

# Flux Quantization

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## Abstract

Flux- and charge-quantization laws for higher gauge fields of Maxwell type — e.g. the common electromagnetic field (the “A-field”) but also the B-, RR-, and C-fields considered in string/M-theory — specify non-perturbative completions of these fields by encoding their solitonic behaviour and hence by specifying the discrete charges carried by the individual branes (the higher-dimensional monopoles or solitons) that source the field fluxes.

This article surveys the general (rational-)homotopy theoretic understanding of flux- and charge-quantization via the Chern-Dold character map generalized to the non-linear (self-sourcing) Bianchi identities that appear in higher-dimensional supergravity theories, notably for B-&RR-fields in  $D = 10$ , for the C-field in  $D = 11$  supergravity, and for the B-field on fivebrane worldvolumes.

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**History** (see more references below).

In 1852 Faraday observes magnetic field flux lines emanating from magnetic poles [Faraday 1852], cf. §2.1 below.

In 1931 Dirac invokes quantum mechanics to argue that, if there were unpaired such (mono-)poles, then the total flux emanating from them — and thus the magnetic charge carried by them — had to come in integer multiples of a unit quantum [Dirac 1931], cf. Ex. 3.9 below.

In 1957 Abrikosov essentially notices that the same electromagnetic flux-&charge-quantization mechanism makes vortex strings in type II superconductors carry units of localized magnetic flux [Abrikosov 1957], cf. p. 6 below.

In 1985 Alvarez understands such solitonic magnetic fields as 2-cocycles in (differential) ordinary cohomology [Alvarez 1985].

In 1988 Gawedzki observes that the B-field flux felt by a string, hence the NS5-brane charge, must similarly be quantized as a 3-cocycle in Deligne cohomology [Gawedzki 1988] [Freed & Witten 1999, §6], cf. Ex. 3.10 below.

In the 1990s, string theorists hypothesize that the flux of RR-fields and hence the charge of D-branes is analogously quantized in a generalized cohomology theory called “topological K-theory” [Minasian & Moore 1997] [Witten 1998]; or more generally in “twisted” such K-theory [Bouwknegt & Mathai 2001], cf. §4.1 below;

and that the flux of the C-field and hence the charge of M-branes is quantized in a “shifted half-integral” cohomology theory [Witten 1997] whose proper mathematical home motivates [Hopkins & Singer 2005] but for a long time remains somewhat mysterious, cf. (47) below.

In the 2020s [Fiorenza et al. 2020] develop a systematic understanding of flux quantization of any higher gauge theory (of Maxwell-type, Def. 2.6 below) in generalized non-abelian cohomology theory, using tools from dg-algebraic rational homotopy theory related to the “FDA”-method in the supergravity literature [Fiorenza et al. 2023], cf. §3 below.

This gives transparent re-derivation of previous flux quantization laws and allows to discuss C-field flux- and M-brane charge-quantization, cf. §4.2 below.

## 1 Overview

**(§2)** A *higher gauge theory* (for review in this volume see [Alfonsi 2024, §2][Borsten et al. 2024], but beware Rem. 1.1 below) of Maxwell-type (Def. 2.6 below) is a (quantum) field theory analogous to vacuum-electromagnetism (on curved spacetimes), but with the analog of the electromagnetic flux density  $F_2$  (which ordinarily is a differential 2-form on 3+1 dimensional spacetime  $X^4$ ) allowed to be a system of differential forms  $\vec{F} = \{F^{(i)}\}_{i \in I}$  of any degree  $\deg_i \geq 1$  on a  $D$ -dimensional spacetime  $X^D$  of any dimension  $D = d + 1 \geq 2$ , and satisfying a higher analog of Maxwell’s equations (6). Such higher gauge theories famously appear as the gauge field-sector in higher-dimensional supergravity (e.g. [Castellani et al. 1991] [Tanii 1996] [de Wit & Louis 1999] [Sezgin 2023]) and hence in super-string/M-theory (e.g. [Duff 1999] [Blumenhagen et al. 2013]), which motivates their deeper investigation.

Like for ordinary Maxwell theory, where one may think of singularities or stable bumps in the electromagnetic flux density  $F_2$  as being sourced by charges carried by (hypothetical) Dirac monopoles or by (observed) Abrikosov vortex strings, respectively (cf. §2.1), so one may think of singularities or stable bumps in these higher flux densities as sourced by singular branes or solitonic branes, respectively (cf. §2.2), for suitably higher dimensional (mem-)branes carrying suitable higher charges.

**(§3)** But for such singular/solitonic branes to be “elementary” objects of individually discernible nature, their total charges, and hence the total fluxes which they source, should take discrete (“quantized”) values (as indeed observed for Abrikosov vortex strings). This is what *flux quantization* is about.

Traditionally one declares the full higher gauge field to be given by gauge potentials  $\widehat{A}$  whose curvature is the flux density  $\vec{F}$  and which is globally subjected to a topological condition that implies the flux and charge quantization. While this is classical for electromagnetism (and Yang-Mills theory), it is not transparent from this point of view how to identify the structure of gauge potentials for general higher gauge theories.

More systematically, one may understand flux quantization as the specification of a generalized (non-abelian) cohomology theory for which the charges are required to be cocycles, and of which the total fluxes are then the differential-geometric (Chern-Dold-) *characters*. From this streamlined point of view the higher gauge potential, and hence the full field content of the higher gauge theory, arises as the homotopy/gauge theoretic witness of the matching of total fluxes with the character of the charges, making the full higher gauge fields be cocycles in a corresponding generalized (nonabelian) *differential cohomology* theory.

**(§4)** Examples of flux quantization beyond Dirac charge quantization in electromagnetism play a key role in string/M-theory:

- the seminal “Hypothesis K” postulating D-brane charge quantization in K-theory,
- the more recent “Hypothesis H” postulating M-brane charge quantization in unstable twisted Cohomotopy.

These hypotheses are not unrelated: Under “double dimensional Kaluza-Klein reduction” along a circle fiber, the C-field fluxes on  $X^{10+1}$  are to give rise to most of the B&RR-field-fluxes in  $X^{9+1}$  (as part of the duality between M-theory and type IIA string theory, which is either conjectural or defining, depending on attitude towards the definition of the elusive M-theory). Therefore Hypothesis H may be understood as providing a non-perturbative/M-theoretic lift of Hypothesis K to the extent that it reduces to the latter under double dimensional reduction, and as a non-perturbative/M-theoretic correction to the extent that it does not quite reduce to Hypothesis K.

**Perspective.** This highlights that a *choice* of flux quantization is (depending on perspective of how that higher gauge theory is ultimately defined): (i) a *hypothesis* about or else (ii) a *specification* of the *non-perturbative completion* of the given higher gauge theory, which is generally an issue that deserves (more) attention.

Traditionally, flux quantization laws have been postulated sporadically and in ad-hoc fashion, in order to patch up “anomalous” theories: Since the ancient past it has been common to define physical theories by stationary action principles embodied by Lagrangian densities, from which perturbative BRST complexes are extracted, whose quantization (e.g. [Henneaux et al. 1992]) is generally afflicted with problems (“anomalies”) some of which are dealt with by ad-hoc flux quantization: For example the original Dirac charge quantization was postulated to cure an anomaly in the quantum theory of an electron propagating in the background field of a magnetic monopole (a “0-brane”), while the enigmatic shifted C-field flux quantization similarly serves to cure an anomaly in the quantum theory of the M2-brane propagating in the background field of an M5-brane ([Witten 1997], §2.2).

More systematically, the available choices of flux-quantization laws  $\mathcal{A}$  are algebro-topologically determined by the form of the higher Gauss law on any Cauchy surface, and any such choice, given by a compatible non-abelian cohomology-theory, determines the non-perturbative phase space stack of flux-quantized gauge fields. This process makes no reference to Lagrangian densities and applies seamlessly to field theories that do not even have a natural Lagrangian description, such as self-dual higher gauge theories.

Typically there is an “evident” choice of flux quantization and this is the choice tacitly made in the literature, where considered at all. But it is important to notice that there are other admissible choices, embodying hypotheses about (or definitions of) non-evident nonperturbative completions of the given higher gauge theory.

**The logic of flux quantization.** The table on p. 4 shows in outline the logic of algebro-topological flux quantization as reviewed here; on the left in generality and on the right for our **running examples**:

1. traditional Dirac charge quantization of the electromagnetic field (experimentally well-supported);
2. traditional D-brane charge quantization in twisted topological K-theory (“Hypothesis K”);
3. more recent M-brane charge quantization in unstable twisted Cohomotopy (“Hypothesis H”).

**The role of  $L_\infty$ -algebras.** As the table on p. 4 indicates, the algebro-topological nature of flux&charge quantization is higher Lie-theoretic (explained in §3), by matching two  $L_\infty$ -algebras associated with a given higher gauge theory of Maxwell type (§2):

(i) **Bianchi-Gauß  $L_\infty$ -algebras.** The *higher Bianchi identities* on duality-symmetric higher fluxes (Def. 2.6 below) and hence their *higher Gauss law* (Prop. 2.14 below) are equivalent to the condition that the flux densities jointly constitute a closed  $L_\infty$ -algebra valued differential form with coefficients in a characteristic  $L_\infty$ -algebra  $\mathfrak{a}$  (Prop. 3.1 below):

$$\begin{array}{l} \text{Space of flux densities} \\ \text{on spacetime, solving} \\ \text{the equations of motion} \end{array} \text{SolSpace}(X^D) \equiv \left\{ \begin{array}{l} \text{electromagnetic flux densities on spacetime} \\ \vec{F} \equiv \left( F^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^D) \right)_{i \in I} \end{array} \left| \begin{array}{l} \text{Bianchi identities} \\ d\vec{F} = \vec{P}(\vec{F}) \\ \star \vec{F} = \vec{\mu}(\vec{F}) \\ \text{self-duality} \end{array} \right. \right\} \text{covariant form}$$

$$\simeq \left\{ \begin{array}{l} \text{magnetic flux densities on Cauchy surface} \\ \vec{B} \equiv \left( B^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^d) \right)_{i \in I} \end{array} \left| \begin{array}{l} \text{Gauß law} \\ d\vec{B} = \vec{P}(\vec{B}) \end{array} \right. \right\} \text{canonical form} \simeq \Omega_{\text{dR}}^1(X^d; \mathfrak{a})_{\text{clsd}} \quad \begin{array}{l} \text{space of closed (flat)} \\ \mathfrak{a}\text{-valued} \\ \text{differential forms} \end{array}$$

(ii) **Whitehead  $L_\infty$ -algebras.** The classifying space  $\mathcal{A}$  of any charge quantization law is rationally characterized by its rational Whitehead  $L_\infty$ -algebra  $\mathfrak{L}\mathcal{A}$  (the “Koszul-dual” of  $\mathcal{A}$ : that  $L_\infty$ -algebra whose Chevalley-Eilenberg algebra  $\text{CE}(\mathfrak{L}\mathcal{A})$  is the Sullivan model of  $\mathcal{A}$ ) and the nonabelian Chern-Dold character map extracts from  $\mathcal{A}$ -cohomology its image in  $\mathfrak{L}\mathcal{A}$ -valued nonabelian de Rham cohomology (35):

$$\begin{array}{ccccccc} & & \xrightarrow{\text{character map on } \mathcal{A}\text{-cohomology}} & & & & \\ & & \downarrow & & & & \\ H^1(X; \Omega\mathcal{A}) & \xrightarrow{\text{rationalization}} & H^1(X; L^{\mathbb{Q}}\Omega\mathcal{A}) & \xrightarrow{\text{extension of scalars}} & H^1(X; L^{\mathbb{R}}\Omega\mathcal{A}) & \xrightarrow{\text{nonabelian de Rham theorem}} & H^1_{\text{dR}}(X; \mathfrak{L}\mathcal{A}) \\ \parallel & & \parallel & & \parallel & & \parallel \\ \pi_0 \text{Map}(X, \mathcal{A}) & \xrightarrow{(\eta_{\mathcal{A}}^{\mathbb{Q}})_*} & \pi_0 \text{Map}(X, L^{\mathbb{Q}}\mathcal{A}) & \xrightarrow{(\eta_{L^{\mathbb{Q}}\mathcal{A}}^{\text{ext}})_*} & \pi_0 \text{Map}(X, L^{\mathbb{R}}\mathcal{A}) & \xrightarrow{\sim} & \text{Hom}_{\text{dgAlg}}(\text{CE}(\mathfrak{L}\mathcal{A}), \Omega_{\text{dR}}^{\bullet}(X))_{\text{cnrd}} \\ & & & & \text{fundamental theorem} & & \\ & & & & \text{of dg-algebraic RHT} & & \end{array}$$

| Logic of flux quantization |   | Higher gauge theory of Maxwell-type  | A-field in $D = 4$   | B&RR-field in $D = 10$   | C-field in $D = 11$                               |
|----------------------------|---|--|--|--|---|
| §2                         | Flux densities  | $\vec{F} \equiv \left( F^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^D) \right)_{i \in I}$   | $F_2$ magnetic<br>$G_2$ electric   | $H_3$ NS5<br>$H_7$ F1<br>$F_{2\bullet}$ $D_{8-2\bullet}$ flux densities on | $G_4$ M5<br>$G_7$ M2                              |
|                            | Self-duality  | $\star \vec{F} = \vec{\mu}(\vec{F})$   | $\star F_2 = G_2$  | $\star H_3 = H_7$<br>$\star F_{2\bullet} = F_{10-2\bullet}$                | $\star G_4 = G_7$                                 |
|                            | Bianchi identities  | $d\vec{F} = \vec{P}(\vec{F})$  | $dF_2 = 0$<br>$dG_2 = 0$   | $dH_3 = 0$<br>$dH_7 = 0$<br>$dF_{2\bullet} = H_3 \wedge F_{2\bullet-2}$    | $dG_4 = 0$<br>$dG_7 = -\frac{1}{2}G_4 \wedge G_4$ |
| §3                         | CE-algebra of characteristic $L_\infty$ -algebra                | $\text{CE}(\mathfrak{a}) \equiv \mathbb{R}[\vec{b}] / (d\vec{b} = \vec{P}(\vec{b}))$   | $df_2 = 0$<br>$dg_2 = 0$   | $dh_3 = 0$<br>$dh_7 = 0$<br>$df_{2\bullet} = h_3 \wedge f_{2\bullet-2}$    | $dg_4 = 0$<br>$dg_7 = -\frac{1}{2}g_4 \wedge g_4$ |
|                            | Solution space of fluxes on $X^D = \mathbb{R}^{0,1} \times X^d$ | Gauß law = $\mathfrak{a}$ -closedness<br>$\Omega_{\text{dR}}(X^d; \mathfrak{a})_{\text{clsd}} \equiv \text{Hom}(\text{CE}(\mathfrak{a}), \Omega_{\text{dR}}(X^d))$ | $\Omega_{\text{dR}}^2(X^d)_{\text{clsd}} \times \Omega_{\text{dR}}^2(X^d)_{\text{clsd}}$<br>can. momenta | 3-twisted de Rham cocycles   | “4-twisted” de Rham cocycles                      |
|                            | Characteristic $L_\infty$ -algebra                              | $\mathfrak{a}$   | $bu(1) \oplus bu(1)$   | $[b_2, v_{2\bullet-1}] = v_{2\bullet+1}$                                   | $[v_3, v_3] = v_6$<br>M-theory gauge algebra      |
|                            | as rational Whitehead $L_\infty$ -algebra                       | $\mathfrak{a} \simeq \mathfrak{LA}$  | $\mathfrak{l}(B^2\mathbb{Z} \times B^2\mathbb{Z})$   | $\mathfrak{l}((\text{KU}_0 // B^2\mathbb{Z}) \times B^7\mathbb{Z})$        | $\mathfrak{l}(S^4)$                               |
| §4                         | Evident choice of classifying space                             | $\mathcal{A}$  | $B^2\mathbb{Z} \times B^2\mathbb{Z}$<br>Dirac's hypothesis   | $(\text{KU}_0 // B^2\mathbb{Z}) \times B^7\mathbb{Z}$<br>Hypothesis K      | $S^4$<br>Hypothesis H                             |
|                            | Corresponding cohomology theory                                 | generalized cohomology   | ordinary cohomology  | twisted K-theory   | unstable CoHomotopy                               |
|                            | Flux-quantized phase space                                      | $\Omega_{\text{dR}}(X^d; \mathfrak{a})_{\text{clsd}} \times \mathcal{A}(X^d)_{L^\infty \mathcal{A}(X^d)}$  | differential cohomology  | differential twisted K-theory  | differential CoHomotopy                           |

The admissible flux quantization laws for a higher gauge theory with Bianchi-Gauß  $L_\infty$ -algebra  $\mathfrak{a}$  are hence those classified by spaces  $\mathcal{A}$  with Whitehead  $L_\infty$ -algebra  $\mathfrak{LA} \simeq \mathfrak{a}$ . Given such a choice, then quantizing a flux density  $\vec{B}$  globally is to lift its  $\mathfrak{a}$ -valued de Rham-class to a class in  $\mathcal{A}$ -valued non-abelian cohomology.

$$\begin{array}{c}
 \begin{array}{ccc}
 \text{charge quantum in } \mathcal{A}\text{-cohomology} & \nearrow & H^1(X^d; \Omega \mathcal{A}) \\
 & & \downarrow \text{ch}_{\mathcal{A}}(X^d) \\
 \text{flux densities on Cauchy surface} & \xrightarrow{\text{total flux}} & H_{\text{dR}}^1(X^d; \mathfrak{a}) \\
 \text{satisfying their higher Gauß law} & & \downarrow \text{sourced flux} \\
 & & \text{a-valued de Rham cohomology}
 \end{array} \\
 \text{choice of } \mathcal{A}\text{-cohomology with } \mathfrak{LA} \simeq \mathfrak{a}
 \end{array}$$

Locally, a *flux-quantized higher gauge field* is (i) a flux density  $\vec{B}$  being a cocycle in  $\mathfrak{a}$ -de Rham cohomology, (ii) a charge  $\chi$  being a cocycle in  $\mathcal{A}$ -cohomology and (iii) a *gauge potential*  $\hat{A}$  being a coboundary between their joint images (thus exhibiting the above identification of their cohomology classes). This makes the flux-quantized higher gauge fields be cocycles in *differential*  $\mathcal{A}$ -cohomology.

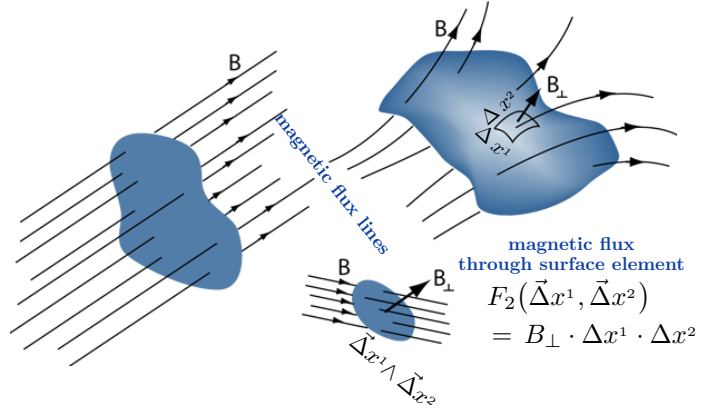
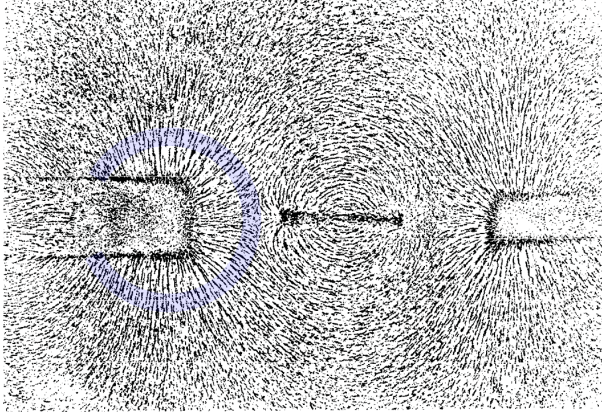
$$\begin{array}{ccc}
 \hat{X}^d & \xrightarrow{\text{charges}} & \mathcal{A} \\
 \downarrow \text{flux densities } \vec{B} & \swarrow \hat{A} & \downarrow \text{ch}_{\mathcal{A}} \\
 \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} & \xrightarrow[\text{shape unit } \eta^f]{} & \int \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} \\
 & & \downarrow \text{differential character}
 \end{array}$$

**Remark 1.1** ( $L_\infty$ -algebra of gauge potentials vs  $L_\infty$ -algebra of flux densities). This means that the  $L_\infty$ -algebras of concern here are *not* the coefficients of the gauge potentials as familiar from Yang-Mills theory, but serve as coefficients for their flux densities (field strengths). Even for higher U(1)-gauge theories (Ex. 3.10) these differ by a degree shift. For RR-fields flux-quantized in K-theory it is mathematically a “coincidence” that  $\mathfrak{su}(n)$ -valued gauge potentials present differential K-theory classes; while for the C-field it remains unclear whether  $\mathfrak{e}_8$ -valued gauge potentials play an analogous role (cf. §4.4 and footnote <sup>4</sup> below). In general, there is no reason to expect that flux-quantized higher gauge potentials are  $L_\infty$ -algebra valued at all.

## 2 Flux Densities and Brane Charges

### 2.1 Electromagnetic flux and its “branes”

Faraday observed “lines of force” – now called flux of the magnetic field – concentrating towards the poles of rod magnets. In modern differential-geometric formulation, the density of these flux lines through any given surface-element is encoded in a differential 2-form  $F_2$ :



From Faraday’s *Diary of experimental investigation*, vol VI, entry from 11th Dec. 1851, as reproduced in [Martin09]; the colored arc is our addition, for ease of comparison with the schematics on the right.

The density and orientation of magnetic field flux lines are encoded in a differential 2-form whose integral over a given surface is proportional to the total magnetic flux through that surface. (Graphics adapted from [Hyperphysics].)

More in detail, with respect to any foliation  $X^4 \simeq \mathbb{R} \times X^3$  of a globally hyperbolic spacetime  $X^4$  by spacelike Cauchy surfaces  $X^3$ , the spatial component of  $F_2$  is the magnetic flux density  $B$ , while the Hodge dual (with respect to  $X^4$ ) of the temporal component is the electric flux density  $E$ .

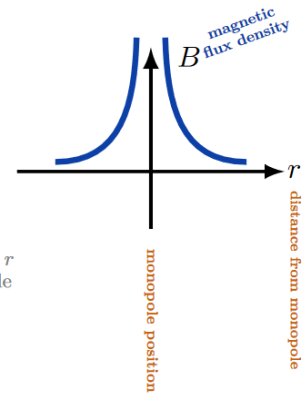
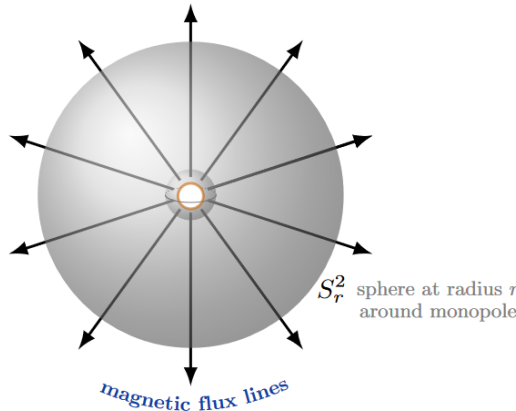
Imagining, with Dirac, that Faraday’s rod magnet could be made *infinitely* long and thin, any one of its poles would look like an isolated mono-pole with flux concentrating towards it from all directions.

