left(r \left(\frac{v}{r}\right)\right)_{r}$$

$$(2)$$

$$= -\frac{\sqrt{AB}}{2r} \left\{ 4 \left( \frac{v}{c} \right)^2 + \frac{1-A}{A} \left( 1 - \left( \frac{v}{c} \right)^2 \right) \right\} T_M^{00}$$





Linearity in  $T_M^{00}$  Divide by  $rT_M^{00}$ 

## Next: simplify and collect: $z = \kappa T_M^{00} r^2$

$$(\kappa T_M^{00} r^2)_t + \left\{ \sqrt{AB} \; \frac{v}{r} \left( \kappa T_M^{00} r^2 \right) \right\}_r = -2\sqrt{AB} \frac{v}{r} \left( \kappa T_M^{00} r^2 \right)$$

$$\left(\frac{v}{r}\right)_t + r\sqrt{AB}\left(\frac{v}{r}\right)\left(\frac{v}{r}\right)_r = -\sqrt{AB}\left\{\left(\frac{v}{r}\right)^2 + \frac{1-A}{2Ar^2}\left(1 - r^2\left(\frac{v}{r}\right)^2\right)\right\}$$

$$r\frac{A'}{A} = \left(\frac{1}{A} - 1\right) - \frac{1}{A}\kappa T_M^{00} r^2$$

$$r\frac{B'}{B} = \left(\frac{1}{A} - 1\right) + \frac{1}{A} \left(\frac{v}{c}\right)^2 \kappa T_M^{00} r^2$$

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$$\left(\frac{v}{r}\right)_t + r\sqrt{AB}\left(\frac{v}{r}\right)\left(\frac{v}{r}\right)_r = -\sqrt{AB}\left\{\left(\frac{v}{r}\right)^2 + \frac{1-A}{2Ar^2}\left(1 - r^2\left(\frac{v}{r}\right)^2\right)\right\}$$

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$$r\frac{B'}{B} = \left(\frac{1}{A} - 1\right) + \frac{1}{A} \left(\frac{v}{c}\right)^2 \kappa T_M^{00} r^2$$

(This is the self-similar variable in the waves from the radiation epoch!)

$$(t,r) \rightarrow (t,\xi)$$

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$$\xi = \frac{t}{t}$$

$$\left(T_M^{00},v\right) \to (z,w)$$

$$z = \kappa T_M^{00} r^2, \quad w = \frac{v}{\xi}$$

$$(t,r) \rightarrow (t,\xi)$$
  $\xi = \frac{1}{t}$ 

$$(T_M^{00}, v) \to (z, w)$$

$$z = \kappa T_M^{00} r^2, \quad w = \frac{v}{\xi}$$

$$\frac{\partial}{\partial r} = \frac{1}{t} \frac{\partial}{\partial r}, \qquad \frac{\partial}{\partial r} f(t, r) = \left( \frac{\partial}{\partial t} - \frac{1}{t^2} \frac{\partial}{\partial \xi} \right) f(t, \xi)$$

## Substituting into (I) and (2) we obtain the following dimensionless eqns:

$$tz_t + \xi \{(-1 + Dw)z\}_{\xi} = -Dwz,$$
 (1)

$$tw_t + \xi (-1 + Dw) w_{\xi} =$$

$$w - D \left\{ w^2 + \frac{1 - \xi^2 w^2}{2A} \left[ \frac{1 - A}{\xi^2} \right] \right\}, \qquad (2)$$

Where:  $D = \sqrt{AB}$ 

#### Take A and D instead of A and B:

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$$\xi A_{\xi} = (A - 1) - z,$$

$$\xi \frac{B_{\xi}}{B} = \frac{1}{A} \left\{ 1 - A + \xi^{2} w^{2} z \right\},$$

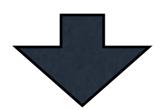
$$\xi(\sqrt{AB})_{\xi} = \sqrt{AB} \left\{ (1 - A) - \frac{(1 - \xi^{2} w^{2})}{2} z \right\}.$$

#### Take A and D instead of A and B:

$$\xi A_{\xi} = (A - 1) - z,$$

$$\xi \frac{B_{\xi}}{B} = \frac{1}{A} \left\{ 1 - A + \xi^{2} w^{2} z \right\},$$

$$\xi(\sqrt{AB})_{\xi} = \sqrt{AB} \left\{ (1 - A) - \frac{(1 - \xi^{2} w^{2})}{2} z \right\}.$$



$$\xi A_{\xi} = (A - 1) - z$$

$$\xi (D)_{\xi} = D \left\{ (1 - A) - \frac{(1 - \xi^2 w^2)}{2} z \right\}$$

# This leads to the following Dimensionless Formulation of the p=0 Einstein Equations:

## Einstein Equations when p=0

$$tz_t + \xi \{(-1 + Dw)z\}_{\xi} = -Dwz,$$

$$tw_t + \xi \left(-1 + Dw\right) w_{\xi} =$$

$$w - D \left\{ w^2 + \frac{1 - \xi^2 w^2}{2A} \left[ \frac{1 - A}{\xi^2} \right] \right\},$$

$$\xi A_{\xi} = (A-1) - z,$$

$$\frac{\xi D_{\xi}}{D} = (1 - A) - \frac{(1 - \xi^2 w^2)}{2} z.$$

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$$\xi A_{\xi} = (A-1) - z,$$

$$\frac{\xi D_{\xi}}{D} = (1 - A) - \frac{(1 - \xi^2 w^2)}{2} z.$$

$$ds^{2} = -Bdt^{2} + \frac{1}{A}dr^{2} + r^{2}d\Omega^{2}, \quad D = \sqrt{AB}, \quad z = \frac{\rho r^{2}}{(1 - v^{2})}, \quad w = \frac{v}{\xi}$$

## 2. The Ansatz and Asymptotics for the Corrections:

$$z(t,\xi) = z_F(\xi) + \Delta z(t,\xi)$$

$$w(t,\xi) = w_F(\xi) + \Delta w(t,\xi)$$

$$A(t,\xi) = A_F(\xi) + \Delta A(t,\xi)$$

$$D(t,\xi) = D_F(\xi) + \Delta D(t,\xi)$$

$$z(t,\xi) = z_F(\xi) + \Delta z(t,\xi)$$

$$w(t,\xi) = w_F(\xi) + \Delta w(t,\xi)$$

$$A(t,\xi) = A_F(\xi) + \Delta A(t,\xi)$$

$$D(t,\xi) = D_F(\xi) + \Delta D(t,\xi)$$

#### The Standard Model is Self-Similar:

$$z_F = \frac{4}{3}\xi^2 + \frac{40}{27}\xi^4 + O(\xi^6)$$

$$w_F = \frac{2}{3} + \frac{2}{9}\xi^2 + O(\xi^4)$$

$$A_F = 1 - \frac{4}{9}\xi^2 - \frac{8}{27}\xi^4 + O(\xi^6)$$

$$D_F = 1 - \frac{1}{9}\xi^2 + O(\xi^4)$$

$$z(t,\xi) = z_F(\xi) + \Delta z(t,\xi)$$

$$w(t,\xi) = w_F(\xi) + \Delta w(t,\xi)$$

$$A(t,\xi) = A_F(\xi) + \Delta A(t,\xi)$$

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$$A_F = 1 - \frac{4}{9}\xi^2 - \frac{8}{27}\xi^4 + O(\xi^6)$$

$$D_F = 1 - \frac{1}{9}\xi^2 + O(\xi^4)$$

$$z(t,\xi) = z_{F}(\xi) + \Delta z(t,\xi) \qquad \Delta z = z_{2}(t)\xi^{2} + z_{4}(t)\xi^{4}$$

$$w(t,\xi) = w_{F}(\xi) + \Delta w(t,\xi) \qquad \Delta w = w_{0}(t) + w_{2}(t)\xi^{2}$$

$$A(t,\xi) = A_{F}(\xi) + \Delta A(t,\xi) \qquad \Delta A = A_{2}(t)\xi^{2} + A_{4}(t)\xi^{4}$$

$$D(t,\xi) = D_{F}(\xi) + \Delta D(t,\xi) \qquad \Delta D = D_{2}(t)\xi^{2}$$

- ullet Note: Corrections only involve even powers of  $\xi$
- The Standard Model is Self-Similar:

$$z_F = \frac{4}{3}\xi^2 + \frac{40}{27}\xi^4 + O(\xi^6)$$

$$w_F = \frac{2}{3} + \frac{2}{9}\xi^2 + O(\xi^4)$$

$$A_F = 1 - \frac{4}{9}\xi^2 - \frac{8}{27}\xi^4 + O(\xi^6)$$

$$D_F = 1 - \frac{1}{9}\xi^2 + O(\xi^4)$$

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + \left\{\frac{2}{9} + w_2(t)\right\}\xi^2 + O(\xi^4),$$

#### • Reiterate:

# We don't use co-moving coordinates, but rather write the SSC eqns in $(t,\xi)$ -coordinates.

$$ds^{2} = -B(t,r)dt^{2} + \frac{1}{A(t,r)}dr^{2} + r^{2}d\Omega^{2}$$

$$\xi = r/t$$
  $D = \sqrt{AB}$ 

## Equations for the Corrections to SM

 When we plug into the equations a remarkable simplification occurs:

$$A_2 = -\frac{1}{3}z_2, \quad A_4 = -\frac{1}{5}z_4, \quad D_2 = -\frac{1}{12}z_2$$

## Equations for the Corrections to SM

 When we plug into the equations a remarkable simplification occurs:

$$A_2 = -\frac{1}{3}z_2, \quad A_4 = -\frac{1}{5}z_4, \quad D_2 = -\frac{1}{12}z_2$$

• This is a coordinate gauge condition reflecting the serendipity of our  $(t,\xi)$ -coordinate system!!

### Plugging Ansatz into Equations...

## Plugging

$$A_2 = -\frac{1}{3}z_2, \quad A_4 = -\frac{1}{5}z_4, \quad D_2 = -\frac{1}{12}z_2$$

and

$$z(t,\xi) = z_F(\xi) + z_2(t)\xi^2 + z_4(t)\xi^4$$

$$w(t,\xi) = w_F(\xi) + w_0(t) + w_2(t)\xi^2$$

$$A(t,\xi) = A_F(\xi) + A_2(t)\xi^2 + A_4(t)\xi^4$$

$$D(t,\xi) = D_F(\xi) + D_2(\xi)\xi^2$$

#### into equations:

$$\begin{aligned} tz_t + \xi \left\{ (-1 + Dw)z \right\}_{\xi} &= -Dwz \\ tw_t + \xi \left( -1 + Dw \right) w_{\xi} &= \\ w - D \left\{ w^2 + \frac{1 - \xi^2 w^2}{2A} \left[ \frac{1 - A}{\xi^2} \right] \right\} \end{aligned}$$

# Gives:

**THEOREM:** The p = 0 waves take the asymptotic form

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + \left\{\frac{2}{9} + w_2(t)\right\}\xi^2 + O(\xi^4),$$

where  $z_2(t), z_4(t), w_0(t), w_2(t)$  evolve according to the equations

$$\begin{aligned}
-t\dot{z}_2 &= 3w_0 \left(\frac{4}{3} + z_2\right), \\
-t\dot{z}_4 &= -5\left\{\frac{2}{27}z_2 + \frac{4}{3}w_2 - \frac{1}{18}z_2^2 + z_2w_2\right\} \\
&-5w_0 \left\{\frac{4}{3} - \frac{2}{9}z_2 + z_4 - \frac{1}{12}z_2^2\right\}, \\
-t\dot{w}_0 &= \frac{1}{6}z_2 + \frac{1}{3}w_0 + w_0^2, \\
-t\dot{w}_2 &= \frac{1}{10}z_4 + \frac{4}{9}w_0 - \frac{1}{3}w_2 + \frac{1}{24}z_2^2 - \frac{1}{3}z_2w_0 \\
&-\frac{1}{3}w_0^2 + 4w_0w_2 - \frac{1}{4}w_0^2z_2.
\end{aligned}$$

#### The Corrections to SM evolve according to

$$-t\dot{z}_{2} = 3w_{0}\left(\frac{4}{3} + z_{2}\right),$$

$$-t\dot{z}_{4} = -5\left\{\frac{2}{27}z_{2} + \frac{4}{3}w_{2} - \frac{1}{18}z_{2}^{2} + z_{2}w_{2}\right\}$$

$$-5w_{0}\left\{\frac{4}{3} - \frac{2}{9}z_{2} + z_{4} - \frac{1}{12}z_{2}^{2}\right\},$$

$$-t\dot{w}_{0} = \frac{1}{6}z_{2} + \frac{1}{3}w_{0} + w_{0}^{2},$$

$$-t\dot{w}_{2} = \frac{1}{10}z_{4} + \frac{4}{9}w_{0} - \frac{1}{3}w_{2} + \frac{1}{24}z_{2}^{2} - \frac{1}{3}z_{2}w_{0}$$

$$-\frac{1}{3}w_{0}^{2} + 4w_{0}w_{2} - \frac{1}{4}w_{0}^{2}z_{2},$$

### Note: RHS is Autonomous!

### Ve can make LHS Automomous too!

$$-z_{2}' = -t\dot{z}_{2} = 3w_{0} \left(\frac{4}{3} + z_{2}\right),$$

$$-z_{4}' = -t\dot{z}_{4} = -5\left\{\frac{2}{27}z_{2} + \frac{4}{3}w_{2} - \frac{1}{18}z_{2}^{2} + z_{2}w_{2}\right\}$$

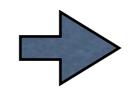
$$-5w_{0} \left\{\frac{4}{3} - \frac{2}{9}z_{2} + z_{4} - \frac{1}{12}z_{2}^{2}\right\},$$

$$-w_{0}' = -t\dot{w}_{0} = \frac{1}{6}z_{2} + \frac{1}{3}w_{0} + w_{0}^{2},$$

$$-w_{2}' = -t\dot{w}_{2} = \frac{1}{10}z_{4} + \frac{4}{9}w_{0} - \frac{1}{3}w_{2} + \frac{1}{24}z_{2}^{2} - \frac{1}{3}z_{2}w_{0}$$

$$-\frac{1}{3}w_{0}^{2} + 4w_{0}w_{2} - \frac{1}{4}w_{0}^{2}z_{2}.$$

$$au = ln(t) \Rightarrow t \frac{d}{dt} = \frac{d}{d\tau} \equiv '$$
 LHS Autonomous



#### Autonomous Eqns for Corrections to SM

$$-z_{2}' = 3w_{0} \left(\frac{4}{3} + z_{2}\right),$$

$$-z_{4}' = -5\left\{\frac{2}{27}z_{2} + \frac{4}{3}w_{2} - \frac{1}{18}z_{2}^{2} + z_{2}w_{2}\right\}$$

$$-5w_{0} \left\{\frac{4}{3} - \frac{2}{9}z_{2} + z_{4} - \frac{1}{12}z_{2}^{2}\right\},$$

$$-w_{0}' = \frac{1}{6}z_{2} + \frac{1}{3}w_{0} + w_{0}^{2},$$

$$-w_{2}' = \frac{1}{10}z_{4} + \frac{4}{9}w_{0} - \frac{1}{3}w_{2} + \frac{1}{24}z_{2}^{2} - \frac{1}{3}z_{2}w_{0}$$

$$-\frac{1}{3}w_{0}^{2} + 4w_{0}w_{2} - \frac{1}{4}w_{0}^{2}z_{2}.$$

$$t_* \le t \le 10^{14} yr$$

$$\ln(t_*) \le \tau \le 14 \cdot \ln(10)$$

Trivializes the large time simulation problem!

### The Equations for the Corrections

$$-z_{2}' = 3w_{0} \left(\frac{4}{3} + z_{2}\right),$$

$$-z_{4}' = -5\left\{\frac{2}{27}z_{2} + \frac{4}{3}w_{2} - \frac{1}{18}z_{2}^{2} + z_{2}w_{2}\right\}$$

$$-5w_{0} \left\{\frac{4}{3} - \frac{2}{9}z_{2} + z_{4} - \frac{1}{12}z_{2}^{2}\right\},$$

$$-w_{0}' = \frac{1}{6}z_{2} + \frac{1}{3}w_{0} + w_{0}^{2},$$

$$-w_{2}' = \frac{1}{10}z_{4} + \frac{4}{9}w_{0} - \frac{1}{3}w_{2} + \frac{1}{24}z_{2}^{2} - \frac{1}{3}z_{2}w_{0}$$

$$-\frac{1}{3}w_{0}^{2} + 4w_{0}w_{2} - \frac{1}{4}w_{0}^{2}z_{2}.$$

## Everything is dimensionless involving only pure numbers!

### The Equations for the Corrections

Leading order 
$$(z_2, w_0) = -5\left\{\frac{2}{27}z_2 + \frac{4}{3}w_2 - \frac{1}{18}z_2^2 + z_2w_2\right\}$$

$$-5w_0\left\{\frac{4}{3} - \frac{2}{9}z_2 + z_4 - \frac{1}{12}z_2^2\right\},$$

$$-w_0' = \frac{1}{6}z_2 + \frac{1}{3}w_0 + w_0^2,$$

$$-w_2' = \frac{1}{10}z_4 + \frac{4}{9}w_0 - \frac{1}{3}w_2 + \frac{1}{24}z_2^2 - \frac{1}{3}z_2w_0$$

$$-\frac{1}{3}w_0^2 + 4w_0w_2 - \frac{1}{4}w_0^2z_2.$$

Note: Leading order Eqns Uncouple!