At the point of the idealized monopole itself, the flux density  $B$  per unit volume would diverge – a “singularity” much in the sense of black holes, which therefore is not to be regarded as part of space(-time): The spacetime domain on which to discuss the fluxes sourced by a magnetic monopole is (more on all this below in §2.2) not Minkowski spacetime  $\mathbb{R}^{3,1}$  itself, but its complement around the worldline  $\mathbb{R}^{0,1}$  of the would-be monopole.

| Electromagnetic flux density.       |                       |
|-------------------------------------|-----------------------|
| $X^4$                               | spacetime 4-fold      |
| $F_2 \in \Omega_{\text{dR}}^2(X^4)$ | Faraday tensor        |
| $= \star(E_{ij} dx^i \wedge dx^j)$  | electric flux density |
| $+ B_{ij} dx^i \wedge dx^j$         | magnetic flux density |

Faraday tensor  $F_2 = B(r) \text{dvol}_{S_r^2}$ , magnetic flux density

Gauß law away from singular locus  $dF_2 = 0$ ,  $\int_{S_r^2} F_2 =$  integrated magnetic flux through any sphere around the monopole, prop. to monopole charge



- ambient space  $\mathbb{R}^3 \setminus \{0\}$
- magnetic monopole

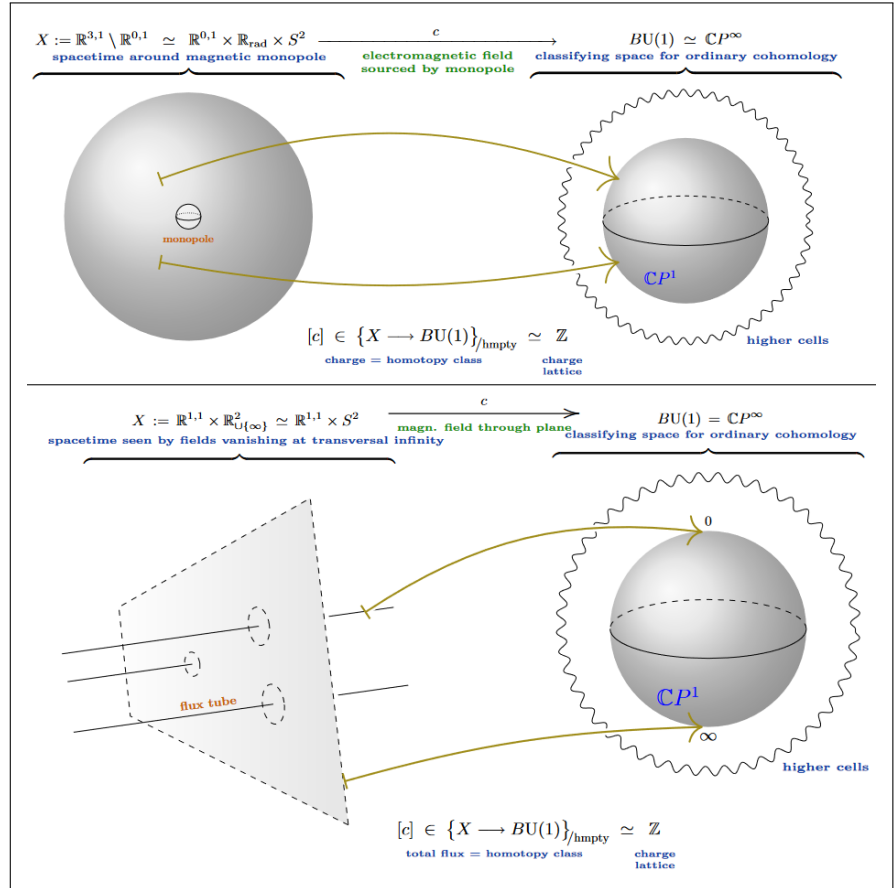
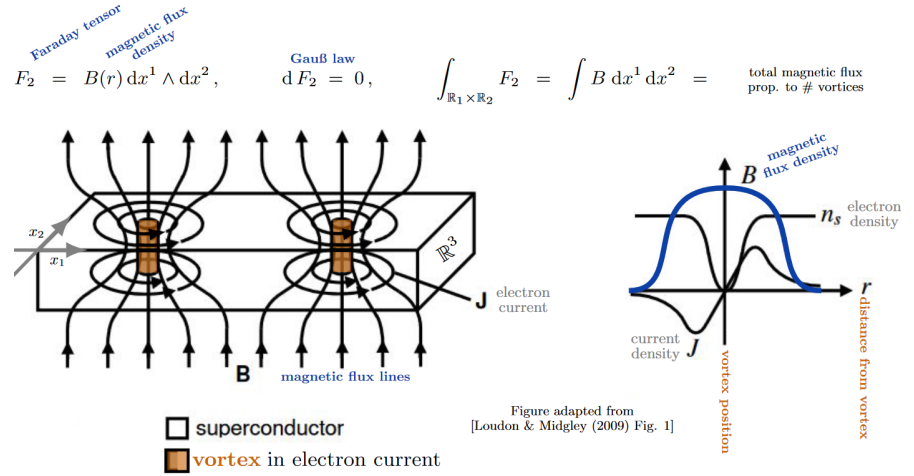
$$\mathbb{R}^{3,1} \setminus \mathbb{R}^{0,1} \underset{\text{homeo}}{\simeq} \mathbb{R}^{0,1} \times (\mathbb{R}^3 \setminus \{0\}) \underset{\text{homeo}}{\simeq} \mathbb{R}^{0,1} \times R_{>0}^1 \times S^2 \underset{\text{hmtp}}{\simeq} S^2.$$

As such, magnetic monopoles are the **singular 0-branes** of electromagnetism (cf. §2.2) – in theory: Whether magnetic monopoles exist in nature remains open; they have not been seen in experiment, but there are decent theoretical arguments that they should exist if the standard model symmetry is a broken “grand unified” symmetry.

However, in the EM-field there are also **solitonic 1-branes** which are experimentally well-established as the *Abrikosov vortices* formed in type II superconductors within a transverse magnetic field [Abrikosov 1957] [Loudon et al. 2009] [Timm 2020, §6.5]. These may be regarded as *strings* approximated by a Nambu-Goto action [Nielsen et al. 1973] [Beekman et al. 2011].

In this case, the “sphere” through which the total magnetic flux density is measured is nominally the  $(x_1, x_2)$ -plane filled by the superconducting material; but since far away from any vortex the magnetic flux has to vanish, this plane appears to the fluxes via its one-point compactification with the “point at infinity” adjoined.

These vortex strings are *solitons* in that the flux density is everywhere finite, and yet the “bumps” in the flux density are topologically stable. Much like a bump in a rug cannot be flattened as long as the boundary of the rug is fixed in place, so the requirement that flux densities “vanish at infinity” keeps the vortex strings in place – or at least this is the case once we take account of Dirac flux quantization below (the corresponding classifying maps for which are previewed on the right).



## 2.2 Singular versus solitonic branes

Generally, imprinted on flux densities may be two kinds of branes, here to be called:

- (i) *singular branes* (black branes), reflected in *diverging flux density* at *singular loci* that are to be removed from spacetime,   
 Beware that in supergravity these are also called “elementary branes” [Duff & Lu 1994], in reference to how black holes carry the same quantum numbers as elementary particles – but here we rather not conflate these two aspects.
- (ii) *solitonic branes*, reflected in *finite flux density* which is localized in that it *vanishes at infinity*, transversally.   
 The terminology “solitonic brane” was introduced by [Duff, Khuri & Lu 1992][Duff, Khuri & Lu 1995] and Duff & Lu 1993, 1994 to mean stable but non-singular brane-like solutions to (supergravity/flux) equations of motion (“solitons”).

This general distinction between singular branes and solitonic branes is important for the correct identification of the implications of choices of flux quantization laws on the corresponding brane charges.

**Spacetime domains for brane fluxes.** More formally, one may encode these two cases by slightly adjusting the nature of the *spacetime domain* on which fluxes are actually defined [Sati & Schreiber 2023a, §2.1]:

- fluxes sourced by singular branes of dimension  $p+1$  inside spacetime  $X^{d+1}$  are actually defined on the complement  $X^{d+1} \setminus Q^{p+1}$  of their singular worldvolume,
- fluxes sourced by solitonic branes of codimension  $d-p$  are actually defined on their transverse space  $T^{d-p}$  equipped with a “point at infinity” on which they are required to vanish.

| Type of brane   | Spacetime domain of flux density  |
|-----------------|---|
| Singular brane  | complement of singular worldvolume locus $Q^{p+1}$ inside spacetime $X^{d+1}$ $X^{d+1} \setminus Q^{p+1}$ |
| Solitonic brane | Transverse space $Y^{d-p}$ to worldvolume equipped with a “point at infinity” $(Y^{d-p}, \infty_Y)$       |

The condition of flux densities vanishing at infinity on some space is naturally formalized by considering the larger category of pointed topological spaces  $(X, x \in X)$  (we discuss further below how to properly speak of differential geometric smoothness in this context) and regarding their given “base point” as being the “point at infinity”, whence we shall write  $(X, \infty_x)$  for the generic pointed space. Then a function “vanishing at infinity” on  $(X, \infty_x)$  is a function on  $X$  that literally vanishes at  $\infty_x$ .

For example:

- The result of *adjoining* to  $\mathbb{R}^n$  its “point at infinity” (this is called its *one-point compactification*, here to be denoted  $\mathbb{R}_{\cup\{\infty\}}^n$ ) is homeomorphic to the  $n$ -sphere with any basepoint:

$$\mathbb{R}_{\cup\{\infty\}}^n \underset{\text{homeo}}{\simeq} S^n.$$

- On the other hand, to consider unconstrained functions on some  $X$  in this context, we may regard all the points of  $X$  as being at finite distance by declaring that the “point at infinity” is disjoint from  $X$ , hence by considering the disjoint union (denoted “ $\sqcup$ ” as opposed to “ $\cup$ ”):

$$X_{\sqcup\{\infty\}} := X \sqcup \{\infty\}.$$

Given two such pointed spaces, their *smash product* “ $\wedge$ ” is their Cartesian product with all points that are at infinity in either factor identified with a single new point at infinity:

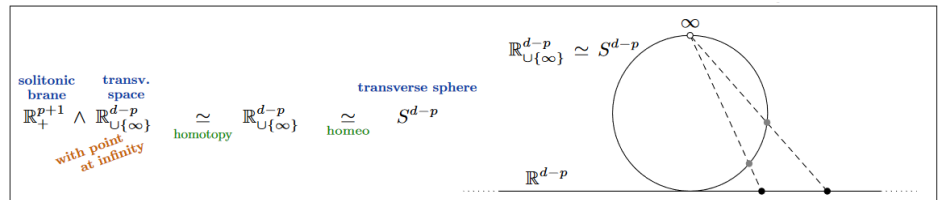
$$(X, \infty_X) \wedge (Y, \infty_Y) := \frac{X \times Y}{X \times \{\infty_Y\} \cup \{\infty_X\} \times Y}$$

**Example 2.1 (Flat branes).** In the case of flat branes – i.e. with Cartesian worldvolumes inside Minkowski spacetime – both these spacetime domains are homotopy equivalent to spheres, but of different dimensions:

(i) The spacetime domain for flat singular branes is homotopy-equivalent to the unit sphere in the transverse space, hence the sphere around the singular brane locus.

|   |
|---|
| <div style="display: flex; justify-content: space-around; align-items: center;"> <div style="text-align: left;"> <p style="font-size: small; color: blue;">bulk</p> <math>\mathbb{R}^{d+1}</math> </div> <div style="text-align: center;"> <p style="font-size: small; color: blue;">singular<br/>brane</p> <math>\setminus \mathbb{R}^{p+1}</math> </div> <div style="text-align: center;"> <math>\simeq</math> </div> <div style="text-align: left;"> <p style="font-size: small; color: blue;">punctured<br/>transverse space</p> <math>(\mathbb{R}^{d-p} \setminus \{0\}) \times \mathbb{R}^{p+1}</math> </div> <div style="text-align: center;"> <math>\simeq</math> </div> <div style="text-align: left;"> <p style="font-size: small; color: blue;">encircling sphere</p> <math>S^{d-p-1}</math> </div> </div> <p style="font-size: x-small; color: green; margin-top: 5px;">homeomorphism <span style="margin-left: 150px;">homotopy equivalence</span></p> |
|---|

(ii) The spacetime domain for flat solitonic branes is homotopy equivalent to the sphere which is the one-point compactification of the transverse space (its stereographic projection).



**Example 2.2 (Flat branes of electromagnetism).** Specifying Ex. 2.1 to the case of ordinary electromagnetic flux (§2.1) it follows from this general reasoning that a flux density 2-form  $F_2$  in  $D = 3 + 1$  may reflect the presence of

- singular 0-branes with spacetime domain  $\mathbb{R}^{3,1} \setminus \mathbb{R}^{0,1} \underset{\text{homeo}}{\simeq} \mathbb{R}^{0,1} \times \mathbb{R}_{>0} \times S^2 \underset{\text{hmtp}}{\simeq} S^2$
- solitonic 1-branes with spacetime domain  $\mathbb{R}_+^{1,1} \wedge \mathbb{R}_{\cup\{\infty\}}^2 \underset{\text{homeo}}{\simeq} \mathbb{R}_+^{1,1} \wedge S^2 \underset{\text{hmtp}}{\simeq} S^2$

which are just the magnetic monopoles (hypothetical) and Abrikosov vortex strings (observed) from §2.1.

**Example 2.3 (Near-horizon geometries of singular branes).**

The idea of regarding singular branes from the complement of their singular locus in spacetime is familiar from the AdS/CFT correspondence:

The near-horizon geometry of any  $> 1/4$  BPS black brane are products of an anti de Sitter spacetime with a (free discrete quotient of) a sphere around the singularity [Acharya et al. 1999]. On a causal chart of AdS spacetime, this is homeomorphic to the flat brane complements.

|                              |  |   |                                    |
|------------------------------|--|---|------------------------------------|
| Near horizon spacetime       | anti-de Sitter spacetime                     |   |                                    |
|                              | AdS <sub>p+2</sub>                           |   | $\times S^{D-p-2} / G$             |
| Metric in horospheric coord. | $\frac{R^2}{z^2} ds_{\mathbb{R}^{p,1}}^2$    | $+$   | $\frac{R^2}{z^2} dz^2$             |
| Causal chart                 | singularity<br>$\mathbb{R}^{p,1}$            | $\times$  | radial direction<br>$\mathbb{R}_+$ |
|                              |  |   | $\times S^{D-p-2} / G$             |
| Metric in natural coord.     | $\frac{r^n}{\ell^n} ds_{\mathbb{R}^{p,1}}^2$ | $+$   | $\frac{\ell^2}{r^2} dr^2$          |
|                              | $\frac{r^n}{\ell^n} \cdot \text{Minkowski}$  | $+$ $\ell^2 ds_{S^{D-p-2}}^2$   |                                    |
|                              |  | transversal space $\mathbb{R}^{D-p-2} \setminus \{0\}$                      |                                    |
|                              |  | $\frac{\ell^2}{r^2} \cdot \text{metric cone } C(S^{D-p-2}) \setminus \{0\}$ |                                    |

In general, the spacetime domain on which to measure flux densities may be a mix of these purely singular and purely solitonic situations, in which case the notions of singular and of solitonic branes blend into each other:

**Example 2.4 (solitonic branes in KK-compactifications).** In mild generalization of Ex. 2.1, consider the case that spacetime is a trivial circle-fiber bundle over a Minkowski spacetime  $X^D \equiv \mathbb{R}^{1,d-1} \times S^1$ . In this case, the flux sourced by solitonic branes of codimension  $n$  as seen on the base spacetime  $\mathbb{R}^{1,d-1}$  is measured on the smash product space

$$(S^1 \times \mathbb{R}^n)_{\cup\{\infty\}} \underset{\text{homeo}}{\simeq} S^1_{\sqcup\{\infty\}} \wedge \mathbb{R}^n_{\cup\{\infty\}} \underset{\text{homeo}}{\simeq} S^1_{\sqcup\{\infty\}} \wedge S^n.$$

This was first understood by [Bergman et al. 1999, §2.2 & §2.3].

**2.3 Higher fluxes and their brane sources**

On this backdrop of ordinary electromagnetic flux (§2.1) and of the general rule for measuring flux sourced by singular branes or solitonic branes (§2.2) it clearly makes sense to consider physical theories of higher gauge fields whose precise nature remains to be discussed, but whose flux densities are reflected in higher-degree differential forms

$$F^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^D), \tag{2}$$

these possibly being of different field species to be labeled by a finite index set  $I \in \text{FinSet}$  and jointly to be denoted as follows:

$$\vec{F} \equiv \left\{ F^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^D) \right\}. \tag{3}$$

Such higher flux densities appear in higher dimensional supergravity, namely as “superpartners” of the gravitino field that cannot be accounted for by the graviton itself. In particular, in  $D = 10$  supergravity and  $D = 11$  supergravity these higher flux densities are known under the (now) fairly standard symbols shown on the right, along with the standard name of the corresponding singular branes (the “higher-dimensional monopoles”), e.g. [Blumenhagen et al. 2013, §18.5].

|                      | Field   | Flux      | Singular source   |
|----------------------|---------|-----------|-------------------|
| $D=4$ Maxwell theory | A-field | $F_2$     | monopole 0-branes |
| $D=10$ supergravity  | B-field | $H_3$     | NS5-brane         |
|                      |         | $H_7$     | F1-branes         |
| $D=11$ supergravity  | C-field | $F_{8-p}$ | Dp-branes         |
|                      |         | $G_4$     | M5-branes         |
|                      |         | $G_7$     | M2-branes         |

(4)

The above flux densities in 11d and 10d are closely related:

**Example 2.5 (Double dimensional reduction of fluxes from 11d to 10d).** Consider the case of C-field flux densities  $G_4$  and  $G_7$  on an 11-dimensional spacetime  $Y^{11}$  which is the total space of a circle-principal bundle and denote by

- $\theta \in \Omega_{\text{dR}}^1(Y^{11})$  any fiberwise Maurer-Cartan form along the fibers (i.e. an Ehresmann connection form),
- $F_2 \in \Omega_{\text{dR}}^2(X^{10})$  the corresponding (first) Chern class-characteristic form, i.e., the curvature form whose pullback to  $Y^{11}$  is  $d\theta = p^*F_2$ .



Assuming that all flux densities are  $S^1$ -invariant (hence focusing on their 0th KK-modes) they decompose into a basic component (a differential form on  $X^{10}$ , pulled back along the projection  $p$ ) and the wedge product of a basic differential form with the Maurer-Cartan form  $\theta$  on the  $S^1$ -fibers:

$$\begin{array}{ccc}
S^1 \hookrightarrow X^{11} & & d\theta = p^*F_2 \\
\downarrow p & & G_4 = p^*F_4 - \theta \wedge p^*H_3 \\
\Downarrow & & \\
X^{10} & & G_7 = p^*H_7 - \theta \wedge p^*F_6
\end{array} \quad (5)$$

This is the process of “double dimensional reduction” [Duff et al. 1987] [Braunack-Mayer et al. 2019, §2.2] – called this way since both spacetime dimension is reduced by Kaluza-Klein reduction on a fiber space, but also the degrees of densities of fluxes “through the fiber space” are decreased – known as part of the duality between M-theory and type IIA string theory: The new component flux densities  $H_3$  and  $H_7$  are interpreted as those of the B-field and the component flux densities  $F_4$  and  $F_6$  (and  $F_2$ ) as those of the RR-field in type IIA supergravity.

Here the flux density  $F_2$ , which in 10d is understood as witnessing singular D6-brane sources, is a gravitational flux from the 11d point of view: If  $X^{10} \equiv \mathbb{R}^{6,1} \times \mathbb{R}_{>0} \times S^2$  is the spacetime domain around a flat singular D6-brane, then the total space of the circle-principal bundle  $Y^{11}$  (a multiple of the complex Hopf fibration) is known as the corresponding “KK-monopole” spacetime.

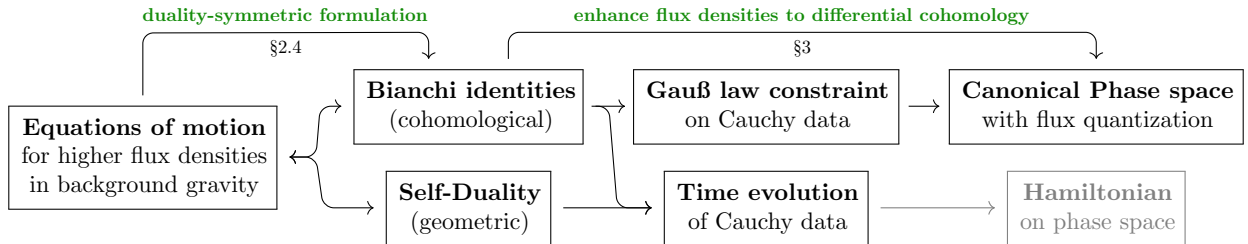
This transmutation, under Kaluza-Klein compactification, of parts of the gravitational field in higher dimensions into gauge fields in lower dimensions is a major subtlety in choosing flux quantization laws: Since these laws apply to higher gauge fields but not directly to the field of gravity, there may appear new possibilities for flux quantization after KK-reduction to lower dimensions which do not come from flux quantization in higher dimensions.

## 2.4 Equations of motion of higher flux

As we now turn to the equations of motion for flux densities (the analogs of Maxwell’s equations for electromagnetic flux), the key move towards identifying possible flux quantization laws (in §3) is to arrange these equations, equivalently, as:

- (i) a purely cohomological system of differential equations known as higher *Bianchi identities*,
  - (ii) a purely geometric system of linear equations expressing a Hodge self-duality,
- the point being that the first item is entirely “algebraic-topological” (homotopy-theoretic), while dependency on geometry, namely on the spacetime metric (the field of gravity) is all isolated in the second item.

It turns out [Sati & Schreiber2023b] that from such duality-symmetric laws of flux, the canonical phase space of the higher gauge theory, including the flux-quantization structure, may be obtained straightforwardly, without going through the traditional and thorny route of BRST-BV analysis based on a stationary action principle given by a Lagrangian density.



This move of isolating “pre-metric flux equations” supplemented by a “constitutive” duality constraint has a curious status in the literature. On the one hand, it is elementary and immediate as an equivalent re-formulation of the usual form of (higher) Maxwell-type equations of motion, and as such has been highlighted a century ago [Cartan1924, §80] and again more recently [Hehl & Obukhov2003] (historical survey in [Hehl et al. 2016]). While the broader community does not seem to have taken much note of “premetric electromagnetism” as such, one may notice that just the same perspective is evidently what in supergravity and string theory is called “duality-symmetric” [Bandos et al. 1998] or “democratic” [Mkrtychyan & Valach 2023] formulations of fluxes in supergravity (see Ex. 2.10 and Ex. 2.12 below).

**Definition 2.6 (Higher Maxwell-type equations).** A system of *higher Maxwell-type equations* on a tuple (2) of flux densities on a spacetime  $X^D$  is

- a system of polynomial ( $\vec{P}$ ) first-order exterior-differential equations  $d\vec{F} = \vec{P}(\vec{F})$  (the higher *Bianchi identities*, crucially admitting polynomial “self-sourcing” of fluxes);
- subject to a system of linear ( $\vec{\mu}$ ) Hodge-self-duality relation  $\star\vec{F} = \vec{\mu}(\vec{F})$  (the “*constitutive relations*”):

higher Maxwell-type  
equations of motion in  
duality-symmetric form

|  |
|--|
| <p><b>Bianchi identities</b><br/> <math>d\vec{F} = \vec{P}(\vec{F})</math><br/> <math>\star F = \vec{\mu}(\vec{F})</math><br/> <b>self-duality</b></p> |
|--|

|  |   |   |     |
|--|---|---|-----|
| <p>flux species<br/> <math>I \in \text{Set}</math>,</p>  | <p>flux degrees<br/> <math>(\text{deg}_i \in \mathbb{N}_{\geq 1})_{i \in I}</math>,</p> | <p>flux densities<br/> <math>\vec{F} \equiv (F^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^D))_{i \in I}</math></p> | (6) |
| <p><math>\vec{P}</math> graded-symm. polynomial , <math>\vec{\mu}</math> invertible matrix</p> |   | <p>flux self-sourcing<br/> vacuum permittivity</p>  |     |

Concretely:

- $\vec{P}$  is an  $I$ -tuple of graded-symmetric polynomials with rational coefficients in  $I$  variables of degrees  $\vec{\text{deg}}$ ,
- $\vec{\mu}$  is a linear endomorphism on the vector space spanned by these variables.

**Remark 2.7.** The equations in Def. 2.6 imply that  $\vec{P}$  and  $\vec{\mu}$  respect degrees in a certain evident way. Moreover, the following property of the Hodge star operator on Lorentzian manifolds (e.g. [Frankel 1997, §14]) implies further constraints on the available higher Maxwell-type equations:

$$\star \star F_{\text{deg}} = -(-1)^{\text{deg}(D-\text{deg})} F_{\text{deg}}, \quad \text{for } F_{\text{deg}} \in \Omega_{\text{dR}}^{\text{deg}}(X^D) \quad (7)$$

This controls notably the existence of genuinely self-dual higher gauge theories, see Ex. 2.11 below.

**Remark 2.8.** Not all higher gauge theories are of the higher Maxwell-form (Def. 2.6): For instance, higher Chern-Simons type theories are different.

**Example 2.9 (Motion of the ordinary electromagnetic fluxes).**

The classical Maxwell equations expressed in terms of differential forms are as shown on the left (e.g. [Frankel 1997, §3.5 & §7.2b]), with their “premetric” form shown on the right.

Here the differential 3-form  $J_3$  embodies the density of an electric current carrying an electric field and inducing a magnetic field.