### The Leading Order Corrections...

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + O(\xi^4),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + O(\xi^2),$$

### ...And Their Equations

$$-z_2' = -t\dot{z}_2 = 3w_0 \left(\frac{4}{3} + z_2\right),$$

$$-w_0' = -t\dot{w}_0 = \frac{1}{6}z_2 + \frac{1}{3}w_0 + w_0^2.$$

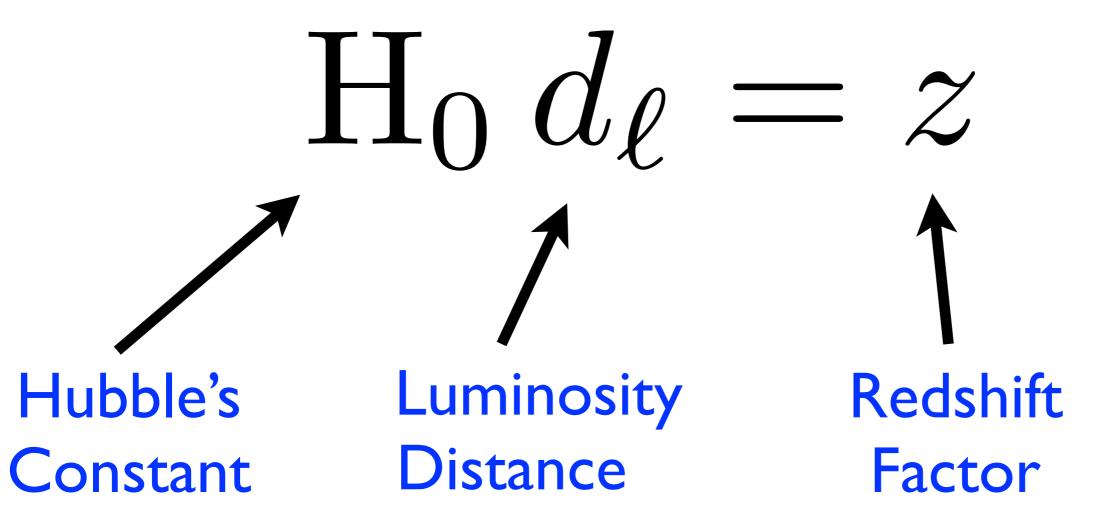
• Keep in mind that  $\xi$  is on the order of fractional distance to the Hubble Length:

$$\xi = r/ct \approx \frac{\text{arclength distance at fixed time}}{\text{distance of light travel since Big Bang}}$$

 For example: At I/I0 way across the visible universe, about I.I billion light-years out:

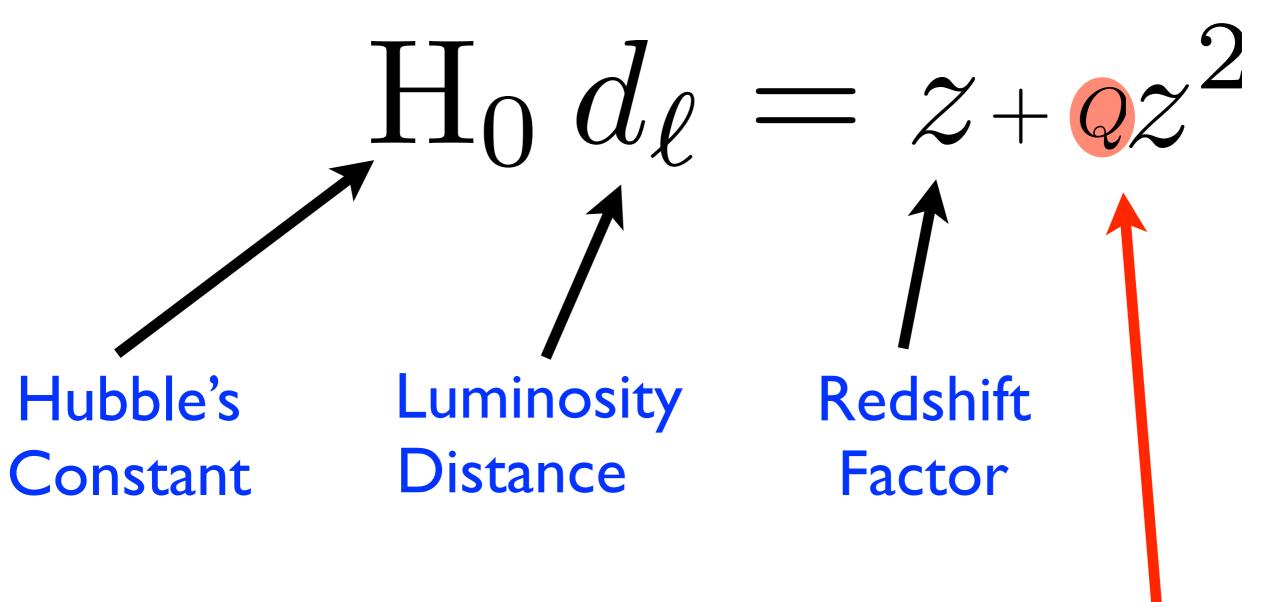
$$\xi^4 \approx \frac{1}{10,000} = .0001$$

### Hubbles Law:



1929: Linear relation between redshift and luminosity

#### Hubbles Law:



1999: There is an anomalous acceleration

• In Fact:  $\xi$  is on the order of the redshift factor, and  $(z_2, w_0)$  determines the quadratic correction to redshift vs luminosity = anomalous acceleration

$$H_0 d_{\ell} = z + \underbrace{Q(z_2, w_0)}_{\uparrow} z^2 + O(z^3)$$

This term accounts for the corrections to the Standard Model Observed in the Supernova Data (Nobel Prize)

• In Fact:  $\xi$  is on the order of the redshift factor, and  $(z_2, w_0)$  determines the quadratic correction to redshift vs luminosity = anomalous acceleration

$$H_0 d_\ell = z + \underbrace{Q(z_2, w_0)}_{\uparrow} z^2 + O(z^3)$$

Determined by the value of the so-called "Deceleration Parameter" q

# • The cubic correction is determined by $(z_2, w_0, w_2)$

$$H_0 d_\ell = z + Q(z_2, w_0) z^2 + C(z_2, w_0, w_2) z^3 + O(z^3)$$

Determined by solving our system of four equations for  $(z_2, z_4, w_0, w_4)$ 

# • The cubic correction is determined by $(z_2, w_0, w_2)$

$$H_0 d_\ell = z + Q(z_2, w_0) z^2 + C(z_2, w_0, w_2) z^3 + O(z^3)$$
A prediction

Beyond experimental precision

### ullet The quadratic correction is determined by our equations for $(z_2,w_0)$

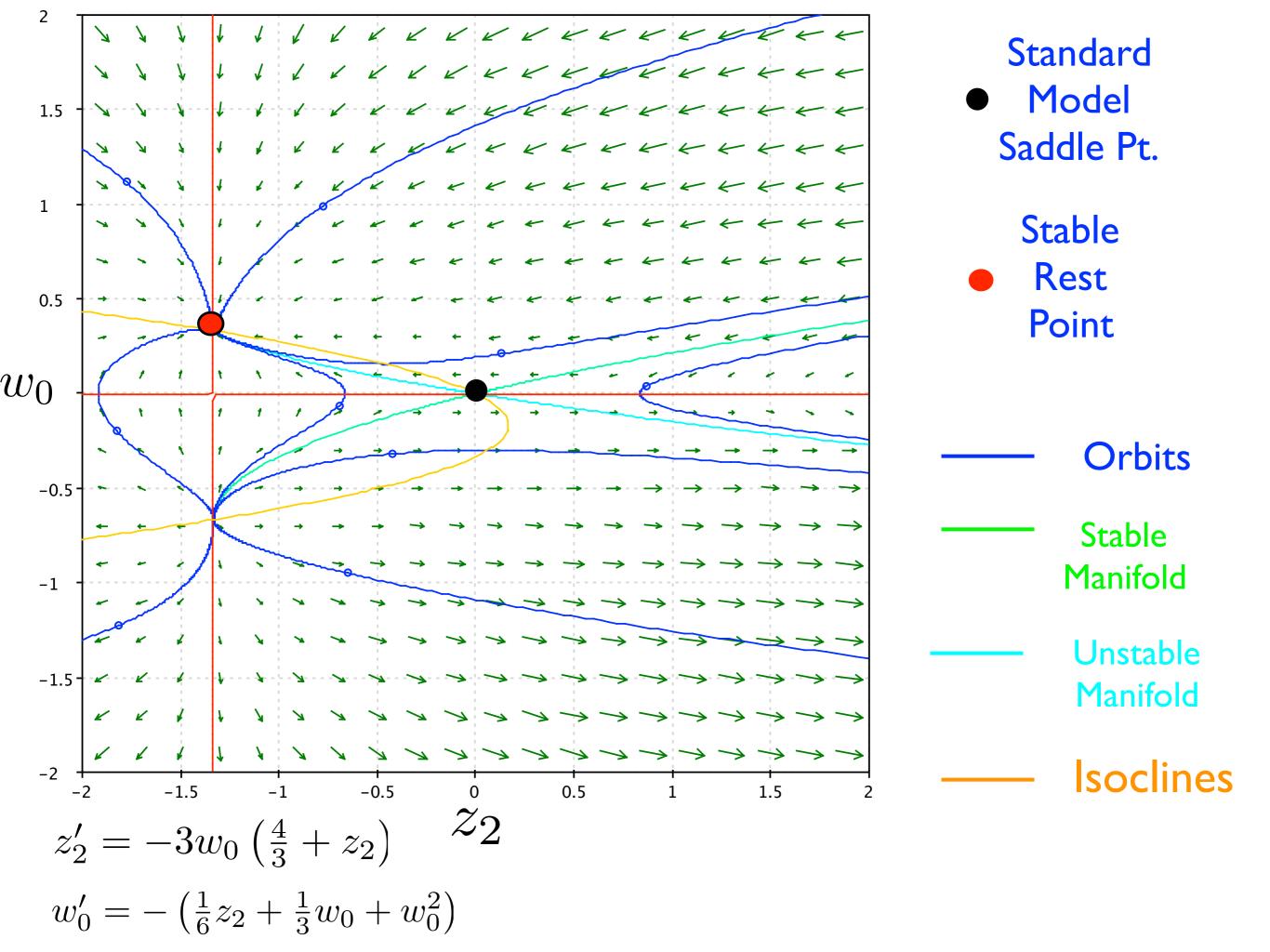
$$H_0 d_{\ell} = z + Q(z_2, w_0) z^2 + O(z^3)$$

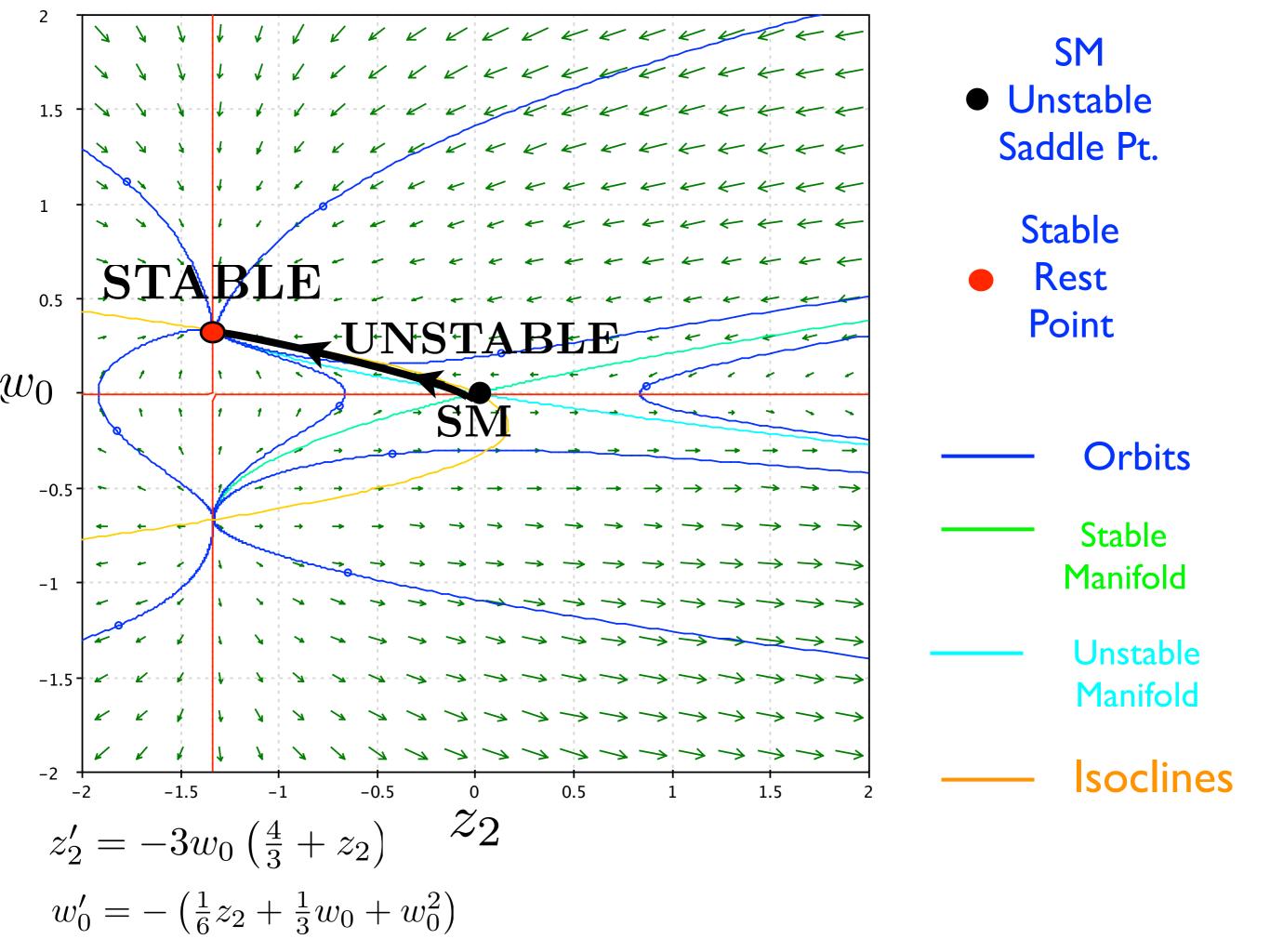
$$-z_2' = -t\dot{z}_2 = 3w_0 \left(\frac{4}{3} + z_2\right),$$
  
$$-w_0' = -t\dot{w}_0 = \frac{1}{6}z_2 + \frac{1}{3}w_0 + w_0^2.$$

### Numerical Simulation

The  $(z_2, w_0)$  phase portrait:

Thanks to: pplane Rice University





3. The Initial Data determined by the Self-Similar Waves from the Radiation Epoch

$$p = \sigma^2 \rho$$
:

$$p = \sigma^2 \rho$$
:

FRW Co-moving:

$$ds^{2} = -dt^{2} + R(t)^{2} \left\{ dr^{2} + r^{2} d\Omega^{2} \right\}$$

$$p = \sigma^2 \rho$$
:

$$ds^{2} = -dt^{2} + R(t)^{2} \left\{ dr^{2} + r^{2} d\Omega^{2} \right\}$$

FRW Self-Similar:

$$\bar{t} = F(\eta)t; \quad \bar{r} = \eta t,$$

$$p = \sigma^2 \rho$$
:

$$ds^{2} = -dt^{2} + R(t)^{2} \left\{ dr^{2} + r^{2} d\Omega^{2} \right\}$$

FRW Self-Similar:

$$\bar{t} = F(\eta)t; \quad \bar{r} = \eta t,$$

$$\xi \equiv \frac{\overline{r}}{\overline{t}} = \frac{\eta}{F(\eta)}; \quad \eta \equiv \frac{\overline{r}}{t}; \quad F(\eta) = \left(1 - \frac{1 - 3\sigma}{9(1 + \sigma)^2}\eta^2\right)^{\frac{3(1 + \sigma)}{2(1 + 3\sigma)}}$$

$$p = \sigma^2 \rho$$
:

FRW Co-moving:

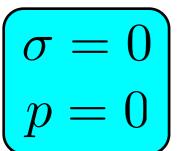
$$ds^{2} = -dt^{2} + R(t)^{2} \left\{ dr^{2} + r^{2} d\Omega^{2} \right\}$$

FRW Self-Similar:

$$\bar{t} = F(\eta)t; \quad \bar{r} = \eta t,$$

$$ds^{2} = -\frac{F(\eta)^{-\frac{1+3\sigma}{3(1+\sigma)}}}{1 - \left(\frac{2}{3(1+\sigma)\eta^{2}}\right)^{2}} d\bar{t}^{2} + \frac{1}{1 - \left(\frac{2}{3(1+\sigma)\eta^{2}}\right)^{2}} d\bar{r}^{2} + \bar{r}^{2} d\Omega^{2}$$

$$\xi \equiv \frac{\bar{r}}{\bar{t}} = \frac{\eta}{F(\eta)}; \quad \eta \equiv \frac{\bar{r}}{t}; \quad F(\eta) = \left(1 - \frac{1 - 3\sigma}{9(1 + \sigma)^2}\eta^2\right)^{\frac{3(1 + \sigma)}{2(1 + 3\sigma)}}$$



$$\sigma = 0$$

$$p = 0$$

$$\begin{bmatrix} \sigma = 0 \\ p = 0 \end{bmatrix} ds^2 = -B_F(\xi)d\bar{t}^2 + \frac{1}{A_F(\xi)}d\bar{r}^2 + \bar{r}^2d\Omega^2$$

$$\sigma = 0$$
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$$\begin{vmatrix} \sigma = 0 \\ p = 0 \end{vmatrix} ds^2 = -B_F(\xi)d\bar{t}^2 + \frac{1}{A_F(\xi)}d\bar{r}^2 + \bar{r}^2d\Omega^2$$

$$A_F(\xi) = 1 - \frac{4}{9}\xi^2 - \frac{8}{27}\xi^4 + O(\xi^6)$$

$$D_F(\xi) \equiv \sqrt{A_F B_F} = 1 - \frac{1}{9}\xi^2 + O(\xi^4).$$

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Note:

$$\xi = \frac{\bar{r}}{\bar{t}} = \frac{\bar{r}}{ct} + O(\xi^2)$$

$$\begin{aligned}
\sigma &= 0 \\
p &= 0
\end{aligned}$$

$$\begin{vmatrix} \sigma = 0 \\ p = 0 \end{vmatrix} ds^2 = -B_F(\xi)d\bar{t}^2 + \frac{1}{A_F(\xi)}d\bar{r}^2 + \bar{r}^2d\Omega^2$$

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Note:

$$\xi = \frac{\overline{r}}{\overline{t}} = \frac{\overline{r}}{ct} + O(\xi^2)$$

Where: 
$$\frac{\bar{r}}{ct} \approx \frac{\text{arclength distance at fixed time}}{\text{distance of light travel since the Big Bang}}$$

...a measure of ``fractional distance to Hubble Radius"

• Conclude: when  $\xi << 1$ 

 $\xi \approx$  fractional distance to the Hubble Radius

in a non-uniform spacetime measures approximately how far out you would think you were if you believed you were at the center of a Friedmann spacetime...

$$\sigma = 0$$
$$p = 0$$

$$\begin{vmatrix} \sigma = 0 \\ p = 0 \end{vmatrix} ds^2 = -B_F(\xi)d\bar{t}^2 + \frac{1}{A_F(\xi)}d\bar{r}^2 + \bar{r}^2d\Omega^2$$

$$A_F(\xi) = 1 - \frac{4}{9}\xi^2 - \frac{8}{27}\xi^4 + O(\xi^6)$$
$$D_F(\xi) \equiv \sqrt{A_F B_F} = 1 - \frac{1}{9}\xi^2 + O(\xi^4).$$

$$z_F(\xi) = \frac{4}{3}\xi^2 + \frac{40}{27}\xi^4 + O(\xi^6)$$
$$w_F \equiv \frac{v}{\xi} = \frac{2}{3} + \frac{2}{9}\xi^2 + O(\xi^4)$$

The p=0 Friedmann Universe in Self-Similar Coordinates

### Thus our equations are for the corrections to the Standard Model:

$$z(t,\xi) = \left(\frac{4}{3} + z_2(t)\right)\xi^2 + \left\{\frac{40}{27} + z_4(t)\right\}\xi^4 + O(\xi^6),$$

$$w(t,\xi) = \left(\frac{2}{3} + w_0(t)\right) + \left\{\frac{2}{9} + w_2(t)\right\}\xi^2 + O(\xi^4)$$

$$(p=0)$$

$$p = \frac{c^2}{3}\rho$$

$$\sigma = \frac{1}{\sqrt{3}}$$

### Self-similar coordinates for Friedmann with

Pure Radiation

$$|\bar{\xi} \neq \xi|$$

$$z_{1/3} \equiv z_{1/3}^{1}(\bar{t},\bar{\xi}) = \frac{3}{4}\bar{\xi}^{2} + \frac{9}{16}\bar{\xi}^{4} + O(\bar{\xi}^{6}),$$

$$v_{1/3} \equiv v_{1/3}^{1}(\bar{t},\bar{\xi}) = \frac{1}{2}\bar{\xi} + \frac{1}{8}\bar{\xi}^{3} + O(\bar{\xi}^{5}),$$

$$A_{1/3} \equiv A_{1/3}^{1}(\bar{t},\bar{\xi}) = 1 - \frac{1}{4}\bar{\xi}^{2} - \frac{1}{8}\bar{\xi}^{4} + O(\bar{\xi}^{6}),$$

$$D_{1/3} \equiv D_{1/3}^{1}(\bar{t},\bar{\xi}) = 1 + O(\bar{\xi}^{4}).$$

The  $p = \frac{c^2}{3}\rho$  Friedmann Universe admits a I-parameter family of Self-Similar spacetimes that perturb the Standard Model during the Radiation Epoch:

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The p=0 Friedmann Universe DOES NOT admit Self-Similar perturbations!

The  $p = \frac{c^2}{3}\rho$  Friedmann Universe is embedded in I-parameter family of Self-Similar spacetimes that perturb the Standard Model during the Radiation Epoch:

The p=0 Friedmann Universe DOES NOT admit Self-Similar perturbations!

(The topic of our PNAS and MEMOIR)

First Discovered by Cahill and Taub: Commun Math Phys., 21, 1-40 (1971)

Extended by others, esp. Carr and Coley, Survey: Physical Review D, 62,044023-1-25 (1999)

Our interest is in the possible connection between these waves and the Anomalous Acceleration.

We extract properties of the waves from a system of ODE's we derived, that defines them:

#### The perturbations are describe by ODE's:

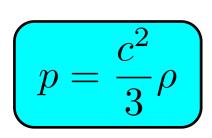
$$\xi A_{\xi} = -\left[\frac{4(1-A)v}{(3+v^2)G-4v}\right]$$

$$\xi G_{\xi} = -G\left\{\left(\frac{1-A}{A}\right)\frac{2(1+v^2)G-4v}{(3+v^2)G-4v}-1\right\}$$

$$\xi v_{\xi} = -\left(\frac{1-v^2}{2\{\cdot\}_D}\right)\left\{(3+v^2)G-4v+\frac{4\left(\frac{1-A}{A}\right)\{\cdot\}_N}{(3+v^2)G-4v}\right\}$$

$$\{\cdot\}_{N} = \{-2v^{2} + 2(3 - v^{2})vG - (3 - v^{4})G^{2}\}$$
  
$$\{\cdot\}_{D} = \{(3v^{2} - 1) - 4vG + (3 - v^{2})G^{2}\}$$

$$\left( G = rac{\xi}{\sqrt{AB}} \; \; ; \; \; \; \xi = rac{r}{t} 
ight)$$



## Self-Similar perturbations of Friedmann for Pure Radiation

(The topic of our PNAS and MEMOIR)

$$z_{1/3}^{a} = \frac{3a^{2}}{4}\bar{\xi}^{2} + \left[\frac{9a^{2}}{16} + 3a^{2}(V_{0} + A_{0})(1 - a^{2})\right]\bar{\xi}^{4} + O(\bar{\xi}^{6})$$

$$v_{1/3}^a = \frac{1}{2}\bar{\xi} + \left[\frac{1}{8} + V_0(1 - a^2)\right]\bar{\xi}^3 + O(\bar{\xi}^5)$$

$$A_{1/3}^{a} = 1 - \frac{a^{2}}{4}\bar{\xi}^{2} - \left[\frac{a^{2}}{8} + a^{2}A_{0}(1 - a^{2})\right]\bar{\xi}^{4} + O(\bar{\xi}^{6})$$

$$D_{1/3}^a = 1 + O(\bar{\xi}^4)$$

$$V_0 = \frac{2}{3}A_0 = \frac{1}{20}$$

# A 1-parameter family of solutions depending on the Acceleration Parameter $0 < a < \infty$

$$z_{1/3}^{a} = \frac{3a^{2}}{4}\bar{\xi}^{2} + \left[\frac{9a^{2}}{16} + 3a^{2}\left(V_{0} + A_{0}\right)\left(1 - a^{2}\right)\right]\bar{\xi}^{4} + O(\bar{\xi}^{6})$$

$$v_{1/3}^a = \frac{1}{2}\bar{\xi} + \left[\frac{1}{8} + V_0(1 - a^2)\right]\bar{\xi}^3 + O(\bar{\xi}^5)$$

$$A_{1/3}^a = 1 - \frac{a^2}{4}\bar{\xi}^2 - \left[\frac{a^2}{8} + a^2A_0(1 - a^2)\right]\bar{\xi}^4 + O(\bar{\xi}^6)$$

$$D_{1/3}^a = 1 + O(\bar{\xi}^4)$$

$$V_0 = \frac{2}{3}A_0 = \frac{1}{20}$$

#### a = 1

#### is the Standard Model for Pure Radiation

$$z_{1/3}^{a} = \frac{3a^{2}}{4}\bar{\xi}^{2} + \left[\frac{9a^{2}}{16} + 3a^{2}(V_{0} + A_{0})(1 - a^{2})\right]\bar{\xi}^{4} + O(\bar{\xi}^{6})$$

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$$D_{1/3}^a = 1 + O(\bar{\xi}^4)$$

$$V_0 = \frac{2}{3}A_0 = \frac{1}{20}$$

### The initial data created by self-similar waves at the end of the Radiation Epoch depends on:

- (I) The temperature  $T_*$  at which p=0
- (2) The value of the acceleration parameter  $\alpha$

OUR GOAL NOW: Use our equations to evolve the initial data at the end of radiation to determine

$$(a,T_*)$$

that gives the correct anomalous acceleration.

I.e.,  $(a, T_*)$  that give the observed quadratic correction to redshift vs luminosity at present time

In the Standard Model p=0 at about

$$t_*pprox$$
 10,000-30,000 yrs

$$T_* \approx 9000^0 K$$

(Depending on theories of Dark Matter)

- Our simulation turns out to be entirely insensitive to the initial  $t_*$ ,  $T_*$
- I.e., we need only compute the value of the acceleration parameter that gives the correct anomalous acceleration.

 Technical Problem: The self-similar waves at the end of radiation are in the wrong gauge due to the fact that time since the Big Bang changes

between 
$$p=0$$
 and  $p=\frac{c^2}{3}\rho$ 

 That is: The initial data for the self-similar waves does not meet the gauge conditions for our p=0 ansatz

$$A_2 = -\frac{1}{3}z_2, \quad A_4 = -\frac{1}{5}z_4, \quad D_2 = -\frac{1}{12}z_2$$

(Resolving this held us back for close to a year!)

 Resolution: We post-process the initial data by a gauge transformation of the form----

$$t = \bar{t} + \frac{1}{2}q(\bar{t} - \bar{t}_*)^2 - t_B$$

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 I.e, The SSC metric form is invariant under arbitrary changes of time, (choice of gauge)match the gauge to match the metrics  Resolution: We post-process the initial data by a gauge transformation of the form----

$$t = \bar{t} + \frac{1}{2}q(\bar{t} - \bar{t}_*)^2 - t_B$$

- I.e, The SSC metric form is invariant under arbitrary changes of time, (choice of gauge)match the gauge to match the metrics
- Check: Same change of gauge is required to match FRW metrics in SSC coordinates...

**THEOREM:** Let the transformation  $\bar{t} \to t$  be defined by

$$t = \bar{t} + \frac{1}{2}q(\bar{t} - \bar{t}_*)^2 - t_B,$$

where q and  $t_B$  are given by

$$t_B = \bar{t}_*(1 - \alpha),$$

$$q = \frac{a^2}{16\bar{\gamma}} = \frac{a^2}{2(1+a^2)},$$

where

$$\alpha = \frac{1}{5} \left( \frac{1 + a^2}{1.3 - a^2} \right).$$

Then, on the constant temperature surface  $T=T_*$ , the initial data from the self-similar waves at the end of the radiation epoch meets the gauge conditions in  $(\bar{t}, \bar{\xi})$ .

• 2nd Technical Problem: The  $T=T_*,\; \rho=\rho_*$  surfaces are distinct from the constant time  $t=t_*$  surfaces

• 2nd Technical Problem: The surfaces are distinct from the constant time  $t=t_{*}$  surfaces

 Resolution: To get the asymptotics correct we have to pull the initial data back to

$$t=t_*$$

The initial data created by self-similar waves on a constant temperature surface at the end of the Radiation Epoch

**THEOREM** The initial data for our p = 0 evolution at time  $t = t_*$  is given as a function of the acceleration parameter a and start temperature  $\rho_* = a_{SB}T_*$  by

$$z_{2}(t_{*}) = \hat{z}_{2},$$

$$z_{4}(t_{*}) = \hat{z}_{4} + 3\hat{w}_{0} \left(\frac{4}{3} + \hat{z}_{2}\right) \gamma,$$

$$w_{0}(t_{*}) = \hat{w}_{0},$$

$$w_{2}(t_{*}) = \hat{w}_{2} + \left(\frac{1}{6}\hat{z}_{2} + \frac{1}{3}\hat{w}_{0} + \hat{w}_{0}^{2}\right) \gamma,$$

where

$$\gamma = \alpha \bar{\gamma} = \alpha \left( \frac{2 - a^2}{4} \right)$$

$$t_* = \sqrt{\frac{3a^2}{4\kappa\rho_*}}, \quad \alpha = 4\frac{2-a^2}{7-4a^2}$$

$$\hat{z}_{2} = \left\{ \frac{3a^{2}\alpha^{2}}{4} - \frac{4}{3} \right\}_{z2} 
\hat{z}_{4} = \left\{ 2\alpha^{3}(1-\alpha)\bar{\gamma}Z_{2} + \alpha^{4}Z_{4} - \frac{40}{27} \right\}_{z4} 
\hat{w}_{0} = \left\{ \frac{\alpha}{2} - \frac{2}{3} \right\}_{v1} 
\hat{w}_{2} = \left\{ \alpha^{2}(1-\alpha)\bar{\gamma}W_{0} + \alpha^{3}W_{2} - \frac{2}{9} \right\}_{v3} 
Z_{2} = \frac{3a^{2}}{4} 
Z_{4} = \left[ \frac{9a^{2}}{16} + 3a^{2}(V_{0} + A_{0})(1-a^{2}) \right] 
V_{0} = \frac{1}{20}, \quad A_{0} = \frac{3}{40} 
W_{0} = \frac{1}{2}, \quad W_{2} = \left[ \frac{1}{8} + V_{0}(1-a^{2}) \right]$$

# 4. Redshift vs Luminosity as a function of our corrections

#### A (long) Calculation gives:

$$H_0 d_{\ell} = z \left\{ 1 + \left[ \frac{1}{4} + E_2 \right] z + \left[ -\frac{1}{8} + E_3 \right] z^2 \right\} + O(z^4)$$

Anomalous Acceleration

Cubic Correction

$$E_2 = \frac{24w_0 + 45w_0^2 + 3z_2}{4(2+3w_0)^2} = E_2(z_2, w_0),$$

$$E_3 = E_3(z_2, w_0, w_3)$$

#### $E_3(z_2, w_0, w_2)$ is quite complicated:

$$H_0 d_{\ell} = z \left\{ 1 + \left[ \frac{1}{4} + E_2 \right] z + \left[ -\frac{1}{8} + E_3 \right] z^2 \right\} + O(z^4)$$

Cubic Correction

#### A calculation gives:

$$E_3 = 2I_2 + I_3,$$

$$I_2 = H_2 + \frac{9w_0}{2(2+3w_0)}$$

$$I_3 = H_3 + 3\left[-1 + \left(\frac{8 - 8H_2 + 3w_0 - 12H_2w_0}{2(2 + 3w_0)^2}\right)\right],$$

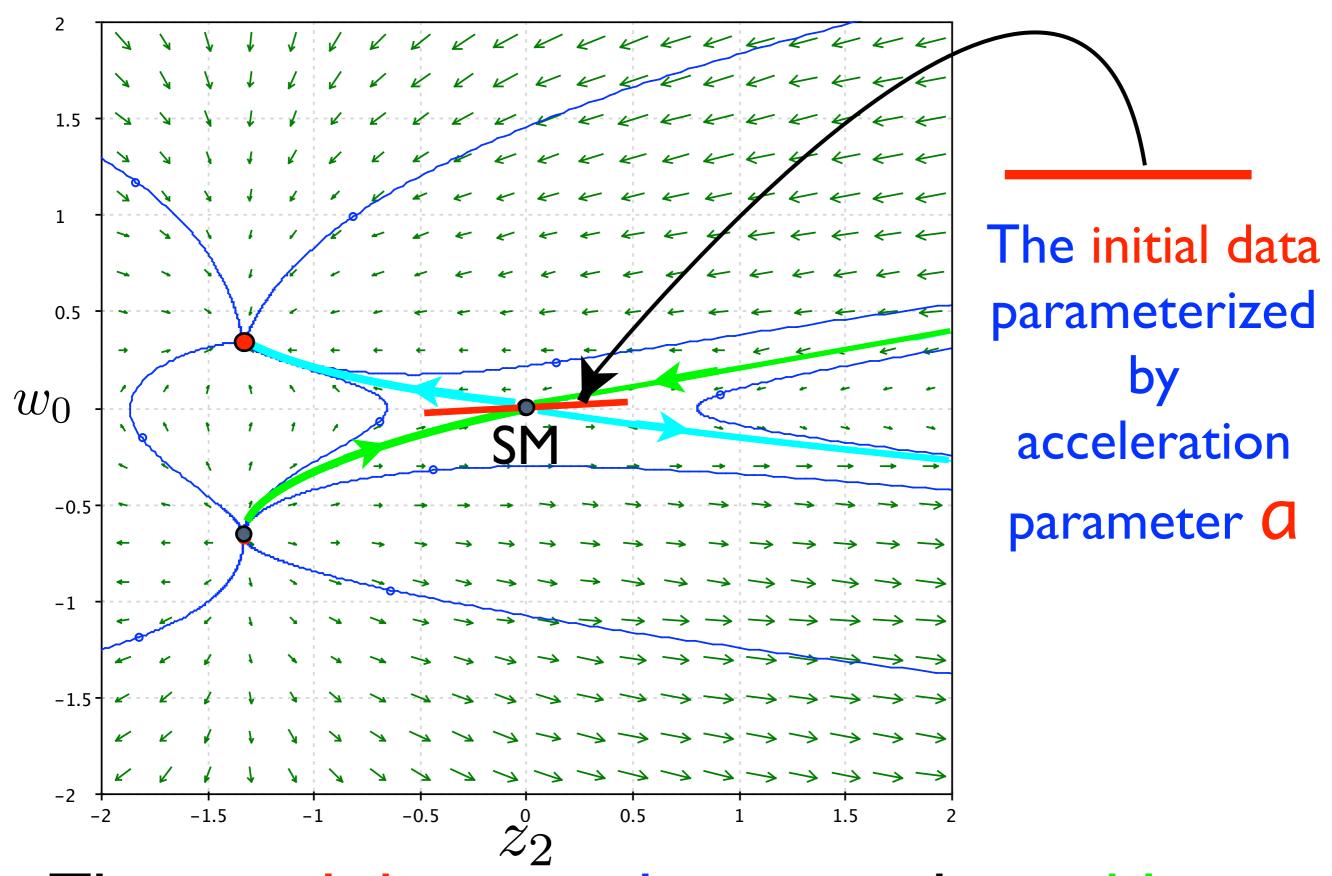
$$H_2 = \frac{1}{4} \left\{ 1 - \frac{1 + 9\left(\frac{2}{3}w_0 + \frac{1}{2}w_0^2 - \frac{1}{12}z_2\right)}{(1 + \frac{3}{2}w_0)^2} \right\},\,$$

$$H_3 = \frac{5}{8} \left\{ 1 - \frac{1 - \frac{18}{5}Q_2 - \frac{81}{5}Q_2^2 + \frac{9}{5}w_0 + \frac{27}{5}Q_3 + \frac{81}{10}Q_3w_0}{\left(1 + \frac{3}{2}w_0\right)^4} \right\}$$