This kind of *external* or *background* source term, where the source is not given by (a polynomial in) the flux densities themselves, does not fit into the Definition 2.6 and will be disregarded for the purpose of the present discussion, meaning that we focus on the special case of Maxwell’s equations “in vacuum”.

|  |  |     |
|--|--|-----|
| $\begin{aligned} dF_2 &= 0 \\ d\star F_2 &= J_3 \end{aligned}$ | $\begin{aligned} dF_2 &= 0 \\ dG_2 &= J_3 \end{aligned}$ | (8) |
|  | $G_2 = \star F_2$  |     |

|  |  |     |
|--|--|-----|
| $\begin{aligned} dF_2 &= 0 \\ d\star F_2 &= 0 \end{aligned}$ | $\begin{aligned} dF_2 &= 0 \\ dG_2 &= 0 \end{aligned}$ | (9) |
|  | $G_2 = \star F_2$                                      |     |

It is clear that, mathematically at least, Ex. 2.9, makes sense more generally for flux densities of any degree. In particular:

**Example 2.10 (Motion of unbounded RR-field fluxes).** The equations of motion of the RR-field fluxes in  $D = 10$  supergravity in the case of vanishing B-field-fluxes are often taken to be as follows (e.g. [Mkrtchyan & Valach 2023]):

|   |  |   |      |
|---|--|---|------|
| $\begin{aligned} dF_{2\bullet+\sigma} &= 0 \\ d\star F_{2\bullet+\sigma} &= 0 \\ \forall 2\bullet+\sigma \leq 5 \\ \star F_5 &= F_5 \text{ if } \sigma = 1 \end{aligned}$ | $dF_{2\bullet+\sigma} = 0 \quad \forall 2\bullet+\sigma$ | $\bullet \in \mathbb{N}$<br>$\sigma = \begin{cases} 0 & \text{for type IIA} \\ 1 & \text{for type IIB} \end{cases}$ | (10) |
|   | $F_{(10-2\bullet-\sigma)} = \star F_{2\bullet+\sigma}$   |   |      |

and, more generally, those with non-vanishing B-field as follows:

|  |   |      |
|--|---|------|
| $\begin{aligned} dF_{2\bullet+\sigma} &= H_3 \wedge F_{2\bullet+\sigma-2} & dH_3 &= 0 \\ d\star F_{2\bullet+\sigma} &= H_3 \wedge \star F_{D-2\bullet-\sigma+2} & d\star H_3 &= \dots \end{aligned}$ | $\begin{aligned} dF_{2\bullet+\sigma} &= H_3 \wedge F_{2\bullet+\sigma-2} & dH_3 &= 0 \\ & & dH_7 &= \dots \end{aligned}$ | (11) |
|  | $F_{D=2\bullet-\sigma} = \star F_{2\bullet+\sigma} \quad H_7 = \star H_3$   |      |

Beware, while these equations are now often stated in this form, and while this is the form that motivates the traditional *Hypothesis K* (§4.1), it is at least subtle to see them in entirety as actually arising from ordinary  $D = 10$  supergravity (namely from KK-compactification of  $D = 11$  supergravity, in the case  $\sigma = 0$ ), since in that context:

- The fluxes  $F_0$  and  $F_{1-}$  are not actually present: They are from *massive* type IIA, which has its own subtleties.
- The flux  $H_7$  has a non-linear Bianchi ( $dH_7 = -F_4 \wedge F_4 + F_2 \wedge F_6$ ) which does not fit the pattern (cf. Ex. 2.13).

Notice that in type IIB, (10) describes a flux density ( $F_5$ ) which is Hodge dual (not just to any other flux in the tuple but) to itself,  $F_5 = \star F_5$ . Generally we have:

**Example 2.11 (Motion of self-dual higher gauge field fluxes).**

Since Def. 2.6 regards *every* higher gauge theory (of Maxwell-type) as being “self-dual” in a sense, the equations of motion of flux densities of actual self-dual higher gauge fields — in the strict sense that one and the same flux density form is required to be Hodge dual to itself — are readily an example of Def. 2.6:

equations of motion of  
self-dual higher gauge field  
in  $D = 4k + 2$

←

$d F_{D/2} = 0$

$F_{D/2} = \star F_{D/2}$

(12)

Due to the properties of the square of the Hodge operator (7), this has non-trivial solutions iff the degree of the flux is odd,  $\text{deg} = 2k + 1$ , and hence iff spacetime dimension is  $D = 4k + 2$ ,  $k \in \mathbb{N}$ .

**Example 2.12 (Motion of C-field fluxes).**

The equations of motion of the C-field in  $D = 11$  supergravity (originally the “3-index A-field” due to [Cremmer et al. 1978] (cf. [Miemiec et al. 2006, p. 32]) are traditionally as shown on the left here, with their equivalent “duality-symmetric” reformulation [Bandos et al. 1998] shown on the right, cf. [Giotopoulos et al. 2024a].

$d G_4 = 0$   
 $d \star G_4 = -\frac{1}{2} G_4 \wedge G_4$

←

$d G_4 = 0$   
 $d G_7 = -\frac{1}{2} G_4 \wedge G_4$

$G_7 = \star G_4$

(13)

**Example 2.13 (Motion of type IIA B&RR-field fluxes).** [Mathai & Sati 2004, §4][Fiorenza et al. 2017, §3]

Under double dimensional reduction (5) of the C-field flux along a circle-bundle, the equations of motion (13) are equivalently expressed in terms of its B&RR-field-components as shown.

dimensionally reduced  
equations of motion  
of C-field flux in 11d  
to B&RR-field flux in 10d  
+ KK-monopole flux EoM

←

|                          |                                       |
|--------------------------|---------------------------------------|
| $d F_2 = 0$              | $d H_3 = 0$                           |
| $d F_4 = H_3 \wedge F_2$ | $d H_7 = -\frac{1}{2} F_4 \wedge F_4$ |
| $d F_6 = H_3 \wedge F_4$ | $+ F_2 \wedge F_6$                    |
| $d F_8 = H_3 \wedge F_6$ |                                       |

|                   |                   |
|-------------------|-------------------|
| $\star F_4 = F_6$ | $\star H_3 = H_7$ |
| $\star F_2 = F_8$ |                   |

11d C-field gravity  
  
C-field gravity

These are the equations of motion of the flux densities of type IIA supergravity in their duality-symmetric formulation [Cremmer et al. 1998, §3].

Several terms above deserve special attention, either for how they appear or for how they do *not* appear:

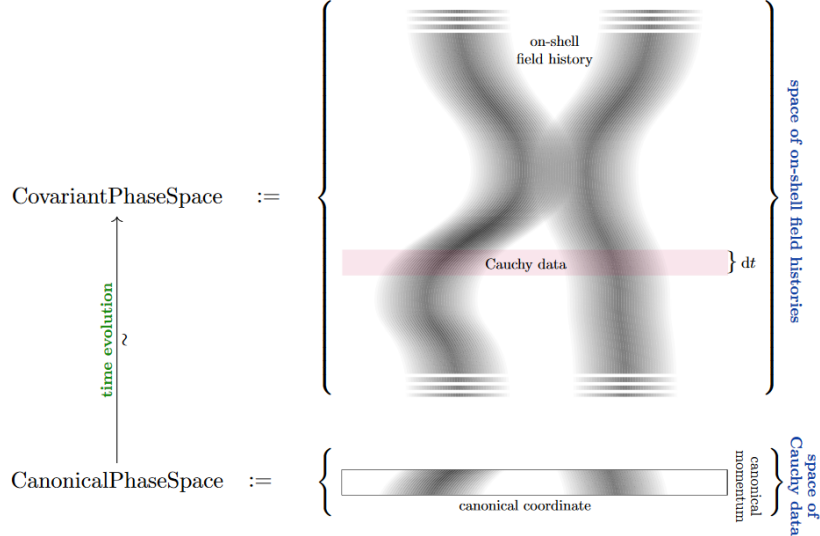
- The Hodge dual  $F_8 = \star F_2$  is *not* part of the C-field in 11d, but is part of the gravitational field (the Hodge star encodes the 10d metric and  $F_2$  is an aspect of the 11d fiber geometry). The expression for  $dH_8$  arises as part of the gravitational field equations [Cremmer et al. 1998, (3.4)].
- The presence in type IIA supergravity of the non-linear Bianchi identity for  $H_7$ , albeit readily verified and “well-known” at least since [Cremmer et al. 1998, (3.4)], is not as widely appreciated as the pattern (11) from Ex. 2.10 – which it *breaks* (cf. also Rem. 3.11 below).
- No flux densities  $F_0$  nor  $F_{10}$  appear in 10d from KK-reduction on a circle, nor are they part of type IIA supergravity. These fluxes are instead part of massive type IIA supergravity whose relation to  $D = 11$  supergravity/M-theory remains less understood.

## 2.5 Solution space of the flux equations

**The phase space.** Abstractly, the phase space of any field theory is nothing but the space of all those field histories that satisfy the given equations of motion (the “on-shell” field histories). Phrased this way, this is sometimes called the *covariant phase space* [Crnković & Witten 1987], to emphasize that no choice of foliation of spacetime by Cauchy surfaces has been or needs to be made.

The more traditionally familiar *canonical phase space* (physics jargon which, beware, is somewhat incompatible with the mathematician’s “canonical”) is instead a parameterization of the covariant phase space by initial value

data on a choice of Cauchy surface. This choice breaks the “manifest covariance” of the covariant phase space. Nevertheless, if a Cauchy surface exists at all (hence on globally hyperbolic spacetimes), then both these phase spaces are equivalent, by definition, the equivalence being the map that generates from initial value data the essentially unique on-shell field history that evolves from it (disregarding gauge transformations for the moment).



**Solution space of on-shell flux densities.** At this point in the discussion the full gauge field content is not yet determined — this will only be implied by a choice of flux quantization below in §3.3 — so far we are only considering the flux densities of the would-be gauge fields. To remember this, we shall call the space of flux densities solving their equations of motion (Def. 2.6) the *solution space*; and we are after its incarnation as a canonical solution space of initial value data on a Cauchy surface. But this goes a long way, since the higher Maxwell-type equations of motion constrain exclusively the flux densities: Once the flux-quantization the canonical phase will simply consist of all flux-quantized gauge potentials compatible with the flux densities in the canonical solution space.

**Proposition 2.14** ([Sati & Schreiber2023b]). *On a globally hyperbolic spacetime  $X^D \simeq \mathbb{R}^{0,1} \times X^d$ , the solution space given higher Maxwell-equations of motion (Def. 2.6) is isomorphic to the solution of (just) the duality-symmetric Bianchi identities restricted (i.e.: pulled back to) to any Cauchy surface  $\iota: X^d \hookrightarrow X^D$ , there to be called the higher Gauss law:*

$$\begin{aligned}
 \text{Space of flux densities on spacetime, solving the equations of motion} \quad \text{SolSpace} &\equiv \left\{ \begin{array}{l} \text{electromagnetic flux densities on spacetime} \\ \vec{F} \equiv \left( F^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^D) \right)_{i \in I} \end{array} \left| \begin{array}{l} \text{Bianchi identities} \\ \text{d} \vec{F} = \vec{P}(\vec{F}) \\ \star F = \vec{\mu}(\vec{F}) \\ \text{self-duality} \end{array} \right. \right\} \text{covariant form} \\
 &\simeq_{\iota^*} \left\{ \begin{array}{l} \text{magnetic flux densities on Cauchy surface} \\ \vec{B} \equiv \left( B^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^d) \right)_{i \in I} \end{array} \left| \begin{array}{l} \text{Gauß law} \\ \text{d} \vec{B} = \vec{P}(\vec{B}) \end{array} \right. \right\} \text{canonical form}
 \end{aligned} \tag{14}$$

**Example 2.15 (Solution- and phase-space of ordinary electromagnetism).** In the case of ordinary vacuum electromagnetism, Prop. 2.14 applied to the ordinary Maxwell equations (Ex. 2.9) says that the initial value data on a Cauchy surface  $X^3$  is given by *independently* specifying magnetic and electric flux densities  $B, E \in \Omega_{\text{dR}}^2(X^3)$  subject only to the ordinary Gauß laws  $\text{d}B = 0, \text{d}E = 0$ . Indeed, the actual phase space of electromagnetism is well-known (e.g. [Blaschke & Gieres 2021, §5]) to have as

- (i) canonical coordinate the gauge potential  $\widehat{A}$ , (ii) canonical momentum the electric flux density  $E$

Thereby  $B \equiv \text{curv}(\widehat{A})$  is indeed independent from  $E$  (and satisfies its Gauß law definitionally, while the Gauß law on  $E$  is a phase space constraint).

**Remark 2.16 (Canonical coordinates/momenta from duality-symmetry).** Notice how, thereby, the traditional split of initial value data into canonical coordinates and canonical momenta (whose definition requires assumption and variation of a Lagrangian density) is preempted here, under Prop. 2.14, already by the pregeometric/duality-symmetric formulation of Maxwell’s equations (Ex. 2.6), in the sense that the spacetime archetypes of the canonical coordinates and momenta on a Cauchy surface (the former seen under the differential) are just the ordinary flux density  $F_2$  (since  $B = \iota^* F_2$ ) and its “duality partner”  $G_2$  (since  $E = \iota^* G_2$ ).

**Remark 2.17 (Gravity “decouples” on canonical phase space).** The inverse isomorphism (14) is given by time evolution of initial value data. Notice that the pseudo-Riemannian metric on  $X^D$  — the background field of gravity — enters only in determining the nature of this isomorphism  $\iota^*$  (the time evolution away from the Cauchy surface), but does not affect the nature of the initial value data (of the canonical phase space) as such.

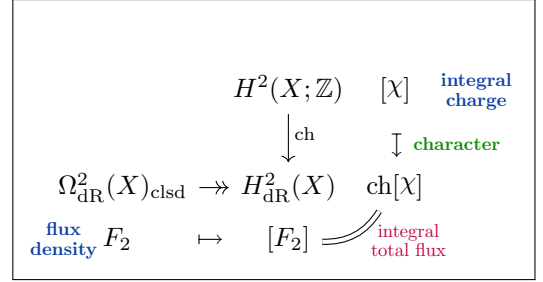
### 3 Flux & Charge Quantization Laws

With the solution space (Prop. 2.14) of higher Maxwell-type equations of motion (Def. 2.6) in hand, the question of flux quantization is to further constrain the flux densities such that the total fluxes and their total source charges take values in some discrete space. The technical issue to be resolved here is that:

- this is a global condition on the flux densities: The local flux densities may take any value (compatible with the equations of motion) and yet the total accumulation of all these local contributions needs to be constrained;
- the evident idea of constraining the ordinary integrals of the flux densities (their “periods”) makes sense only for closed differential forms and hence does not work for non-linear Bianchi identities (such as those of the C-field, Ex. 2.12, and the B&RR-field, Ex. 2.13).

To resolve this, one may first observe that:

- the integrals/periods of ordinary closed differential  $n$ -forms  $f_n$  over  $n$ -manifolds are in natural correspondence with their de Rham-classes,  $[F_n] \in H_{\text{dR}}^n(-)$ , which in turn are equivalently their “deformation classes”, namely their *concordance* classes:  $H_{\text{dR}}^n(-) \simeq \Omega_{\text{dR}}^n(-)_{\text{clsd}} / \text{cncrdnc}$ ;
- so that integrality of closed flux density  $F_n$  is witnessed by an integral cohomology class  $[\chi] \in H^n(X; \mathbb{Z})$  whose “de Rham character” image  $\text{ch}[\chi] \in H_{\text{dR}}^n(X)$  coincides with the deformation class  $[F_n]$ ;



and, second, one may observe that this perspective generalizes [Fiorenza et al. 2023][Sati & Schreiber2023b]:

Higher Maxwell-type equations have a **characteristic  $L_\infty$ -algebra  $\mathfrak{a}$** : The flux densities are equivalently  $\mathfrak{a}$ -valued differential forms, and the Gauß law (14) is equivalently the condition that these be *closed* (i.e.: flat, aka “Maurer-Cartan element”; in Italian SuGra literature: “satisfying an FDA”).

Also every topological space  $\mathcal{A}$  (under mild conditions) has a characteristic  $L_\infty$ -algebra: Its  $\mathbb{R}$ -rational **Whitehead bracket  $L_\infty$ -algebra  $\mathfrak{A}$** .

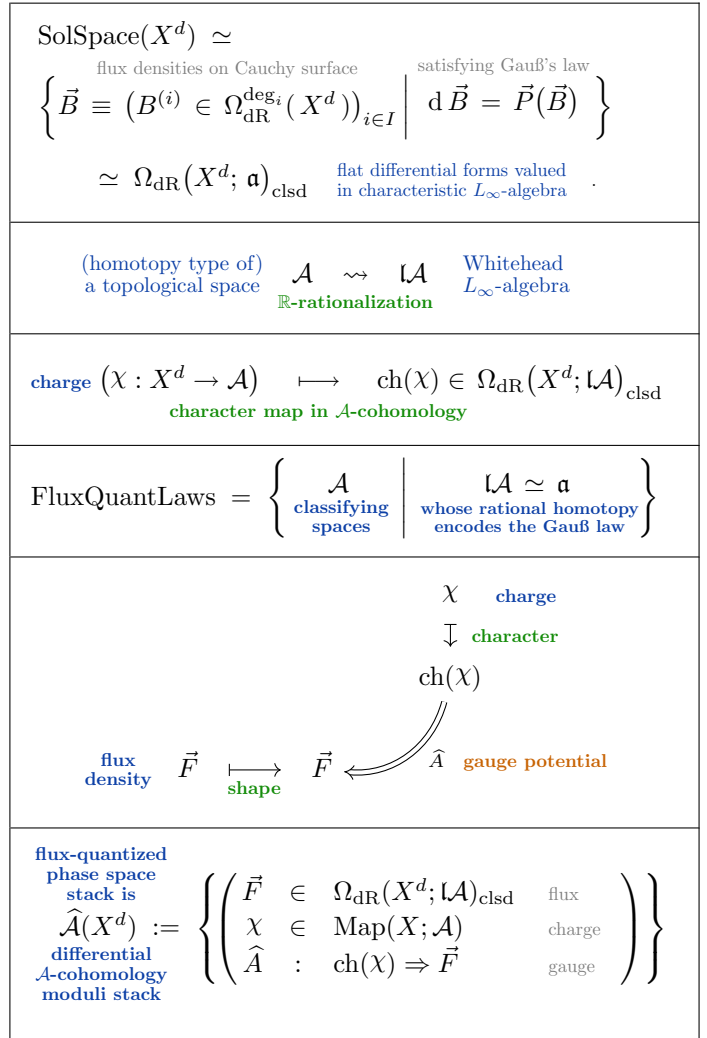
The **nonabelian Chern-Dold character map** turns  $\mathcal{A}$ -valued maps into closed  $\mathfrak{A}$ -valued differential forms, generalizing the Chern character for  $\mathcal{A} = \text{KU}_0$ .

The **possible flux quantization laws** for a given higher gauge field are those spaces  $\mathcal{A}$  whose Whitehead  $L_\infty$ -algebra is the characteristic one.

Given a flux quantization law  $\mathcal{A}$ , the corresponding **higher gauge potentials** are deformations of the flux densities into characters of a  $\mathcal{A}$ -valued map, witnessing the flux densities as reflecting discrete charges quantized in  $\mathcal{A}$ -cohomology.

(It is not obvious that this reduces to the usual notion of gauge potentials, but it does.)

These non-perturbatively completed higher gauge fields form a *smooth higher groupoid*: the “canonical **differential  $\mathcal{A}$ -cohomology moduli stack**”. Since these are now the flux-quantized on-shell fields, this is the **phase space** of the flux-quantized higher gauge theory (p. 11).



This flux-quantized phase space hence subsumes the “solitonic” fields with non-trivial charge sectors  $\chi$ , and as such is a non-perturbative completion of the traditional phase spaces (which correspond to a fixed charge sector only, typically to  $\chi = 0$ ).

Incidentally, it follows that the choice of flux quantization law  $\mathcal{A}$  not only defines the solitonic content of the theory but completely characterizes it:

The shape (topological realization) of this phase space stack is the **space of topological fields**,

which implies that the ordinary homology of the phase space stack constitutes the **topological observables** on the higher gauge theory.

Hence if we focus only on the solitonic or *topological field*-content of the phase space, then we see plain  $\mathcal{A}$ -cohomology moduli of the Cauchy surface and the full phase space stack only serves to justify this object.

We now explain all this in more detail.

|   |
|---|
| $\int \widehat{\mathcal{A}}(X^d) \simeq \mathcal{A}(X^d) = \text{Map}(X^d, \mathcal{A})$  |
| $H_\bullet(\widehat{\mathcal{A}}(X^d); \mathbb{C}) \simeq H_\bullet(\mathcal{A}(X^d); \mathbb{C})$  |
| <p>flux-quantized<br/>topological<br/>phase space</p> <p><math>\mathcal{A}(X^d) := \{\chi \in \text{Map}(X, \mathcal{A})\}</math></p> <p>non-abelian<br/><math>\mathcal{A}</math>-cohomology<br/>moduli space</p> |

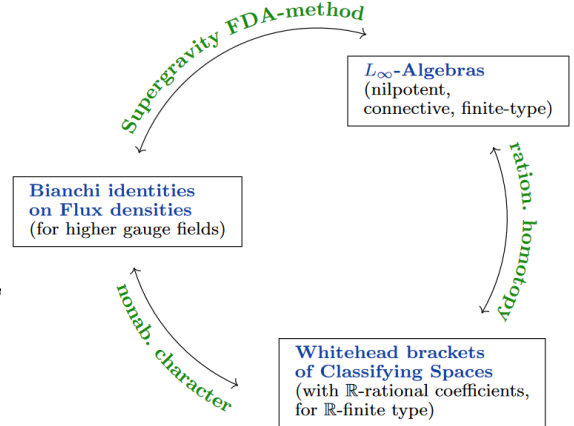
### 3.1 Total flux as Nonabelian de Rham cohomology

We explain how higher Bianchi identities (6) and their corresponding higher Gauss laws (14) are equivalently the closure (flatness) condition on differential forms valued in a characteristic  $L_\infty$ -algebra (Prop. 3.1 below), so that total flux is a class in  $\mathcal{A}$ -valued nonabelian de Rham cohomology (Def. 3.3 below).

The notion of  $L_\infty$ - or *strong homotopy Lie algebra* is finally becoming more widely appreciated in physics, where they appear in various guises (cf. [Stasheff 2016]). Here we are concerned with  $L_\infty$ -algebras which are (i) nilpotent, (ii) connective (iii) of finite type, in their joint incarnation as higher flux density coefficients and as higher Whitehead brackets (all to be explained in a moment), which one might refer to as the

#### Flux Homotopy Lie algebra triality:

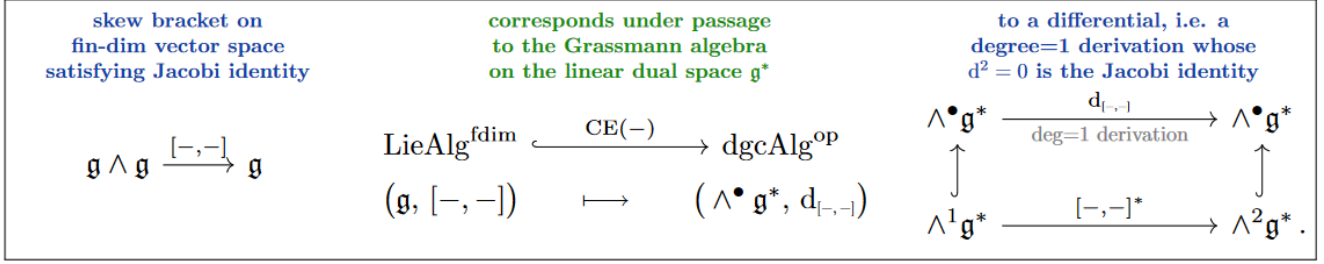
- With *rational homotopy*, we are referring here specifically the fundamental theorem of dg-algebraic rational homotopy theory, mainly due to Quillen, Sullivan and Bousfield & Gugenheim, as reviewed in [Fiorenza et al. 2023, §5].
- The *FDA method* in supergravity refers to the observations of [van Nieuwenhuizen 1983][D’Auria & Fré 1982][Castellani et al. 1991], as explained and contextualized in [Fiorenza et al. 2015b][Fiorenza et al. 2018][Huerta et al 2019], reviewed in [Fiorenza et al. 2019].
- The *nonabelian character* is the generalization of the Chern-Dold character map from topological K-theory and Whitehead-generalized cohomology to generalized non-abelian cohomology, constructed in [Fiorenza et al. 2023].



In particular, this means that  $L_\infty$ -algebras as used here are *not* directly to be understood as generalizations of the gauge Lie algebras familiar from Yang-Mills theory, which are coefficients of the gauge potentials, but instead as the coefficients of their flux densities (cf. Rem. 1.1).

**$L_\infty$ -Algebras.** Since we are assuming  $L_\infty$ -algebras to be connective and of finite type (meaning that they are degreewise finite-dimensional and concentrated in non-negative degrees) we may *define* them through their Chevalley-Eilenberg (CE) algebras in the following manner, which is not only convenient for dealing with the otherwise intricate sign rules, but also essential to their alternative perspectives in the above triality:

**Chevalley-Eilenberg algebras of Lie algebras.** Namely, for  $\mathfrak{g}$  a finite-dimensional Lie algebra (our ground field is the real numbers, throughout) with Lie bracket a skew-symmetric linear map  $[-, -] : \mathfrak{g} \otimes \mathfrak{g} \rightarrow \mathfrak{g}$ , its linear dual vector space  $\mathfrak{g}^*$  is equipped with the dual bracket  $[-, -]^* : \mathfrak{g}^* \rightarrow \mathfrak{g}^* \wedge \mathfrak{g}^*$  which extends uniquely to a degree=1 derivation on the graded Grassmann algebra  $\wedge^\bullet \mathfrak{g} := \bigoplus_{n \in \mathbb{N}} \underbrace{\mathfrak{g}^* \wedge \cdots \wedge \mathfrak{g}^*}_{n \text{ factors}}$ :



One readily checks that this derivation squares to zero iff the bracket satisfies its Jacobi identity(!):

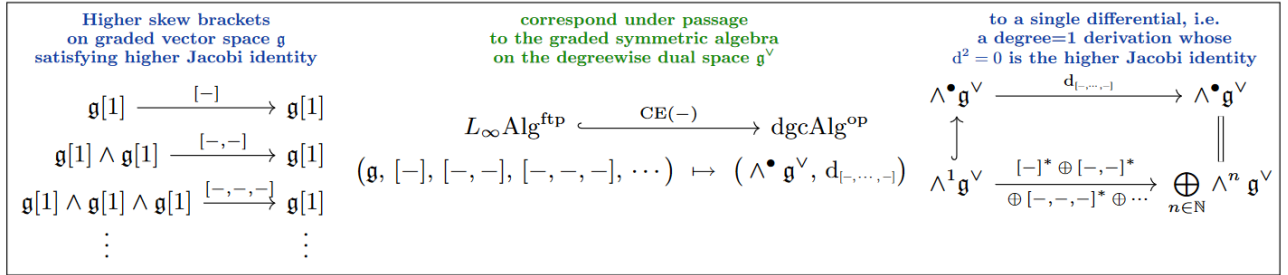
$$\text{Jacobi identity for } [-,-] \quad \Leftrightarrow \quad d_{[-,-]} \circ d_{[-,-]} = 0.$$

The resulting differential graded-commutative (dgc) algebra  $(\wedge^\bullet \mathfrak{g}^*, d_{[-,-]})$  is known as the *Chevalley-Eilenberg complex*  $\text{CE}(\mathfrak{g})$  whose cochain cohomology computes the Lie algebra cohomology of  $\mathfrak{g}$  (with trivial coefficients) — but the key point at the moment is that its construction is a *fully faithful* embedding the category of finite-dimensional Lie algebras into the opposite of that of dgc-algebras.