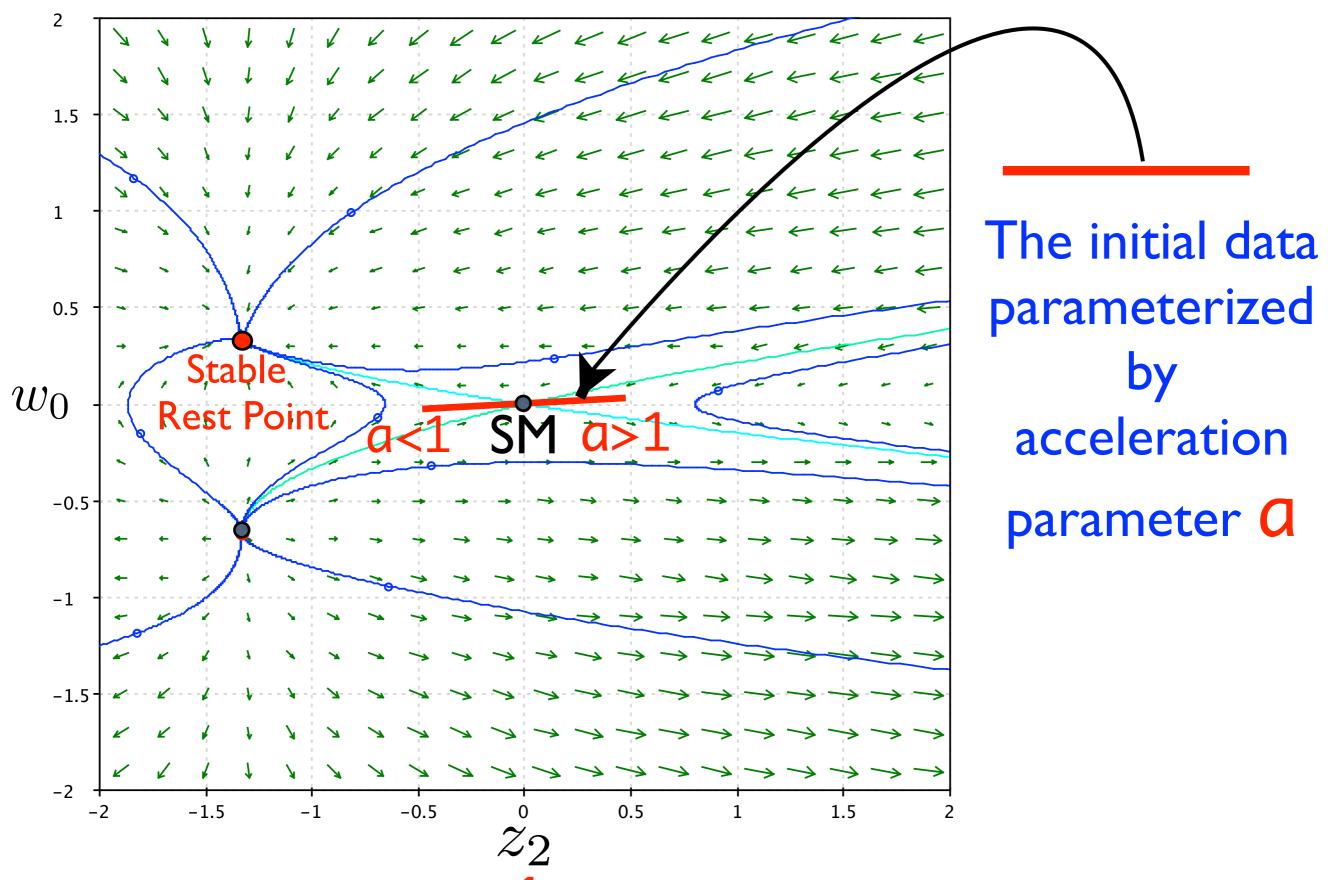
$$Q_2 = \frac{2}{3}w_0 + \frac{1}{2}w_0^2 - \frac{1}{12}z_2$$

$$Q_3 = \frac{2}{9}w_0 + w_0^2 + \frac{1}{2}w_0^3 + w_2 - \frac{1}{18}z_2 - \frac{1}{3}z_2w_0$$

(Each term represents a different effect...)



The initial data cuts between the stable and unstable manifold of SM



Under-densities <a>a</a> 1 are within the domain of attraction of the Stable Rest Point

# 3. Comparison with the Standard Model

• Redshift vs Luminosity for k=0 Friedmann can be obtained from exact formulas:  $p = \sigma \rho$ 

$$H_0 d_{\ell} = \frac{2}{1+3\sigma} \left\{ (1+z) - (1+z)^{\frac{1-3\sigma}{2}} \right\}.$$

• In the case  $p = \sigma = 0$ , we get

$$H_0 d_{\ell} = z + \frac{1}{4}z^2 - \frac{1}{8}z^3 + O(z^4)$$

• C.f. our formula:

$$H_0 d_{\ell} = z \left\{ 1 + \left[ \frac{1}{4} + E_2 \right] z + \left[ -\frac{1}{8} + E_3 \right] z^2 \right\} + O(z^4)$$

# Cosmology now assumes a Cosmological Constant with

#### Seventy Percent Dark Energy

$$H_0 d_{\ell} = (1+z) \int_0^z \frac{dy}{(1+z)\sqrt{1+\Omega_M y}}.$$
  $\Omega_M + \Omega_{\Lambda} = 1$ 

Taylor expanding gives:

$$H_0 d_{\ell} = z + \frac{1}{2} \left( -\frac{\Omega_M}{2} + 1 \right) z^2 + \frac{1}{6} \left( -1 - \frac{\Omega_M}{2} + \frac{3\Omega_M^2}{4} \right) z^3 + O(z^4)$$

In the case  $\Omega_M = .3$ ,  $\Omega_{\Lambda} = .7$  this gives

$$H_0 d_{\ell} = z + .425 z^2 - .1804 z^3 + O(z^4)$$

#### CONCLUDE: k = 0, p = 0 Friedmann

with and without Dark Energy  $\Omega_M + \Omega_{\Lambda} = 1$ 

$$\Omega_M + \Omega_\Lambda = 1$$

$$H_0d_\ell=z+.425\,z^2-.1804\,z^3+O(z^4)$$

The Anomalous Acceleration 
$$H_0d_\ell=z+.25\,z^2-.125\,z^3+O(z^4)$$

Standard Model with Dark Energy

$$\Omega_{\Lambda} = .7$$

Standard Model Without Dark Energy

$$\Omega_{\Lambda} = 0$$

# IN FACT: As the Dark Energy Parameter ranges from 0 to 1, the Anomalous Acceleration ranges from .25 to .5

$$H_0 d_{\ell} = z + \frac{1}{2} \left( -\frac{\Omega_M}{2} + 1 \right) z^2 + \frac{1}{6} \left( -1 - \frac{\Omega_M}{2} + \frac{3\Omega_M^2}{4} \right) z^3 + O(z^4)$$

Range: .25 to .5

as

$$0 \le \Omega_M \le 1$$

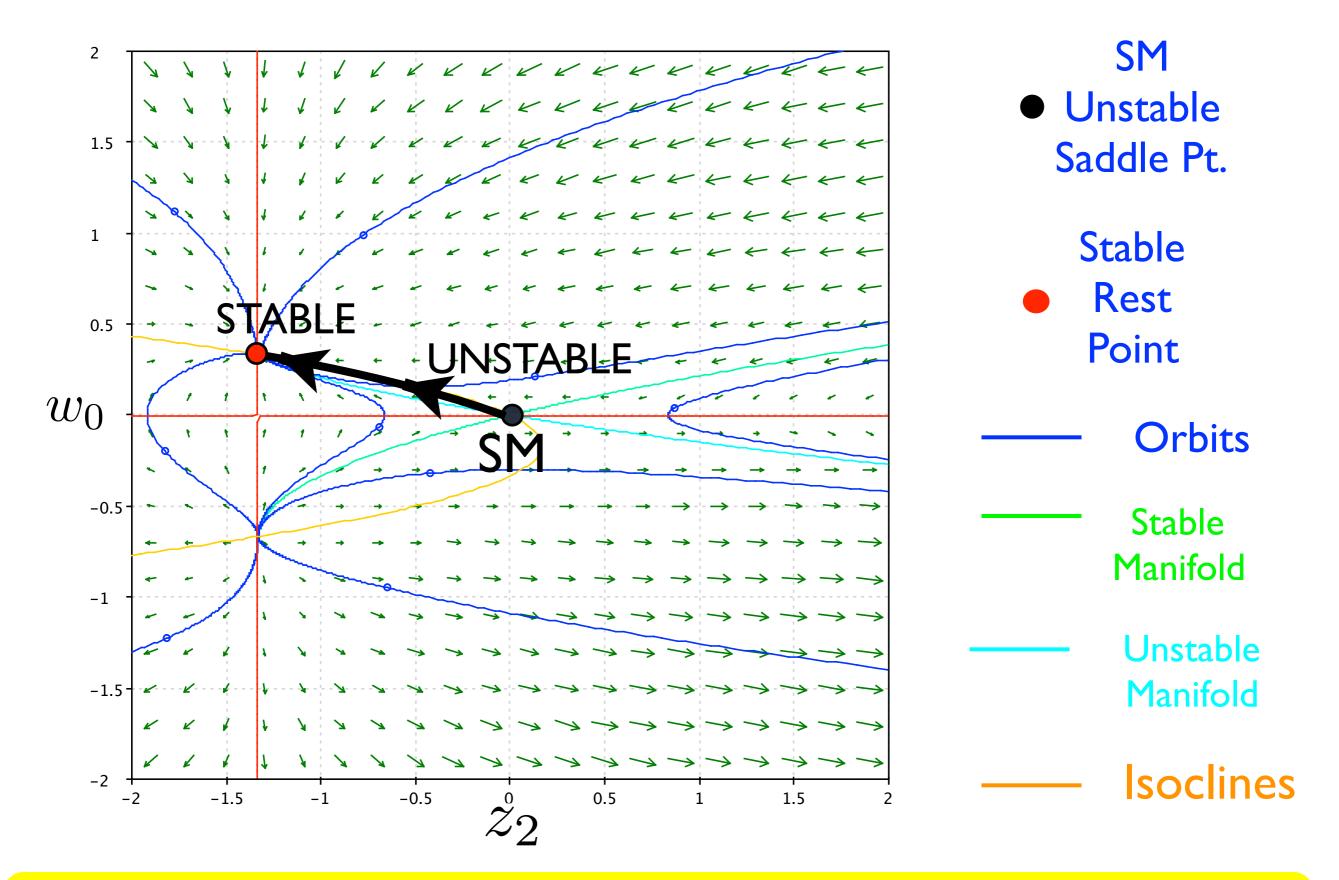
# We get the Same Conclusion in the Wave Theory!

$$H_0 d_{\ell} = z \left\{ 1 + \left[ \frac{1}{4} + E_2 \right] z + \left[ -\frac{1}{8} + E_3 \right] z^2 \right\} + O(z^4)$$

Range: .25 to .5

$$E_2 = \frac{24w_0 + 45w_0^2 + 3z_2}{4(2+3w_0)^2}$$

along the orbit
from the Standard Model
to the
Stable Rest Point



The Anomalous Acceleration ranges from .25 to .5 along orbit from SM to Stable Rest Point  $\approx Dark\ Energy$ 

### 5. Determination of the value of the Acceleration Parameter that matches the Anomalous Acceleration

We simulate our equations starting from the self-similar wave data at the end of radiation  $T = T_*$ , to find the value of  $(a, T_*)$  that gives the same Anomalous Acceleration as seventy percent Dark Energy when  $H = H_0$ :

$$H_0 d_{\ell} = z + \underbrace{.425}_{\Omega_{\Lambda}} z^2 - .1804 z^3 + O(z^4)$$
 Dark Energy  $\Omega_{\Lambda} = .7$ 

$$H_0 d_\ell = z + [.25 + E_2]z^2 + [-.125 + E_3]z^3 + O(z^4)$$

$$-z_2' = -t\dot{z}_2 = 3w_0 \left(\frac{4}{3} + z_2\right),$$

$$-z_4' = -t\dot{z}_4 = -5\left\{\frac{2}{27}z_2 + \frac{4}{3}w_2 - \frac{1}{18}z_2^2 + z_2w_2\right\}$$

$$-5w_0 \left\{\frac{4}{3} - \frac{2}{9}z_2 + z_4 - \frac{1}{12}z_2^2\right\},$$

$$-w_0' = -t\dot{w}_0 = \frac{1}{6}z_2 + \frac{1}{3}w_0 + w_0^2,$$

$$-w_2' = -t\dot{w}_2 = \frac{1}{10}z_4 + \frac{4}{9}w_0 - \frac{1}{3}w_2 + \frac{1}{24}z_2^2 - \frac{1}{3}z_2w_0$$

$$-\frac{1}{3}w_0^2 + 4w_0w_2 - \frac{1}{4}w_0^2z_2.$$

Our Wave Model

$$E_2 = \frac{24w_0 + 45w_0^2 + 3z_2}{4(2+3w_0)^2}$$

THE ANSWER: The value of the acceleration for the wave perturbation of SM that produces a quadradic correction of .425 at the present value of  $H_0$  is:

$$\underline{a} = 0.99999957 = 1 - (4.3 \times 10^{-7})$$

$$H_0 d_\ell = z + .425z^2 + .359z^3$$

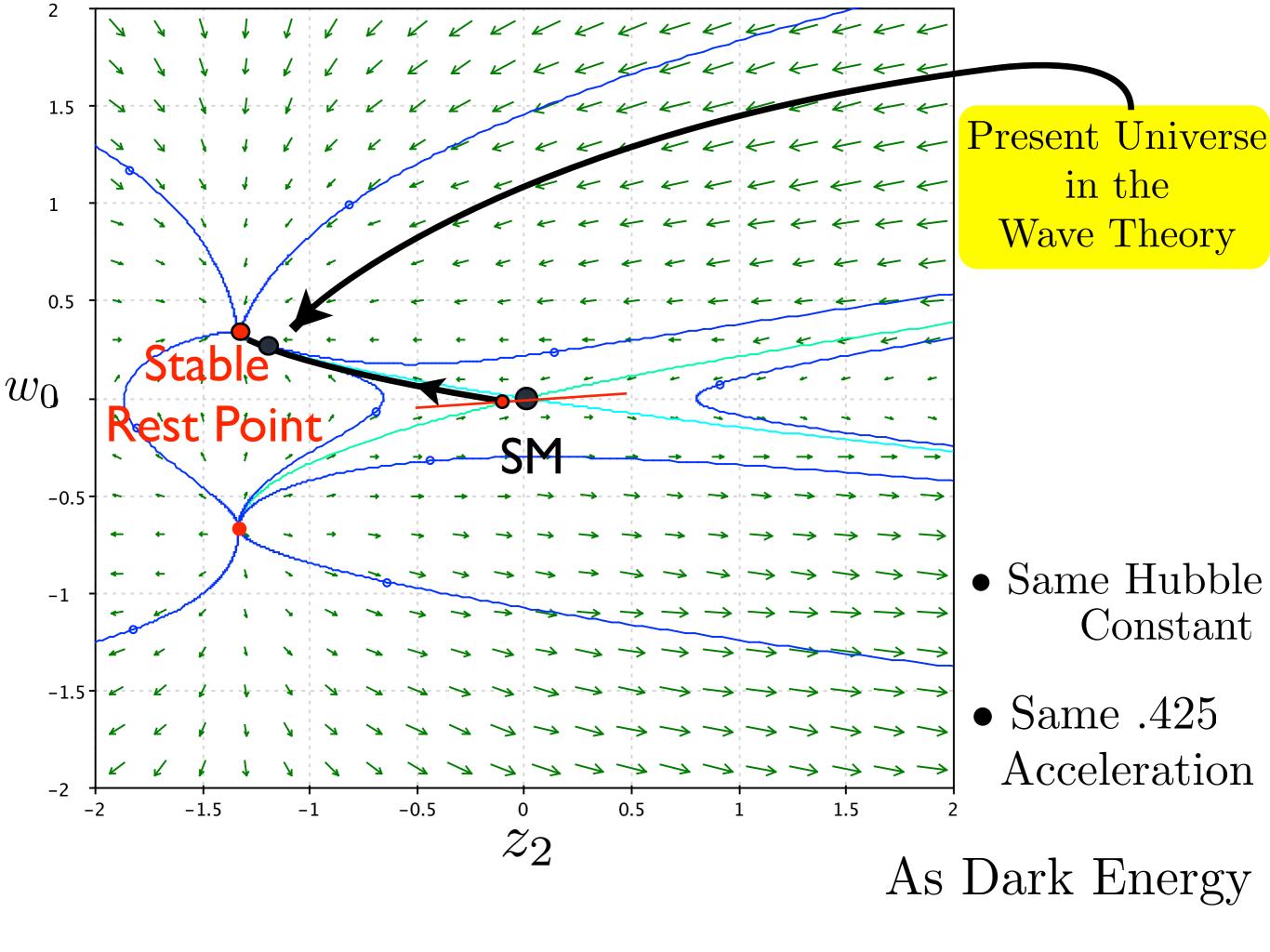
THE ANSWER: The value of the acceleration for the wave perturbation of SM that produces a quadradic correction of .425 at the present value of  $H_0$  is:

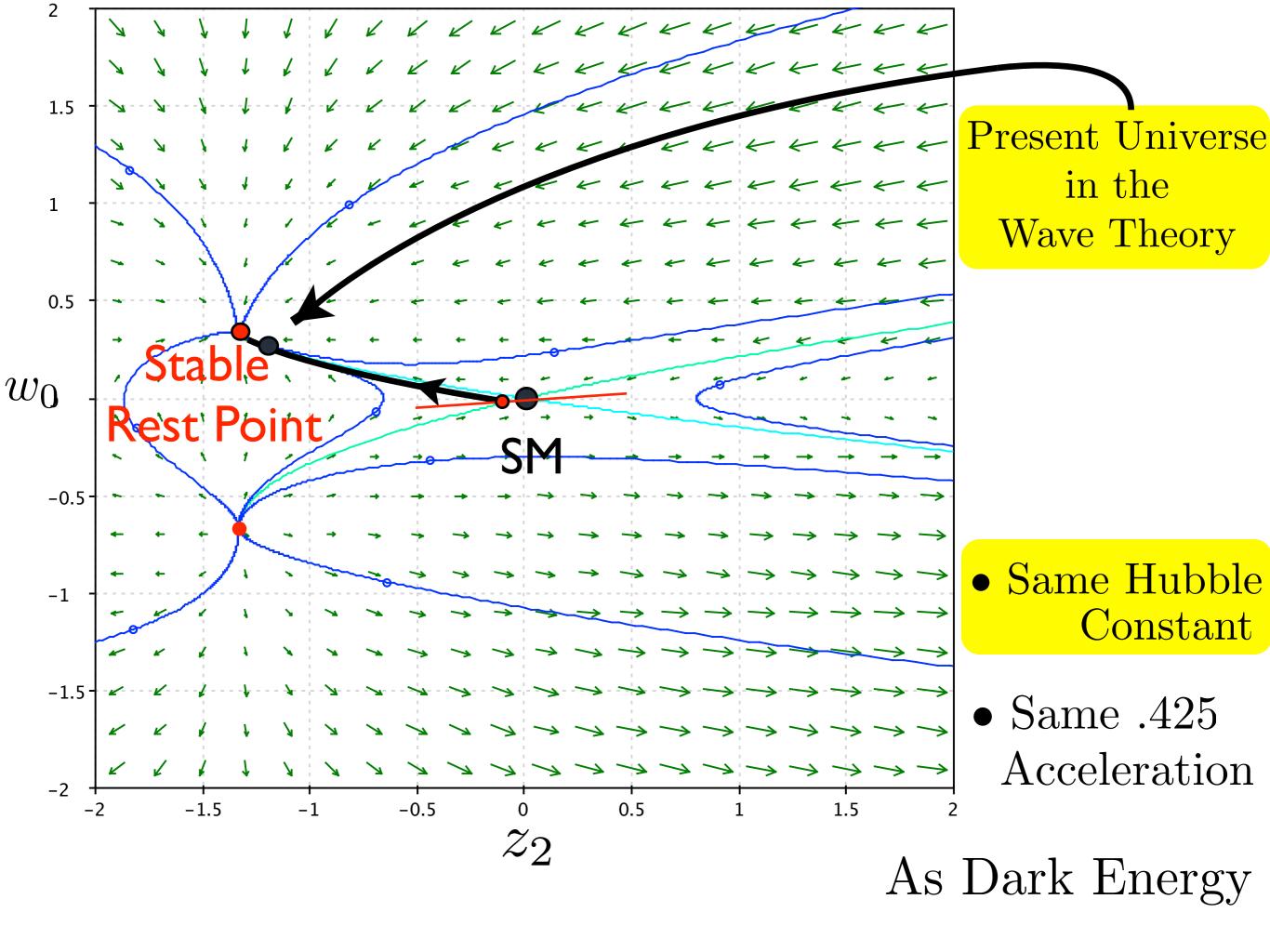
$$\underline{a} = 0.99999957 = 1 - (4.3 \times 10^{-7})$$

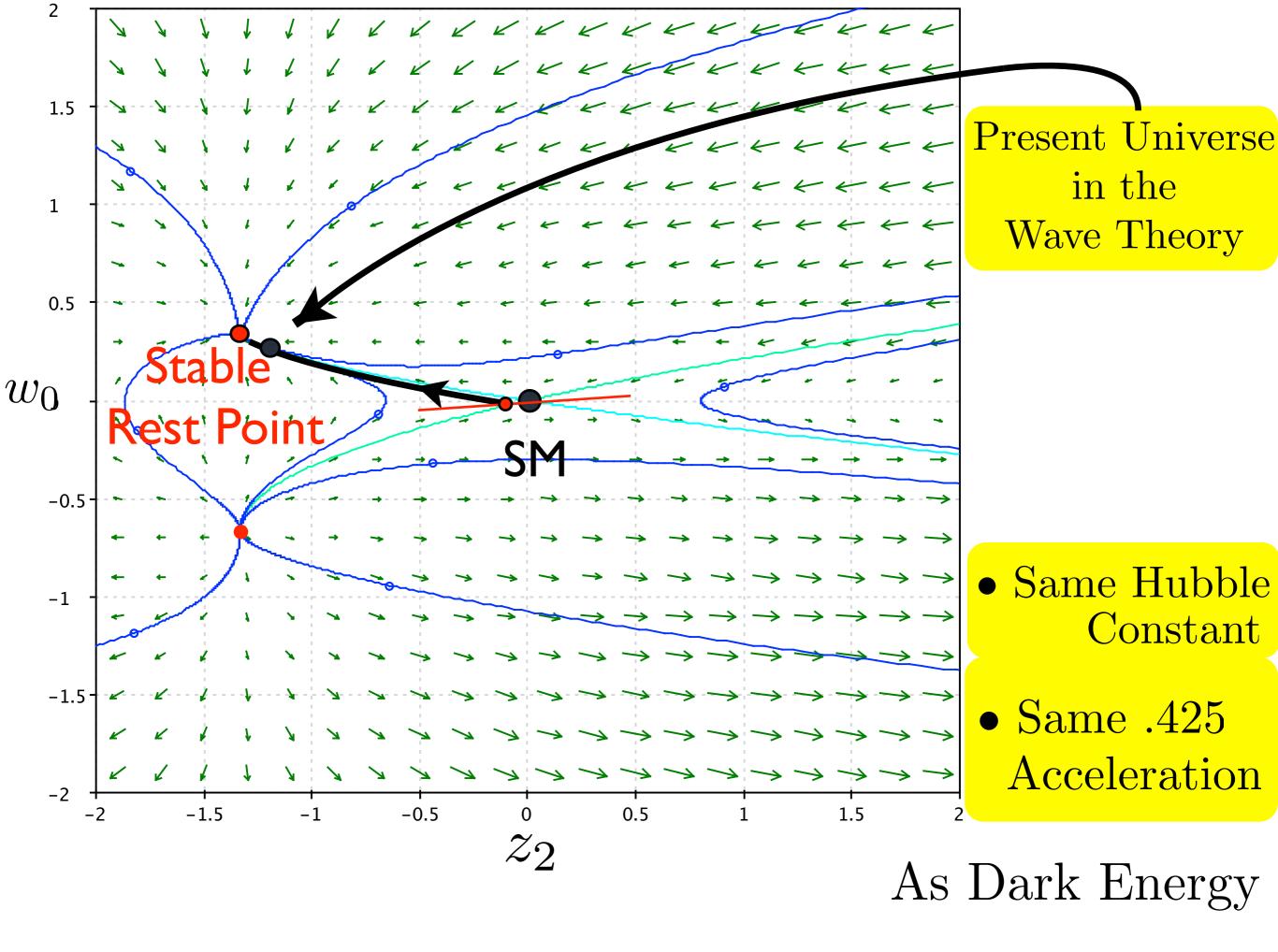
$$H_0 d_\ell = z + .425z^2 + .359z^3$$

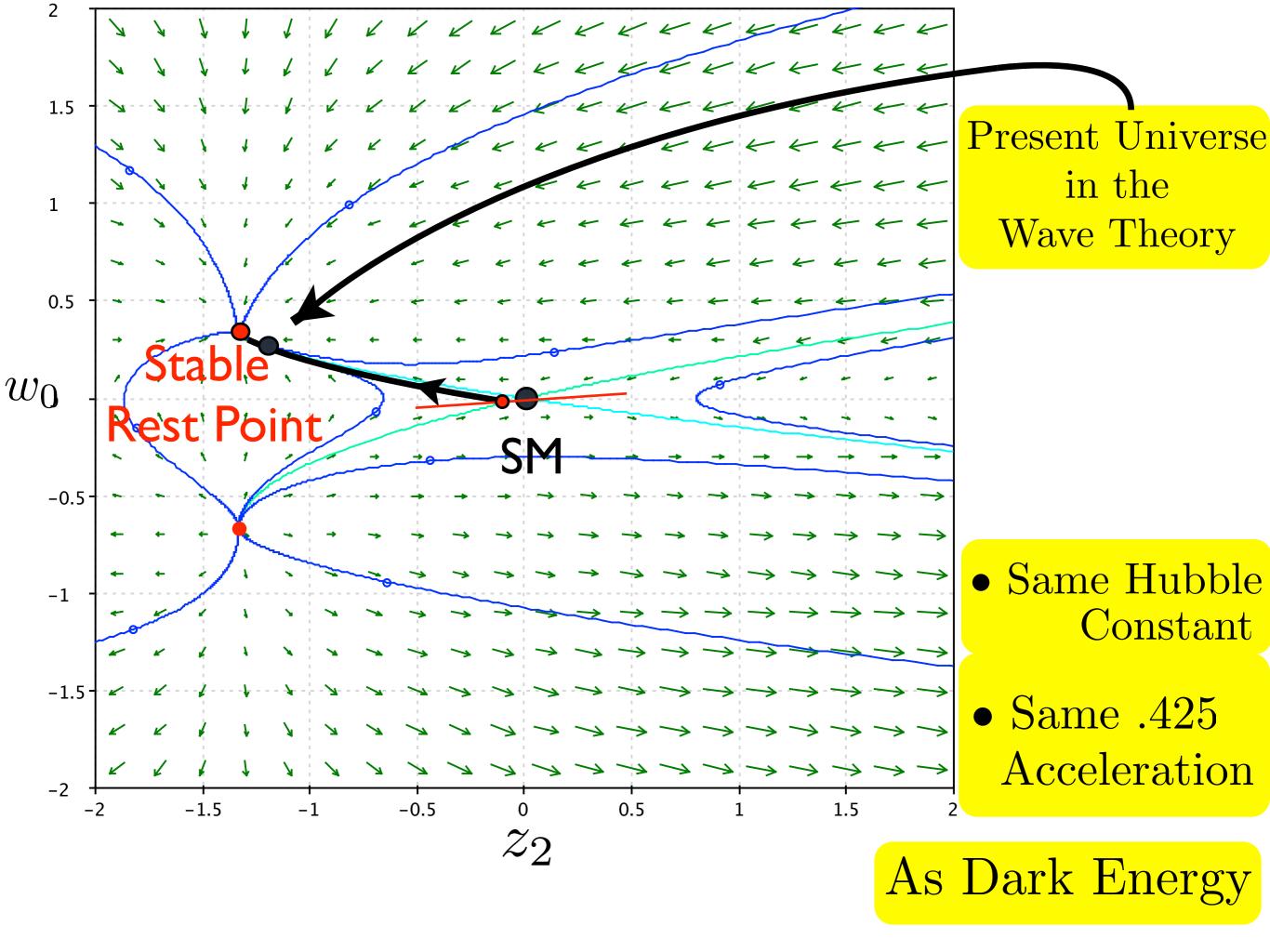
This corresponds to an relative underdensity of

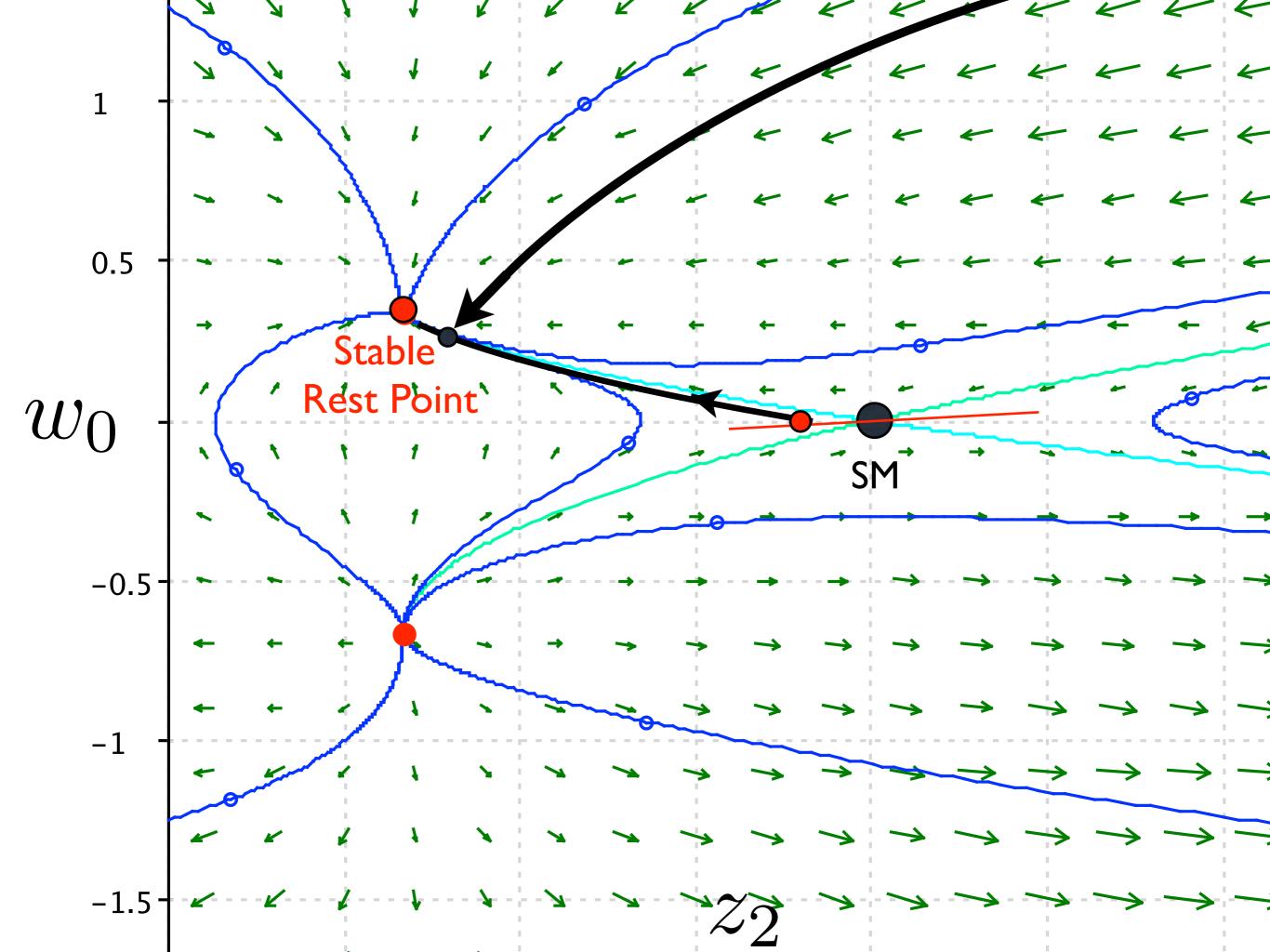
$$\frac{\rho_{SM} - \rho_{ssw}}{\rho_{SM}} = 7.45 \times 10^{-6}$$

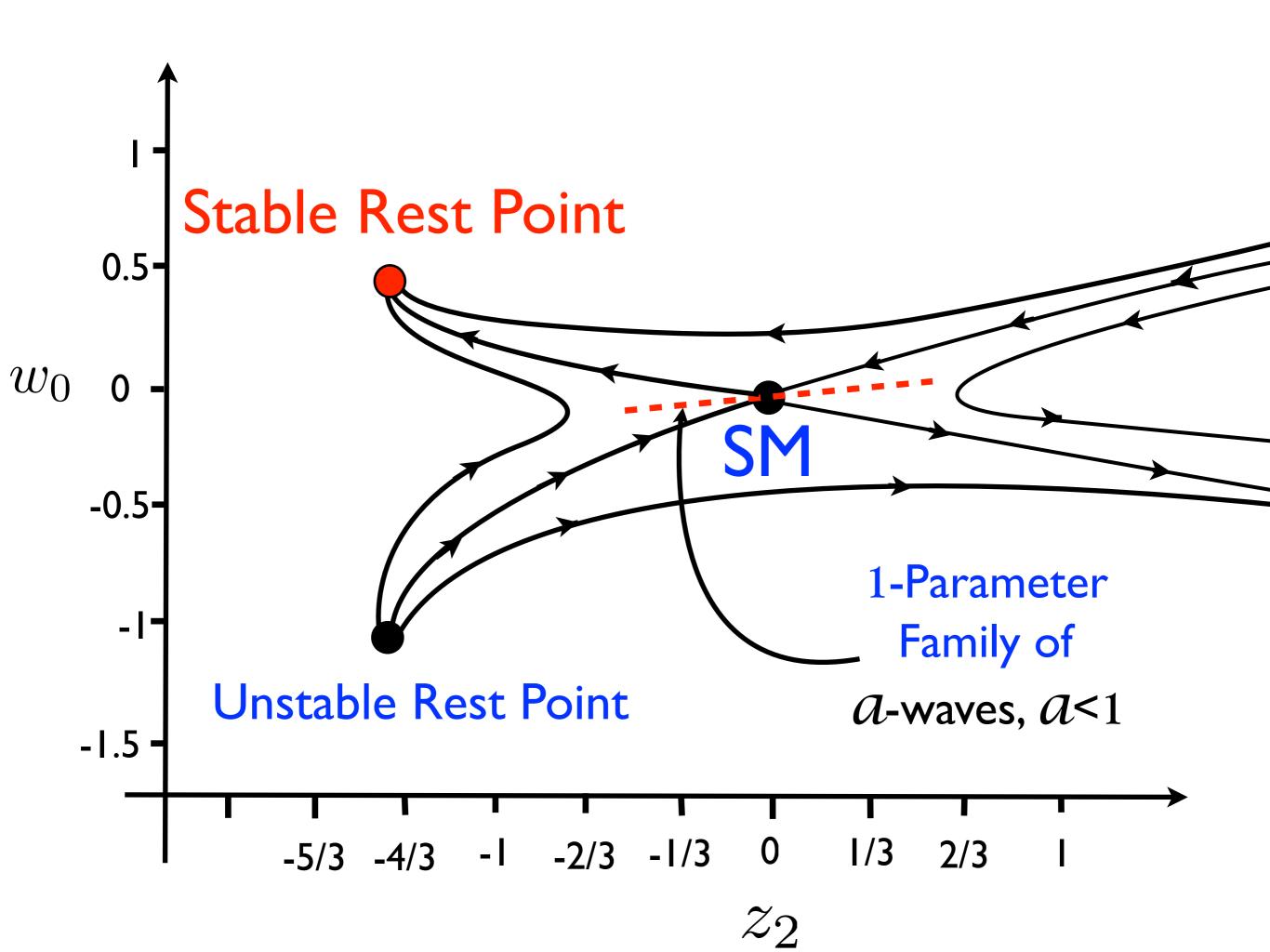












The relative underdensity at the end of radiation:

$$\frac{\rho_{SM} - \rho_{ssw}}{\rho_{SM}} = 7.45 \times 10^{-6}$$

Numerical Simulation gives the relative under-density at present time as:

$$\frac{\rho_{ssw}}{\rho_{SM}} = .144 \approx \frac{1}{7}$$

Conclude: An under-density of one part in  $10^6$ at the end of radiation produces a seven-fold under-density at present time!

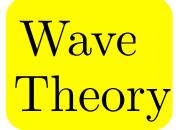
## Conclude: The Standard Model is Unstable to Perturbation by this family of Waves...