**$L_\infty$ -Algebras of finite type.** With ordinary Lie algebras viewed as special dgc-algebras this way, it is immediate to generalize them to the case where  $\mathfrak{g}$  may be a graded vector space of degreewise finite dimension (“of finite type”): Namely, writing

$$(\mathfrak{g}^\vee)_n := (\mathfrak{g}_n)^*, \quad \wedge^\bullet \mathfrak{g}^\vee := \text{Sym}(\mathfrak{g}^\vee[1])$$

we can use *verbatim* the same construction. A degree=1 derivation on  $\wedge^\bullet \mathfrak{g}^\vee$  is determined by its restriction to  $\wedge^1 \mathfrak{g}^\vee$ , where it is a sum of co- $n$ -ary linear maps, whose linear duals are identified with  $n$ -ary degree=-1 brackets on  $\mathfrak{g}[1]$ :



Here the simple condition that  $d_{[\dots]}$  be a differential implies a tower of conditions on these brackets, generalizing the Jacobi identity on an ordinary Lie algebra and known as the conditions that make  $(\mathfrak{g}, [-], [-,-], [-,-,-], \dots)$  an  $L_\infty$ -algebra:

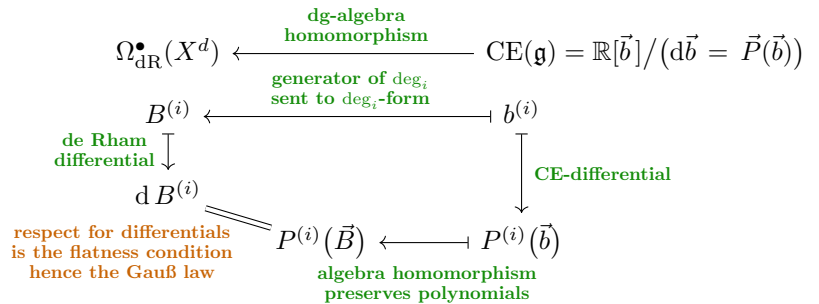
$$\text{Higher Jacobi identity for } [-], [-,-], [-,-,-], \dots \quad \Leftrightarrow \quad d_{[\dots]} \circ d_{[\dots]} = 0. \quad (15)$$

In other words, we may identify  $L_\infty$ -algebras of finite type as the formal dual to dgc-algebras whose underlying graded-commutative algebra is free on a graded vector space. Several examples are indicated in (17).

**Flat  $L_\infty$ -algebra valued differential forms** now have an immediate definition from this perspective: They are the dg-algebra homomorphism from their CE-algebras into de Rham algebras (aka “Maurer-Cartan elements”):

$$\left. \begin{array}{l} \mathfrak{a} \in L_\infty \text{Alg}^{\text{ftp}} \\ X \in \text{SmothMfd} \end{array} \right\} \vdash \quad \Omega_{\text{dR}}^1(X; \mathfrak{a})_{\text{clsd}} := \text{Hom}_{\text{dgAlg}}(\text{CE}(\mathfrak{a}), \Omega_{\text{dR}}^\bullet(X)). \quad (16)$$

Namely, a graded algebra homomorphism from a CE-algebra sends the algebra generators  $\vec{b}$  to differential forms  $\vec{B}$ , and its respect for the differentials imposes on these differential forms exactly the closure/flatness condition. Examples are shown in (18).



| $L_\infty$ -algebra   | $\mathfrak{g}$                      | $\mathfrak{g}^\vee[1]$  | $d_{[-, \dots, -]}$   |
|---|-------------------------------------|---|---|
| line<br>Lie algebra   | $\mathfrak{u}(1)$                   | $\mathbb{R}\langle\omega_1\rangle$  | $d\omega_1 = 0$   |
| special unitary<br>Lie algebra  | $\mathfrak{su}(2)$                  | $\mathbb{R}\langle\omega_1^{(1)}, \omega_1^{(2)}, \omega_1^{(3)}\rangle$                      | $d\omega_1^{(i)} = \epsilon_{ijk}\omega_1^{(j)} \wedge \omega_1^{(k)}$  |
| line<br>Lie 2-algebra   | $b\mathfrak{u}(1)$                  | $\mathbb{R}\langle\omega_2\rangle$  | $d\omega_2 = 0$   |
| string<br>Lie 2-algebra<br>cf. [Fiorenza et al. 2014, §A]   | $\mathfrak{string}(3)$              | $\mathbb{R}\langle\omega_1^{(1)}, \omega_1^{(2)}, \omega_1^{(3)}, \omega_2\rangle$            | $d\omega_1^{(i)} = -\frac{1}{2}\epsilon_{ijk}\omega_1^{(j)} \wedge \omega_1^{(k)}$<br>$d\omega_2 = \epsilon_{ijk}\omega_1^{(i)} \wedge \omega_1^{(j)} \wedge \omega_1^{(k)}$        |
| line<br>Lie 3-algebra   | $b^2\mathfrak{u}(1)$                | $\mathbb{R}\langle\omega_3\rangle$  | $d\omega_3 = 0$   |
| T-duality<br>Lie 3-algebra<br>[Fiorenza et al. 2018, §7]  | $b\mathcal{T}_1$                    | $\mathbb{R}\langle\omega_2^{(i)}, \omega_2^{(B)}, h_3\rangle$                                 | $d\omega_2^{(i)} = 0$<br>$d\omega_2^{(B)} = 0$<br>$dh_3 = \omega_2^{(i)} \wedge \omega_2^{(B)}$   |
| line<br>Lie 4-algebra   | $b^3\mathfrak{u}(1)$                | $\mathbb{R}\langle\omega_4\rangle$  | $d\omega_4 = 0$   |
| M-theory gauge<br>Lie 7-algebra<br>[Sati 2010, §4]<br>[Sati & Voronov 2022, §2.2]   | $\mathfrak{I}S^4$                   | $\mathbb{R}\langle\omega_4, \omega_7\rangle$  | $d\omega_4 = 0$<br>$d\omega_7 = -\omega_4 \wedge \omega_4$  |
| cyclified<br>M-theory gauge<br>Lie 7-algebra<br>[Fiorenza et al. 2017, Ex. 3.3]<br>[Braunack-Mayer et al. 2019, Ex. 2.47] | $\mathfrak{I}(\mathcal{L}S^4//S^1)$ | $\mathbb{R}\langle\omega_2, \omega_4, \omega_6\rangle$<br>$\mathbb{R}\langle h_3, h_7\rangle$ | $dh_3 = 0$<br>$d\omega_2 = 0$<br>$d\omega_4 = h_3 \wedge \omega_2$<br>$d\omega_6 = h_3 \wedge \omega_4$<br>$dh_7 = -\frac{1}{2}\omega_4 \wedge \omega_4 + \omega_2 \wedge \omega_6$ |

(17)

|   |  |
|---|--|
| $\begin{array}{ccc} \text{CE}(b\mathfrak{u}(1)) & \longrightarrow & \Omega_{\text{dR}}^\bullet(X) \\ \omega_2 & \longmapsto & F \\ \downarrow d_{b\mathfrak{u}(1)} & & \downarrow d_{\text{dR}} \\ 0 & \longmapsto & 0 \end{array}$   | $\begin{array}{ccc} \text{CE}(\mathfrak{I}(\mathcal{L}S^4//S^1)) & \longrightarrow & \Omega_{\text{dR}}^\bullet(X) \\ h_3 & \longmapsto & H_3 \\ \downarrow d_{\mathfrak{I}(\mathcal{L}S^4//S^1)} & & \downarrow d_{\text{dR}} \\ 0 & \longmapsto & 0 \end{array}$       |
| $\begin{array}{ccc} \text{CE}(\mathfrak{string}(3)) & \longrightarrow & \Omega_{\text{dR}}^\bullet(X) \\ \omega_1^{(i)} & \longmapsto & A^i \\ \downarrow d_{\mathfrak{string}(3)} & & \downarrow d_{\text{dR}} \\ -\frac{1}{2}\epsilon_{ijk}\omega_1^{(j)} \wedge \omega_1^{(k)} & \longmapsto & -\frac{1}{2}A^i \wedge A^j \\ \omega_2 & \longmapsto & B \\ \downarrow d_{\mathfrak{string}(3)} & & \downarrow d_{\text{dR}} \\ \epsilon_{ijk}\omega_1^{(i)} \wedge \omega_1^{(j)} \wedge \omega_1^{(k)} & \longmapsto & \epsilon_{ijk}A^i \wedge A^j \wedge A^k \end{array}$ | $\begin{array}{ccc} \omega_2 & \longmapsto & F_2 \\ \downarrow d_{\mathfrak{I}(\mathcal{L}S^4//S^1)} & & \downarrow d_{\text{dR}} \\ 0 & \longmapsto & 0 \end{array}$  |
| $\begin{array}{ccc} \text{CE}(\mathfrak{I}S^4) & \longrightarrow & \Omega_{\text{dR}}^\bullet(X) \\ \omega_4 & \longmapsto & G_4 \\ \downarrow d_{\mathfrak{I}S^4} & & \downarrow d_{\text{dR}} \\ 0 & \longmapsto & 0 \end{array}$   | $\begin{array}{ccc} \omega_4 & \longmapsto & F_4 \\ \downarrow d_{\mathfrak{I}(\mathcal{L}S^4//S^1)} & & \downarrow d_{\text{dR}} \\ h_3 \wedge \omega_2 & \longmapsto & H_3 \wedge F_2 \end{array}$   |
| $\begin{array}{ccc} \omega_7 & \longmapsto & 2G_7 \\ \downarrow d_{\mathfrak{I}S^4} & & \downarrow d_{\text{dR}} \\ -\omega_4 \wedge \omega_4 & \longmapsto & -G_4 \wedge G_4 \end{array}$  | $\begin{array}{ccc} \omega_6 & \longmapsto & F_6 \\ \downarrow d_{\mathfrak{I}(\mathcal{L}S^4//S^1)} & & \downarrow d_{\text{dR}} \\ h_3 \wedge \omega_4 & \longmapsto & H_3 \wedge F_4 \end{array}$   |
|   | $\begin{array}{ccc} h_7 & \longmapsto & H_7 \\ \downarrow d_{\mathfrak{I}(\mathcal{L}S^4//S^1)} & & \downarrow d_{\text{dR}} \\ -\frac{1}{2}\omega_4 \wedge \omega_4 + \omega_2 \wedge \omega_6 & \longmapsto & -\frac{1}{2}F_4 \wedge F_4 + F_2 \wedge F_6 \end{array}$ |

(18)



**Flux densities satisfying Gauß law are flat  $L_\infty$ -valued differential forms.** Remarkably, it follows that polynomials  $\vec{P}$  defining Bianchi identities (6) and Gauss laws (14) are equivalently structure constants of  $L_\infty$ -algebras  $\mathfrak{a}$ , such that the Bianchi/Gauß law is the closure/flatness condition on  $\mathfrak{a}$ -valued forms:

$$\begin{array}{c}
\text{Sheaf of closed } L_\infty\text{-algebra-valued differential forms} \\
\Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} = \text{Hom}_{\text{dgAlg}}(\text{CE}(\mathfrak{a}), \Omega_{\text{dR}}^\bullet(-)) = \left\{ \vec{B} \equiv (B^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(-)) \mid d\vec{B} = \vec{P}(\vec{B}) \right\} \\
\text{insert spacetime manifold here} \\
\downarrow \\
\text{Chevalley-Eilenberg algebra of } L_\infty\text{-algebra} \\
\text{CE}(\mathfrak{a}) = \mathbb{R}[\{b_{\text{deg}_i}^{(i)}\}_{i \in I}] / (d\vec{b} = \vec{P}(\vec{b})) \\
\downarrow \\
\text{free differential graded-commutative algebra on these graded generators} \\
\downarrow \\
\text{equipped with these higher Lie brackets} \\
\mathfrak{a} = \mathbb{R}\langle \{v_{\text{deg}_i-1}^{(i)}\}_{i \in I} \rangle \quad [v^{(i)}, \dots, v^{(i_n)}] = \sum_{i \in I} P_{i \dots i_n}^{(i)} v^{(i)} \\
\downarrow \\
\text{vector space spanned by these graded generators} \\
\downarrow \\
L_\infty\text{-algebra}
\end{array} \tag{19}$$

With Prop. 2.14, this means:

**Proposition 3.1** (Flux solutions as flat  $L_\infty$ -valued forms). *Given a higher gauge theory of Maxwell-type (Def. 2.6) with Bianchi identities given by graded-symmetric polynomials  $\vec{P}$  (6) its space of flux densities solving the higher Maxwell equations is identified with the space of closed differential forms with coefficients in the  $L_\infty$ -algebra  $\mathfrak{a}$  on  $I$  deg-graded generators with structure constants  $\vec{P}$ :*

$$\begin{aligned}
\text{Space of flux densities on spacetime, solving the equations of motion} \quad \text{SolSpace}(X^D) &\equiv \left\{ \vec{F} \equiv \left( F^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^D) \right)_{i \in I} \mid \begin{array}{l} \text{Bianchi identities} \\ d\vec{F} = \vec{P}(\vec{F}) \\ \star \vec{F} = \vec{\mu}(\vec{F}) \\ \text{self-duality} \end{array} \right\} \text{ covariant form} \\
&\simeq_{\iota^*} \left\{ \vec{B} \equiv \left( B^{(i)} \in \Omega_{\text{dR}}^{\text{deg}_i}(X^d) \right)_{i \in I} \mid \begin{array}{l} \text{Gauß law} \\ d\vec{B} = \vec{P}(\vec{B}) \end{array} \right\} \text{ canonical form} \\
&\simeq \Omega_{\text{dR}}^1(X^d; \mathfrak{a})_{\text{clsd}} \quad \text{space of closed (flat) } \mathfrak{a}\text{-valued differential forms}
\end{aligned} \tag{20}$$

**Example 3.2.** The characteristic  $L_\infty$ -algebra of ordinary vacuum electromagnetism is the direct sum  $\text{bu}(1) \oplus \text{bu}(1)$  of two copies of the line Lie 2-algebra, which by the previous example and Prop. 3.1 corresponds to:

$$\text{SolSpace}_{\text{EM}}(X^3) \simeq \Omega_{\text{dR}}^1(X^3; \text{bu}(1) \times \text{bu}(1))_{\text{clsd}} \simeq \Omega_{\text{dR}}^2(X^3) \times \Omega_{\text{dR}}^2(X^3).$$

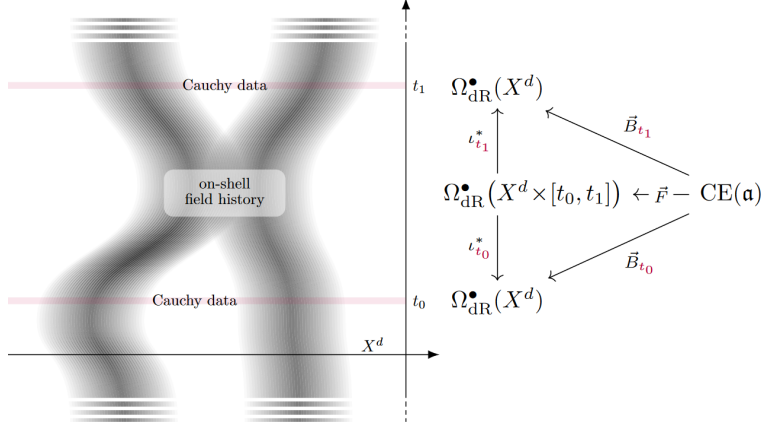
An element here corresponds to a pair  $(E, B)$ , where

- the magnetic flux density is the curvature  $B = \text{curv}(\hat{A})$  of the gauge potential which plays the role of the “canonical coordinate” on the field space;
- the electric flux density  $E$  serves as the corresponding “canonical momentum”.

**Total flux in non-abelian de Rham cohomology.** While, with Prop. 3.1, the Gauß law of the given higher gauge theory of Maxwell-type constrains the flux densities on any Cauchy surface  $X^d \hookrightarrow X^D$  to constitute a flat  $L_\infty$ -algebra valued differential form, the actual value of these differential forms depends on the Cauchy surface, which is an arbitrary choice. We should, therefore, regard as the *total flux* that aspect of the flux densities which is invariant under choice of Cauchy surfaces.

But since the Gauß law is (by Prop. 2.14) nothing but the restriction to the Cauchy surface of the Bianchi identities (on the duality-symmetric flux densities of Def. 2.6), the argument of Prop. 3.1 shows that this invariant aspect is the equivalence classes of flux densities under *concordance*:

$$\begin{array}{c}
\text{deformation of flux densities} \\
\vec{B}_0 \sim \vec{B}_1 \quad :\Leftrightarrow \quad \exists \vec{F} \in \Omega_{\text{dR}}^1(X^d \times [0, 1])_{\text{clsd}} \text{ with } \begin{cases} \iota_0^* \vec{F} = \vec{B}_0 \\ \iota_1^* \vec{F} = \vec{B}_1 \end{cases}
\end{array} \tag{21}$$



**Definition 3.3 (Non-abelian de Rham cohomology** [Fiorenza et al. 2023, Def. 6.3]). Given an  $L_\infty$ -algebra  $\mathfrak{a}$  and a smooth manifold  $X^d$ , we say that a pair of flat *closed  $\mathfrak{a}$ -valued differential forms*  $\vec{B}_0, \vec{B}_1 \in \Omega_{\text{dR}}^1(X^d; \mathfrak{a})_{\text{clsd}}$  (16) are *cohomologous* iff they are concordant: iff there exists a closed  $\mathfrak{a}$ -valued differential form  $\vec{F}$  on the cylinder over  $X^d$  whose pullback to the  $k$ th boundary component equals  $\vec{B}_k$ :

$$\vec{B}_0 \sim \vec{B}_1 \quad \Leftrightarrow \quad \exists \vec{F} \in \Omega_{\text{dR}}^1(X^d \times [0, 1]; \mathfrak{a})_{\text{clsd}} \quad \text{with} \quad \begin{cases} \vec{B}_1 = i_1^* \vec{F}, \\ \vec{B}_0 = i_0^* \vec{F}. \end{cases} \quad (22)$$

The quotient set by this equivalence relation is  $\mathfrak{a}$ -valued *nonabelian de Rham cohomology* of  $X^d$ :

$$H_{\text{dR}}^1(X^d; \mathfrak{a}) := \Omega_{\text{dR}}^1(X^d; \mathfrak{a})_{\text{clsd}} / \sim. \quad (23)$$

**Remark 3.4 (Conservation of total flux).**

Regarding the image of flux densities in non-abelian de Rham cohomology as expressing their *total flux* it follows immediately that:

*Total flux is conserved under time evolution.*

$$\begin{array}{ccc} \text{closed } \mathfrak{a}\text{-valued} & & \text{nonabelian } \mathfrak{a}\text{-valued} \\ \text{differential forms} & \longrightarrow & \text{de Rham cohomology} \\ \Omega_{\text{dR}}^1(X^d; \mathfrak{a})_{\text{clsd}} & \longrightarrow & H_{\text{dR}}^1(X^d; \mathfrak{a}) \\ \vec{B} & \longmapsto & [\vec{B}] \\ \text{flux densities} & & \text{total flux} \end{array}$$

**Remark 3.5 (Nonabelian de Rham cohomology as dg-homotopy classes).**

For comparison to the flux quantization rules discussed below in §3.2 it is useful to understand this equivalently [Fiorenza et al. 2023, Thm. 6.5] as the set of dg-homotopy classes of the corresponding dgc-homomorphisms:

$$[\vec{B}] \in H_{\text{dR}}^1(X^d; \mathfrak{a}) := \pi_0 \left\{ \begin{array}{ccc} & \text{cocycle (dga-hom)} & \\ & \vec{B} & \\ \text{deformation class} & & \text{de Rham cohomology} \\ \text{of flux densities} & & \\ \left[ \vec{B} \right] \in & H_{\text{dR}}^1(X^d; \mathfrak{a}) & := \pi_0 \left\{ \begin{array}{ccc} \Omega_{\text{dR}}^bullet(X^d) & \xrightarrow{\text{coboundary (concordance)}} & \text{CE}(\mathfrak{a}) \\ \uparrow & & \downarrow \\ & \vec{B}' & \\ & \text{another cocycle} & \end{array} \right\}. \end{array} \right. \quad (24)$$

**Example 3.6** ([Fiorenza et al. 2023, Prop. 6.4]). In the case of ordinary electromagnetism and abelian higher gauge fields, hence for  $\mathfrak{a} = b^n \mathfrak{u}(1)$  the line Lie  $n + 1$ -algebra, Def. 3.3 reduces to the ordinary notions:

- $\Omega_{\text{dR}}^1(X^d; b^n \mathbb{R})_{\text{clsd}} \simeq \Omega^{n+1}(X^d)_{\text{clsd}}$  are ordinary closed differential forms;
- concordance between these coincides with the coboundary relation in the ordinary de Rham complex;
- $H_{\text{dR}}^1(X^d; b^n \mathbb{R}) \simeq H_{\text{dR}}^{n+1}(X^d)$  is ordinary de Rham cohomology;

and since the latter also gives the periods of closed differential forms, this recovers indeed the usual notion of total (integrated) flux. Concretely, with  $F_2$  the flux density around a magnetic monopole of charge  $q$  (1), the total flux is as shown:

$$\begin{aligned} F_2 &:= q \text{dvol}_{S^2} \in \Omega_{\text{dR}}^2(S^2) \xrightarrow{p_{S^2}^*} \Omega_{\text{dR}}^2(\mathbb{R}^3 \setminus \{0\}) \simeq \Omega_{\text{dR}}^2(\mathbb{R}^3 \setminus \{0\}; b\mathbb{R}) \\ & \quad H_{\text{dR}}^2(\mathbb{R}^3 \setminus \{0\}) \simeq H_{\text{dR}}^2(S^2) \simeq \mathbb{R} \\ & \quad F_2 \quad \mapsto \quad [F_2] \quad \mapsto \quad \int_{S^2} F_2 = q. \end{aligned}$$

With on-shell flux densities thus understood as cocycles in nonabelian de Rham cohomology, we find their flux quantization laws among the corresponding torsion-ful nonabelian cohomology theories:

### 3.2 Flux quantization laws as Nonabelian cohomology

We explain how the  $\mathfrak{a}$ -valued nonabelian de Rham cohomology of the previous subsection receives character maps from generalized nonabelian cohomology theories whose classifying spaces  $\mathcal{A}$  have compatible rational Whitehead  $L_\infty$ -algebra  $\mathfrak{L}\mathcal{A} \simeq \mathfrak{a}$  – whence  $\mathcal{A}$  encodes a flux quantization law for Bianchi identities characterized by  $\mathfrak{a}$ , and lifting through the  $\mathcal{A}$ -character map corresponds to choices of charge quanta which source given total flux.

**Classifying spaces for generalized cohomology.** It is a classical fact of algebraic topology — which may have remained somewhat underappreciated in mathematical physics — that reasonable generalized cohomology theories have classifying spaces  $\mathcal{A}$ , in that the sets of cohomology classes assigned to a given domain space (which we take to be a smooth manifold  $X^d$ ) are in natural bijection with the homotopy classes  $\pi_0 \text{Map}(X, \mathcal{A})$  of continuous maps from  $X$  into  $\mathcal{A}$ . (Throughout, it is only the homotopy type of  $\mathcal{A}$  that matters.)

The archetypical examples are Eilenberg-MacLane spaces like  $K(\mathbb{Z}, n)$  which classify ordinary cohomology such as integral cohomology, in any degree  $n$ . As  $n$  ranges, these EM-spaces happen to be loop spaces of each other, up to weak homotopy equivalence:  $K(\mathbb{Z}, n) \simeq \Omega K(\mathbb{Z}, n+1)$ .

Generalizing from this classical example, one considers Whitehead-generalized cohomology theories which are classified by any sequences of pointed topological spaces  $\{E_n\}_{n \in \mathbb{N}}$  equipped with weak homotopy equivalences  $E_n \simeq \Omega E_{n+1}$ , called a *spectrum of spaces*.

This implies that each  $E_n$  is an infinite-loop space, which makes them be “abelian  $\infty$ -groups”, reflecting the fact that the homotopy classes of maps into these spaces indeed have the structure of abelian groups.

Perhaps the most familiar example of such *abelian* generalized cohomology is topological K-theory, whose classifying space  $KU_0$  may be identified with the space of Fredholm operators on an infinite-dimensional separable complex Hilbert space.