#### Comparison with Dark Energy:

$$H_0 d_{\ell} = z + .425z^2 - .180z^3$$

Dark
Energy

$$H_0 d_{\ell} = z + .425z^2 + .359z^3$$



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Dark Energy

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Wave Theory

The Wave Theory predicts a
Larger Anomalous Acceleration
far from the center than
Dark Energy

#### Comparison with Dark Energy:

$$H_0 d_\ell = z + .425z^2 - .180z^3$$

Dark
Energy

$$H_0 d_{\ell} = z + .425z^2 + .359z^3$$

Wave Theory

Age of universe about the same:

$$t_0 \approx (.95)t_{DE}$$

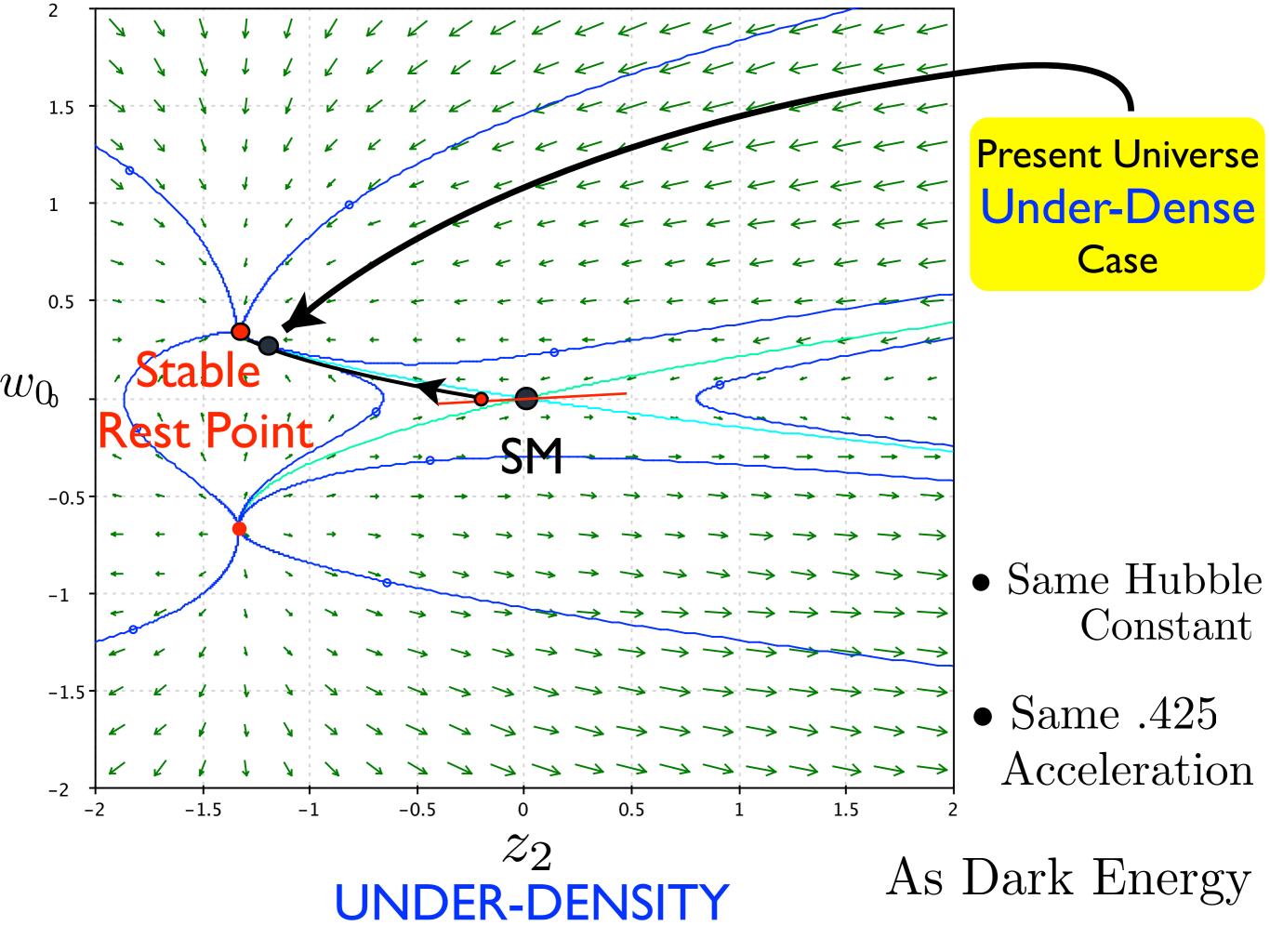
 $t_{DE} \approx 13.8 \text{ Billion years} \approx (1.45) t_{SM}$ 

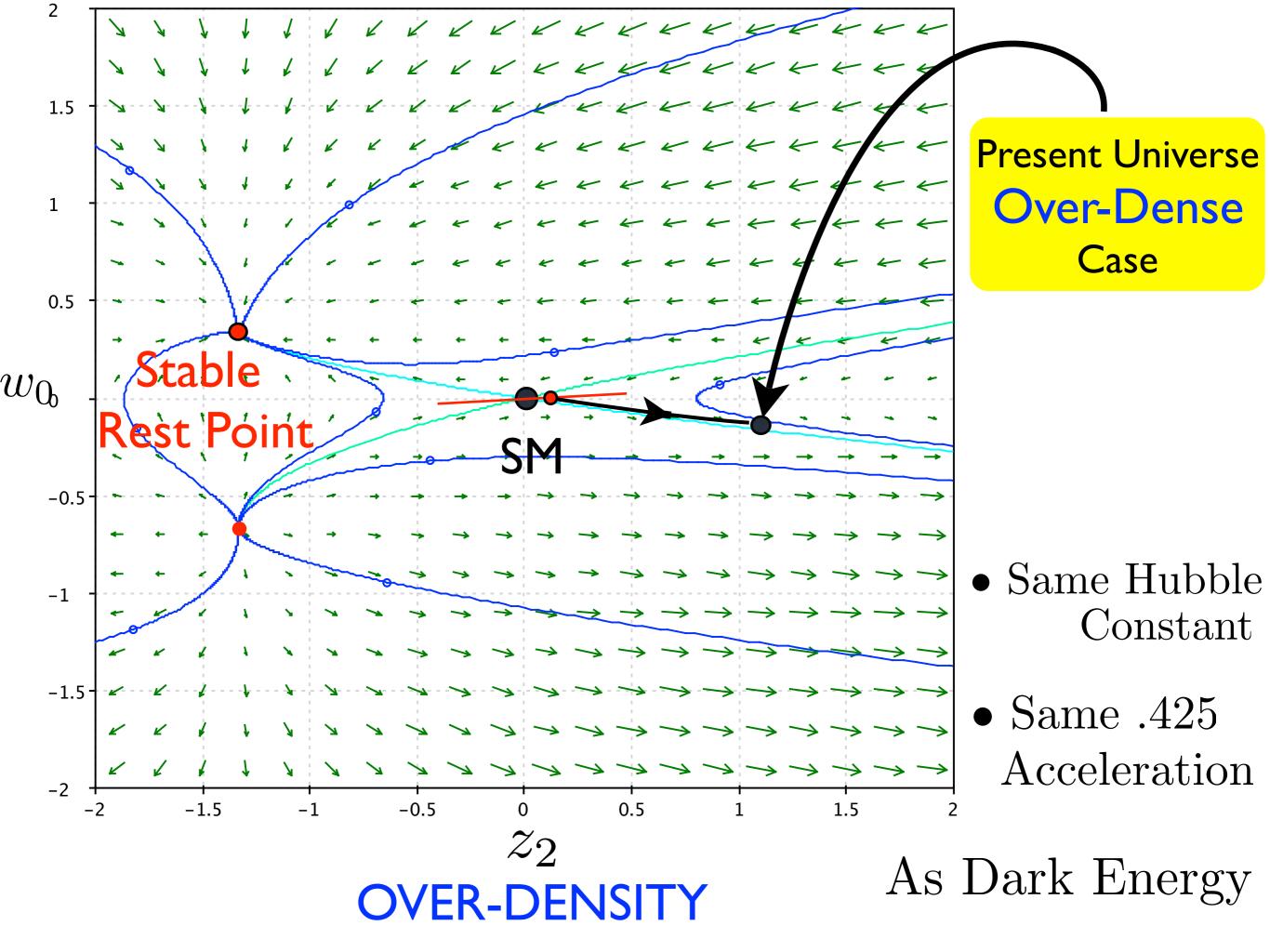
# In Fact: A slight over-density will also create the Anomalous Acceleration

$$\bar{a} = 1.0000006747 = 1 + (6.747 \times 10^{-7})$$

$$H_0 d_\ell = z + .425z^2 - 2.756z^3$$

A different cubic correction





## Conclude: The Standard Model Unstable to Perturbation by this Family of Waves, and under-densities create an **Anomalous Acceleration**

**Theorem:** Let  $t = t_0$  denote present time since the Big Bang in the wave model and  $t = t_{DE}$  present time since the Big Bang in the Dark Energy model. Then there exists a unique value of the acceleration parameter  $a = 0.99999959 \approx 1 - 4.3 \times 10^{-7}$  corresponding to an under-density relative to the SM at the end of radiation, such that the subsequent p = 0 evolution starting from this initial data evolves to time  $t = t_0$  with  $H = H_0$  and Q = .425, in agreement with the values of H and Q at  $t = t_{DE}$  in the Dark Energy model. The cubic correction at  $t = t_0$  in the wave theory is then C = 0.359, while Dark Energy theory gives C = -0.180 at  $t = t_{DE}$ . The times are related by  $t_0 \approx (.95)t_{DE}$ 

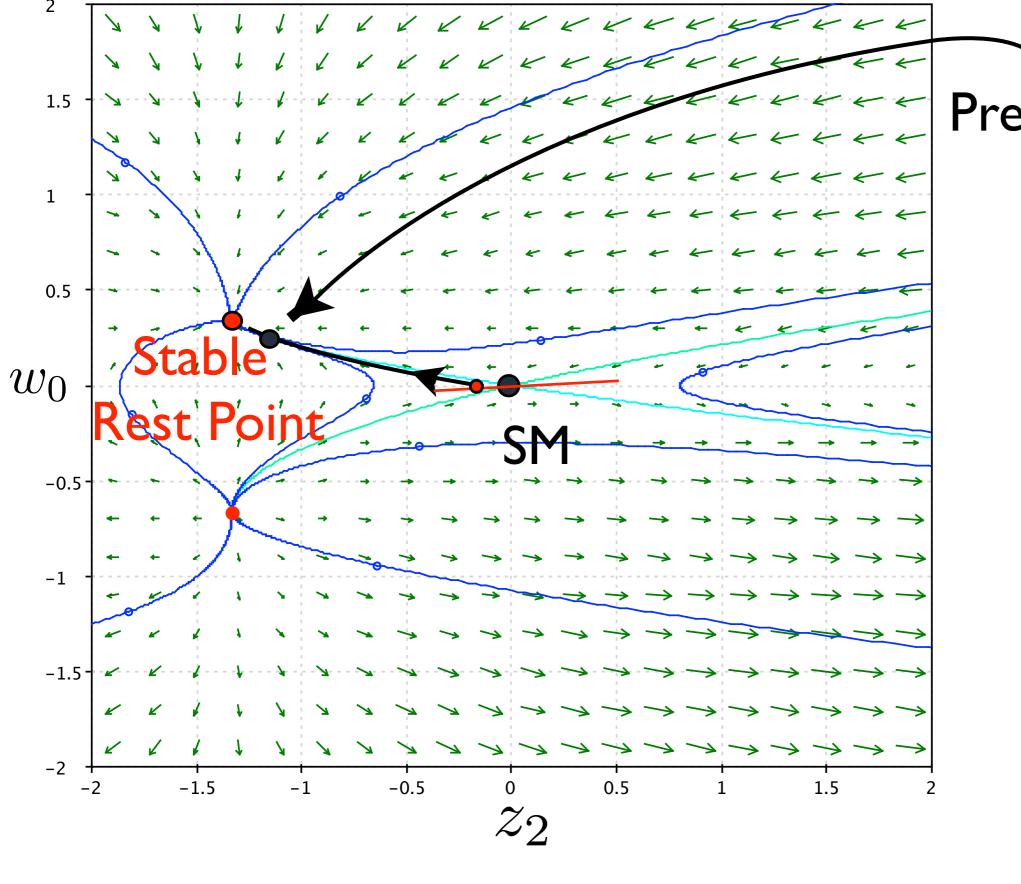
## 6. The Flat Uniformly Expanding Spacetime at the Center of the Wave

(Under-Dense Case: <u>a</u> < 1)

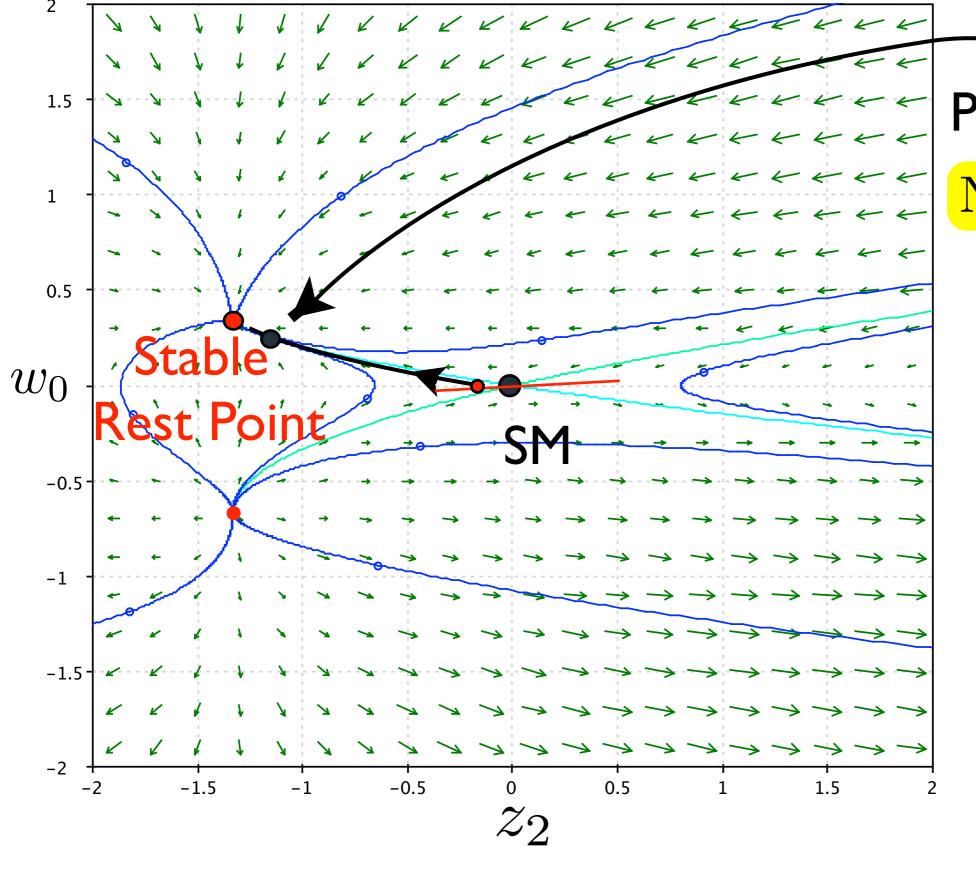
Consider the evolution of the spactime at the center obtained by neglecting all errors of order

 $O(\xi^4)$ 

## The spacetime near the center evolves toward the Stable Rest Point

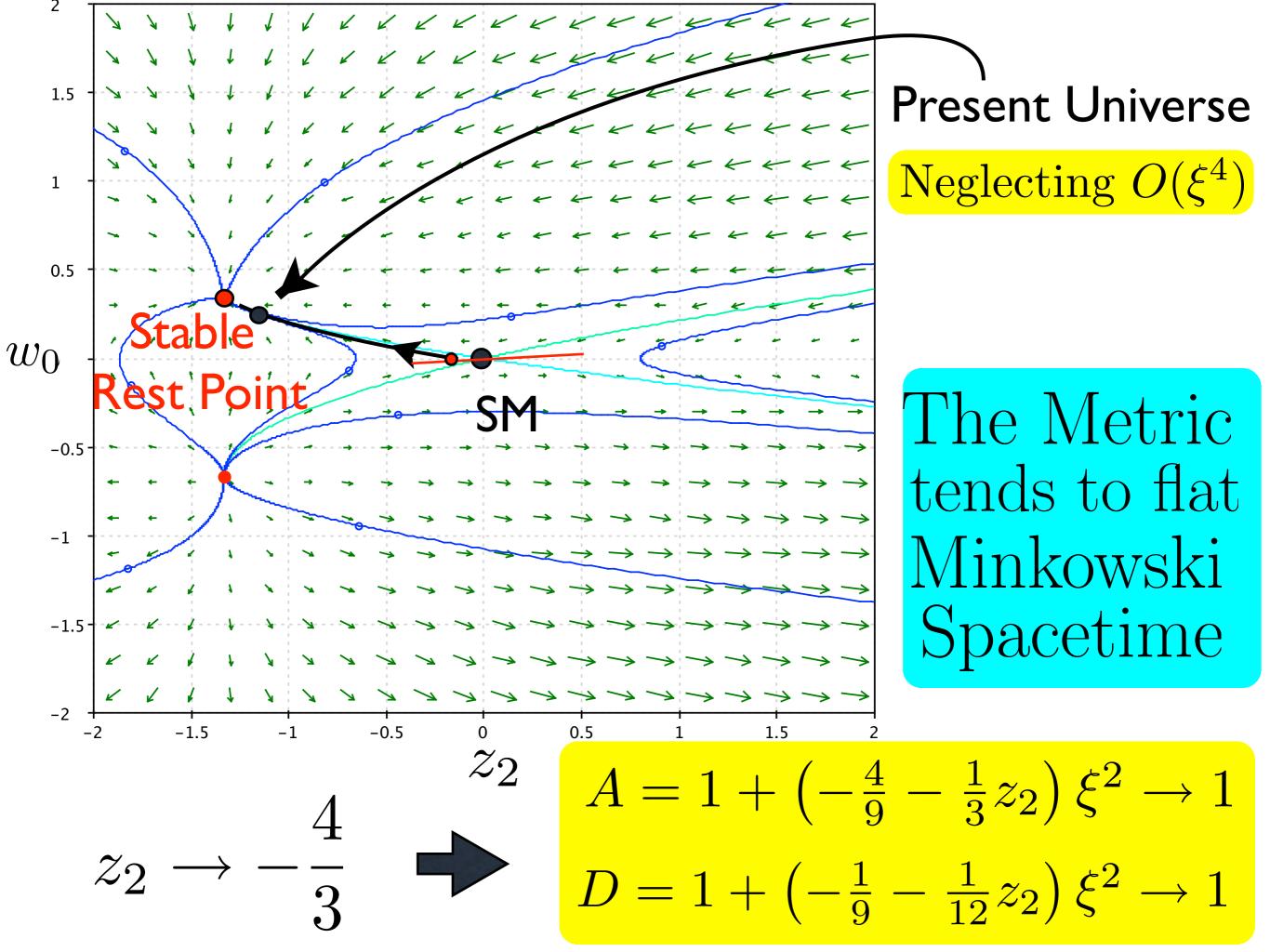


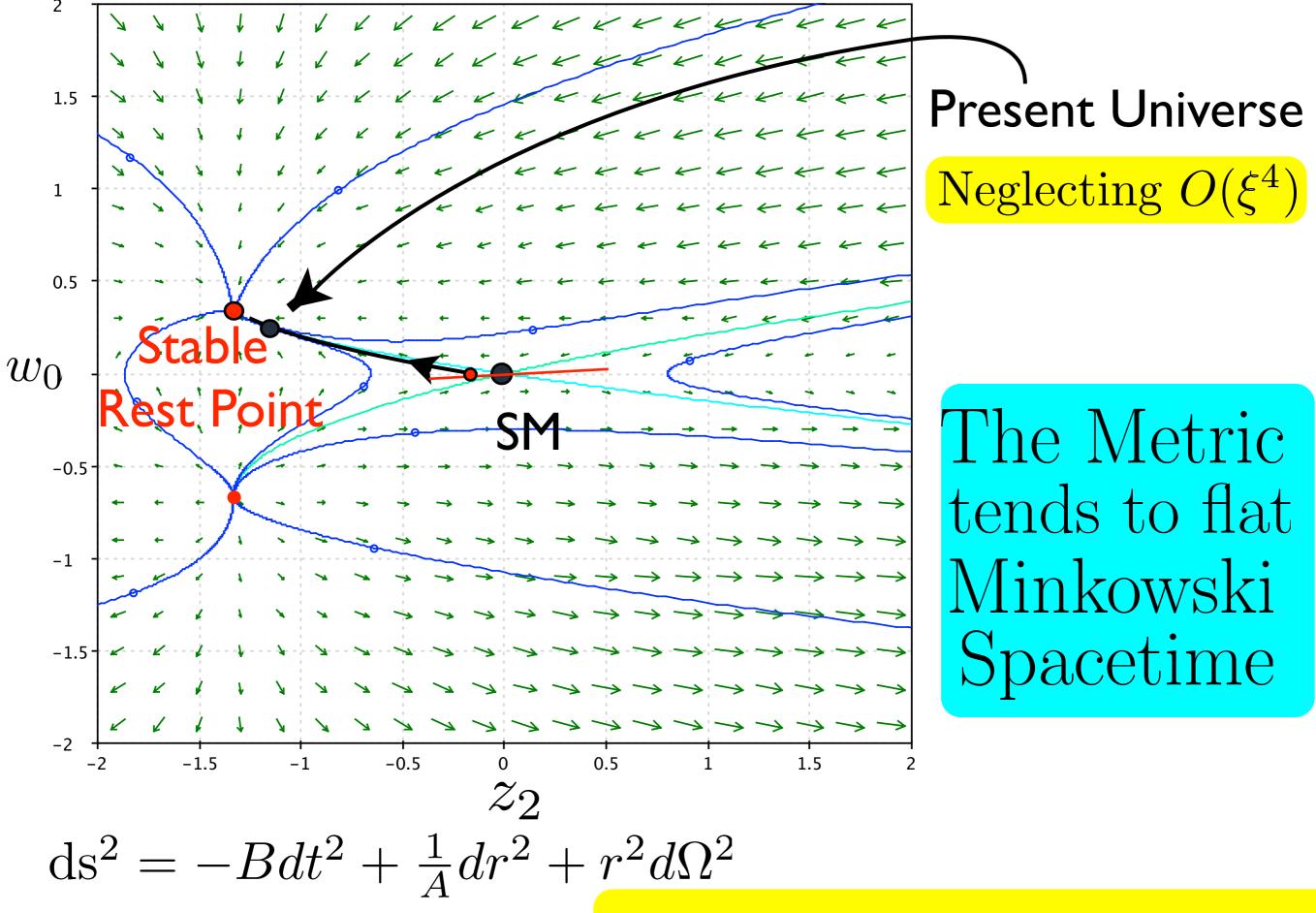
#### Present Universe



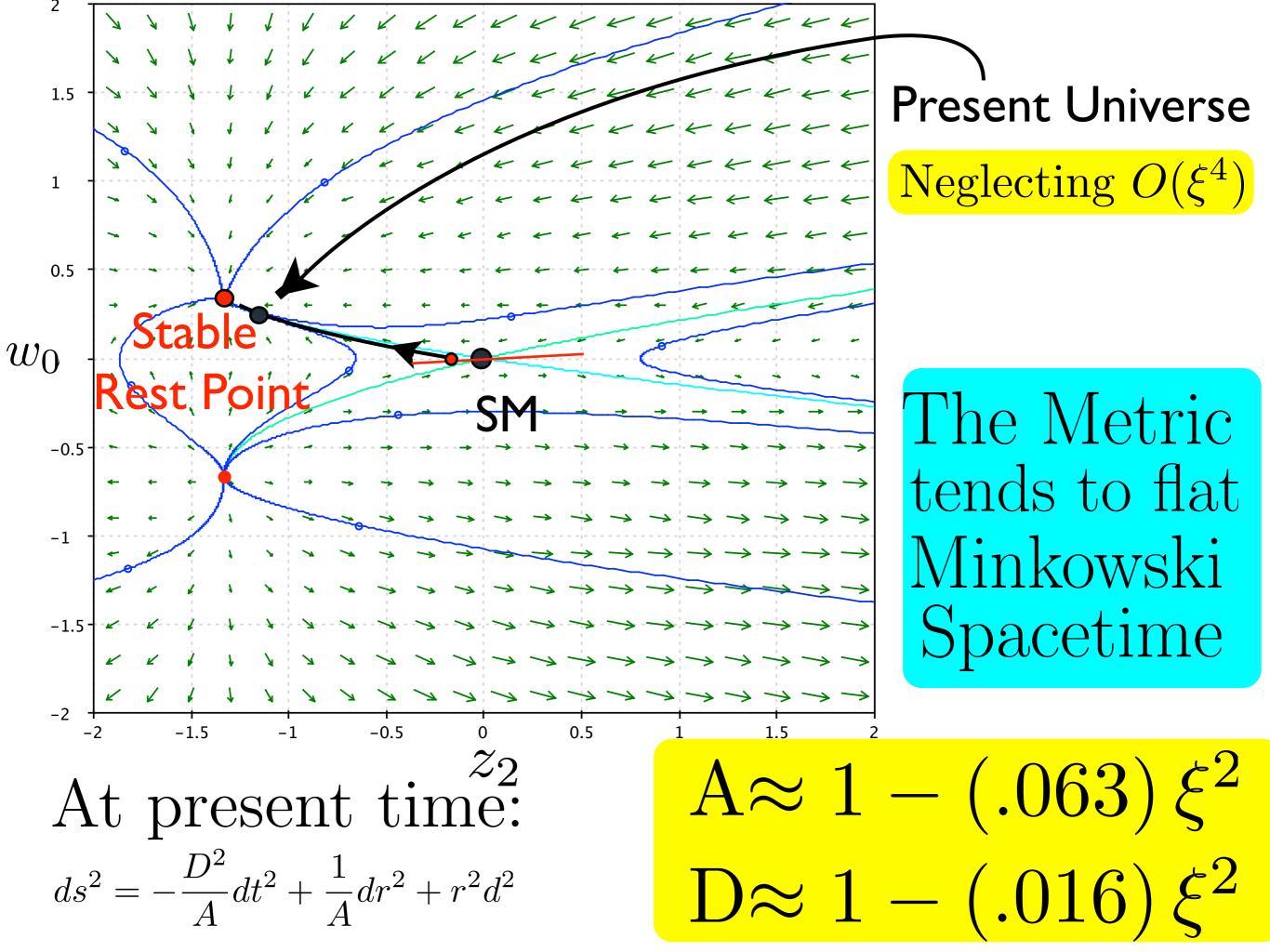
#### Present Universe

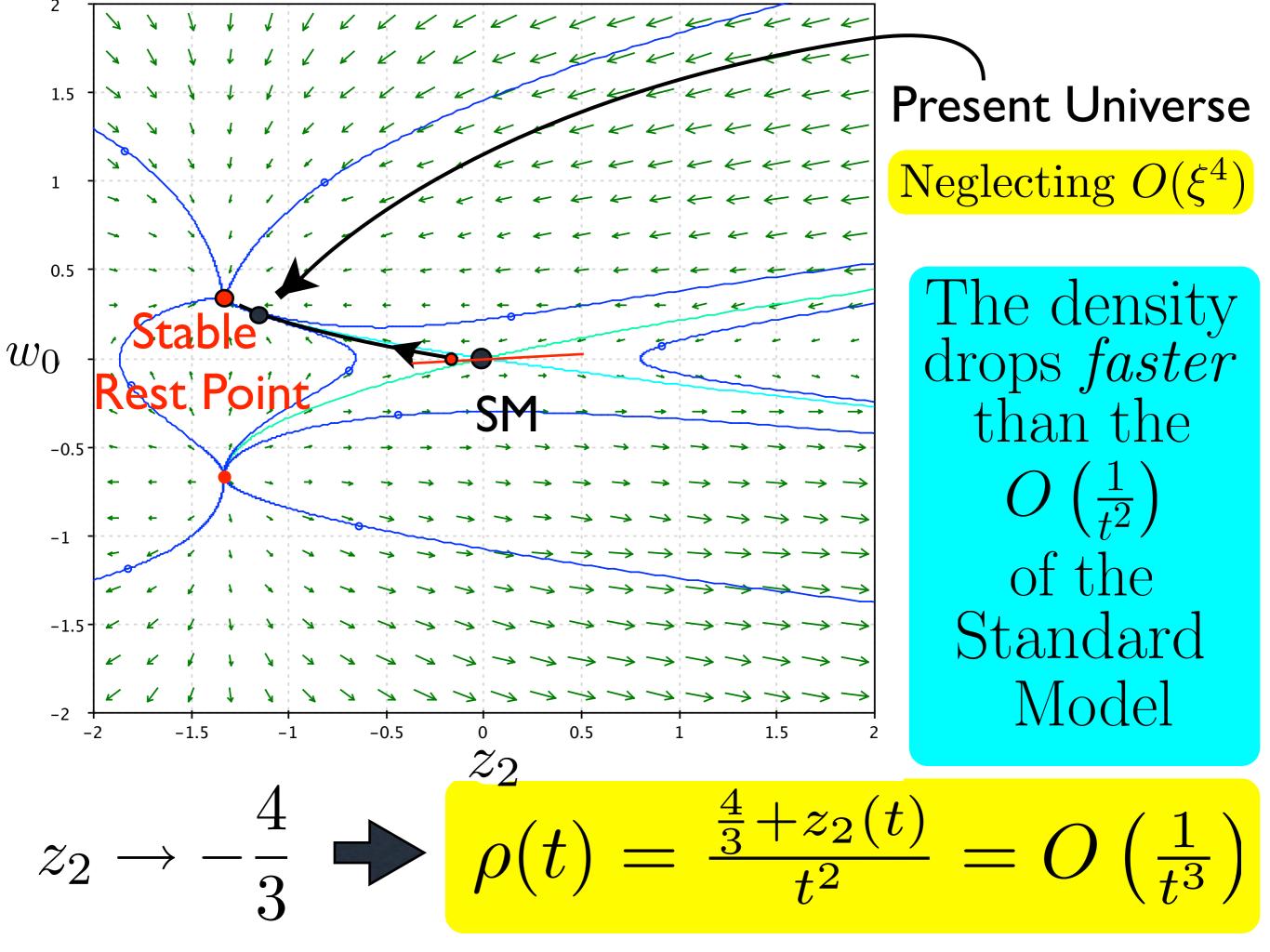
Neglecting  $O(\xi^4)$ 





 $ds^2 = -dt^2 + dr^2 + r^2 d\Omega^2$ 





#### Neglecting $O(\xi^4)$ errors: The spacetime near the center evolves toward the Stable Rest Point

- The metric tends to Flat Minkowski Spacetime which is not co-moving with the fluid
- BUT: The evolution creates a uniformly expanding density near the center, which, neglecting relativistic corrections,

is CENTER-INDEPENDENT

(like Friedmann Spacetimes)

**THEOREM:** Neglecting  $O(\xi^4)$ , as the orbit tends to the

Stable Rest Point, the density drops *FASTER* than SM,

$$\rho(t) = \frac{k_0}{t^{3(1+\bar{w})}},$$

$$\rho_{SM}(t) = \frac{4}{3t^2},$$

where  $\bar{w}(t)$  and  $k_0(t)$  change exponentially slowly.

**CONCLUDE:** The wave creates a

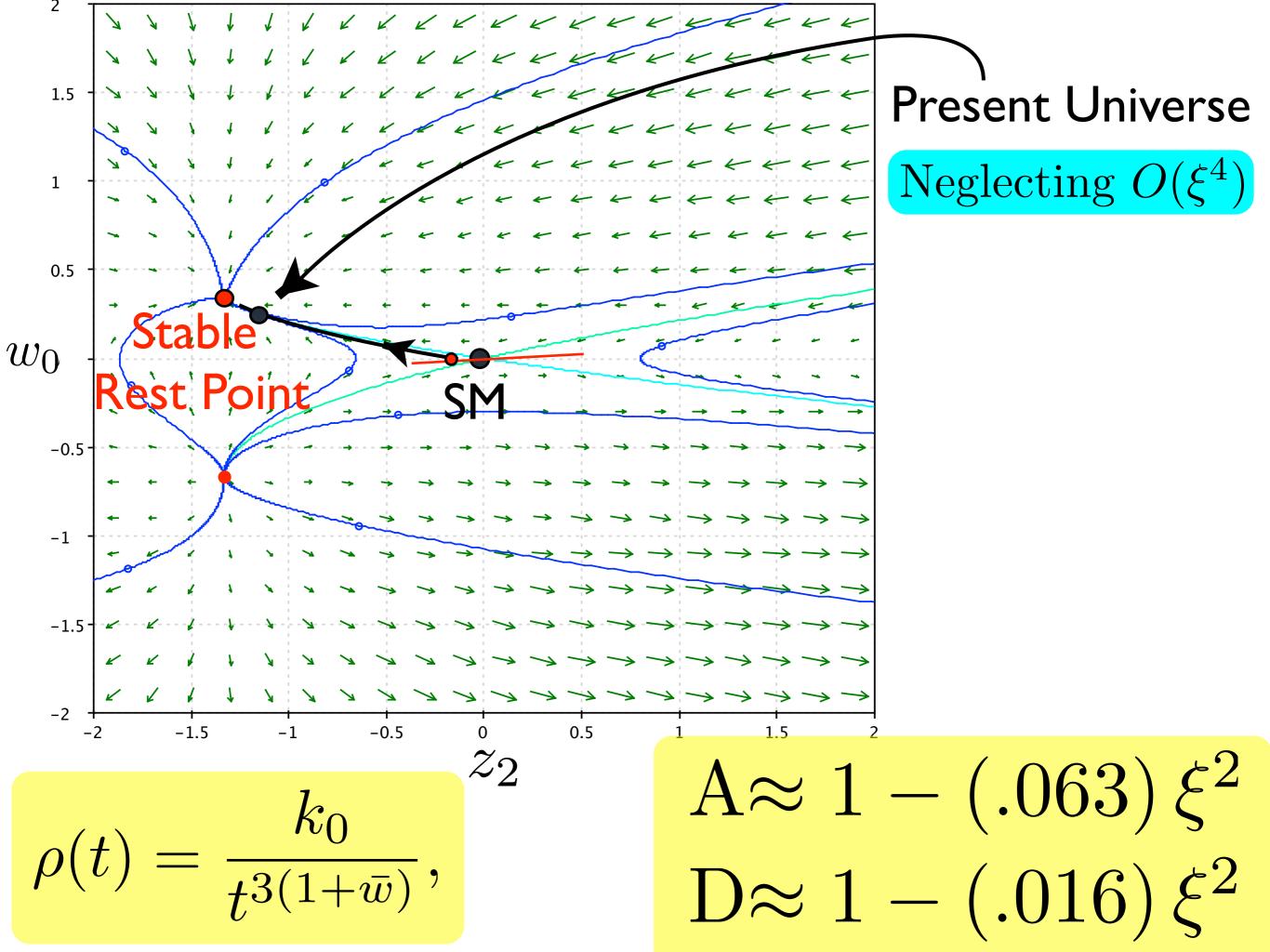
#### UNIFORMLY EXPANDING SPACETIME

with an

### ANOMALOUS ACCELERATION in a

#### LARGE, FLAT, CENTER-INDEPENDENT

region near in the center of the wave.



# 7. The Universality of the Phase Portrait

A radially symmetric function f(r) is a smooth function in Euclidean coordinates x at r=0 if and only if

$$g(x) = f(|x|)$$

is a smooth function of x at x = 0.

#### Equating the n'th derivative of g(x)from left and right at r = 0 gives

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Theorem: f is smooth iff odd derivatives vanish, i.e.,

$$f(r) = f(0) + f_2 r^2 + f_4 r^4 + \cdots$$

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...only even powers of r.

#### Consider now a metric in SSC:

$$ds^{2} = -B(t,r)dt^{2} + \frac{1}{A(t,r)}dr^{2} + r^{2}d\Omega^{2}$$

#### Along radial geodesics at fixed time:

$$\frac{dr}{ds} = A(t,r)$$

## Thus r is smooth with respect to arc-length at r=0 if and only if

$$A = 1 + A_2(t)r^2 + A_4(t)r^4 + \cdots$$

Since 
$$\xi = \frac{r}{t} = 0$$
 at  $r = 0$ ,

#### the smoothness condition is

$$A = 1 + A_2(t)\xi^2 + A_4(t)\xi^4 + \cdots$$

Since 
$$\xi = \frac{r}{t} = 0$$
 at  $r = 0$ ,

#### the smoothness condition is

$$A = 1 + A_2(t)\xi^2 + A_4(t)\xi^4 + \cdots$$

Conclude: our ansatz is just expressing smoothness at r=0 in SSC coordinates.