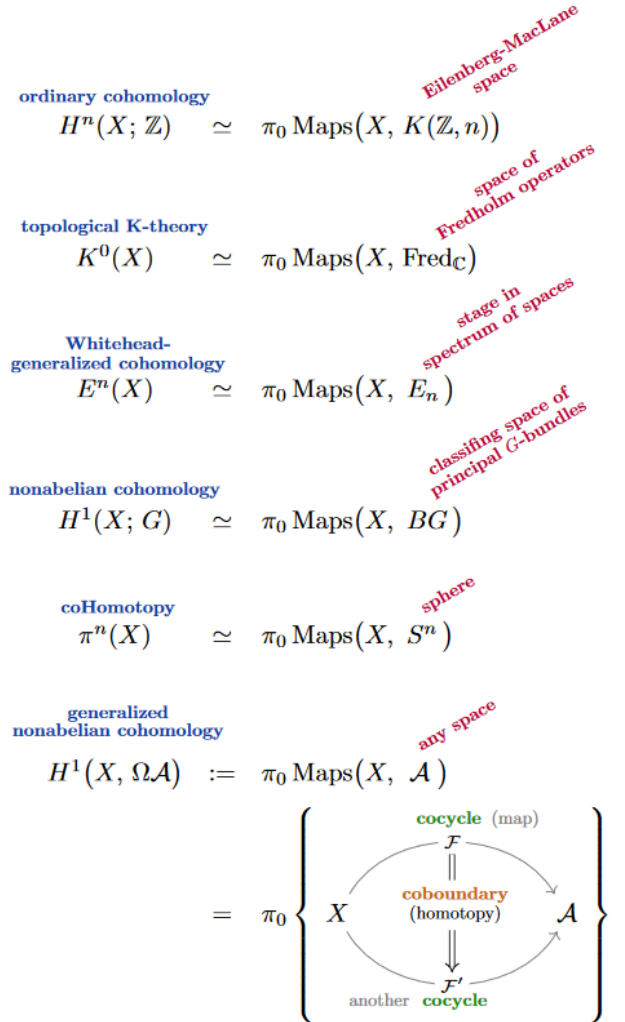
While Whitehead-generalized cohomology theory has received so much attention that it is now widely understood as the default or even the exclusive meaning of “generalized cohomology”, historically long preceding it is the *non-abelian cohomology* of Chern-Weil theory, classified by the original classifying spaces  $BG$  of compact Lie groups  $G$ .

Unless  $G$  happens to be abelian itself, this nonabelian cohomology does not assign abelian cohomology groups, nor even any groups at all, but just pointed cohomology sets. Nevertheless, as the historical name “nonabelian cohomology” clearly indicates, these systems of cohomology sets may usefully be regarded as constituting a kind of cohomology theory, too.

In this vein one may observe [Fiorenza et al. 2023, §2] that (the homotopy type of) every connected space  $\mathcal{A}$  is equivalently the classifying space of an infinity-group  $\Omega\mathcal{A}$ , namely of its own loop space regarded as an  $A_\infty$ -space under concatenation of loops), so that homotopy classes of maps into any connected space are examples of an evident generalization of Chern-Weil-style nonabelian cohomology.

A fundamental and historical example of such “truly-generalized” nonabelian cohomology is CoHomotopy, whose classifying spaces are the (homotopy types) of spheres. Notice that “generalized nonabelian cohomology” is really “not necessarily abelian”. It subsumes all the other cases: For  $E_\bullet$  a spectrum we have:  $E^n(X) \simeq H^1(X; \Omega E_n)$

**Character maps on generalized cohomology.** Moreover, it is classical that, over smooth manifolds, reasonable cohomology theories have their non-torsion content reflected in de Rham cohomology via *character maps*:



$$\begin{array}{llll}
\text{Ordinary integral cohomology} & H^n(X; \mathbb{Z}) & \xrightarrow{\text{de Rham map}} & H_{\text{dR}}^n(X) \simeq \text{Hom}_{\text{dgAlg}_{\mathbb{R}}}(\mathbb{R}[\omega_n], H_{\text{dR}}^\bullet(X)) & \text{differential forms in degree } n \\
\text{Traditional nonabelian cohomology} & H^1(X; G) & \xrightarrow{\text{Chern-Weil homomorphism}} & \text{Hom}_{\text{dgAlg}_{\mathbb{R}}}(\text{inv}^\bullet(\mathfrak{g}), H_{\text{dR}}^\bullet(X)) & \text{differential forms for } \mathfrak{g}\text{-invariant polynomials} \\
\text{Topological K-theory} & K^0(X) & \xrightarrow{\text{Chern character}} & \text{Hom}_{\text{dgAlg}_{\mathbb{R}}}(\mathbb{R}[\omega_0, \omega_2, \omega_4, \dots], H_{\text{dR}}^\bullet(X)) & \text{differential forms in every even degree} \\
\text{abelian Whitehead-generalized cohomology} & E^n(X) & \xrightarrow{\text{Chern-Dold character}} & \text{Hom}_{\text{dgAlg}_{\mathbb{R}}}(\wedge^\bullet(\pi_\bullet(E) \otimes_{\mathbb{Z}} \mathbb{R})^\vee, H_{\text{dR}}^{\bullet+n}(X)) & \text{differential forms for rational homotopy groups of the classifying space} \\
\text{Generalized non-abelian cohomology} & H^1(X; \Omega\mathcal{A}) & \xrightarrow{\text{nonabelian character}} & H_{\text{dR}}^1(X; \mathfrak{L}\mathcal{A}) := \text{Hom}_{\text{dgAlg}_{\mathbb{R}}}(\text{CE}(\mathfrak{L}\mathcal{A}), \Omega_{\text{dR}}^\bullet(X)) / \sim & \text{differential forms with coefficients in Whitehead } L_\infty\text{-algebra}
\end{array} \tag{25}$$

The nonabelian character in the generality of generalized non-abelian cohomology, such as CoHomotopy, is due to [Fiorenza et al. 2023, Def. IV.2], constructed via the fundamental theorem of dg-algebraic rational homotopy theory. We next survey how this works.

The key point is that rational homotopy theory characterizes the non-torsion content of (the homotopy type of) a (classifying) space  $\mathcal{A}$  by an  $L_\infty$ -algebra-approximation  $\mathfrak{L}\mathcal{A}$  to its loop space  $\infty$ -group  $\Omega\mathcal{A}$ .

**Proposition 3.7. Quillen-Sullivan-Whitehead  $L_\infty$ -algebra** cf. [Fiorenza et al. 2023, Prop. 4.23, 5.6 & 5.13]

For a topological space  $\mathcal{A}$  which is

- simply connected:  $\pi_0\mathcal{A} = *$ ,  $\pi_1\mathcal{A} = 1$ ;
- of rational finite type:  $\dim_{\mathbb{Q}}(H^n(\mathcal{A}; \mathbb{Q})) < \infty$

there is a polynomial dgc-algebra over  $\mathbb{R}$ , unique up to dga-isomorphism, whose

- generators are the  $\mathbb{R}$ -rational homotopy groups of  $\mathcal{A}$ ,

$$\text{CE}(\mathfrak{L}\mathcal{A}) = \left( \wedge^\bullet(\pi_\bullet(\Omega\mathcal{A}) \otimes_{\mathbb{Z}} \mathbb{R})^\vee, d_{\text{CE}(\mathfrak{L}\mathcal{A})} \right)$$

- cochain cohomology is the ordinary real cohomology of  $\mathcal{A}$

$$H^\bullet(\text{CE}(\mathfrak{L}\mathcal{A})) = H^\bullet(\mathcal{A}; \mathbb{R}).$$

This dgc-algebra is known as the *minimal Sullivan model* of  $\mathcal{A}$ . By (15) it is the Chevalley-Eilenberg algebra of an  $L_\infty$ -algebra which we denote  $\mathfrak{L}\mathcal{A}$ : The Whitehead bracket algebra structure on the  $\mathbb{R}$ -rational homotopy groups of the loop space (think of "l(-)" as standing for "Lie" or for "loops"):

$$\mathfrak{L}\mathcal{A} = \pi_\bullet(\Omega\mathcal{A}) \otimes_{\mathbb{Z}} \mathbb{R}. \tag{26}$$

Some **examples** for how to use Prop. 3.7 to compute Sullivan models and hence  $\mathbb{R}$ -Whitehead  $L_\infty$ -algebras  $\mathfrak{L}\mathcal{A}$  of spaces are spelled out on p. 21.

Many of the Whitehead  $L_\infty$ -algebras of familiar spaces do not have established names as  $L_\infty$ -algebras. An interesting exception is the Whitehead  $L_\infty$ -algebra of the 4-sphere, which happens to coincide [Sati & Voronov 2022, (13)] with what in  $D = 11$  supergravity-theory is known (quite independently) as the gauge algebra of the C-field [Cremmer et al. 1998, (2.6)][Sati 2010, §4]:

| Homotopy type<br>(topological space)<br>$\mathcal{A}$ | Sullivan model<br>("FDA")<br>$\text{CE}(\mathfrak{L}\mathcal{A})$  | Whitehead $L_\infty$ -algebra<br>(strong homotopy Lie algebra)<br>$\mathfrak{L}\mathcal{A}$                                     |
|---|--|---|
| $S^4$<br><br>4-sphere                                 | $\mathbb{R} \left[ \begin{array}{l} \omega_7, \\ \omega_4 \end{array} \right] / \left( \begin{array}{l} d\omega_7 = -\frac{1}{2}\omega_4 \wedge \omega_4 \\ d\omega_4 = 0 \end{array} \right)$<br><br>abstract Bianchi identity of<br>duality-symmetric C-field fluxes | $\mathbb{R} \left\langle \begin{array}{l} v_6, \\ v_3 \end{array} \right\rangle, [v_3, v_3] = v_6$<br><br>C-field gauge algebra |

(27)

**Remark 3.8 (Prefactors in Sullivan algebras).** As stated so far, the ubiquitous prefactor  $-1/2$  is pure convention, due to the freedom of rescaling generators by rational (or even real) numbers while retaining dga-isomorphy. However, this factor is fixed by requiring certain integrality properties of the generators, see [Fiorenza et al. 2021a, Prop. 4.6]. This becomes relevant when regarding the lift back from  $\mathfrak{L}S^4$  to  $\mathcal{A} \equiv S^4$  as a flux quantization law, because then it implies that the C-field flux densities  $G_4$  and  $G_7$  in the image of the normalized generators  $\omega_4$  and  $\omega_7$  satisfy expected integrality conditions [Fiorenza et al. 2021a, Thm. 4.8]. We discuss this further below in §4.2.

|  |  |
|--|--|
| <p><b>Circle:</b> <math>\mathcal{A} \equiv S^1 \simeq B\mathbb{Z}</math>.<br/> <math>(\pi_*(S^1) \otimes_{\mathbb{Z}} \mathbb{R})^\vee \simeq \mathbb{R}\langle \omega_1 \rangle</math>, <math>H^\bullet(S^1; \mathbb{R}) \simeq \mathbb{R}[\omega_1]</math><br/> Since <math>\mathbb{R}[\omega_1]</math> is already the correct cohomology ring, it must be that <math>d_{S^1} = 0</math> and hence</p> $\text{CE}(\mathfrak{t}S^1) \simeq \mathbb{R}[\omega_1] / (d\omega_1 = 0)$  | <p>While the circle is not simply connected, it is a “nilpotent space”, and Sullivan’s theorem actually applies in this generality.<br/> Nilpotent spaces have nilpotent fundamental group (e.g.: abelian) such that all higher homotopy groups are nilpotent modules (e.g.: trivial modules).</p>   |
| <p><b>2-Sphere:</b> <math>\mathcal{A} \equiv S^2</math>.<br/> <math>(\pi_*(S^2) \otimes_{\mathbb{Z}} \mathbb{R})^\vee \simeq \mathbb{R}\langle \omega_2, \omega_3 \rangle</math>, <math>H^\bullet(S^2; \mathbb{R}) \simeq \mathbb{R}[\omega_2] / (\omega_2^2)</math><br/> The differential on <math>\mathbb{R}[\omega_2, \omega_3]</math> needs to remove <math>\omega_2^2</math> and <math>\omega_3</math> from cohomology, hence it must be that:</p> $\text{CE}(\mathfrak{t}S^2) \simeq \mathbb{R} \left[ \begin{array}{c} \omega_3, \\ \omega_2 \end{array} \right] / \left( \begin{array}{l} d\omega_3 = -\frac{1}{2}\omega_2 \wedge \omega_2 \\ d\omega_2 = 0 \end{array} \right)$   | <p>The homotopy group corresponding to the generator <math>\omega_3</math> is that represented by the <i>complex Hopf fibration</i></p> $S^3 \xrightarrow{h_C} S^2.$   |
| <p><b>3-Sphere:</b> <math>\mathcal{A} \equiv S^3</math>.<br/> <math>(\pi_*(S^3) \otimes_{\mathbb{Z}} \mathbb{R})^\vee \simeq \mathbb{R}\langle \omega_3 \rangle</math>, <math>H^\bullet(S^3; \mathbb{R}) \simeq \mathbb{R}[\omega_3]</math><br/> Since <math>\mathbb{R}[\omega_3]</math> is already the correct cohomology ring, it must be that <math>d_{S^3} = 0</math> and hence</p> $\text{CE}(\mathfrak{t}S^3) \simeq \mathbb{R}[\omega_3] / (d\omega_3 = 0)$   | <p>While <math>S^3 \simeq \text{SU}(2)</math>, we see that <math>\mathfrak{t}\text{SU}(2)</math> is different from <math>\mathfrak{su}(2)</math>. But the former captures the cocycles of the latter:</p> $\begin{array}{ccc} \mathfrak{su}(2) & \longrightarrow & \mathfrak{t}\text{SU}(2) \\ \text{CE}(\mathfrak{su}(2)) & \longleftarrow & \text{CE}(\mathfrak{t}\text{SU}(2)) \\ \text{tr}(-, [-, -]) & \longleftarrow & \omega_3 \end{array}$ |
| <p><b>4-Sphere:</b> <math>\mathcal{A} \equiv S^4</math>.<br/> <math>(\pi_*(S^4) \otimes_{\mathbb{Z}} \mathbb{R})^\vee \simeq \mathbb{R}\langle \omega_4, \omega_7 \rangle</math>, <math>H^\bullet(S^4; \mathbb{R}) \simeq \mathbb{R}[\omega_4] / (\omega_4^2)</math><br/> The differential on <math>\mathbb{R}[\omega_4, \omega_7]</math> needs to remove <math>\omega_4^2</math> and <math>\omega_7</math> from cohomology, hence it must be that:</p> $\text{CE}(\mathfrak{t}S^4) \simeq \mathbb{R} \left[ \begin{array}{c} \omega_7, \\ \omega_4 \end{array} \right] / \left( \begin{array}{l} d\omega_7 = -\frac{1}{2}\omega_4 \wedge \omega_4 \\ d\omega_4 = 0 \end{array} \right)$   | <p>The homotopy group corresponding to the generator <math>\omega_7</math> is that represented by the <i>quaternionic Hopf fibration</i></p> $S^7 \xrightarrow{h_{\mathbb{H}}} S^4$  |
| <p><b>Complex Projective space:</b> <math>\mathcal{A} \equiv \mathbb{C}P^n</math>.<br/> <math>(\pi_*(\mathbb{C}P^n) \otimes_{\mathbb{Z}} \mathbb{R})^\vee \simeq \mathbb{R}\langle \omega_2, \omega_{2n+1} \rangle</math>,<br/> <math>H^\bullet(\mathbb{C}P^n; \mathbb{R}) \simeq \mathbb{R}[\omega_2] / (\omega_2^{n+1})</math><br/> The differential on <math>\mathbb{R}[\omega_2, \omega_{2n+1}]</math> needs to remove <math>\omega_2^{n+1}</math> from cohomology, hence it must be that:</p> $\text{CE}(\mathfrak{t}\mathbb{C}P^n) \simeq \mathbb{R} \left[ \begin{array}{c} \omega_{2n+1}, \\ \omega_2 \end{array} \right] / \left( \begin{array}{l} d\omega_{2n+1} = -\omega_2^{n+1} \\ d\omega_2 = 0 \end{array} \right)$ | <p>This is related to the above sequence of examples by the fact that <math>\mathbb{C}P^n</math> is an <math>S^1</math>-quotient of <math>S^{2n+1}</math>:</p> $\begin{array}{ccc} S^1 & \hookrightarrow & S^{2n+1} \\ & & \downarrow \\ & & \mathbb{C}P^n \end{array}$  |
| <p><b>Infinite Projective space:</b> <math>\mathcal{A} \equiv \mathbb{C}P^\infty \simeq \text{BU}(1) \simeq B^2\mathbb{Z}</math>.<br/> <math>(\pi_*(\mathbb{C}P^\infty) \otimes_{\mathbb{Z}} \mathbb{R})^\vee \simeq \mathbb{R}\langle \omega_2 \rangle</math>, <math>H^\bullet(\mathbb{C}P^\infty; \mathbb{R}) \simeq \mathbb{R}[\omega_2]</math><br/> Since <math>\mathbb{R}[\omega_2]</math> is already the correct cohomology ring, it must be that <math>d_{\mathbb{C}P^\infty} = 0</math>:</p> $\text{CE}(\mathfrak{t}\mathbb{C}P^\infty) \simeq \mathbb{R}[\omega_2] / (d\omega_2 = 0)$   | <p>This is the Lie 2-algebra of the shifted circle group:</p> $\mathfrak{t}\text{BU}(1) \simeq \mathfrak{b}\mathfrak{u}(1)$  |
| <p><b>Eilenberg-MacLane space:</b> <math>\mathcal{A} \equiv B^n\text{U}(1) \simeq B^{n+1}\mathbb{Z}</math>.<br/> <math>(\pi_*(B^{n+1}\mathbb{Z}) \otimes_{\mathbb{Z}} \mathbb{R})^\vee \simeq \mathbb{R}\langle \omega_{n+1} \rangle</math>, <math>H^\bullet(B^{n+1}\mathbb{Z}; \mathbb{R}) \simeq \mathbb{R}[\omega_{n+1}]</math><br/> Since <math>\mathbb{R}[\omega_{n+1}]</math> is already the correct cohomology ring, it must be that <math>d_{B^{n+1}\mathbb{Z}} = 0</math>:</p> $\text{CE}(\mathfrak{t}B^{n+1}\mathbb{Z}) \simeq \mathbb{R}[\omega_{n+1}] / (d\omega_{n+1} = 0)$   | <p>This is the Lie <math>(n+1)</math>-algebra of the circle <math>(n+1)</math>-group:</p> $\mathfrak{t}B^n\text{U}(1) \simeq \mathfrak{b}^n \mathfrak{u}(1)$   |
| <p><b>Classifying space:</b> <math>\mathcal{A} \equiv BG</math> of cpt. 1-conn. Lie group.<br/> <math>H^\bullet(BG; \mathbb{R}) \simeq \text{inv}^\bullet(\mathfrak{g})</math> the invar. polynomials on Lie alg. (Chern-Weil theory)<br/> Since <math>H^\bullet(BG; \mathbb{R})</math> is already a free graded-symmetric ring it must be that <math>d_{BG} = 0</math> (cf. [Fiorenza et al. 2023, Lem. 8.2]):</p> $\text{CE}(\mathfrak{t}BG) \simeq \text{inv}^\bullet(\mathfrak{g}) / (d_{BG} = 0)$   | <p><math>\mathfrak{t}BG</math> captures all the curvature invariants hence all the invariant flux densities of <math>\mathfrak{g}</math>-connections <math>A \in \Omega_{\text{dR}}^1(X) \otimes \mathfrak{g}</math>,</p> <p>e.g. <math>\text{CE}(\mathfrak{t}\text{BSU}(2)) \longrightarrow \Omega_{\text{dR}}^\bullet(X)</math></p> $\text{tr}(-, -) \mapsto \delta_{ij} F_A^{(i)} \wedge F_A^{(j)}$   |

**Rational homotopy theory: Discarding torsion in nonabelian cohomology.** From the perspective (above) that any topological space  $\mathcal{A}$  serves as the classifying space of a generalized nonabelian cohomology theory, the idea of rational homotopy theory (survey in [Hess 2006]; [Fiorenza et al. 2023, §4]) becomes that of extracting the *non-torsion* content of such a cohomology theory, which we will see is, over smooth manifolds, that shadow of it that is reflected in the non-abelian de Rham cohomology (Def. 3.3) of  $\mathcal{A}$ -valued differential forms.

|  |                              |             |                    |
|--|------------------------------|-------------|--------------------|
| <span style="color: green;">regard spaces as classifying spaces</span> $\left\{ \begin{array}{l} \curvearrowright \\ \curvearrowright \end{array} \right.$ | <b>Homotopy theory</b>       | Rational    | Sullivan model     |
|  | <b>Nonabelian cohomology</b> | Non-torsion | de Rham cohomology |

(28)

Now, in a sense, the signature of any  $\mathcal{A}$ -cohomology theory is its (reduced) cohomology groups on spheres, equal to the homotopy groups of the classifying space:

$$\text{reduced } \mathcal{A}\text{-cohomology of the } n\text{-sphere } \tilde{H}^1(S^n; \Omega\mathcal{A}) \equiv \pi_0 \text{Map}^{*/}(S^n, \mathcal{A}) \equiv \pi_n(\mathcal{A}) \text{ } n\text{th homotopy group of classifying space}$$

Assuming throughout (for ease of exposition) that  $\mathcal{A}$  is simply-connected, the remaining non-trivial homotopy groups are abelian  $\pi_{n \geq 2}(i) \in \text{AbGrp}$ . Discarding torsion elements (nilpotent group elements) from these groups is achieved by tensoring with the abelian group of rational numbers:

$$\begin{array}{ccc} \text{reduced } \mathcal{A}\text{-cohomology of the } n\text{-sphere } \tilde{H}^1(S^n; \Omega\mathcal{A}) \simeq \pi_n(i) & \xrightarrow{\text{rationalization}} & \pi_n(i) \otimes_{\mathbb{Z}} \mathbb{Q} \text{ rationalized } \mathcal{A}\text{-cohomology of the } n\text{-sphere} \\ [c] \text{ with } k \cdot [c] = 0 & \longmapsto & [c] \otimes 1 = [c] \otimes k \cdot \frac{1}{k} = k \cdot [c] \otimes \frac{1}{k} = 0 \end{array}$$

This is a “projection operation” (jargon: “localization”), in that doing it twice has no further effect:

$$\begin{array}{ccc} \text{double rationalization } \pi_n(i) \otimes_{\mathbb{Z}} \mathbb{Q} \otimes_{\mathbb{Z}} \mathbb{Q} & \xleftarrow{\text{isomorphic}} & \pi_n(i) \otimes_{\mathbb{Z}} \mathbb{Q} \text{ single rationalization} \\ [c] \otimes \frac{p_1}{q_1} \otimes \frac{p_2}{q_2} = [c] \otimes \frac{p_1}{q_1} \otimes q_1 \frac{p_2}{q_1 q_2} & \longleftrightarrow & [c] \otimes \frac{p_1 p_2}{q_1 q_2} \end{array}$$

Hence to have a classifying space for the non-torsion part of  $\mathcal{A}$ -cohomology means to ask for:

**The rationalization of  $\mathcal{A}$ :**

|  |   |      |
|--|---|------|
| A topological space  | $L^{\mathbb{Q}}\mathcal{A}$   | (29) |
| all whose homotopy groups have the structure of $\mathbb{Q}$ -vector spaces                    | $\pi_n(L^{\mathbb{Q}}\mathcal{A}) \in \text{Mod}_{\mathbb{Q}}$  |      |
| equipped with a map from $\mathcal{A}$   | $\mathcal{A} \xrightarrow{\eta_{\mathcal{A}}^{\mathbb{Q}}} L^{\mathbb{Q}}\mathcal{A}$   |      |
| which induces isomorphisms on rationalized homotopy groups and is universal with this property | $\pi_n(i) \otimes_{\mathbb{Z}} \mathbb{Q} \xrightarrow[\sim]{\eta_{\mathcal{A}}^{\mathbb{Q}} \otimes_{\mathbb{Z}} \mathbb{Q}} \pi_n(L^{\mathbb{Q}}\mathcal{A}) \otimes_{\mathbb{Z}} \mathbb{Q}$ |      |

Notice that infinitely many spaces  $\mathcal{A}$  share the same rationalization, whence the choice of such an  $\mathcal{A}$  as a flux quantization law below is genuine further information.