## Our ansatz expresses smoothness at r=0 in SSC coordinates...

$$z(t,\xi) = z_{F}(\xi) + \Delta z(t,\xi) \qquad \Delta z = z_{2}(t)\xi^{2} + z_{4}(t)\xi^{4}$$

$$w(t,\xi) = w_{F}(\xi) + \Delta w(t,\xi) \qquad \Delta w = w_{0}(t) + w_{2}(t)\xi^{2}$$

$$A(t,\xi) = A_{F}(\xi) + \Delta A(t,\xi) \qquad \Delta A = A_{2}(t)\xi^{2} + A_{4}(t)\xi^{4}$$

$$D(t,\xi) = D_{F}(\xi) + \Delta D(t,\xi) \qquad \Delta D = D_{2}(t)\xi^{2}$$

$$z_F = \frac{4}{3}\xi^2 + \frac{40}{27}\xi^4 + O(\xi^6)$$

$$w_F = \frac{2}{3} + \frac{2}{9}\xi^2 + O(\xi^4)$$

$$A_F = 1 - \frac{4}{9}\xi^2 - \frac{8}{27}\xi^4 + O(\xi^6)$$

$$D_F = 1 - \frac{1}{9}\xi^2 + O(\xi^4)$$

## Thus our phase portrait applies to any SSC solution of the Einstein equations that is smooth at r=0...

#### Thus our phase portrait applies to any SSC p=0 solution of the Einstein equations that is smooth at r=0... Stable Rest $w_0$ ( **Parameter** Family of Unstable Rest Point a-waves,

Lematre-Tolman-Bondi (LTB) coordinates are used in other under-density models...

In LTB the radial coordinate is co-moving with the fluid...

Transforming from SSC to LTB introduces a coordinate singularity at r=0...

**Lemma** Assume that  $\rho(t,r)$  is a scalar density function which extends to a smooth function  $\rho(t,|x|)$  in SSC coordinates, so that it is given near r=0 by

$$\rho(t,r) = f_0(t) + f_2(t)r^2 + \cdots$$

Let

$$\hat{\rho}(\hat{t}, \hat{r}) = \rho(t(\hat{t}, \hat{r}), r(\hat{t}, \hat{r}))$$

denote the representation of the function  $\rho(t, r)$  in LTB coordinates. Then the third partial derivative of  $\hat{\rho}$  with respect to  $\hat{r}$  at  $(\hat{t}, 0)$  is given by

$$\frac{\partial^3 \hat{\rho}}{\partial \hat{r}^3} = \frac{\partial \rho}{\partial t} \frac{\partial^3 t}{\partial \hat{r}^3} + 3 \frac{\partial^2 \rho}{\partial r^2} \frac{\partial r}{\partial \hat{r}} \frac{\partial^2 r}{\partial \hat{r}^2}.$$

Conclude: In LTB coordinates its not so easy to expand about the center because coordinates can be singular with respect to the geometry at r=0...

When we submitted to RSPA, the editors asked us to address a long list of papers on under-density theories based on Lematre-Tolman-Bondi (LTB)

Coordinates...

- . [35] M. Kowalski et al. [Supernova Cosmology Project Collaboration], "Improved cosmological constraints from new, old and combined supernova datasets", Astrophys. J. 686 (2008) 749 [arXiv:0804.4142 [astro-ph]].
- . [36] D. Rubin, E. V. Linder, M. Kowalski, G. Aldering, R. Amanullah, K. Barbary, N. V. Connolly and K. S. Dawson et al., "Looking beyond Lambda with the Union supernova compilation", Astrophys. J. 695 (2009) 391 [arXiv:0807.1108 [astro-ph]].
- . [37] T. Biswas and A. Notari, "Swiss-cheese inhomogeneous cosmology and the dark energy problem", JCAP 0806 (2008) 021 [astro-ph/0702555].
- . [38] V. Marra, E. W. Kolb, S. Matarrese, "Light-cone averages in a swiss-cheese Universe", Phys. Rev. D77 (2008) 023003. [arXiv:0710.5505 [astro-ph]].
- . [39] V. Marra, E. W. Kolb, S. Matarrese, A. Risotto, "On cosmological observables in a swiss-cheese universe", Phys. Rev. D76 (2007) 123004. [arXiv:0708.3622 [astro-ph]].
- . [40] R. A. Vanderveld, E. E. Flanagan, I. Wasserman, "Luminosity distance in Swiss cheese cosmology with randomized voids: I. Single void size", Phys. Rev. D78 (2008) 083511. [arXiv:0808.1080 [astro-ph]].

- . [41] W. Valkenburg, "Swiss Cheese and a Cheesy CMB", JCAP 0906 (2009) 010. [arXiv:0902.4698 [astro-ph.CO]].
- . [42] E. E. Flanagan, N. Kumar, I. Wasserman and R. A. Vanderveld, "Luminosity distance in Swiss cheese cosmology with randomized voids: II. Magnification probability distributions", Phys. Rev. D 85 (2012) 023510 [arXiv:1109.1873 [gr-qc]].
- . [43] D. L. Wiltshire, "Exact solution to the averaging problem in cosmology", Phys. Rev. Lett. 99 (2007) 251101. [arXiv:0709.0732 [gr-qc]].
- . [44] D. L. Wiltshire, "Cosmic clocks, cosmic variance and cosmic averages", New J. Phys. 9, 377 (2007). [gr-qc/0702082].
- . [45] D. L. Wiltshire, "Dark energy without dark energy", [arXiv:0712.3984 [astro-ph]]. (Presented at DARK 2007, Sydney, Australia.)
- . [46] B. M. Leith, S. C. C. Ng, D. L. Wiltshire, "Gravitational energy as dark energy: Concordance of cosmological tests", Astrophys. J. 672 (2008) L91-L94. [arXiv:0709.2535 [astro-ph]].
- . [47] D. L. Wiltshire, "Cosmological equivalence principle and the weak-field limit", Phys. Rev. D78 (2008) 084032. [arXiv:0809.1183 [gr-qc]].
- . [48] D. L. Wiltshire, "Average observational quantities in the timescape cosmology", Phys. Rev. D80 (2009) 123512. [arXiv:0909.0749 [astro-ph.CO]].
- . [49] P. R. Smale, D. L. Wiltshire, "Supernova tests of the timescape cosmology", Mon. Not. Roy. Astron. Soc. 413 (2011) 367-385. [arXiv:1009.5855 [astro-ph.CO]].
- . [50] D. L. Wiltshire, "What is dust? Physical foundations of the averaging problem in cosmology", Class. Quant. Grab. 28 (2011) 164006. [arXiv:1106.1693 [gr-qc]].
- . [51] P. R. Smale, "Gamma ray burst distances and the timescale cosmology", Mon. Not. Roy. Astron. Soc. 418 (2011) 2779 [arXiv:1107.5596 [astro-ph.CO]].

- . [52] D. L. Wiltshire, "Gravitational energy as dark energy: Cosmic structure and apparent acceleration", [arXiv:1102.2045 [astro-ph.CO]]. (Presented at the "Conference on Two Cosmological Models", Mexico DF, 2010.)
- . [53] J. A. G. Duley, M. A. Nazer and D. L. Wiltshire, "Timescape cosmology with radiation fluid", Class. Quant. Grav. 30 (2013) 175006 [arXiv:1306.3208 [astro-ph.CO]].
- . [54] D. L. Wiltshire, "Cosmic structure, averaging and dark energy", arXiv:1311.3787 [astro-ph.CO].
- . [55] M. A. Nazer and D. L. Wiltshire, "Cosmic microwave background anisotropies in the timescape cosmology" arXiv:1410.3470 [astro-ph.CO]. To appear in Physical Review D.
- . [56] J. D. Barrow and J. Stein-Schabes, "Inhomogeneous cosmologies with cosmological constant", Phys. Lett. A 103 (1984) 315.
- . [57] S. Rasanen, "Backreaction in the Lemaitre–Tolman–Bondi model", JCAP 0411 (2004) 010 [gr-qc/0408097].
- . [58] D. J. H. Chung and A. E. Romano, "Mapping luminosity-redshift relationship to LTB cosmology", Phys. Rev. D 74 (2006) 103507 [astro-ph/0608403].
- . [59] A. Paranjape and T. P. Singh, "The possibility of cosmic acceleration via spatial averaging in Lemaitre–Tolman–Bondi Models", Class. Quant. Grav. 23 (2006) 6955 [astro-ph/0605195].
- . [60] R. A. Vanderveld, E. E. Flanagan, I. Wasserman, "Mimicking dark energy with Lemaitre–Tolman–Bondi models: Weak central singularities and critical points", Phys. Rev. D74 (2006) 023506. [astro-ph/0602476].
- . [61] S. Alexander, T. Biswas, A. Notary and D. Vaid, "Local void vs dark energy: Confrontation with WMAP and type Ia supernovae", JCAP 0909 (2009) 025 [arXiv:0712.0370 [astro-ph]].
- . [62] J. Garcia-Bellido, T. Haugboelle, "Confronting Lemaitre-Tolman-Bondi models with observational cosmology", JCAP 0804 (2008) 003. [arXiv:0802.1523 [astro-ph]].

In LTB coordinates its not so easy to expand about the center because coordinates can be singular with respect to the geometry at r=0...

### **CONCLUSIONS:**

Our Proposal: The AA is due to a local underdensity on the scale of the supernova data, created by a self-similar wave from the radiation epoch that triggers an instability in the SM when the pressure drops to zero.

We have made no assumptions regarding the space-time far from the center of the perturbations. The consistency of this model with other observations in astrophysics would require additional assumptions.

### **CONCLUSIONS:**

- This is arguably the simplest explanation for the anomalous acceleration within Einstein's original theory of GR, without requiring Dark Energy.
- It demonstrates that any local center of the Standard Model of Cosmology is unstable to perturbation by exact solutions from the Radiation Epoch.
- These perturbations are stabilized by a nearby stable rest point that generates the same accelerations as Dark Energy.
- It makes testable predictions.

### **QUESTIONS:**

- On what scale would such waves apply?
- If these came from time-asymptotic wave patterns created in an earlier epoch, would we expect secondary transitional waves far from the center?
- How does cosmology address the instability?
   Can Dark Energy help? (NO!)
- Implications of a preferred center?
- Is this more fine-tuned than Dark Energy?

### **QUESTIONS:**

- Was it reasonable to expect to observe the redshift vs luminosity relation of the SM if its unstable looking outward from any center? (Aren't unstable solutions usually considered un-observable in Physics?)
- Given that that phase portrait applies to any smooth spherical perturbation, shouldn't we expect to observe an anomalous acceleration in nearby galaxies?

We reiterate: The purpose of our paper is not to solve all the problems of Cosmology in one grand solution. Rather, our purpose is to introduce and deconstruct a new instability in the Friedmann space-time of the Standard Model of Cosmology, to identify mechanisms that trigger it, to show how it naturally could account for the anomalous acceleration within Einstein's original theory without Dark Energy, and then to derive new predictions from it.

### Prokopek...2013 (Astrophysicist, Utrecht University) There are large scale anomalies in the data indicating a lack of uniformity on the largest length scale

The main large angular scale anomalies are [4, 5]:

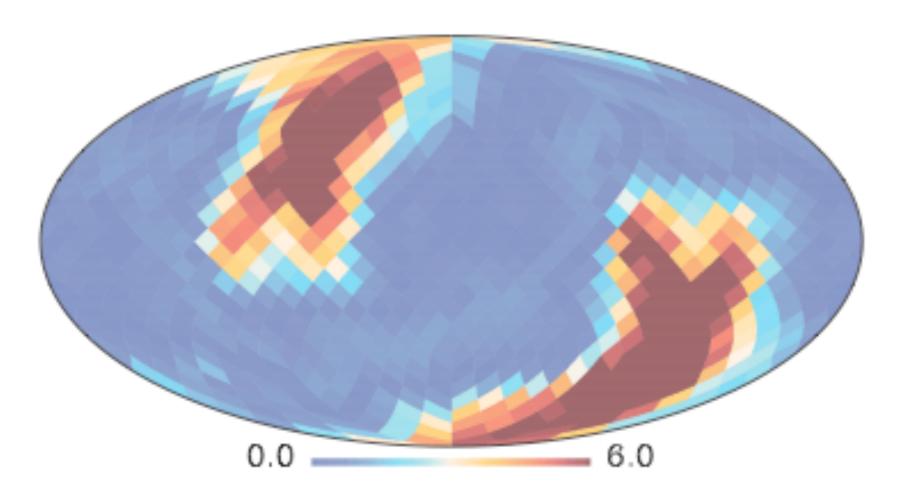
a high quadrupole-octupole alignment (if accidental, it would occur in about 3% cases);

5

- a low variance in the lower galactic ecliptic plane and a low skewness in the southern plane;
- a northern/southern ecliptic hemisphere asymmetry (the northern hemisphere correlation function is featureless and lacks power on large angular scales);
- phase correlations on large angular scales shown in figure 2, whose significance is more than three standard deviations and which imply that there are non-Gaussian features on large angular scales;

### Prokopek...2013 (Astrophysicist, Utrecht University)

- a dipolar asymmetry, which includes a dipolar modulation and a dipolar power asymmetry;
- a parity asymmetry (which is related to the dipolar modulation) that comes in two disguises:
   a parity reflection asymmetry and a mirror asymmetry, both of which show significant statistical evidence for low multipoles;
- a very cold spot (on angular scale of about 5° with significance of more than four standard deviations);
- a lack of power on one hemisphere on angular scales corresponding to the multipoles ℓ ∈ [5, 25]
   that has statistical significance of almost three standard deviations.



FINAL COMMENT

# Every aspect of this work came from Applied Mathematics,

Whatever its implications to Physics, it stands on its own as a self-contained model in Applied Mathematics

# Mathematics is part of physics... ...the part of physics where experiments are cheap.

-Arnold, Paris, 1997

## 

### WORLD VIEW A personal take on events



#### Big Bang blunder bursts the multiverse bubble

Premature hype over gravitational waves highlights gaping holes in models for the origins and evolution of the Universe, argues **Paul Steinhardt**.

Then a team of cosmologists announced at a press conference in March that they had detected gravitational waves generated in the first instants after the Big Bang, the origins of the Universe were once again major news. The reported discovery created a worldwide sensation in the scientific community, the media and the public at large (see *Nature* **507**, 281–283; 2014).

According to the team at the BICEP2 South Pole telescope, the detection is at the 5–7 sigma level, so there is less than one chance in two million of it being a random occurrence. The results were hailed as proof of the Big Bang inflationary theory and its progeny, the multiverse. Nobel prizes were predicted and scores of theoretical models spawned. The announcement also influenced decisions about

academic appointments and the rejections of papers and grants. It even had a role in governmental planning of large-scale projects.

The BICEP2 team identified a twisty (B-mode) pattern in its maps of polarization of the cosmic microwave background, concluding that this was a detection of primordial gravitational waves. Now, serious flaws in the analysis have been revealed that transform the sure detection into no detection. The search for gravitational waves must begin anew. The problem is that other effects, including light scattering from dust and the synchrotron radiation generated by electrons moving around galactic magnetic fields within our own Galaxy, can also produce these twists.

The BICEP2 instrument detects radiation at only one frequency, so cannot distinguish the cos-

mic contribution from other sources. To do so, the BICEP2 team used measurements of galactic dust collected by the Wilkinson Microwave Anisotropy Probe and Planck satellites, each of which operates over a range of other frequencies. When the BICEP2 team did its analysis, the Planck dust map had not yet been published, so the team extracted data from a preliminary map that had been presented several months earlier. Now a careful reanalysis by scientists at Princeton University and the Institute for Advanced Study, also in Princeton, has concluded that the BICEP2 B-mode pattern could be the result mostly or entirely of foreground effects without any contribution from gravitational waves. Other dust models considered by the BICEP2 team do not change this negative conclusion, the Princeton team showed (R. Flauger, J. C. Hill and D. N. Spergel, preprint at http://arxiv.org/abs/1405.7351; 2014).

The sudden reversal should make the scientific community contemplate the implications for the future of cosmology experimentation

and theory. The search for gravitational waves is not stymied. At least eight experiments, including BICEP3, the Keck Array and Planck, are already aiming at the same goal.

This time, the teams can be assured that the

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world will be paying close attention. This time, acceptance will require measurements over a range of frequencies to discriminate from foreground effects, as well as tests to rule out other sources of confusion. And this time, the announcements should be made after submission to journals and vetting by expert referees. If there must be a press conference, hopefully the scientific community and the media will demand that it is accompanied by a complete set of documents, including details of the systematic analysis and sufficient data to enable objective verification.

The BICEP2 incident has also revealed a truth about inflationary theory. The common view is that it is a highly predictive theory. If that was the case and the detection of gravitational waves was the 'smoking gun' proof of inflation, one would think that non-detection means that the

theory fails. Such is the nature of normal science. Yet some proponents of inflation who celebrated the BICEP2 announcement already insist that the theory is equally valid whether or not gravitational waves are detected. How is this possible?

The answer given by proponents is alarming: the inflationary paradigm is so flexible that it is immune to experimental and observational tests. First, inflation is driven by a hypothetical scalar field, the inflaton, which has properties that can be adjusted to produce effectively any outcome. Second, inflation does not end with a universe with uniform properties, but almost inevitably leads to a multiverse with an infinite number of bubbles, in which the cosmic and physical properties vary from bubble to bubble. The part of the multiverse that we observe corresponds to a piece

of just one such bubble. Scanning over all possible bubbles in the multiverse, everything that can physically happen does happen an infinite number of times. No experiment can rule out a theory that allows for all possible outcomes. Hence, the paradigm of inflation is unfalsifiable.

This may seem confusing given the hundreds of theoretical papers on the predictions of this or that inflationary model. What these papers typically fail to acknowledge is that they ignore the multiverse and that, even with this unjustified choice, there exists a spectrum of other models which produce all manner of diverse cosmological outcomes. Taking this into account, it is clear that the inflationary paradigm is fundamentally untestable, and hence scientifically meaningless.

Cosmology is an extraordinary science at an extraordinary time. Advances, including the search for gravitational waves, will continue to be made and it will be exciting to see what is discovered in the coming years. With these future results in hand, the challenge for theorists will be to identify a truly explanatory and predictive scientific paradigm describing the origin, evolution and future of the Universe.

**Paul Steinhardt** is professor of physics at Princeton University. e-mail: steinh@princeton.edu

THE INFLATIONARY

PARADIGM IS

**FUNDAMENTALLY** 

AND HENCE

**SCIENTIFICALLY**