For example, the rationalization of an integral Eilenberg-MacLane space  $B^{\mathbb{Z}} \equiv K(\mathbb{Z}, n)$  classifies ordinary rational cohomology, mapping to ordinary de Rham cohomology:

$$\begin{array}{ccccccc} \text{integral EM-space } B^n \mathbb{Z} & \xrightarrow[\eta_{B^n \mathbb{Z}}^{\mathbb{Q}}]{\text{rationalization}} & L^{\mathbb{Q}} B^n \mathbb{Z} \simeq B^n \mathbb{Q} & \xrightarrow[B^n((-) \otimes_{\mathbb{Q}} \mathbb{R})]{\text{extension of scalars}} & B^n \mathbb{R} & & \text{real EM space} \\ \pi_0 \text{Map}(X, B^n \mathbb{Z}) & \xrightarrow[\pi_0 \text{Map}(X, \eta_{B^n \mathbb{Z}}^{\mathbb{Q}})]{} & \pi_0 \text{Map}(X, B^n \mathbb{Q}) & \xrightarrow[\pi_0 \text{Map}(X, B^n((-) \otimes_{\mathbb{Q}} \mathbb{R}))]{} & \pi_0 \text{Map}(X, B^n \mathbb{R}) & & \\ \wr & & \wr & & \wr & & \\ H^n(X; \mathbb{Z}) & \xrightarrow[\text{integral ordinary cohomology}]{\text{cohomology operation}} & H^n(X; \mathbb{Q}) & \xrightarrow[\text{rational ordinary cohomology}]{\text{cohomology operation}} & H^n(X; \mathbb{R}) & \xrightarrow[\text{real ordinary cohomology}]{\text{de Rham isomorphism}} & H^n_{\text{dR}}(X) \\ & & & & & & \text{de Rham cohomology} \\ & & & & & & \uparrow \\ & & & & & & \text{ordinary character map} \end{array} \quad (30)$$

We may regard this as the archetype of a *character map* and ask for its generalization to any  $\mathcal{A}$ -cohomology theory. The pivotal observation of [Fiorenza et al. 2023] is that for this purpose one may invoke the fundamental theorem

of dg-algebraic rational homotopy theory:

**The Fundamental Theorem of dg-Algebraic Rational Homotopy Theory** (review in [Fiorenza et al. 2023, Prop. 5.6]) says that the homotopy theory of rational spaces (simply-connected with fin-dim rational cohomology) is all encoded by their Whitehead  $L_\infty$ -algebras (26) over the rational numbers. In particular, for  $X$  a CW-complex, the homotopy classes of maps into the rationalization  $L^\mathbb{Q}\mathcal{A}$  (29) of a space  $\mathcal{A}$  is identified with dg-homotopy classes of homomorphisms from the rational Sullivan model of  $\mathcal{A}$  to the “piecewise  $\mathbb{Q}$ -polynomial de Rham complex” of the topological space  $X$ :

$$\text{Map}(X, L^\mathbb{Q}\mathcal{A})_{/\text{homotopy}} \simeq \text{Hom}_{\text{dgAlg}}\left(\text{CE}(\iota^\mathbb{Q}\mathcal{A}), \Omega_{\text{PLdR}}^\bullet(X)\right)_{/\text{concordance}}, \quad (31)$$

Observing that the right-hand side looks close to the definition of  $\mathcal{L}\mathcal{A}$ -valued de Rham cohomology (Def. 3.3), in order to actually connect to such smooth differential forms one needs to extend the ground field scalars from the rational numbers to the real numbers:

**Rational homotopy theory over the Reals** ([Bousfield & Gugenheim 1976], reviewed in [Fiorenza et al. 2023, Def. 5.7, Rem. 5.2, Prop. 5.8]). The construction (29) also works over  $\mathbb{R}$  (but is then not a “localization”) to give

**The  $\mathbb{R}$ -rationalization of  $\mathcal{A}$ :**

|  |   |      |
|--|---|------|
| A topological space  | $L^\mathbb{R}\mathcal{A}$   |      |
| equipped with a map  | $L^\mathbb{Q}\mathcal{A} \xrightarrow{\eta_{L^\mathbb{Q}\mathcal{A}}^{\text{ext}}} L^\mathbb{R}\mathcal{A}$   | (32) |
| which on homotopy groups<br>is extension of scalars<br>suitably universal as such. | $\pi_n(L^\mathbb{Q}\mathcal{A}) \xrightarrow[= (-) \otimes_{\mathbb{Q}} \mathbb{R}]{\pi_n(\eta_{L^\mathbb{Q}\mathcal{A}}^{\text{ext}})} \pi_n(L^\mathbb{R}\mathcal{A})$ |      |

With this “derived extension of scalars” [Fiorenza et al. 2023, Lem. 5.3] and for  $X$  a smooth manifold, the fundamental theorem (31) does relate to smooth differential forms via a non-abelian de Rham theorem [Fiorenza et al. 2023, Lem. 6.4, Thm. 6.5]:

$$\begin{array}{ccc}
\begin{array}{c} \text{non-abelian} \\ \text{rational cohomology} \\ H^1(X; L^\mathbb{Q}\Omega\mathcal{A}) \end{array} & \xrightarrow{\text{derived extension of scalars}} & \begin{array}{c} \text{non-abelian} \\ \text{real cohomology} \\ H^1(X; L^\mathbb{R}\Omega\mathcal{A}) \end{array} \\
\parallel & & \parallel \\
\pi_0\text{Map}(X, L^\mathbb{Q}\mathcal{A}) & \xrightarrow{\pi_0\text{Map}(X, \eta_{L^\mathbb{Q}\mathcal{A}}^{\text{ext}})} & \pi_0\text{Map}(X, L^\mathbb{R}\mathcal{A}) \\
\uparrow \wr & & \uparrow \wr \\
\text{Hom}_{\text{dgAlg}}(\text{CE}(\iota^\mathbb{Q}\mathcal{A}), \Omega_{\text{PLdR}}^\bullet(X))_{/\text{cncd}} & \xrightarrow[\text{extension of scalars}]{} & \text{Hom}_{\text{dgAlg}}(\text{CE}(\iota\mathcal{A}), \Omega_{\text{dR}}^\bullet(X))_{/\text{cncd}} \equiv H_{\text{dR}}^1(X; \mathcal{L}\mathcal{A}) \\
& & \text{non-abelian de Rham cohomology}
\end{array} \quad (33)$$

$\swarrow$  *non-abelian de Rham theorem*

**In abelian (i.e., Whitehead-generalized) cohomology theories** both the rationalization step and the subsequent extension of scalars to  $\mathbb{R}$  can be more easily described as forming the smash product of the coefficient spectrum with the rational Eilenberg-MacLane spectrum  $H\mathbb{R}$  [Fiorenza et al. 2023, Ex. 5.7]. This is how the Chern-Dold character map over  $\mathbb{R}$  is tacitly used in all the literature on abelian (Whitehead-generalized) differential cohomology theory (e.g. [Bunke & Nikolaus 2019, Def. 4.2]):

$$\begin{array}{ccc}
\text{Spectra} & \xrightarrow[\text{rationalization localization}]{(-) \wedge_{H\mathbb{Q}}} & \text{Spectra} & \xrightarrow[\text{extension of scalars}]{(-) \wedge_{H\mathbb{Q}} H\mathbb{R}} & \text{Spectra} \\
& \xrightarrow[\text{rationalization over } \mathbb{R}]{(-) \wedge_{H\mathbb{R}}} & & & 
\end{array} \quad (34)$$

The point of the non-abelian de Rham theorem (33) is to generalize the realification (34) of Whitehead-generalized cohomology to generalized non-abelian cohomology, such as to Cohomotopy; and the key result that makes this work is the fundamental theorem of dg-algebraic homotopy theory (31). This, ultimately, is the “reason” why  $L_\infty$ -valued differential forms relate fluxes to their flux-quantization laws.

**The general non-abelian character map** is now immediate [Fiorenza et al. 2023, Def. IV.2]: It is the cohomology operation induced by  $\mathbb{R}$ -rationalization of classifying spaces (32), seen under the non-abelian de Rham theorem (33):

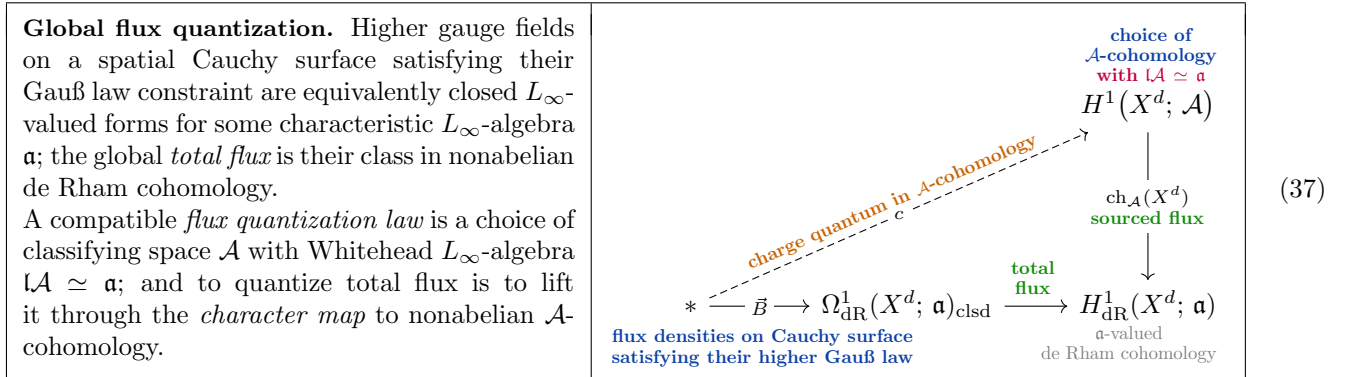
$$\begin{array}{c}
 \xrightarrow{\text{character map on } \mathcal{A}\text{-cohomology}} \\
 \begin{array}{ccccccc}
 H^1(X; \Omega\mathcal{A}) & \xrightarrow{\text{rationalization}} & H^1(X; L^{\mathbb{Q}}\Omega\mathcal{A}) & \xrightarrow{\text{extension of scalars}} & H^1(X; L^{\mathbb{R}}\Omega\mathcal{A}) & \xrightarrow{\text{nonabelian de Rham theorem}} & H^1_{\text{dR}}(X; \mathcal{L}\mathcal{A}) \\
 \parallel & & \parallel & & \parallel & & \parallel \\
 \pi_0\text{Map}(X, \mathcal{A}) & \xrightarrow{(\eta_{\mathcal{A}}^{\mathbb{Q}})_*} & \pi_0\text{Map}(X, L^{\mathbb{Q}}\mathcal{A}) & \xrightarrow{(\eta_{L^{\mathbb{Q}}\mathcal{A}}^{\text{ext}})_*} & \pi_0\text{Map}(X, L^{\mathbb{R}}\mathcal{A}) & \xrightarrow{\sim} & \text{Hom}_{\text{dgAlg}}(\text{CE}(\mathcal{L}\mathcal{A}), \Omega^{\bullet}_{\text{dR}}(X))_{/\text{cncrd}} \\
 & & & & & \text{fundamental theorem} & \\
 & & & & & \text{of dg-algebraic RHT} & 
 \end{array}
 \end{array} \quad (35)$$

All the classical abelian character maps (25) are special cases of this generalized nonabelian character [FSS23-Char, §7], but now examples in generalized nonabelian cohomology are also included; for instance, there is a character map on Cohomotopy-theory [Fiorenza et al. 2023, Ex. 6.11].

**Flux quantization in generalized nonabelian cohomology.** With the generalized nonabelian character map (35) in hand, we may finally state the general concept of global flux quantization. Recalling from §3.1 that the total flux of the higher gauge fields characterized by the  $L_{\infty}$ -algebra  $\mathfrak{a}$  in encoded in the  $\mathcal{L}\mathcal{A}$ -valued nonabelian de Rham cohomology of a Cauchy surface, it follows that for every choice of classifying space  $\mathcal{A}$  with  $\mathcal{L}\mathcal{A} \simeq \mathfrak{a}$  the nonabelian character map (35) may be understood as assigning to discrete charges embodied by  $\mathcal{A}$ -cohomology classes the corresponding total flux (thereby losing torsion-information encoded in the charges but not in the fluxes).

$$\begin{array}{ccc}
 \text{non-abelian} & & \text{non-abelian} \\
 \text{cohomology} & \xrightarrow{\text{non-abelian character}} & \text{de Rham cohomology} \\
 H^1(X^d; \Omega\mathcal{A}) & \xrightarrow{\text{ch}_{\mathcal{A}}} & H^1_{\text{dR}}(X^d; \mathcal{L}\mathcal{A}) \\
 c & \mapsto & [\vec{B}] \\
 \text{total charge} & & \text{total flux}
 \end{array} \quad (36)$$

Since the total charges in  $H^1(X^d; \Omega\mathcal{A})$  on the left form a discrete set, we may think of global flux quantization in  $\mathcal{A}$ -cohomology as *lifting* of total fluxes through this character map:



Notice that such a lift is not *just* a (quantization/discretization-) *condition* on the total fluxes, but also *extra structure*, namely a choice of torsion-component of the total charge reflected in total fluxes, as see in  $\mathcal{A}$ -cohomology:

Since the character map generally...

$$\begin{array}{ccccc}
 & \text{kernel consisting of} & & & \\
 & \text{all compatible charges} & & & \\
 H^1(X^d; \Omega\mathcal{A})_{[\vec{B}]} & \hookrightarrow & H^1(X^d; \Omega\mathcal{A}) & \longrightarrow & * \\
 \downarrow & & \downarrow & & \downarrow \\
 \dots \text{fails to be surjective, i.e., has a cokernel:} & & \text{ch}_{\mathcal{A}}(X^d) & & \\
 \Rightarrow \text{flux quantization is a condition on fluxes} & & \text{sourced flux} & & \\
 \dots \text{fails to be injective, i.e., has a kernel:} & & \downarrow & & \\
 \Rightarrow \text{flux quantization is a choice of "torsion"} & & * & \longrightarrow & H^1_{\text{dR}}(X; \mathcal{L}\mathcal{A}) / H^1(X^d; \Omega\mathcal{A}) \\
 & & \text{given total flux} & & \text{cokernel consisting of} \\
 & & & & \text{total fluxes violating} \\
 & & & & \text{the flux quantization law}
 \end{array}$$

However, in a higher gauge theory it is unnatural to have extra structure given by an equality of gauge equivalence classes, instead one should consider a gauge transformation between actual fields. Doing so leads to emergence of the gauge potentials and of the higher phase space stack of the theory, in the next subsection §3.3.



**Example 3.9 (Flux quantization laws for ordinary electromagnetism).** By Ex. 2.15, the characteristic  $L_\infty$ -algebra of vacuum electromagnetism is two copies of the line Lie 2-algebra  $bu(1)$ . This is the Whitehead  $L_\infty$ -algebra of the classifying space  $BU(1) \simeq B^2\mathbb{Z}$  and hence of its rationalization  $B^2\mathbb{Q}$ . Therefore — among many further variants — there are the following choices of flux quantization laws for ordinary electromagnetism:

|  |   |
|--|---|
| $\underbrace{B^2\mathbb{Q}}_{\text{mag}} \times \underbrace{B^2\mathbb{Q}}_{\text{el}}$                                    | This choice imposes essentially <i>no</i> flux quantization (it does rule out irrational total fluxes) and as such was the tacit choice since [Maxwell 1865] until [Dirac 1931].  |
| $\underbrace{B^2\mathbb{Z}}_{\text{mag}} \times \underbrace{B^2\mathbb{Q}}_{\text{el}}$                                    | This choice imposes integrality of magnetic charge but no further condition on electric flux — common choice since [Dirac 1931], for instance in [Alvarez 1985, p. 299] [Brylinski 1993, §7.1][Freed 2000, Ex. 2.1.2].              |
| $\underbrace{B^2\mathbb{Z}}_{\text{mag}} \times \underbrace{B^2\mathbb{Z}}_{\text{el}}$                                    | This choice imposes integrality of both magnetic and electric charge — considered in [Freed et al. 2007b][Freed et al. 2007c][Becker et al. 2017, Rem. 2.3] [Lazaroiu & Shahbazi 2022][Lazaroiu & Shahbazi 2023]                    |
| $\underbrace{B^2\mathbb{Z}}_{\text{mag}} \times \underbrace{BK}_{\text{el}} \times \underbrace{B^2\mathbb{Z}}_{\text{el}}$ | For a finite group $K \rightarrow \text{Aut}(\mathbb{Z})$ — this choice induces non-commutativity between EL/EL- and EL/M-fluxes, an example of a “non-evident” flux quantization condition considered in [Sati & Schreiber 2023c]. |

### 3.3 Phase spaces as Differential nonabelian cohomology

With higher Maxwell-type equations of flux given §2.4 and with a compatible flux/charge quantization law  $\mathcal{A}$  chosen §3.2, the full on-shell field content of the higher gauge theory (including the gauge potentials) and hence its phase space (p. 11) appears as the corresponding “moduli stack” of nonabelian differential cohomology  $\hat{\mathcal{A}}$  evaluated on any Cauchy surface.

**Smooth  $\infty$ -groupoids.** In order to describe this, we need to make free use of the notions of *smooth  $\infty$ -groupoids* presented by simplicial presheaves on Cartesian spaces. Exposition and pointers may be found in this collection [Schreiber 2024], a concise compilation of the technical details is given in [Fiorenza et al. 2023, §1], for more see [Sati & Schreiber 2021b, §3]:

In the present context, the point of smooth  $\infty$ -groupoids is that they provide the *joint* home for both *fluxes and charges*, namely for

- (i) the sheaves of closed  $\mathfrak{a}$ -valued differential forms  $\Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}}$ , which one may regard as the 0-truncated smooth moduli-stacks (namely: *smooth sets*, see [Giotopoulos & Sati 2023]) of flux densities
- (ii) the homotopy types of classifying spaces  $\mathcal{A}$ , which one may regard as the geometrically discrete moduli stacks of charges.

Once regarded in this joint context, these two moduli-stacks become comparable, as previewed in the diagram on the right, via an object to be denoted  $\int \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}}$  (described in a moment) and it is thereby that flux quantization in  $\mathcal{A}$ -cohomology may be imposed as a *local* structure that is equivalent to that of the non-perturbative higher gauge fields themselves.

$$\begin{array}{ccccc}
 \text{smooth sets} & & \text{smooth } \infty\text{-groupoids} & & \infty\text{-groupoids} \\
 \text{SmthSet} & \xleftarrow[\text{(no gauge transf.)}]{\text{smooth structure}} & \text{SmthGrpd}_\infty & \xleftarrow[\text{(discrete smooth struc.)}]{\text{higher gauge transf.}} & \text{Grpd}_\infty \\
 \parallel & & \parallel & & \parallel \\
 L^{\text{iso}} \text{PSh}(\text{CartSp}) & & L^{\text{heq}} \text{PSh}(\text{CartSp}, \text{PSh}(\Delta)_{\text{Kan}}) & & \text{PSh}(\Delta)_{\text{Kan}} \\
 \cup & & \cup & & \cup \\
 \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} & \xleftarrow[\text{shape unit}]{\eta^f} & \int \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} & \xleftarrow[\text{differential character}]{\text{ch}_\mathcal{A}} & \mathcal{A} \\
 \text{moduli of flux densities} & & \text{deformations of flux densities} & & \text{moduli of charges}
 \end{array} \tag{38}$$

Notice that when *presenting* smooth  $\infty$ -groupoids by the projective model structure on simplicial presheaves over the site of Cartesian spaces [Fiorenza et al. 2023, Ex. 1.20], these moduli objects appear canonically as fibrant objects, so that the only further step for computing the required derived mapping spaces into them is to cofibrantly resolve the domain manifold  $X^d$ . This is achieved by passage to the Čech nerve  $\widehat{X}^d$  of any *good* open cover  $\{U_j \xrightarrow{\iota_j} X^d\}_{j \in J}$  [Fiorenza et al. 2023, Ex. 1.24], as indicated on the right.

$$\begin{array}{ccc}
\text{Čech groupoid} & \xrightarrow{\text{local homotopy equivalence}} & \text{smooth manifold} \\
\widehat{X}^d & \xrightarrow{\text{lheq}} & X^d \\
\text{Plot}(\mathbb{R}^n \times \Delta^2, \widehat{X}^d) & \longrightarrow & \text{Plot}(\mathbb{R}^n \times \Delta^2, X^d) \\
\begin{array}{c}
(\mathbf{x}, j) \\
\swarrow (\mathbf{x}, i, j) \quad \searrow (\mathbf{x}, j, k) \\
(\mathbf{x}, i) \xrightarrow{(\mathbf{x}, i, k)} (\mathbf{x}, k) \\
\parallel \\
(\mathbf{x}, i, j, k) \\
\downarrow
\end{array} & \mapsto & \begin{array}{c}
\mathbf{x} \\
\parallel \\
\mathbf{x} \\
\parallel \\
\mathbf{x}
\end{array}
\end{array}$$

where  $\mathbb{R}^n \xrightarrow{(\mathbf{x}, i, j)} U_i \cap U_j \hookrightarrow U_i \hookrightarrow X$  etc.

$$\begin{array}{c}
\mathbb{R}^n \xrightarrow{(\mathbf{x}, i, j)} U_i \cap U_j \hookrightarrow U_i \hookrightarrow X \\
\downarrow \quad \downarrow \quad \downarrow \\
\mathbf{x} \xrightarrow{(\mathbf{x}, i)} \mathbf{x} \xrightarrow{\text{smooth}} \mathbf{x}
\end{array}$$

**Higher deformations of flux densities.** Recall (22) that a coboundary in  $\mathfrak{a}$ -valued de Rham cohomology is a “concordance” of flux densities, to be thought of as a path of smooth variations of the flux densities, subject to their Bianchi identities. But in higher gauge theories there are also non-trivial deformations-of-deformations varying over the higher dimensional  $n$ -simplices  $\Delta_{\text{geo}}^n$ , forming the following simplicial object:

$$\begin{array}{ccc}
\text{deformation paths} & & \\
\text{of flux densities} & & \\
\Omega_{\text{dR}}^1(- \times [0, 1]; \mathfrak{a})_{\text{clsd}} & \equiv & \left\{ \vec{B}_0 \xrightarrow{\vec{B}_{[0,1]}} \vec{B}_1 \right\} \\
\downarrow \text{take endpoint of} & \uparrow & \downarrow \text{take starting point} \\
& & \text{deformation path} \\
\Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} & \equiv & \{ \vec{B} \} \\
\text{flux densities satisfying} & & \\
\text{their Bianchi identities} & & 
\end{array}$$

$$\int \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} = \left( \begin{array}{ccc}
\begin{array}{c} 1 \\ \triangle \\ 0 \quad 2 \end{array} & \text{deformation paths} & \Omega_{\text{dR}}^1(- \times \Delta_{\text{geo}}^3; \mathfrak{a})_{\text{clsd}} \\
& \text{of deformation paths} & \\
& \text{of deformation paths} & \\
& \text{of flux densities} & \\
\downarrow & \downarrow & \downarrow \\
(-)_{[1,2,3]} & (-)_{[0,2,3]} & (-)_{[0,1,3]} & (-)_{[0,1,2]} \\
\downarrow & \downarrow & \downarrow & \downarrow \\
\begin{array}{c} 1 \\ \triangle \\ 0 \quad 2 \end{array} & \text{deformation paths} & \Omega_{\text{dR}}^1(- \times \Delta_{\text{geo}}^2; \mathfrak{a})_{\text{clsd}} & \equiv & \left\{ \begin{array}{c} \vec{B}_1 \\ \nearrow \vec{B}_{[0,1]} \quad \searrow \vec{B}_{[1,2]} \\ \parallel \\ \vec{B}_{[0,1,2]} \\ \downarrow \\ \vec{B}_0 \xrightarrow{\vec{B}_{[0,2]}} \vec{B}_2 \end{array} \right\} \\
& \text{of deformation paths} & \\
& \text{of deformation paths} & \\
& \text{of flux densities} & \\
\downarrow & \downarrow & \downarrow & \downarrow \\
(-)_{[0,1]} & (-)_{[0,2]} & (-)_{[1,2]} \\
\downarrow & \downarrow & \downarrow \\
0 \text{ --- } 1 & \text{deformation paths} & \Omega_{\text{dR}}^1(- \times \Delta_{\text{geo}}^1; \mathfrak{a})_{\text{clsd}} & \equiv & \left\{ \vec{B}_0 \xrightarrow{\vec{B}_{[0,1]}} \vec{B}_1 \right\} \\
& \text{of flux densities} & \\
\downarrow & \downarrow & \downarrow \\
& & \text{take endpoint of} & \uparrow & \text{take starting point} \\
& & \text{deformation path} & & \text{of deformation path} \\
& & (-)_{-1} & & (-)_{-0} \\
& & \downarrow & & \downarrow \\
& & \text{flux densities satisfying} & & \\
& & \text{their Bianchi identities} & & \\
& & \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} & \equiv & \{ \vec{B} \}
\end{array} \right)$$

This is a Kan-simplicial presheaf [Fiorenza et al. 2023, Def. 9.1, Prop. 5.10] that we may think of as the *shape* or *smooth path*  $\infty$ -groupoid [Sati & Schreiber 2021b, p. 144] of the 0-truncated moduli stack of flux densities. It is in

this object that flux densities become comparable to their charges:

- (i) There is an evident inclusion  $\Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} \xrightarrow{\text{shape unit}} \int \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}}$  [Fiorenza et al. 2023, (9.3)], which we may identify as the *shape unit* of the moduli of flux densities;
- (ii) given an identification  $\mathfrak{a} \simeq \mathbb{L}\mathcal{A}$  with a Whitehead  $L_\infty$ -algebra (37), then the fundamental theorem of dg-algebraic rational homotopy theory (31) furthermore says [Fiorenza et al. 2023, Lem. 9.1] that we have a (homotopy-) equivalence to the  $\mathbb{R}$ -rationalization  $L^{\mathbb{R}}\mathcal{A}$  of  $\mathcal{A}$  (32), so that rationalization gives a *differential character map* [Fiorenza et al. 2023, Def. 9.2]:

$$\mathcal{A} \xrightarrow{\text{rationalization}} L^{\mathbb{Q}}\mathcal{A} \xrightarrow{\text{extension of scalars}} L^{\mathbb{R}}\mathcal{A} \xrightarrow[\sim]{\text{fundamental thm. of RHT piecewise smooth version}} \int \Omega_{\text{dR}}^1(-; \mathbb{L}\mathcal{A})_{\text{clsd}}$$

$\underbrace{\hspace{15em}}_{\text{differential character map}}$

**Local flux quantization: Gauge potentials in differential cohomology.** This way one may now *locally* implement flux quantization, by taking the higher gauge field fields on  $X^d$  to be *homotopies* deforming flux densities  $\vec{B}$  into the differential character of local charges  $\chi$ .

On equivalence classes, this reproduces the quantization of total fluxes (37) and thereby lifts it to a local structure. Indeed, the higher gauge fields defined this way are the cocycles of the nonabelian *differential*  $\mathcal{A}$ -cohomology [Fiorenza et al. 2023, Def. 9.3].

$$\begin{array}{ccc} \widehat{X}^d & \xrightarrow{\text{charges } \chi} & \mathcal{A} \\ \downarrow \text{flux densities } \vec{B} & \swarrow \text{gauge potentials } \hat{A} & \downarrow \text{differential character } \text{ch} \\ \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} & \xrightarrow{\text{shape unit } \eta^f} & \int \Omega_{\text{dR}}^1(-; \mathfrak{a})_{\text{clsd}} \end{array} \quad (39)$$

In terms of physics these homotopies turn out to reflect the expected higher *gauge potentials* — which is not entirely obvious from the definition but follows by examination:

**Example 3.10 (Higher U(1)-gauge potentials in ordinary differential cohomology).** The data  $\hat{A} : \chi \Rightarrow \vec{B}$  in (39) is equivalent [Fiorenza et al. 2023, Prop. 9.5]...

- (A) ...for the case  $\mathfrak{a} = \mathfrak{bu}(1)$  and  $\mathcal{A} \equiv \text{BU}(1) = B^2\mathbb{Z}$  (Ex. 3.9):  
to that of connections on U(1)-principal bundles – which of course is the traditional data for the gauge potential of ordinary electromagnetism [Wu & Yang 1975], cf. [Wu & Yang 2006][Eguchi et al. 1980, Ex. 5.5][Rudolph & Schmidt 2017, §6.1];
- (B) ...for the case of  $\mathfrak{a} = \mathfrak{b}^2\mathfrak{u}(1)$  and  $\mathcal{A} = B^2\text{U}(1) = B^3\mathbb{Z}$ :  
to that of 3-cocycles in Deligne cohomology (often equivalently regarded as connections on “bundle gerbes”), this being the traditional understanding of the B-field gauge potential in string theory [Gawedzki 1988][Freed & Witten 1999, §6][Carey et al. 2004][Bonora et al. 2008];
- (C) ...for the case of  $\mathfrak{a} = \mathfrak{b}^3\mathfrak{u}(1)$  and  $\mathcal{A} = B^3\text{U}(1) = B^4\mathbb{Z}$ :  
to that of 4-cocycles in Deligne cohomology (also regarded as connections on “bundle 2-gerbes”), which was one of the proposed models for the C-field gauge potential (in the case where the class  $\frac{1}{2}p_1[TY^{11}]$  of space-time is even, otherwise the expected half-integral shift has been added “by hand”) [Aschieri & Jurčo 2004][Hopkins & Singer 2005][Diaconescu et al. 2007][Fiorenza et al. 2015a].

**Remark 3.11 (Shortcoming of higher U(1)-charge quantization).** For a long time, these examples 3.10 used to be the state of the art in understanding flux quantization of higher gauge fields. But notice that in all three items the flux-quantization of the duality-partner fields (and hence of the canonical momenta) have been ignored. For item (A) this can readily be rectified, since here the partner (electric) field can be flux-quantized in the same way (and later has been, Ex. 3.9), but in items (B) and (C) it is actually impossible to model the dual fields (with flux densities  $H_7$  and  $G_7$ , respectively) as higher U(1)-gauge fields (nor even as generalized higher abelian gauge fields, Ex. 3.12), since their Bianchi identities are non-linear (by Ex. 2.13 and Ex. 2.12, respectively), cf. §4.2.

**Example 3.12 (abelian Whitehead-generalized differential cohomology).** For  $E_\bullet$  a spectrum of spaces and  $\mathcal{A} \equiv E_n$ , the data  $\hat{A} : \chi \Rightarrow \vec{B}$  in (39) is equivalent [Fiorenza et al. 2023, Ex. 9.1] to cocycles in the “canonical” version of the differential cohomology theory  $\hat{E}^n(-)$ , as originally introduced in [Hopkins & Singer 2005], cf. [Bunke 2012, p. 88]; for exposition see also in this collection the contribution [Debray 2024].

For applications to flux quantization, the most prominent example of such abelian generalized differential cohomology remain flavors of differential K-theory, to which we come in §4.1.

## 4 Examples in String-/M-Theory

While flux quantization is an issue in any higher gauge theory, the examples where it has received most (essentially all) of the attention are those of evident relevance in string theory — which is what we focus on in the following.

While string theory is an attempt to understand the all-important but elusive non-perturbative behaviour of Yang-Mills theories (notably quantum chromodynamics) by regarding quarks confined by color flux tubes as endpoints of open strings stuck on intersecting branes in an unobserved higher dimensional spacetime (cf. [Polyakov 2012] [Hari Dass 2024]), ironically also string physics itself has really been understood only perturbatively (namely by replacing Feynman diagrams in ordinary worldline perturbative quantum field theory with worldsheet  $n$ -point functions of a 2d SCFT).

However, since flux quantization laws (as discussed in §3) are hypotheses/prescriptions for otherwise missing non-perturbative degrees of freedom of the string’s background fields, their investigation goes towards the heart of the open problem finding a non-perturbative completion of string theory itself, famous under the working title *M-theory* [Hořava & Witten 1996] [Duff 1996] [Duff 1999].

For instance, the traditional *Hypothesis K* (§4.1) that RR-field fluxes are quantized in topological K-theory has been motivated/justified [Witten 1998, §3] as describing – or in fact pre-scribing – the stable end results of the tachyon condensation of open string modes stretching between D-brane/anti D-brane pairs, a process which cannot be followed by string perturbation theory, but which is expected (“Sen’s conjecture” [Sen 1998]) to find the non-perturbative true vacuum state where D-brane/anti D-brane pairs have mutually annihilated as far as possible. Indeed, at least in practice, RR-field flux quantization in topological K-theory has become the widely-accepted definition of stable D-brane vacua, and as such must be understood as a partial proposal for the nature of non-perturbative string theory.

On the other hand, strongly-coupled string theory at large-scale/low-energy is also famously argued to be described by D=11 supergravity, whence it stands to reason that flux quantization of the supergravity C-field in 11d should go further still towards the full non-perturbative definition of string theory (hence of M-theory). While the details are subtle and generally deserve more attention, the systematic understanding of non-linear flux quantization reviewed above provides a systematic mathematical theory that clearly delineates the available choices of non-perturbative completions and allows one to rigorously derive their consequences.

### 4.1 RR-field flux quantization in 10d

Recall the Gauss law of the unbounded RR-field flux densities (Ex. 2.10, Prop. 2.14) as commonly expected in *massive* type IIA supergravity and ignoring (as commonly done, but see [Braunack-Mayer et al. 2019] for possible justification) the non-linear Bianchi identity (Ex. 2.13) of the dual B-field flux  $H_7$  (whence we now notationally suppress  $H_7$  altogether, as usual) its admissible flux quantization laws have classifying spaces  $\mathcal{A}$  whose  $\mathbb{R}$ -Sullivan algebra looks as follows:

$$\text{SolSpace} = \left\{ \begin{array}{l} H_3 \in \Omega_{\text{dR}}^3(X^9) \\ F_{2\bullet} \in \Omega_{\text{dR}}^{2\bullet}(X^9) \end{array} \middle| \begin{array}{l} dH_3 = 0 \\ dF_{2\bullet} = H_3 \wedge F_{2\bullet-2} \end{array} \right\} \Rightarrow \text{CE}(\mathcal{A}) = \mathbb{R}[h_3, f_{2\bullet}] / \left( \begin{array}{l} dh_3 = 0 \\ df_{2\bullet} = h_3 \wedge f_{2\bullet-2} \end{array} \right). \quad (40)$$

Now it so happens that a space  $\mathcal{A}$  with this property is given [Freed et al. 2007a, p. 6] [Braunack-Mayer et al. 2019, Lem. 2.31] by the classifying space  $\text{KU}_0$  for complex topological K-theory in degree=0, homotopy-quotiented by an action of the projective unitary group  $\text{PU}(\mathcal{H})$  on an essentially unique separably infinite-dimensional complex Hilbert space  $\mathcal{H}$ :

$$\text{KU}_0 \xrightarrow{\text{hofib}_p} \overbrace{\text{KU}_0 // \text{PU}(\mathcal{H})}^{\mathcal{A} \equiv} \xrightarrow{p} B\text{PU}(\mathcal{H}) \simeq B^3\mathbb{Z}, \quad (41)$$

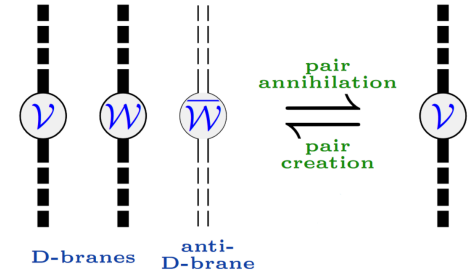
Accordingly, the generalized nonabelian cohomology theory classified by  $\mathcal{A} \equiv \text{KU}_0 // \text{PU}(\mathcal{H})$  decomposes over the ordinary integral cohomology in degree=3, with fibers being the abelian Whitehead-generalized cohomology of topological K-theory. As such it may and traditionally is understood as an abelian but *twisted* cohomology theory: *Twisted topological K-theory* [Atiyah & Segal 2004, Def. 3.3] [Freed et al. 2007a, (2.6)] [Sati & Schreiber 2021b, Ex. 4.5.4] [Fiorenza et al. 2023, Ex. 3.4]. For a review of twisted K-theory see also the contribution [Rosenberg 2024] in this collection.

$$\begin{array}{ccc} & & \text{KU}_0 // \text{PU}(\mathcal{H}) \\ & \nearrow \text{RR-field charges} & \downarrow \text{classifying fibration} \\ X^9 & \dashrightarrow & B\text{PU}(\mathcal{H}) \\ & \searrow \text{background B-field charges} & \downarrow \wr \\ & & B^2\text{U}(1) \end{array} \quad (42)$$

Moreover, under this decomposition, the nonabelian character map on (41) is [Fiorenza et al. 2023, Prop. 10.1] the *twisted Chern character* (the archetypical example which gives its name to the more general “character map”). Therefore, choosing (41) as the flux quantization law for the unbounded RR-fields means to hypothesize/declare that RR-field flux and hence D-brane charge is quantized in twisted topological K-theory, with the twisted de Rham cohomology-classes of the RR-field flux densities just being the image of these K-theory classes under the twisted Chern character. This is the *Hypothesis K* (our terminology) originally due to [Minasian & Moore 1997] [Witten 1998], with further perspectives added by [Freed & Hopkins 2000][Bouwknegt & Mathai 2001] and others.

**Remark 4.1** (Comparison to the literature). The original motivation close to the above logic via the character map may be found in [Witten 1998], in the paragraph wrapping p. 9-10. When comparing to [Minasian & Moore 1997] beware that these authors, and the literature following them, take the  $\mathbb{R}$ -rational D-brane charge to be expressed by the Chern character *multiplied* with the square root of the A-hat genus of the tangent bundle of spacetime. However, since this term is multiplicatively invertible, this is not intrinsic to the notion of D-brane charge and may be disregarded for the purpose of charge quantization (cf. [Freed & Hopkins 2000, fn. 12]); its role is rather as a technical convenience making the Chern character natural under push-forward [Brodzki et al. 2008, §2].

An influential argument why *Hypothesis K* (41) should be singled out among (the infinitude of) other compatible choices of RR-field flux quantization was the observation [Witten 1998, §3] that the equivalence relation on virtual vector bundles which characterizes topological K-theory  $KU_0(X)$  on compact Hausdorff spaces plausibly mimics the expected mechanism (“Sen’s conjecture” [Sen 1998]) of D-brane/anti D-brane-annihilation via tachyon condensation, by which pairs  $(\mathcal{W}, \overline{\mathcal{W}})$  of isomorphic but opposite Chan-Paton bundles on the worldvolume of coincident D-branes should mutually annihilate.



We highlight that this is a heuristic argument: There is no string-theoretic computation that actually verifies this intuition (cf. commentary by [Erl er 2013, p. 32]). In fact, [Witten 1998, fn. 2] already points out that, on closer inspection, it is less clear how the picture should work. This is noteworthy in view of the fact that the starting point (40) of *Hypothesis K* is a little shaky (as discussed after Ex. 2.12): There is room to doubt that *Hypothesis K* is quite the correct flux-quantization law for the RR-field, after all; see also the concerns raised by [de Boer et al. 2002, §4.5.2 & §4.6.5][Fredenhagen & Quella 2005, p. 1]<sup>1</sup>[Evs lin 2006, §8][Erl er 2013, p. 32].<sup>2</sup>

This motivates having a closer look at the flux quantization of the M-theoretic avatar/origin of the RR-fields: The C-field.

## 4.2 C-Field flux quantization in 11d

Given the C-field’s Gauss law (Ex. 2.12, Prop. 2.14) its admissible flux quantization laws have a classifying space  $\mathcal{A}$  whose  $\mathbb{R}$ -Sullivan algebra is as shown on the right here:

$$\text{SolSpace} = \left\{ \begin{array}{l} B_4 \in \Omega_{\text{dR}}^4(X^{10}) \\ B_7 \in \Omega_{\text{dR}}^7(X^{10}) \end{array} \middle| \begin{array}{l} dB_4 = 0 \\ dB_7 = -\frac{1}{2}B_4 \wedge B_4 \end{array} \right\} \Rightarrow \text{CE}(\mathfrak{L}\mathcal{A}) = \mathbb{R}[b_4, b_7] / \left( \begin{array}{l} db_4 = 0 \\ db_7 = -\frac{1}{2}b_4 \wedge b_4 \end{array} \right).$$

We review now two possible choices of such flux quantization laws  $\mathcal{A}$  for the C-field that have been considered in the literature (here we denote them by  $\mathcal{A}_{\text{DFM}}$  and  $\mathcal{A}_{\text{FSS}}$ , respectively), both of which, while quite distinct from each other, being an “evident” choice from their respective natural perspective.<sup>3</sup>

Recall again that, besides these “evident” choices, there is an infinitude of admissible variant flux quantization laws which differ in their torsion content. In the present case, any such choice is a hypothesis/definition concerning aspects of the elusive M-theory. Careful investigation of the implications of the “evident” flux quantization laws of the C-field may not only serve to decide if either is “correct” (which is not always straightforward to decide, as long as a plausibly complete definition of M-theory remains missing), but also to understand how variant flux quantization laws would have to be chosen if the “evident” ones are deemed to have undesired implications.

<sup>1</sup>[Fredenhagen & Quella 2005, p. 1]: “It might surprise that despite all the progress that has been made in understanding branes on group manifolds, there are usually not enough D-branes known to explain the whole charge group predicted by (twisted) K-theory. [...] it is fair to say that a satisfactory answer is still missing.”

<sup>2</sup>[Erl er 2013, p. 32]: “It would also be interesting to see if these developments can shed light on the long-specified relation between string field theory and the K-theoretic description of D-brane charge. We leave these questions for future work.”

<sup>3</sup>There are other proposals that advocate seeking a generalized cohomology underlying the fields in M-theory, see [Sati 2005][Sati 2006].

**DFM-like flux-quantization.** If one takes the point of view that a higher  $U(1)$ -flux quantization law as in Ex. 3.10 is the most natural starting point, which naively demands  $G_4$  to be quantized in integral 4-cohomology with classifying space  $B^4\mathbb{Z} \equiv K(\mathbb{Z}, 4)$ , then one is naturally led to consider the deformation of this situation which just adds-on the condition that half the cup-square of this 4-class be trivialized in rational cohomology.

In terms of classifying spaces, this means to pass to the homotopy fiber, here to be denoted  $\mathcal{A}_{\text{DFM}}$ , of the map that classifies (minus) half the cup-square cohomology operation on integral 4-cohomology. This has the required Sullivan model, as shown (these kinds of computations are reviewed in [Fiorenza et al. 2023, §1, §5]):

$$\begin{array}{l}
\text{classifying map} \quad B^4\mathbb{Z} \xrightarrow{-\frac{1}{2}\text{sq}} B^8\mathbb{Q} \\
\text{Sullivan model} \quad \frac{\mathbb{R}[g_4]}{(d g_4=0)} \xleftarrow{-\frac{1}{2}g_4 \wedge g_4 \leftarrow q_8} \frac{\mathbb{R}[q_8]}{(d q_8=0)} \\
\text{cohomology operation} \quad H^4(-; \mathbb{Z}) \xrightarrow{-\frac{1}{2}\text{sq}_* = -\frac{1}{2}(-) \cup^2} H^8(-; \mathbb{Q})
\end{array}$$

$$\begin{array}{ccc}
\begin{array}{c} \text{homotopy fiber} \\ \text{of classifying map} \\ \text{of fractional} \\ \text{cup-square operation} \end{array} & \begin{array}{c} \text{presented as} \\ \text{principal 8-bundle} \\ \text{over classifying space} \end{array} & \begin{array}{c} \text{image in} \\ \text{Sullivan models} \end{array} \\
\begin{array}{ccc} \mathcal{A}_{\text{DFM}} & \longrightarrow & * \\ \downarrow & \swarrow \text{(hpb)} & \downarrow 0 \\ B^4\mathbb{Z} & \xrightarrow{-\frac{1}{2}\text{sq}} & B^8\mathbb{Q} \end{array} & \xrightarrow{\text{hmt p}} & \begin{array}{ccc} \mathcal{A}_{\text{DFM}} & \longrightarrow & EB^7\mathbb{Q} \\ \downarrow & \text{(pb)} & \downarrow p \\ B^4\mathbb{Z} & \xrightarrow{-\frac{1}{2}\text{sq}} & B^8\mathbb{Q} \end{array} & \xrightarrow{\text{CE}(I(-))} & \begin{array}{ccc} \frac{\mathbb{R}[g_4, g_7]}{(d g_4=0, d g_7=-\frac{1}{2}g_4 \wedge g_4)} & \xleftarrow{-\frac{1}{2}g_4 \wedge g_4 \leftarrow q_8} & \frac{\mathbb{R}[g_7, q_8]}{(d q_8=0, d g_7=q_8)} \\ \uparrow & \text{(po)} & \uparrow \\ \frac{\mathbb{R}[g_4]}{(d g_4=0)} & \xleftarrow{-\frac{1}{2}g_4 \wedge g_4 \leftarrow q_8} & \frac{\mathbb{R}[q_8]}{(d q_8=0)} \end{array} \quad (43)
\end{array}$$

This flux quantization law corresponds essentially to the model of the C-field considered in [Diaconescu et al. 2007]<sup>4</sup> [Moore 2005] following [Hopkins & Singer 2005] – when specialized to the case where the Pontrjagin classes of space-time vanish (such as for near horizon geometries of flat singular branes, §2.2, by [Sati & Schreiber 2021a, Prop. 22]), namely it enforces integer  $G_4$ -flux quantization much as in Ex. 3.10 while implementing  $J_8 := \frac{1}{2}G_4 \wedge G_4$  as an electric source term, but essentially no  $G_7$ -flux quantization is enforced.

By the long exact sequence of homotopy groups associated with the fiber sequence (43) it follows that the homotopy groups of  $\mathcal{A}_{\text{DFM}}$  are concentrated in degrees 4 and 7, which implies that the flat singular branes (cf. §2.2) it predicts are exactly integer numbers of M5-branes and any (rational) number of M2-branes (cf. §2.3)

$$\pi_n(\mathcal{A}_{\text{DFM}}) = \begin{cases} \mathbb{Z} & | & n = 4 \\ \mathbb{Q} & | & n = 7 \\ 0 & | & \text{otherwise} \end{cases} \Rightarrow \begin{array}{l} \text{charges of flat} \\ \text{singular M5-branes} \end{array} = H^1(\mathbb{R}^{10,1} \setminus \mathbb{R}^{5,1}; \Omega \mathcal{A}_{\text{DFM}'}) \simeq \pi_4(\mathcal{A}_{\text{DFM}'}) \simeq \mathbb{Z}, \\ \begin{array}{l} \text{charges of flat} \\ \text{singular M2-branes} \end{array} = H^1(\mathbb{R}^{10,1} \setminus \mathbb{R}^{2,1}; \Omega \mathcal{A}_{\text{DFM}'}) \simeq \pi_7(\mathcal{A}_{\text{DFM}'}) \simeq \mathbb{Q}.
\end{array}$$

Of course, with charge-quantization in generalized non-abelian cohomology (§3.2) it is straightforward to fix this, by forming instead the homotopy fiber of the integral  $-\text{sq} : B^4\mathbb{Z} \rightarrow B^8\mathbb{Z}$  and using the freedom in isomorphism of Sullivan models to rescale the generator  $g_7$  by 2:

$$\begin{array}{ccc}
\begin{array}{c} \text{homotopy fiber} \\ \text{of classifying map} \\ \text{of integral} \\ \text{cup-square operation} \end{array} & \begin{array}{c} \text{presented as} \\ \text{bundle 6-gerbe} \\ \text{over classifying space} \end{array} & \begin{array}{c} \text{image in} \\ \text{Sullivan models} \end{array} \\
\begin{array}{ccc} \mathcal{A}_{\text{DFM}'} & \longrightarrow & * \\ \downarrow & \swarrow \text{(hpb)} & \downarrow 0 \\ B^4\mathbb{Z} & \xrightarrow{-\text{sq}} & B^8\mathbb{Z} \end{array} & \xrightarrow{\text{hmt p}} & \begin{array}{ccc} \mathcal{A}_{\text{DFM}'} & \longrightarrow & EB^6U(1) \\ \downarrow & \text{(pb)} & \downarrow p \\ B^4\mathbb{Z} & \xrightarrow{-\text{sq}} & B^7U(1) \end{array} & \xrightarrow{\text{CE}(I(-))} & \begin{array}{ccc} \frac{\mathbb{R}[g_4, 2g_7]}{(d g_4=0, d 2g_7=-g_4 \wedge g_4)} & \xleftarrow{-g_4 \wedge g_4 \leftarrow q_8} & \frac{\mathbb{R}[2g_7, q_8]}{(d q_8=0, d 2g_7=q_8)} \\ \uparrow & \text{(po)} & \uparrow \\ \frac{\mathbb{R}[g_4]}{(d g_4=0)} & \xleftarrow{-g_4 \wedge g_4 \leftarrow q_8} & \frac{\mathbb{R}[q_8]}{(d q_8=0)} \end{array} \quad (44)
\end{array}$$

This adjusted flux-quantization law “ $\mathcal{A}_{\text{DFM}'}$ ” now enforces the desired M2-charge quantization:

$$\pi_n(\mathcal{A}_{\text{DFM}'}) = \begin{cases} \mathbb{Z} & | & n = 4 \\ \mathbb{Z} & | & n = 7 \\ 0 & | & \text{otherwise} \end{cases} \Rightarrow \begin{array}{l} \text{charges of flat} \\ \text{singular M5-branes} \end{array} = H^1(\mathbb{R}^{10,1} \setminus \mathbb{R}^{5,1}; \Omega \mathcal{A}_{\text{DFM}'}) \simeq \pi_4(\mathcal{A}_{\text{DFM}'}) \simeq \mathbb{Z}, \\ \begin{array}{l} \text{charges of flat} \\ \text{singular M2-branes} \end{array} = H^1(\mathbb{R}^{10,1} \setminus \mathbb{R}^{2,1}; \Omega \mathcal{A}_{\text{DFM}'}) \simeq \pi_7(\mathcal{A}_{\text{DFM}'}) \simeq \mathbb{Z}.
\end{array}$$

(cf. the claim in [Moore 2005, §5]).

But it still does not *predict* the half-integral shift of the M5-brane charge, nor the correction of the electric source by the  $I_8$ -term in the case that the Pontrjagin classes of spacetime do not vanish, even though these effects can be added “by hand”. In order to see these effects arise automatically we turn to yet another possible flux-quantization law of the C-field:

<sup>4</sup>Beware that much of the content of [Diaconescu et al. 2007] is concerned with first adjoining an  $E_8$  Yang-Mills field to the C-field and then imposing gauge equivalences which make this field disappear again up to gauge equivalence; see (3.11) there.

**FSS flux-quantization.** Another perspective is to regard the baseline of all flux quantization to be that classified by the point  $\mathcal{A}_0 \equiv *$  (for the trivial theory) and to obtain non-trivial classifying spaces from this maximally unbiased starting point by iterated attachment of cells in the sense of CW-complexes (e.g. [Hatcher 2002, p. 5]). The *minimal* choice of C-field flux quantization in this sense, requiring the minimum number 1 of cell attachments, is to take  $\mathcal{A}_{\text{FSS}} \equiv S^4$  to be the (homotopy type of) the 4-sphere [Sati 2013, §2.5] (which is a valid choice of C-field flux quantization, by the examples on p. 21).

The generalized nonabelian cohomology theory classified by the  $n$ -spheres is known as *Cohomotopy* [Borsuk 1936][Pontrjagin 1938][Spanier 1949][Peterson 1956], being dual to the unstable homology theory constituted by the homotopy groups of spaces.

Therefore, the hypothesis that the proper classifying space for C-field flux quantization is the 4-sphere may be called *Hypothesis H*, for “Homotopy cohomology theory” [Fiorenza et al. 2020][Sati & Schreiber 2020][Grady & Sati 2021][Fiorenza et al. 2021a][Fiorenza et al. 2021b][Sati & Schreiber 2021a], review in [Fiorenza et al. 2023, §12].

**Hypothesis H on flat spacetimes.** The non-torsion homotopy groups of  $S^4$  are exactly in degrees 4 and in degree 7 (whose generator is the quaternionic Hopf fibration, cf. [Fiorenza et al. 2020, p. 4]). This implies that Hypothesis H gives the expected integral charge quantization for flat singular M5-branes (cf. §2.2 & §2.3) *and* for flat singular M2-branes – the “Page charge” (51).

| Unstable homology theory                        | Unstable (nonabelian) cohomology theory         |
|---|---|
| homotopy  | cohomotopy                                      |
| $\pi_n(X) \equiv \pi_0 \text{Map}^{*/}(S^n, X)$ | $\pi^n(X) \equiv \pi_0 \text{Map}^{*/}(X, S^n)$ |

|   |
|---|
| $\begin{aligned} \pi^4(\mathbb{R}^{10,1} \setminus \mathbb{R}^{5,1}) &= \pi^4(\mathbb{R}^{5,1} \times \mathbb{R}_+ \times S^4) \\ &= \pi^4(S^4) = \pi_4(S^4) = \mathbb{Z} \end{aligned}$                        |
| $\begin{aligned} \pi^4(\mathbb{R}^{10,1} \setminus \mathbb{R}^{2,1}) &= \pi^4(\mathbb{R}^{2,1} \times \mathbb{R}_+ \times S^7) \\ &= \pi^4(S^7) = \pi_7(S^4) = \mathbb{Z} \oplus \mathbb{Z}_{12} \end{aligned}$ |

$$\begin{array}{ccc}
 \begin{array}{c} \text{quaternionic} \\ \text{Hopf fibration} \\ S^3 \hookrightarrow S^7 \\ \downarrow h_{\mathbb{H}} \\ S^4 \end{array} & [S^7 \xrightarrow{h_{\mathbb{H}}} S^4] = 1 \in \mathbb{Z} \hookrightarrow \pi_7(S^4) & \begin{array}{c} S^7 \in \pi^7(X) \\ \downarrow h_{\mathbb{H}} \\ S^4 \in \pi^4(X) \end{array} \\
 & & \begin{array}{c} \text{pure} \\ \text{M}_2\text{-brane charges} \\ \text{mapping into} \\ \text{full} \\ \text{M-brane charges} \end{array} \quad (45)
 \end{array}$$

On the other hand,  $S^4$  also has plenty of torsion homotopy groups [Sati & Schreiber 2023a, (22)]: Under Hypothesis H, each of them is a prediction of novel “fractional M-brane” species (such as of fractional M2-branes, cf. [Aharony et al. 2008, §2.2], of order 12, even in flat space) which do not manifest as BPS-states of supergravity.

Notice that the prediction of stable non-BPS branes carrying torsion charges is a generic property of flux quantization laws (in fact: their defining property, cf. (28)), and is well-familiar in the context of Hypothesis K (§4.1): cf. [Braun 2000][Brunner et al. 2002a][Brunner et al. 2002b].

The specific torsion content in the C-field that is implied by Hypothesis H has the following consequences:

**Divided powers of M5-brane charge.** Noticing that the generator  $S^4 \rightarrow B^4\mathbb{Z}$  of  $\pi_4(B^4\mathbb{Z}) \simeq \mathbb{Z}$  induces a cohomology operation  $\pi^4(-) \rightarrow H^4(-; \mathbb{Z})$ , there is an integer-cohomology class  $\gamma_4$  underlying the Cohomotopical C-field charge. This has the following properties:

- (i)  $\gamma_4 \cup \gamma_4$  is divisible by 2, thus making the C-field’s electric source  $-\frac{1}{2}G_4 \wedge G_4$  be integral [Grady & Sati 2021, Prop. 2.7 (iii)], cf. [Diaconescu et al. 2007, p. 29];
- (ii)  $\gamma_4 \cup \gamma_4 \cup \gamma_4$  is divisible by 6, thus making the 11d CS-term, locally  $\frac{1}{6}C_3 \wedge G_4 \wedge G_4$ , be globally well-defined [Grady & Sati 2021, p. 12], cf. [Witten 1997, p. 10].

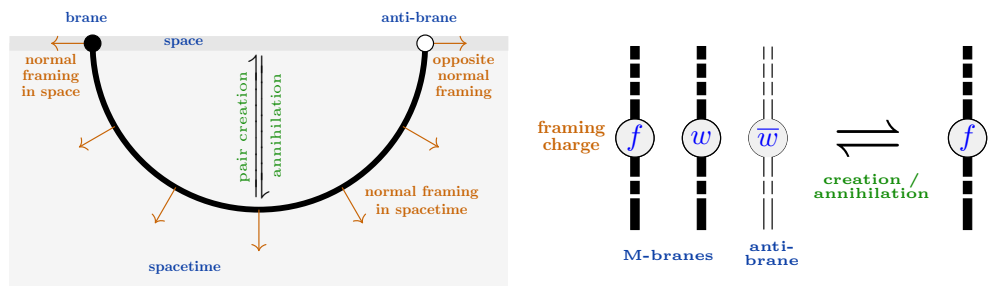
**Worldvolumes in Cobordism cohomology.** The Pontrjagin theorem [Pontrjagin 1938] [Kosinski 1993, §IX] identifies unstable  $n$ -Cohomotopy with (unstable, framed) Cobordism,

$$\pi^n(X^d) \xrightarrow[\sim]{\text{Pontrjagin theorem}} \text{Cob}_{\text{Fr}}^n(X^d)$$

cobordism classes of normally framed submanifolds of codimension= $n$ .

suggestive of the *worldvolumes of solitonic M-branes* (§2.2) carrying Cohomotopy charge, cf. [Sati & Schreiber 2020, §2.1][Sati & Schreiber 2023a, p. 13][Sati & Schreiber 2022, §2.4]. Here the cobordism relation reflects (anti-)brane annihilation

(and creation) much as expected in K-theory (cf. p. 29). *Stabilized* Cohomotopy is equivalent to algebraic K-theory over the “absolute base field  $\mathbb{F}_1$ ” [Chu et al. 2012, Thm. 5.9].



**Hypothesis H on gravitational backgrounds.** The equations of motion of  $D = 11$  supergravity are in fact subject to “higher curvature corrections” which shift the  $G_4$ -flux by  $\frac{1}{2}(\frac{1}{2}p_1)$  (cf. [Tsimpis 2004, p. 8]) and shift the Bianchi identity for  $G_7$  by a term proportional to  $I_8 \equiv \frac{1}{48}(p_2 - (\frac{1}{2}p_1)^2)$  (cf. [Souères & Tsimpis 2017, §4]), where  $p_n$  denotes the  $n$ th Pontrjagin form of the Levi-Civita connection (the gravitational field) on spacetime (cf. [Fiorenza et al. 2023, Ex. 8.1 (ii)]). In order to reflect such extra coupling of the C-field flux to gravitational background charges, the flux-quantizing cohomology theory must be twisted, somehow, by the tangent bundle of spacetime, induced by an action of a Spin-group on the classifying space  $\mathcal{A}$ .

While Spin-groups do not seem to act usefully on  $\mathcal{A}_{\text{DFM}}$ , preventing a *systematic* coupling of this model to gravitational charges, it is noteworthy that Spin(5) acts, of course, canonically on  $\mathcal{A}_{\text{FSS}} \equiv S^4 \simeq S(\mathbb{R}^5)$ .

This means in particular that for  $D = 11$  supergravity on 8-manifolds  $X^{10} \simeq T^2 \times X^8$  equipped with tangential Spin(5)-structure  $\tau$ , there is a canonical notion of tangentially *twisted 4-Cohomotopy* [Fiorenza et al. 2020, §2] [Fiorenza et al. 2023, Ex. 3.8], given by homotopy classes of sections of the 4-spherical fibration associated to the tangential structure – cf. the analogous case of twisted K-theory (42).

$$\pi^\tau(X^{10}) \equiv \left\{ \begin{array}{c} \text{tangentially} \\ \text{twisted} \\ \text{4-Cohomotopy} \\ T^2 \times X^8 \end{array} \right\} \begin{array}{c} \xrightarrow{\tau} B\text{Spin}(5) \\ \xrightarrow{\text{tangential twist}} B\text{Spin}(8) \end{array} \left. \begin{array}{c} \xrightarrow{c_3} S^4 // \text{Spin}(5) \\ \downarrow \text{classifying fibration for} \\ \text{twisted 4-Cohomotopy} \\ \downarrow \\ B\text{Spin}(5) \\ \swarrow \\ B\text{Spin}(8) \end{array} \right\} \quad (46)$$

/rel  
hmtp

Now C-field flux quantization in twisted Cohomotopy does imply the expected half-integral shifted quantization of the  $G_4$ -flux, as follows [Fiorenza et al. 2020, §3.4, Prop. 3.13]:

$$\begin{array}{ccc} \text{cocycle in} & X \xrightarrow{c_3} S^4 // \text{Spin}(5) \simeq B\text{Spin}(4) & \\ \text{tangentially twisted} & \swarrow \text{Fr}(X) & \\ \text{4-cohomotopy} & B\text{Spin}(d) & \\ \\ \text{induced charge in} & H^\bullet(X; \mathbb{R}) \xleftarrow{c_3^*} H^\bullet(B\text{Spin}(4); \mathbb{R}) = \mathbb{R}[p_1, \chi_4] & \\ \text{real cohomology} & [p_1(\nabla)] \xleftarrow{\quad} p_1 \text{ first Pontrjagin class} & \\ & [G_4] \xleftarrow{\quad} \frac{1}{2}\chi_4 \text{ fractional Euler class} & \\ \\ \text{induced charge in} & H^\bullet(X; \mathbb{Z}) \xleftarrow{c_3^*} H^\bullet(B\text{Spin}(4); \mathbb{Z}) = \mathbb{Z}[\frac{1}{2}p_1, \frac{1}{2}\chi_4 + \frac{1}{4}p_1] & \\ \text{integral cohomology} & \underbrace{[G_4 + \frac{1}{4}p_1(\nabla)]}_{[\tilde{G}]} \xleftarrow{\quad} \frac{1}{2}\chi_4 + \frac{1}{4}p_1 \text{ universal integral} & \\ & & \text{characteristic class} \end{array} \quad (47)$$

However, for this twisting to preserve the distinction between M2- and M5-brane charges and hence the quaternionic Hopf fibration (45), one actually has [Fiorenza et al. 2020, Prop. 2.20][Fiorenza et al. 2022, Prop. 2.2] to regard Spin(5) as the quaternionic unitary group Sp(2) which acts canonically also on  $S^7 = S(\mathbb{H}^2)$ .

While both groups are abstractly isomorphic, Spin(5)  $\simeq$  Sp(2), they are *not* isomorphic *as subgroups* of Spin(8), *but* they are mapped into each other under the triality automorphism  $\text{tri} : \text{Spin}(8) \xrightarrow{\sim} B\text{Spin}(8)$ . [Fiorenza et al. 2020, Prop. 2.17].

This means (i) that one actually needs Sp(2)-structure on spacetime to couple both M5- as well as M2-brane charges to gravitational charges, (ii) which differs from Spin(5)-structure by triality and (iii) the gravitational charges in degree=4 are invariant under this transformation, but the degree=8 charges pick up the  $I_8$ -term [Fiorenza et al. 2020, Lem. 2.19].

For  $\tau$  an Sp(2)-structure on spacetime, the flux densities in the image of the twisted character on  $\tau$ -twisted 4-Cohomotopy are shown on the right<sup>5</sup> [Fiorenza et al. 2020, Prop. 3.20][Fiorenza et al. 2022, Thm. 2.14].

$$\begin{array}{ccc} B\text{Sp}(2) & \xrightarrow{\sim} & B\text{Spin}(5) \\ \downarrow & & \downarrow \\ B\text{Spin}(8) & \xrightarrow[\sim]{B\text{tri}} & B\text{Spin}(8) \\ \\ H^\bullet(B\text{Sp}(2); \mathbb{R}) & \xleftarrow{(B\text{tri})^*} & H^\bullet(B\text{Spin}(5); \mathbb{R}) \\ \frac{1}{2}p_1 & \xleftarrow{\quad} & \frac{1}{2}p_1 \\ (\frac{1}{4}p_1)^2 - 24 \cdot I_8 & \xleftarrow{\quad} & \frac{1}{4}p_2 \end{array}$$

$$\begin{array}{l} dG_4 = 0, \quad [\tilde{G}_4] = [G_4 + \frac{1}{4}p_1] \in H^4(X^d; \mathbb{Z}) \\ dG_7 = -\frac{1}{2}\tilde{G}_4 \wedge (\tilde{G}_4 - \frac{1}{2}p_1) - 12 \cdot I_8 \end{array} \quad (48)$$

<sup>5</sup>A gravitational shift in the Bianchi identity for  $G_7$  is expected but has remained undetermined [Tsimpis 2004, (4.16)] [Souères & Tsimpis 2017, (4.11)]. On the factor 12 in (48) cf. [Fiorenza et al. 2020, pp. 12-13 & §3.8][Sati & Schreiber 2021a, Rem. 7 & 8][Sati & Schreiber 2023a, Rem. 4.1]; and notice that this term disappears in (52) below.



### 4.3 B-Field flux quantization in 6d

The (5+1)-dimensional worldvolume  $\Sigma^6$  of an M5-brane sigma-model is to carry a 3-form flux  $H_3$  which in simple (decoupled) situations satisfies the equations of motion of a self-dual higher gauge field (Ex. 2.11) [Claus et al. 1998] [Witten 2010]. The effect of the self-duality on the phase space is, with Prop. 2.14, that the evident Gauß law  $dH_3 = 0$  is imposed on a *single* flux density, in contrast to the non-self-dual case (Ex. 2.15). Therefore a traditional flux quantization of the self-dual field is in a *single* copy [Freed et al. 2007c, p. 32] of ordinary differential cohomology in degree=3 (cf. Ex. 3.10, [Sati & Schreiber 2023b, §3.2]), in contrast to the two copies seen for non-self dual abelian gauge fields Ex. 3.9.

But in general the  $H_3$ -flux on the Fivebrane is not actually closed, rather it is sourced by the pullback of  $G_4$ , and its self-duality is subtle [Howe & Sezgin 1997, (36, 40)] [Howe, Sezgin & West 1997][Sorokin 2000, (5.57, 5.82)][Giotopoulos et al. 2024b].

Hence the quantization of  $H_3$  depends on the quantization of  $G_4$ . Assuming Hypothesis H for the latter and since the homotopy fiber of the  $\mathbb{H}$ -Hopf fibration is  $S^3$ , it is natural to quantize  $H_3$  in 3-Cohomotopy twisted by the C-field's Cohomotopy-charge, via the  $\mathbb{H}$ -Hopf fibration [Fiorenza et al. 2020, Prop. 3.20][Fiorenza et al. 2021b, (10)]. In fact this extends from Cauchy surfaces  $\Sigma^5$  to the worldvolume  $\Sigma^6$  [Giotopoulos et al. 2024b],

$$\begin{array}{c} \boxed{\text{B-field flux on}} \\ \boxed{\text{5-brane worldvolume}} \\ \boxed{\phi : \Sigma^6 \rightarrow Y^{11}} \end{array} \left\{ \begin{array}{l} \rightarrow \boxed{dH_3 = \phi^* \tilde{G}_4} \\ \rightarrow \boxed{\text{subtle self-duality}} \end{array} \right. \quad (49)$$

$$\text{C-field-twisted 3-Cohomotopy } \pi^{\phi^* c_3}(\Sigma^5) \equiv \left\{ \begin{array}{c} \Sigma^5 \xrightarrow[b_2]{\text{B-field charges on 5-brane worldvolume}} S^7 // \text{Sp}(2) \\ \downarrow \phi \quad \text{fivebrane worldvolume embedding} \quad \mathbb{H}\text{-Hopf fibration as classifying fibration for twisted 3-Cohomotopy} \\ X^{10} \xrightarrow[c_3]{\text{C-field charges on spacetime (46)}} S^4 // \text{Sp}(2) \end{array} \right\} \Big/_{\text{rel hntp}} \quad (50)$$

Moreover, for the topological analysis of WZW terms we may consider  $\Sigma^7$  an “extended worldvolume”, namely a closed 7-manifold which is a cobordism from  $\emptyset$  to the fivebrane worldvolume  $\Sigma^6$  and back to  $\emptyset$ , and over which the worldvolume fields  $(\phi, H_3)$  extend, cf. [Fiorenza et al. 2021a, §2]. Then the M2-brane background flux exerting Lorentz force on the 5-brane is as on the right [Page 1983, (8)][Duff et al. 1991, (43)],  $\text{flux density of Page charge of M2-branes } 2\tilde{G}_7 := 2(\phi^* G_7 + \frac{1}{2} H_3 \wedge \phi^* \tilde{G}_4) \in \Omega_{\text{dR}}^7(\Sigma^7) \quad (51)$  which we are showing scaled by 2 in accord with (44) and since this is how it actually appears as the *Hopf-WZ term* in the Fivebrane’s worldvolume theory [Intriligator 2000, (2.8)].

The combined C-field- & B-field flux quantization ought to imply (as for any Dirac charge quantization condition) the flux density (51) to have integer-valued integral over  $\Sigma^7$ :

- (i) its integral is the total magnetic charge of singular M2-branes enclosed by  $\Sigma^7$ , and these ought to appear in integer numbers
- (ii) its exponentiated integral mod  $\mathbb{Z}$  is an anomaly in the Hopf-WZ term in the fivebrane’s action functional, whose vanishing is the “level quantization”-condition for 7d Chern-Simons theory with Lagrange density (51).

The obstruction, under Hypothesis H, to (51) being integral, hence the M5-’s Hopf-WZ-term anomaly, turns out to be  $24I_8 = \chi_8$  of the given  $\text{Sp}(2)$ -structure. [Fiorenza et al. 2021a, Thm. 4.8, Ex. 3.2].

$$\begin{array}{ccc} \text{M-Fivebrane str. } \widehat{B\text{Sp}(2)} & \longrightarrow & * \\ \downarrow \tau & \text{(hpb)} & \downarrow \\ X^{10} & \xrightarrow{24I_8} & B^8\mathbb{Z} \end{array} \Rightarrow \begin{array}{ccc} \Sigma^7 & \xrightarrow{b_2} & S^7 // \widehat{\text{Sp}(2)} \dashrightarrow B^7\mathbb{Z} \\ \phi \downarrow & & \downarrow h_{\mathbb{H}} // \widehat{\text{Sp}(2)} \\ X^{10} & \xrightarrow{c_3} & S^4 // \widehat{\text{Sp}(2)} \end{array} \quad \left\{ \begin{array}{l} \text{quantized Hopf-WZ/Page-charge} \\ \exists \end{array} \right.$$

In particular, the Page charge (51) of *flat* M2-branes (Ex. 2.1) is integral, being (the Whitehead-integral formula for) the *Hopf invariant* of the cohomotopical C-field charge  $c_3 : S^7 \rightarrow S^4$ , cf. also [Sati & Schreiber 2023a, §2.7].

In conclusion, the Hypothesis  $\hat{\text{H}}$  that the C-field is flux-quantized in tangentially  $\widehat{\text{Sp}(2)}$ -twisted 4-Cohomotopy, and that the Fivebrane’s worldvolume B-field is flux-quantized in the correspondingly twisted 3-Cohomotopy implies the expected quantization of  $G_4$ -flux/M5-charge and of  $G_7$ -flux/M2-charge.

$$\begin{array}{ccc} \Sigma^7 & \xrightarrow[b_2]{\text{worldvolume B-field charges}} & S^7 // \widehat{\text{Sp}(2)} \\ \downarrow \phi & & \downarrow h_{\mathbb{H}} // \widehat{\text{Sp}(2)} \\ T^2 \times X^8 & \xrightarrow[c_3]{\text{C-field charges}} & S^4 // \widehat{\text{Sp}(2)} \\ \downarrow \tau & \text{M-Fivebrane str.} & \downarrow \\ & \xrightarrow{\text{tan. bundle / gravitational charges}} & B\widehat{\text{Sp}(2)} \\ & & \downarrow \\ & & B\text{Spin}(8) \end{array} \Rightarrow \left\{ \begin{array}{l} dG_4 = 0 \\ dH_3 = \phi^* \tilde{G}_4 \\ dG_7 = -\frac{1}{2} \tilde{G}_4 \wedge (\tilde{G}_4 - \frac{1}{2} p_1) \\ [G_4 + \frac{1}{4} p_1] \in H^4(X^{10}; \mathbb{Z}) \\ \underbrace{[\phi^*(2G_7) + H_3 \wedge \phi^* \tilde{G}_4]}_{2\tilde{G}_7} \in H^7(\Sigma^7; \mathbb{Z}) \end{array} \right. \quad (52)$$

## 4.4 Green-Schwarz mechanism in 11d

Just as the B-field (49) on the M-Fivebrane worldvolume (and we will see in a moment that this is not a coincidence), so the B-field flux density  $H_3$  in heterotic supergravity is famously not closed, but is sourced by the second Chern form of a  $G$ -Yang-Mills gauge potential  $\hat{A}$  minus the first fractional Pontrjagin form of the spin-connection  $\hat{\omega}$  on spacetime. This sourcing of  $H_3$ -flux is (the reflection of) the *Green-Schwarz mechanism* [Green & Schwarz 1984] [Candelas et al. 1985, p. 49], which implies that for  $G = E_8 \times E_8$  the heterotic superstring on such a background is anomaly-free — iff the de Rham coboundary (53) is flux quantized to a coboundary on *integral* cohomology classes ([Witten 2000, (2.11)]).

There is a higher Lie group, the String<sup>c<sub>2</sub></sup> 2-group (see pointers in [Fiorenza et al. 2014, App.]) whose classifying space is the *universal* space carrying a homotopy of this form [Sati 2011, §2][Sati et al. 2012, §2.2].

This means that the gauge-, gravitational- and B-field charges of heterotic supergravity may *jointly* be understood as a single higher gauge field with higher structure group the String<sup>c<sub>2</sub></sup> 2-group, in differential (stacky) refinement (“twisted differential String-structures”) of the resulting diagram on the right.

$$dH_3 = c_2(\hat{A}) - \frac{1}{2}p_1(\hat{\omega}) \quad (53)$$

This 2-group gauge theoretic flux quantization of (the Green-Schwarz mechanism in) heterotic supergravity has been discussed in [Sati et al 2009, p. 13] [Sati et al. 2012] [Fiorenza et al. 2014, §3.7, §3.8] [Fiorenza et al. 2015a], a corresponding construction (for the translation see [Capotosti 2016]) in terms of bundle gerbes (and for the special case  $c_2 = 0$ ) is in [Waldorf 2013]. The terminology, at least, of 2-group gauge theory for GS-mechanisms has recently become popular in the non-mathematical physics literature, which also speaks of “1-form symmetries” (e.g. [Benini et al. 2019]) or “categorical symmetries” (e.g. [Cordova et al. 2022], cf. [Schreiber & Škoda 2009]).

However, as with the RR-fields in type II supergravity §4.1, the flux-quantization and hence *partial* non-perturbative completion of the fields in heterotic supergravity somewhat begs the question: How does it connect to the expected *full* non-perturbative completion via M-theory?

Indeed the GS-mechanism has been argued to lift to M-theory, where the heterotic spacetime  $X^{10}$  appears as a *pair* of “MO<sub>9</sub>-planes” in  $Y^{11}$  each carrying *one* of the two copies of  $E_8$  gauge fields [Hořava & Witten 1996], cf. e.g. [Dumitru 2022, §1.3], which is thought to, somehow, be the restriction of an  $E_8$ -gauge field  $\hat{A}_{E_8}$  on all of  $Y^{11}$  itself, modifying the relation (49) to [Witten 1997, (2.2)][Diaconescu et al. 2007, (3.9)]:

$$\text{M-theoretic avatar of GS-mechanism (53)} \quad dH_3 = G_4 - \frac{1}{4}p_1(\hat{\omega}) + c_2(\hat{A}_{E_8}) \xrightarrow[S(U(1)^2) \subset E_8]{\text{heterotic line bundle}} dH_3 = G_4 - \frac{1}{4}p_1(\hat{\omega}) - F_2 \wedge F_2. \quad (54)$$

While the ontology of  $\hat{A}_{E_8}$  had remained mysterious [Evslin & Sati 2003], for quasi-realistic phenomenology it is “heterotic line bundles” that matter [Anderson et al. 2012], reducing the structure group along  $S(U(1)^n) \hookrightarrow SU(n) \hookrightarrow E_8$  with  $2 \leq n \leq 5$ . For  $n = 2$  the resulting Green-Schwarz type Bianchi identity (54) does have a flux quantization compatible with the C-field flux quantization  $\mathcal{A}_{\text{FSS}}$  in §4.2:

The quaternionic Hopf fibration (45) factors through the *twistor fibration*  $t_{\mathbb{H}} : \mathbb{C}P^3 \rightarrow S^4$ , the tangentially twisted flux quantization law  $\mathcal{A} \equiv \mathbb{C}P^3$  is admissible for M-theory with heterotic line bundles and implies all the desired total charge quantizations [Fiorenza et al. 2022, Thm. 2.14, (6)].

For MO<sub>9</sub>-s but also for (M-Fivebranes probing) ADE-singularities in heterotic M-theory, this flux quantization restricts to (53), by [Sati & Schreiber 2020, Thm. 1.1].

Remarkably, the mechanism behind this flux-quantized Green-Schwarz mechanism lifts, at least for flat branes, to C-field fluxes seen not just in rational but in any *complex-oriented* Whitehead-generalized cohomology theory [Sati & Schreiber 2023a, §2.9]. This includes complex K-theory but also elliptic cohomology and complex Cobordism and might finally explain the role these cohomology theories play for flux quantization in M-theory.

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